

Nuclear collisions as seen through photons

Jean-François Paquet

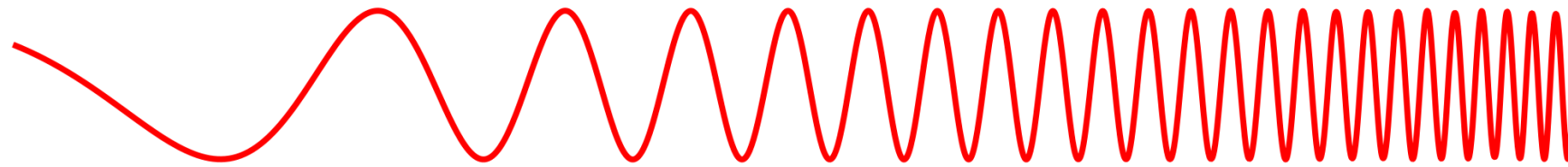
November 14, 2023



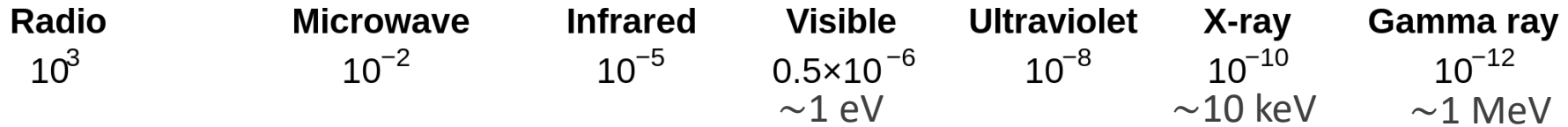
INSTITUTE for NUCLEAR THEORY

The higher end of the electromagnetic spectrum

Penetrates Earth's Atmosphere?

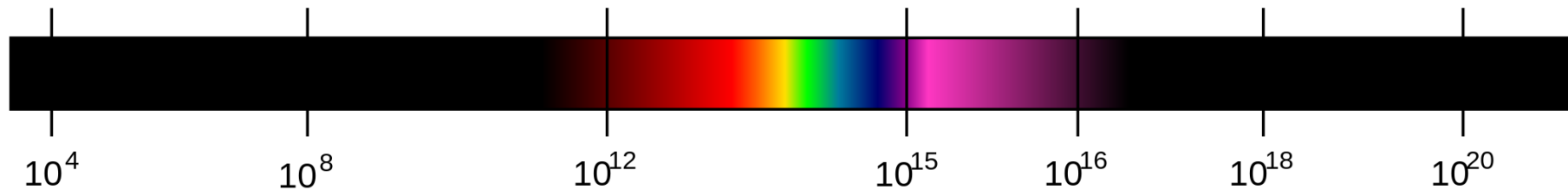


Radiation Type
Wavelength (m)



~1 GeV

Frequency (Hz)



Temperature of objects at which this radiation is the most intense wavelength emitted

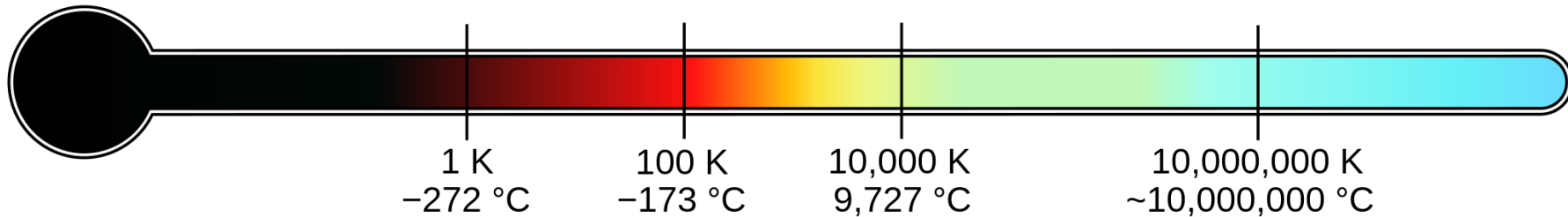
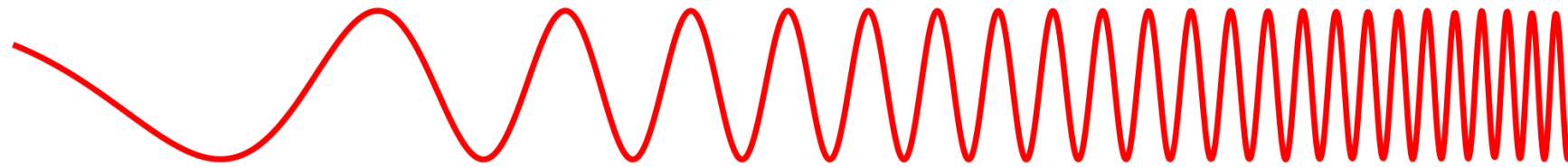


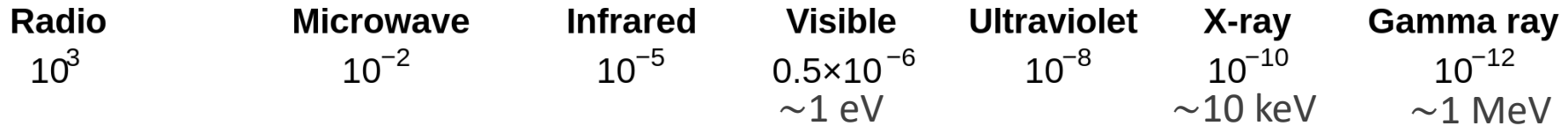
Image modified from Wikimedia

The higher end of the electromagnetic spectrum

Penetrates Earth's Atmosphere?

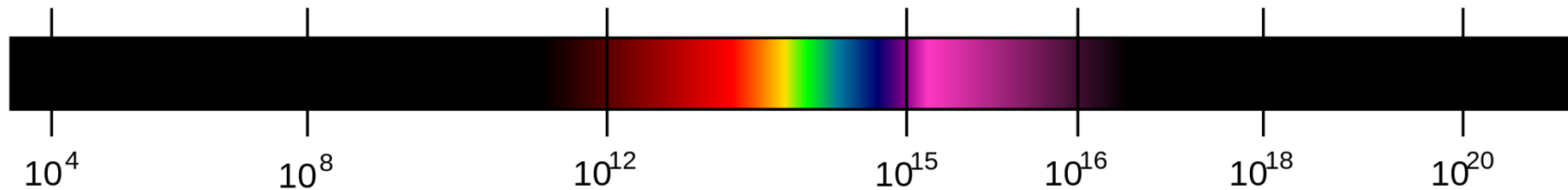


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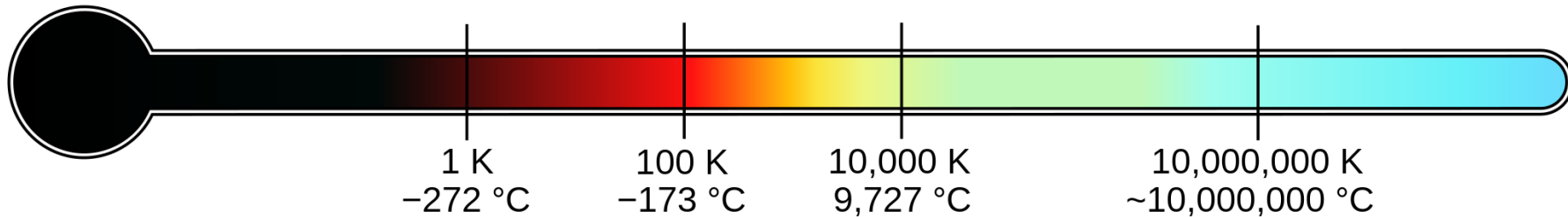


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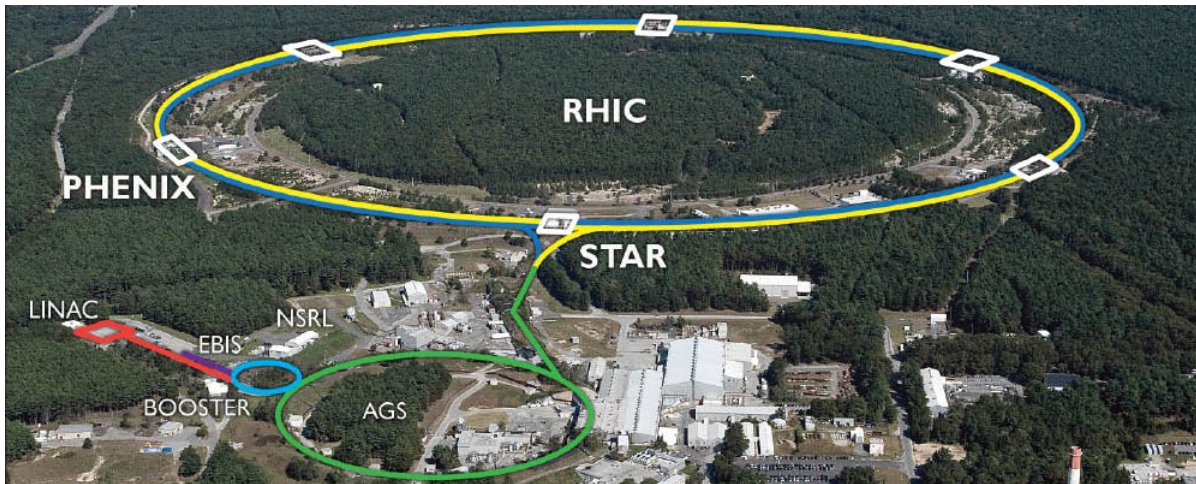


$T \sim 10^{12} \text{K}$

Image modified from Wikimedia

RHIC and LHC

Relativistic Heavy Ion Collider (RHIC)
[Brookhaven National Lab, Long Island, NY]



$$\sqrt{s_{NN}} \sim 10^2 \text{ GeV}$$

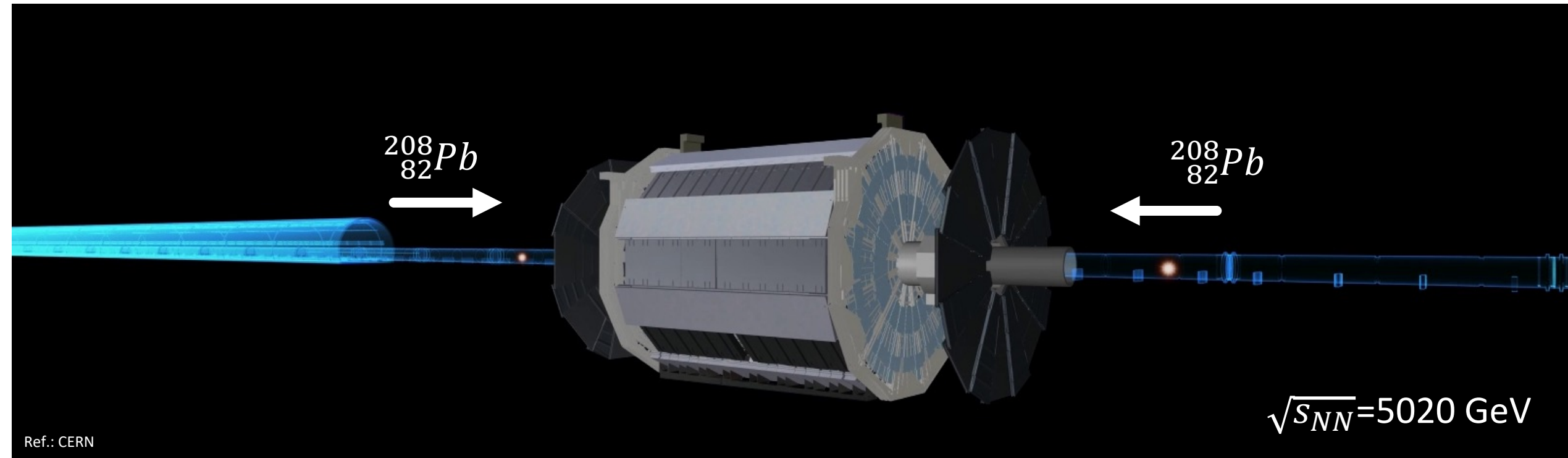
Large Hadron Collider (LHC)
[CERN, Geneva, Switzerland/France]



$$\sqrt{s_{NN}} \sim 10^3 \text{ GeV}$$

Nuclear collisions

- Kinetic energy of nuclei \sim 100-2500 times mass of nuclei



Pb-Pb at rest: $\sqrt{s_{NN}} \approx 2 \text{ GeV}$

Hadronic decay photons in nuclear collisions

- Kinetic energy of nuclei \sim 100-2500 times mass of nuclei

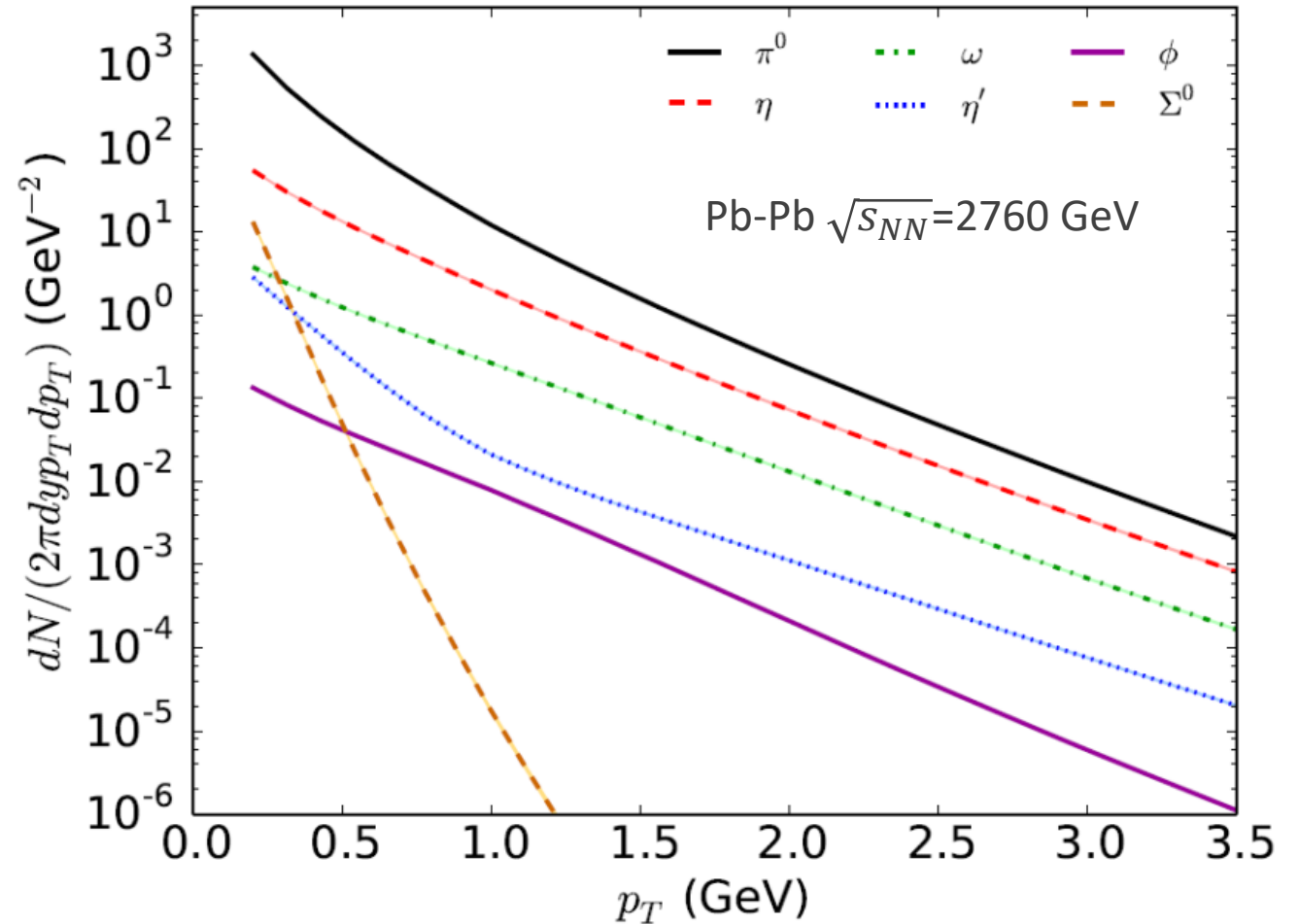


Pb-Pb at rest: $\sqrt{s_{NN}} \approx 2 \text{ GeV}$

Hadronic decay photons in nuclear collisions

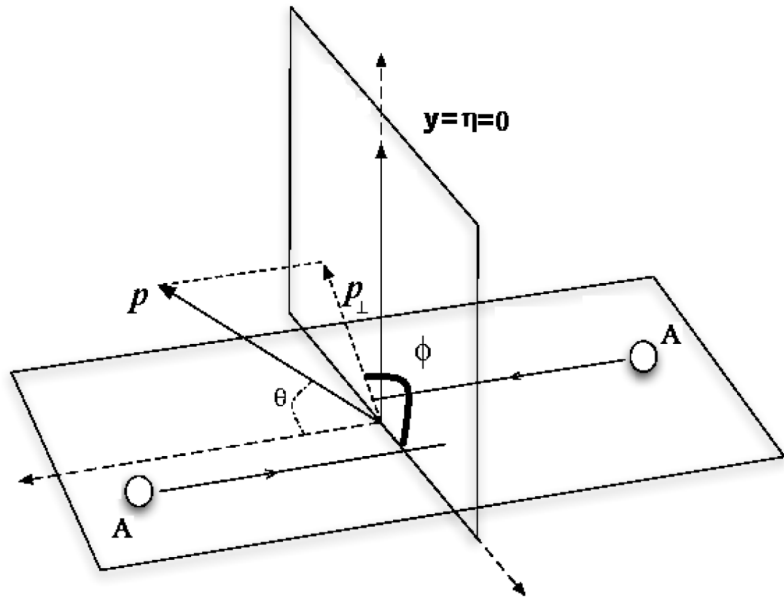
particle	mass (MeV)	decay	BR
π^0	134.98	$\gamma\gamma$	$9.882 \cdot 10^{-1}$
		$e^+e^-\gamma$	$1.174 \cdot 10^{-2}$
η	547.85	$\gamma\gamma$	$3.941 \cdot 10^{-1}$
		$\pi^0\gamma\gamma$	$2.560 \cdot 10^{-4}$
		$\pi^+\pi^-\gamma$	$4.220 \cdot 10^{-2}$
		$e^+e^-\gamma$	$6.899 \cdot 10^{-3}$
		$\mu^+\mu^-\gamma$	$3.090 \cdot 10^{-4}$
η'	957.66	$\rho^0\gamma$	$2.908 \cdot 10^{-1}$
		$\omega\gamma$	$2.746 \cdot 10^{-2}$
		$\gamma\gamma$	$2.198 \cdot 10^{-2}$
		$\mu^+\mu^-\gamma$	$1.080 \cdot 10^{-4}$
ω	782.65	$\pi^0\gamma$	$8.350 \cdot 10^{-2}$
		$\eta\gamma$	$4.600 \cdot 10^{-4}$
		$\pi^0\pi^0\gamma$	$7.000 \cdot 10^{-5}$
ρ^0	775.49	$\pi^+\pi^-\gamma$	$9.900 \cdot 10^{-3}$
		$\pi^0\gamma$	$6.000 \cdot 10^{-4}$
		$\eta\gamma$	$3.000 \cdot 10^{-4}$
		$\pi^0\pi^0\gamma$	$4.500 \cdot 10^{-5}$

Ref: F. Bock, PhD thesis

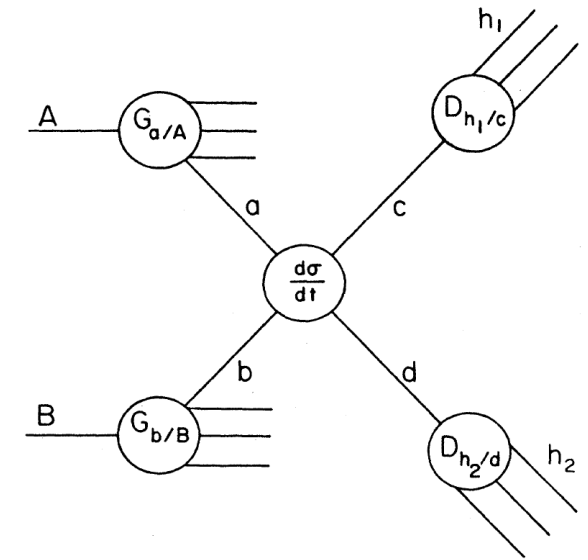


Ref: Chun Shen, PhD thesis

Figure adapted from K. Tuchin (2013) AHEP

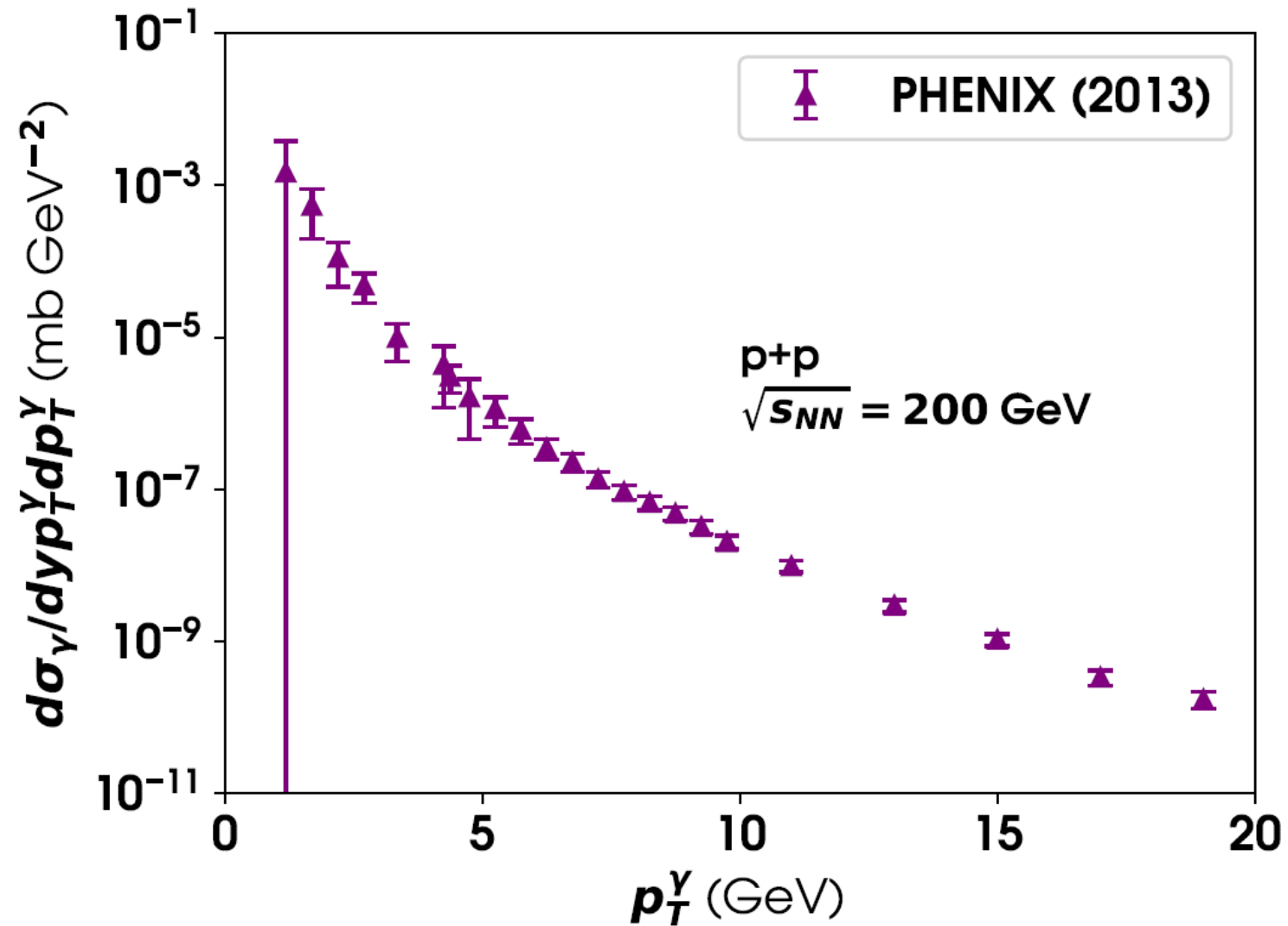


PROTON-PROTON COLLISIONS



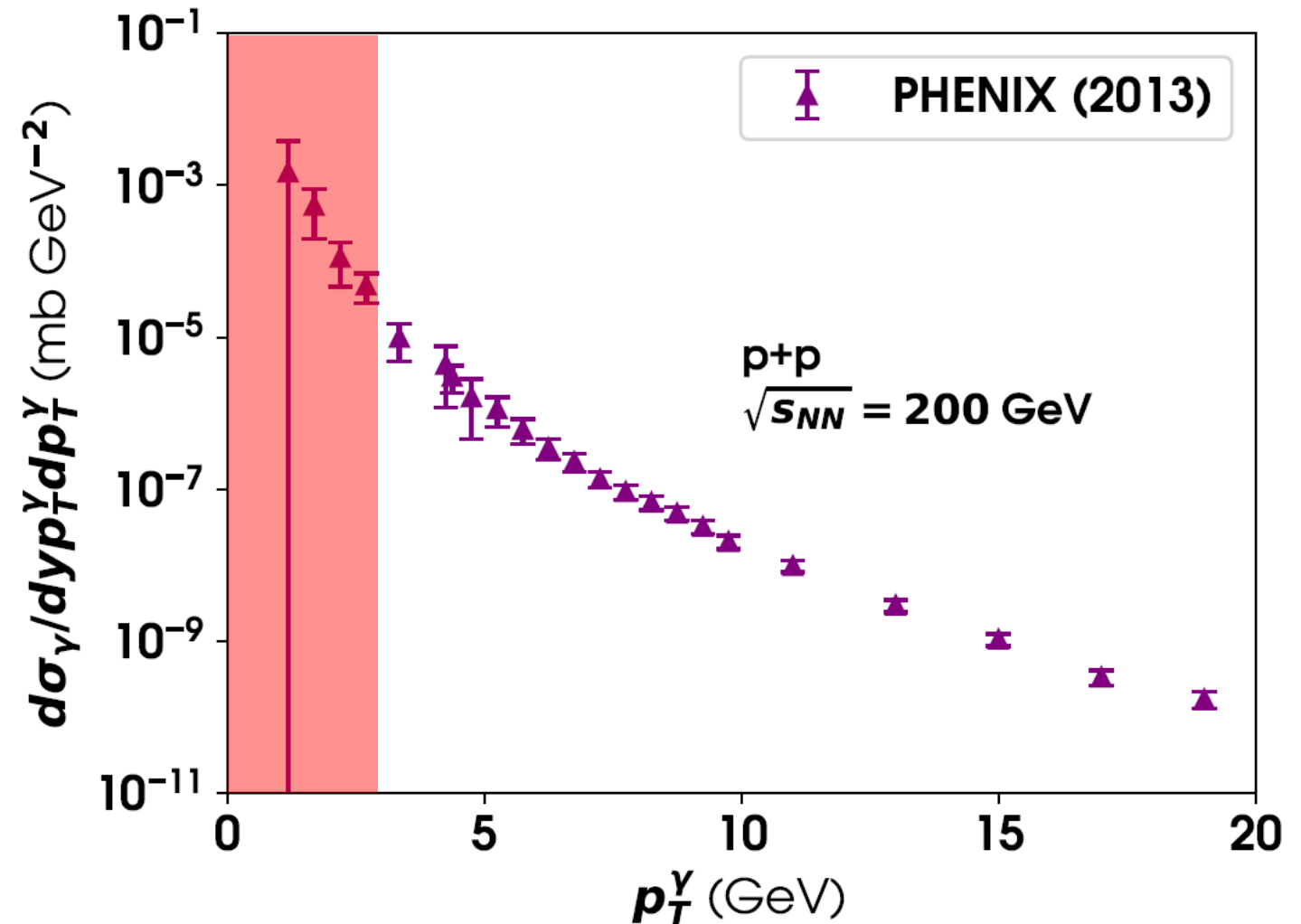
Ref: Owens (1987) RMP

Direct (non-decay) photons in proton-proton collisions



Direct photons in proton-proton collisions: “low” energy

- Low p_T photons:
 - Few measurements (in proton-proton collisions)
 - Difficult to compute from first principles
 - Non-perturbative effects likely significant

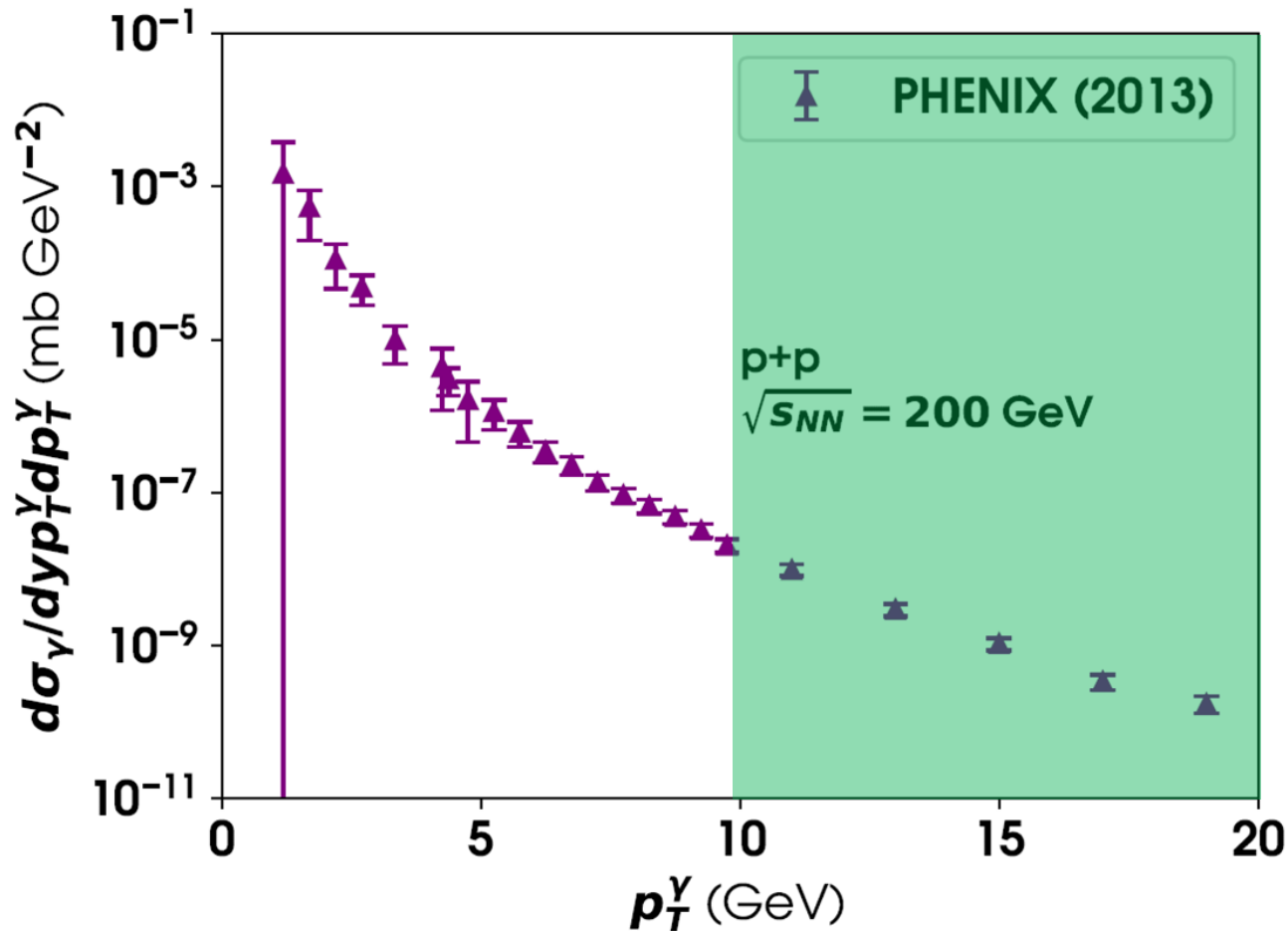


Direct photons in p-p collisions: high energy

Nuclear Physics B327 (1989) 105–143
North-Holland, Amsterdam

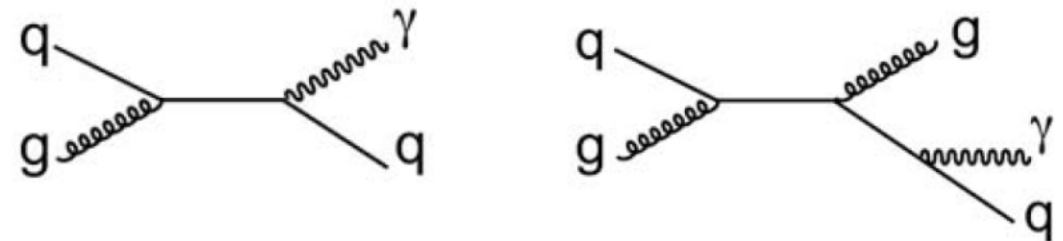
QCD CORRECTIONS TO PARTON-PARTON SCATTERING PROCESSES

F. AVERSA*, P. CHIAPPETTA, M. GRECO*, J.Ph. GUILLET**



- Can be calculated in collinear-factorization based perturbative QCD, up to next-to-leading order

$$\frac{d\sigma_{\gamma}^{pp}}{dp_T} = f_a \otimes f_b \otimes d\hat{\sigma}_{ab \rightarrow \gamma/c+d} [\otimes D_{\gamma/c}]$$



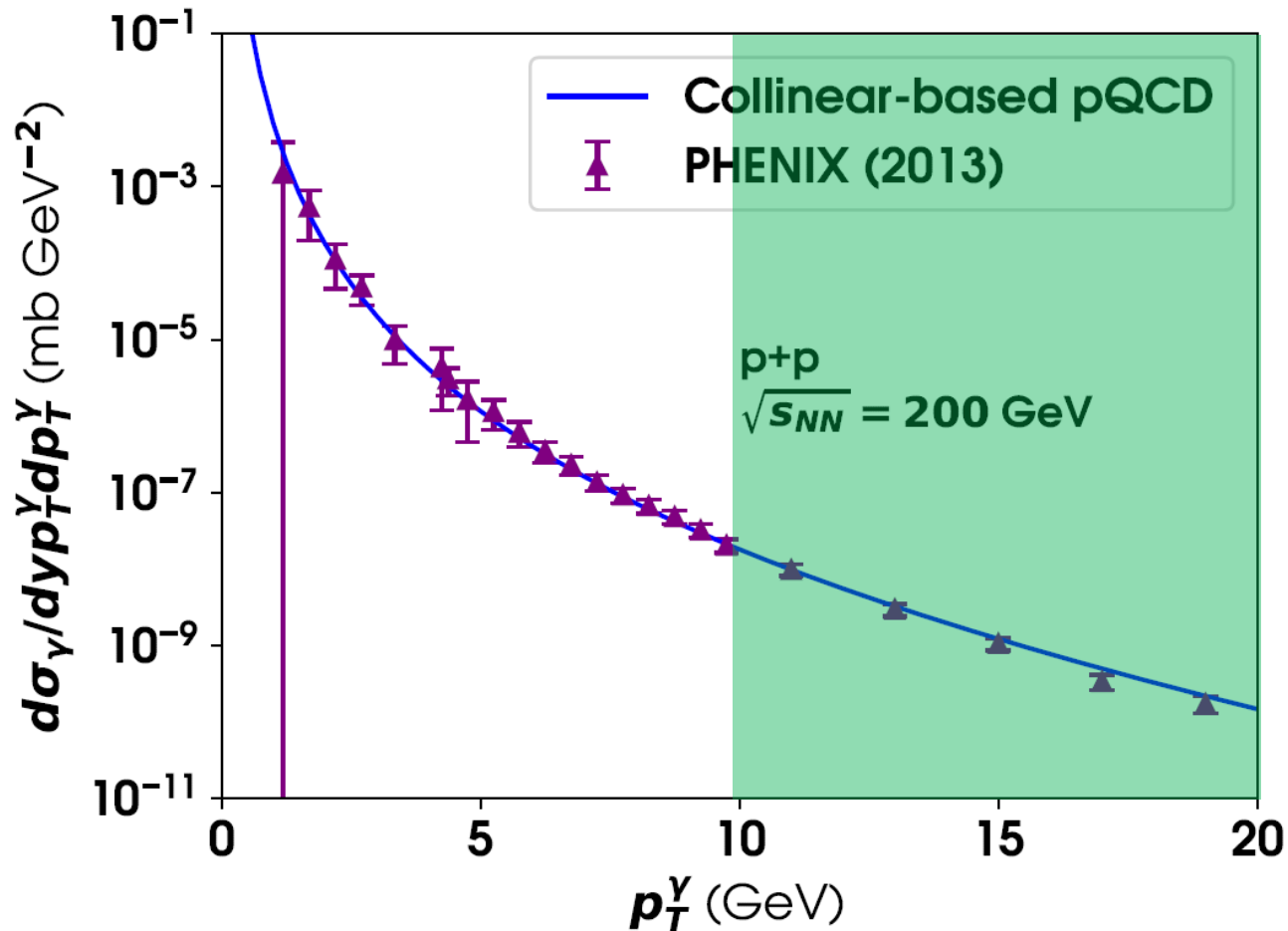
Frag fct: Bourhis, Fontannaz, Guillet (1998) EPJ

Direct photons in p-p collisions: high energy

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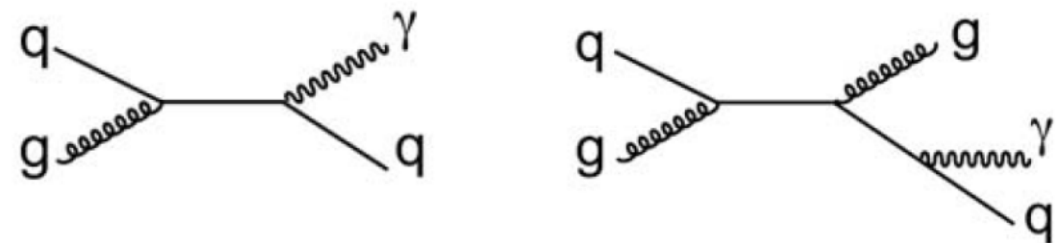
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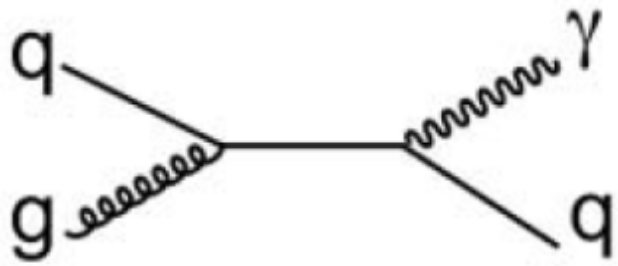
$$\frac{d\sigma_{\gamma}^{pp}}{dp_T} = f_a \otimes f_b \otimes d\hat{\sigma}_{ab \rightarrow \gamma/c+d} [\otimes D_{\gamma/c}]$$



Frag fct: Bourhis, Fontannaz, Guillet (1998) EPJ

Direct photons in proton-proton collisions: channels

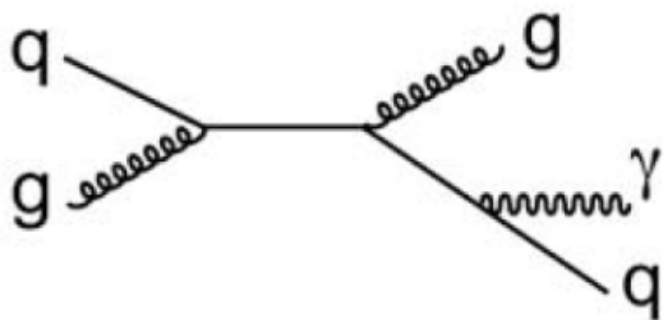
- Hard partonic collisions
 - “Isolated”



(Can be
calculated at
NNLO)

$$d\sigma_{\gamma}^{pp}/dp_T = f_a \otimes f_b \otimes d\hat{\sigma}_{ab \rightarrow \gamma/c+d}$$

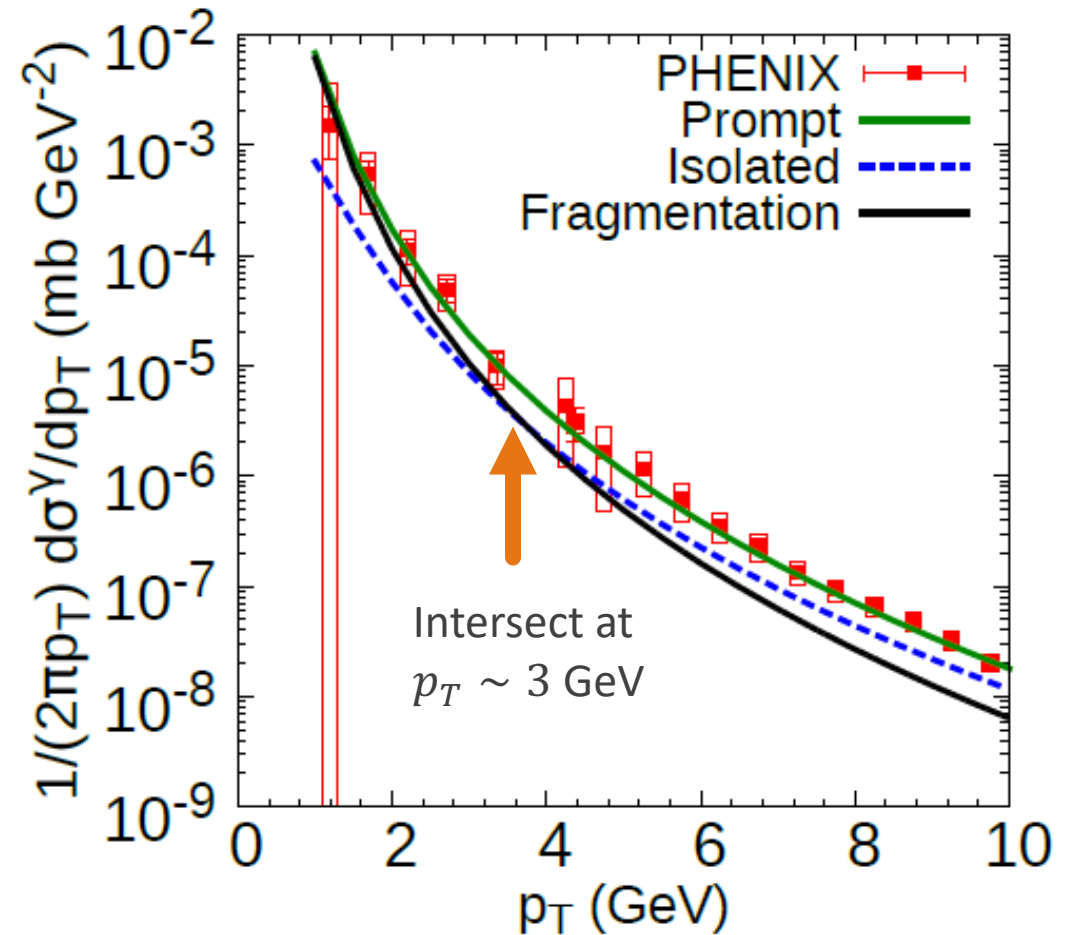
- Fragmentation



(Fragmentation
function
unmeasured at
NNLO and poorly
constrained at NLO)

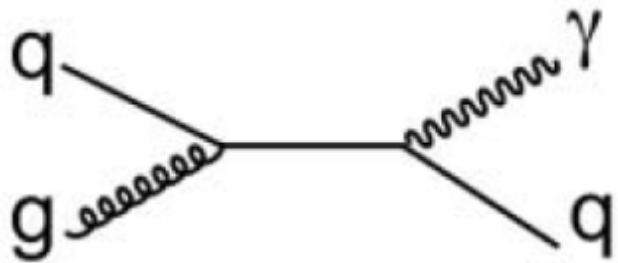
$$d\sigma_{\gamma}^{pp}/dp_T = f_a \otimes f_b \otimes d\hat{\sigma}_{ab \rightarrow \gamma/c+d} \otimes D_{\gamma/c}$$

p+p $\sqrt{s}=0.2$ TeV



Direct photons in proton-proton collisions: channels

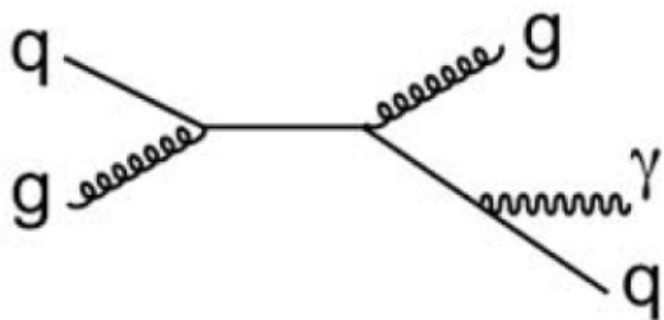
- Hard partonic collisions
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(Can be calculated at NNLO)

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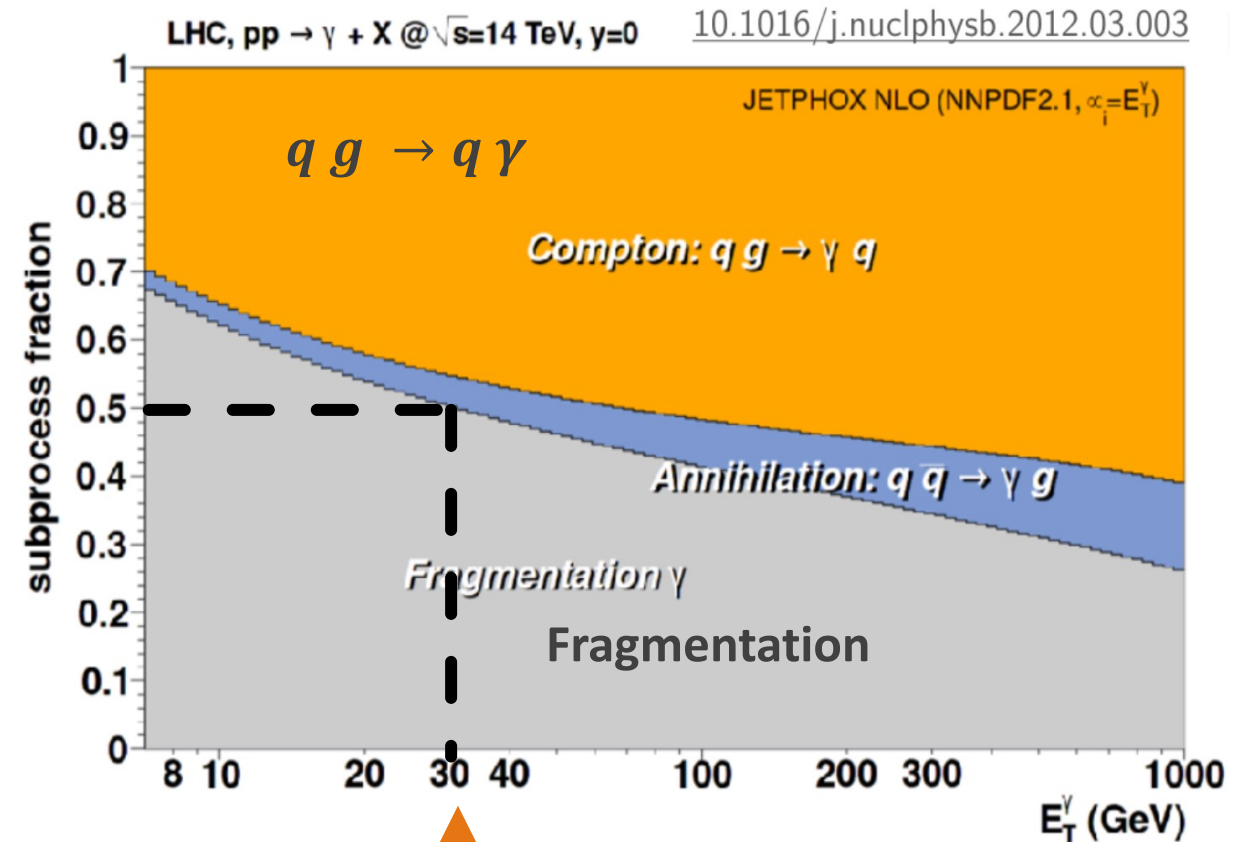
- Fragmentation



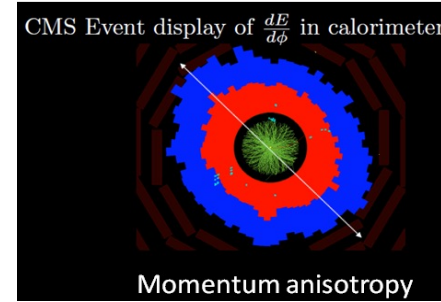
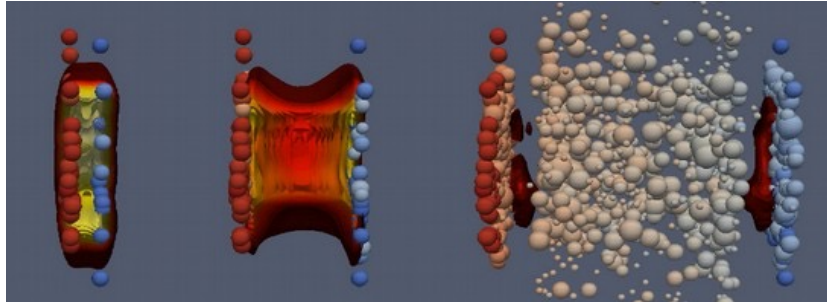
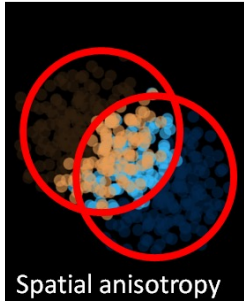
(Fragmentation function unmeasured at NNLO and poorly constrained at NLO)

$$d\sigma_{\gamma}^{pp}/dp_T = f_a \otimes f_b \otimes d\hat{\sigma}_{ab \rightarrow \gamma/c+d} \otimes D_{\gamma/c}$$

p+p $\sqrt{s}=14$ TeV

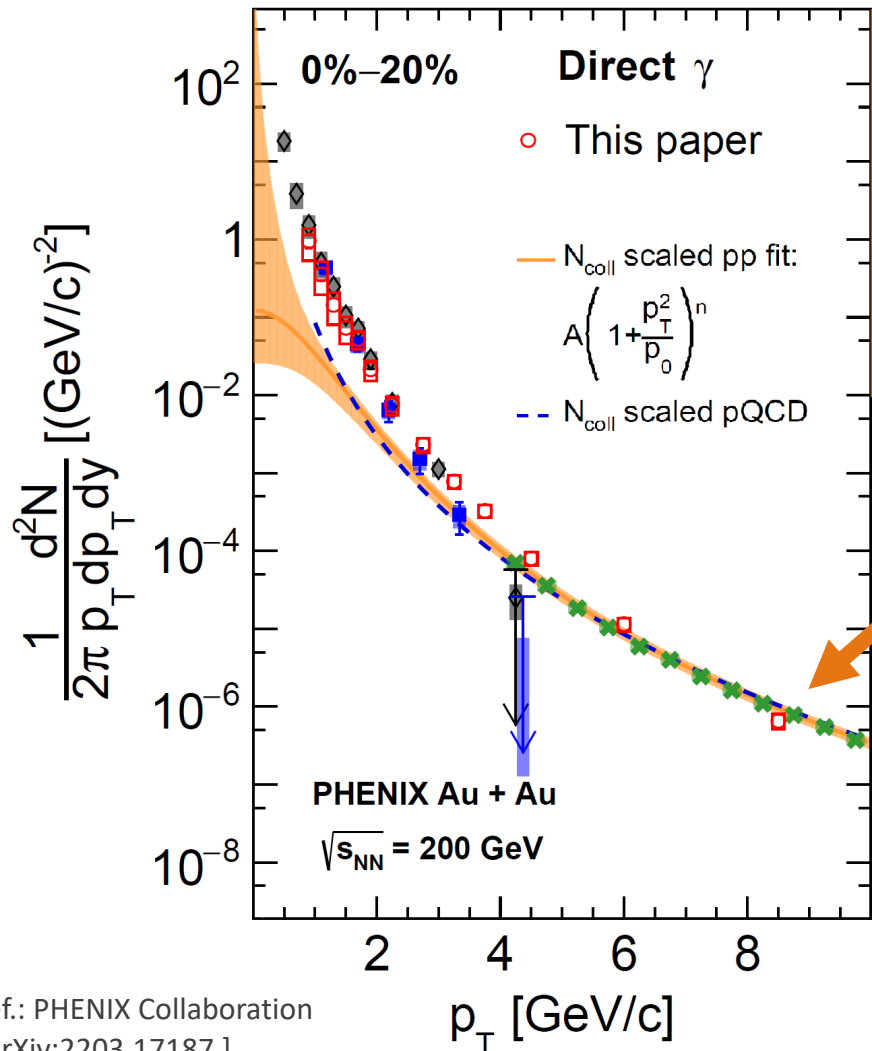


Intersect at $p_T \sim 30$ GeV



HEAVY-ION COLLISIONS: PHOTONS FROM “PROTON+PROTON-LIKE” MECHANISMS

Photon energy spectrum in heavy-ion collisions



- Systematic excess of low energy photons in nucleus collisions

(also observed by STAR [RHIC] and ALICE [LHC] Collaborations)

Orange band:
 Result for incoherent
 superposition of proton-
 proton collisions

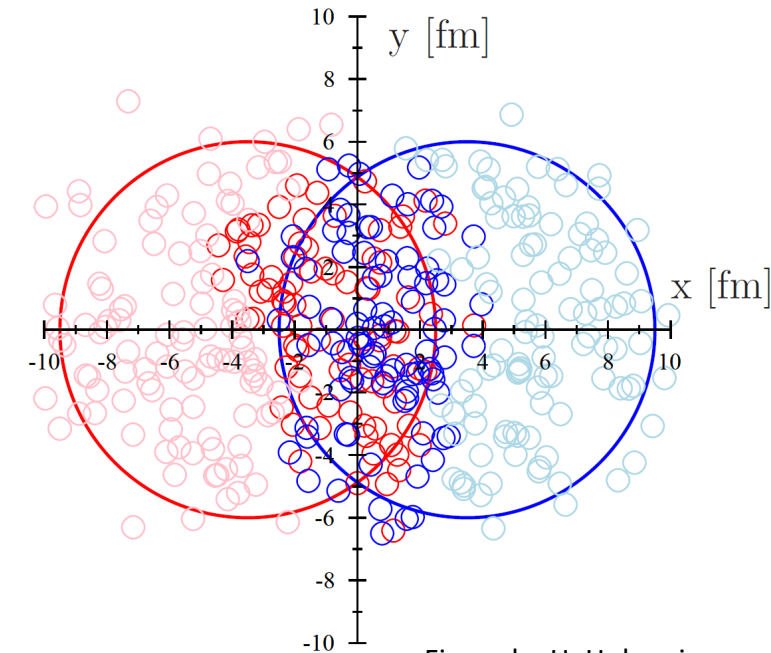
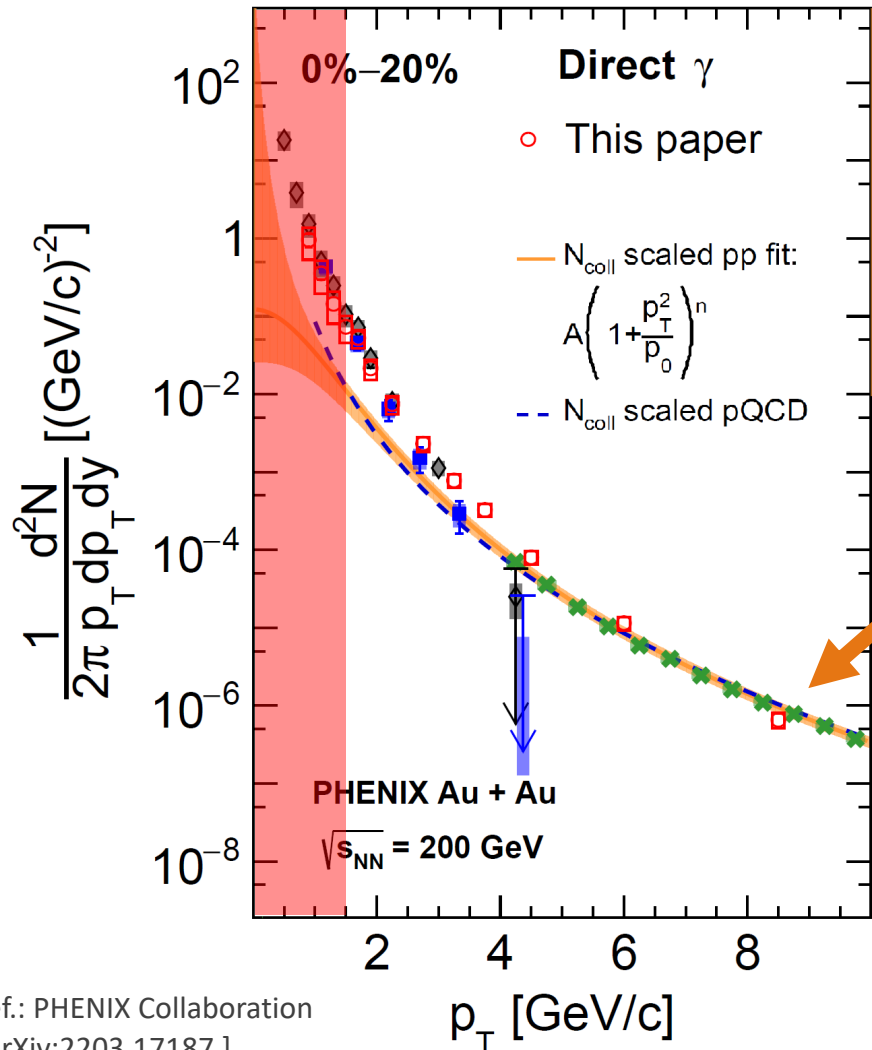


Figure by H. Holopainen

Photon energy spectrum in heavy-ion collisions



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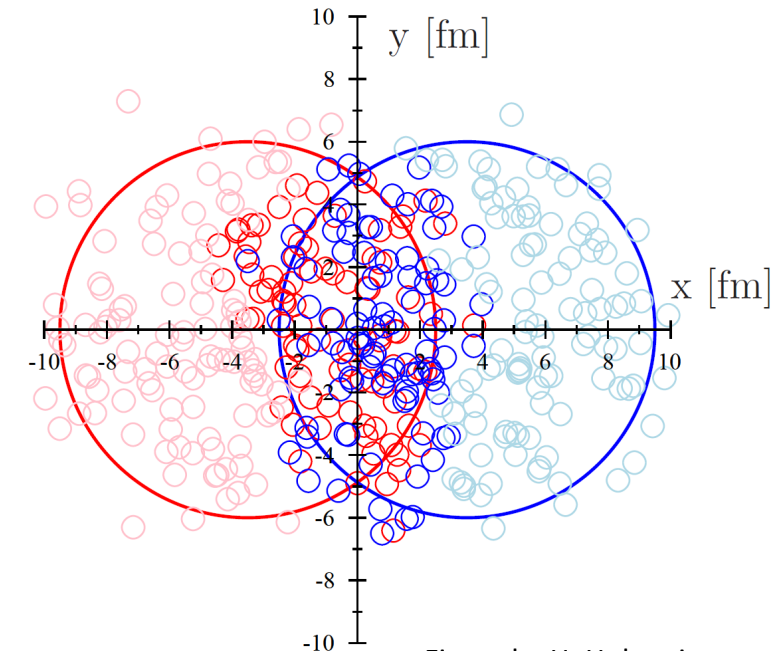
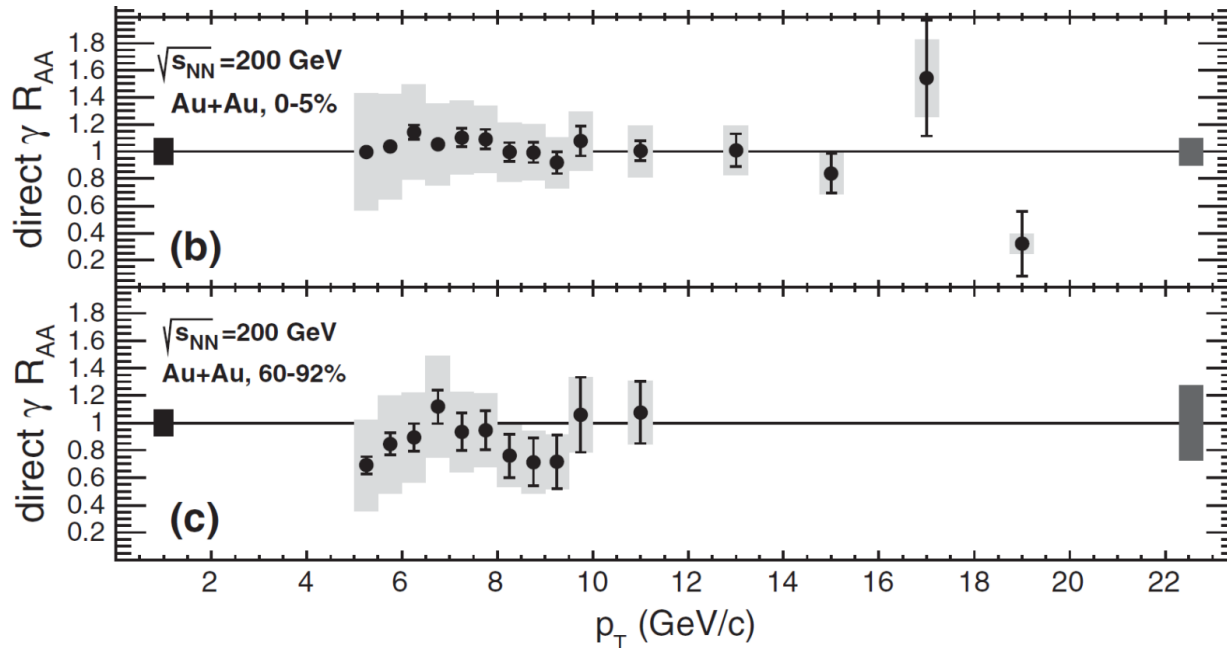


Figure by H. Holopainen

Photons in heavy-ion collisions: high p_T

- **Prompt photons** produced as superposition of nucleon-nucleon collisions (“binary scaling”)

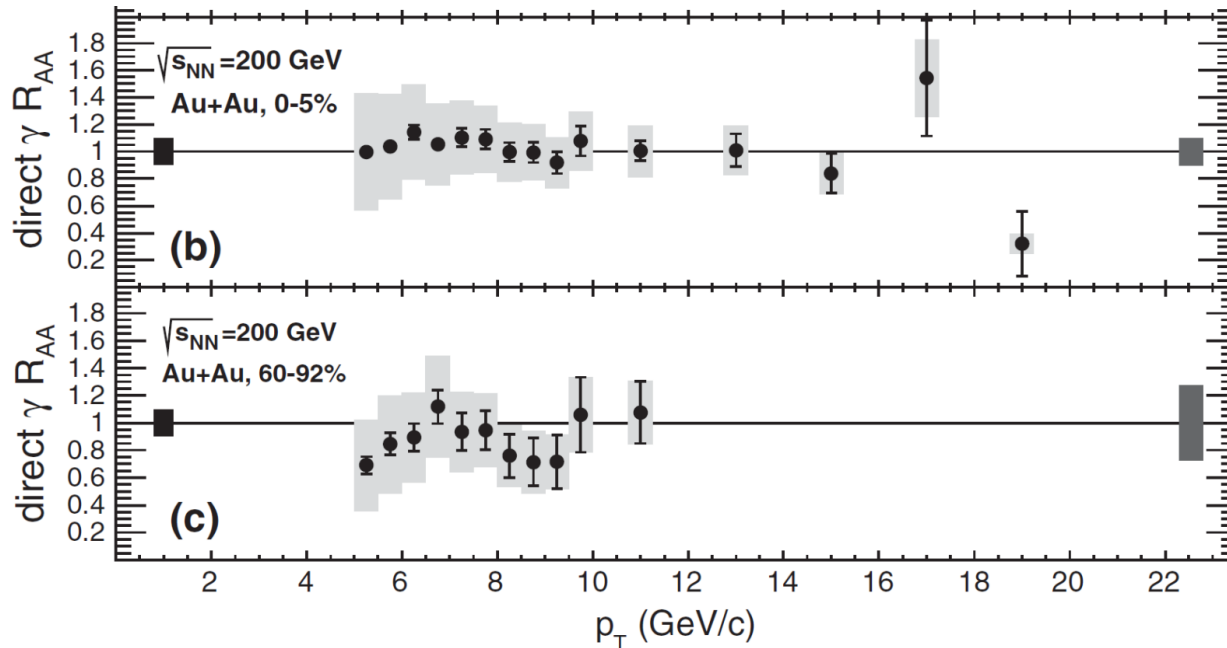


Ref.: PHENIX Collaboration (2012) PRL

$$R_{AA}^{\gamma} = \frac{\frac{dN_{\gamma}^{AA}}{dp_T}}{\left(\frac{N_{binary}}{\sigma_{pp}^{inel}}\right) \frac{d\sigma_{\gamma}^{pp}}{dp_T}} \approx 1 \quad (\text{at high } p_T)$$

Photons in heavy-ion collisions: high p_T

- **Prompt photons** produced as superposition of nucleon-nucleon collisions (“binary scaling”)



Ref.: PHENIX Collaboration (2012) PRL

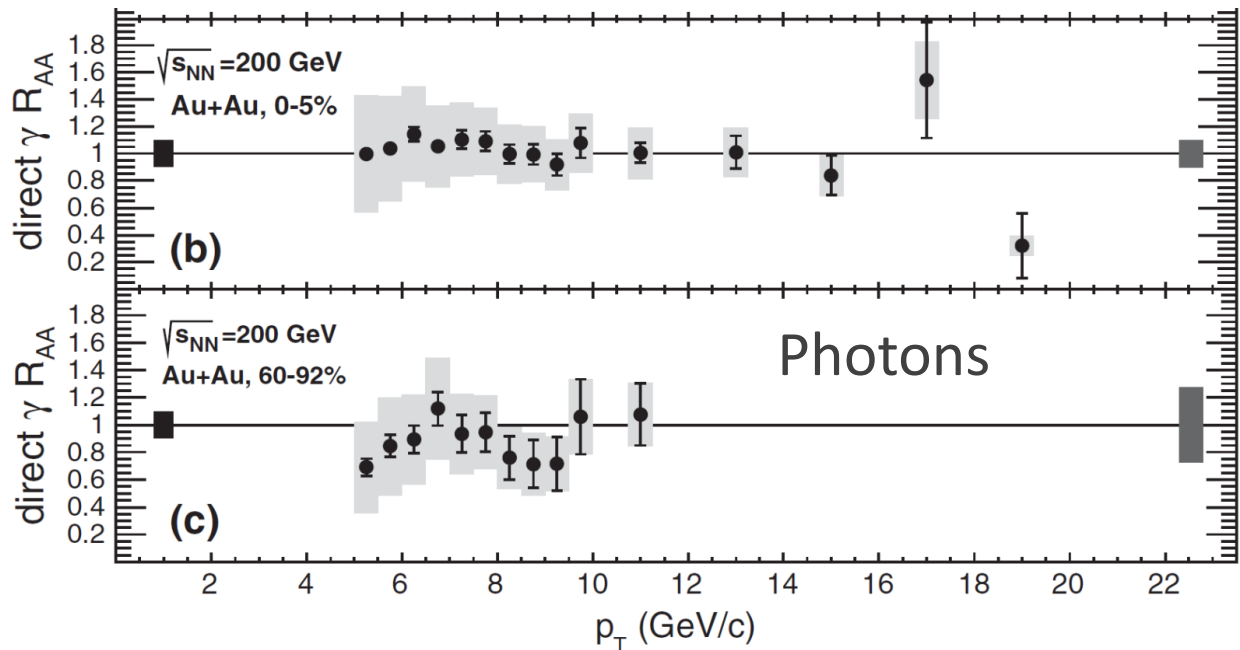
$$R_{AA}^{\gamma} = \frac{\frac{dN_{\gamma}^{AA}}{dp_T}}{\left(\frac{N_{binary}}{\sigma_{pp}^{inel}}\right) \frac{d\sigma_{\gamma}^{pp}}{dp_T}} \approx 1 \quad (\text{at high } p_T)$$

Deviations from $R_{AA}^{\gamma} = 1$ originate from:

- Isospin effect (parton content of neutrons vs protons)
- Nuclear effects on parton distribution functions
- Parton energy loss [more about this later]

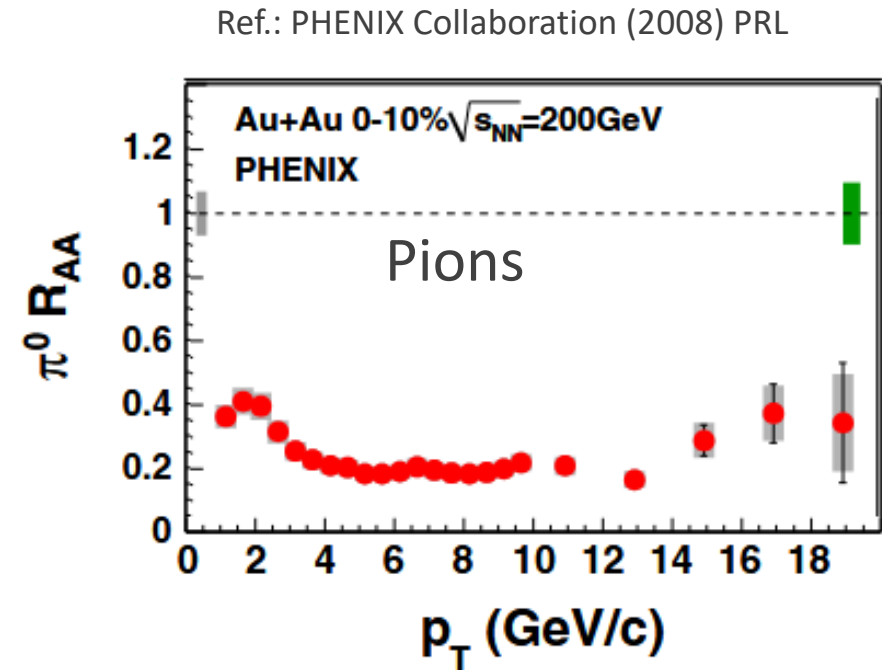
Photons vs π^0 in heavy-ion collisions

- **Prompt photons** produced as superposition of nucleon-nucleon collisions (“binary scaling”)



Ref.: PHENIX Collaboration (2012) PRL

$$R_{AA}^{\gamma} = \frac{\frac{dN_{\gamma}^A}{dp_T}}{\left(\frac{N_{binary}}{\sigma_{pp}^{inel}}\right) \frac{d\sigma_{\gamma}^{pp}}{dp_T}} \approx 1$$



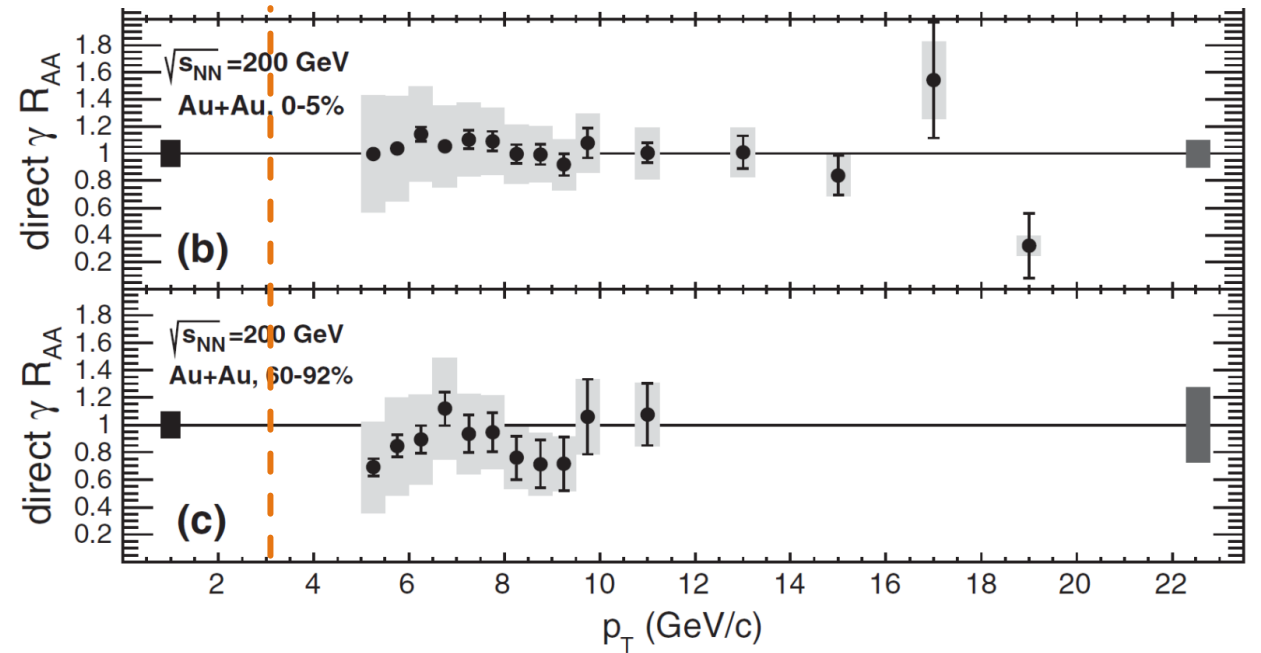
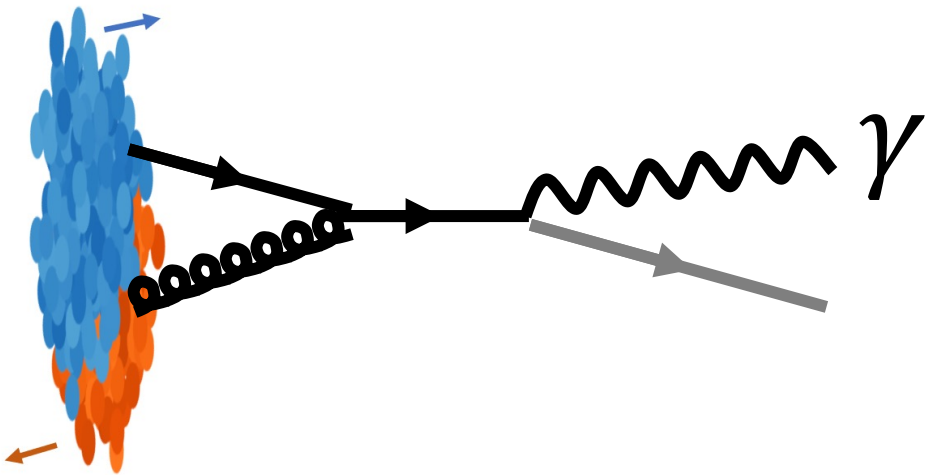
Ref.: PHENIX Collaboration (2008) PRL

Isolated photons in heavy-ion collisions

$$q + \bar{q} \rightarrow g + \gamma$$

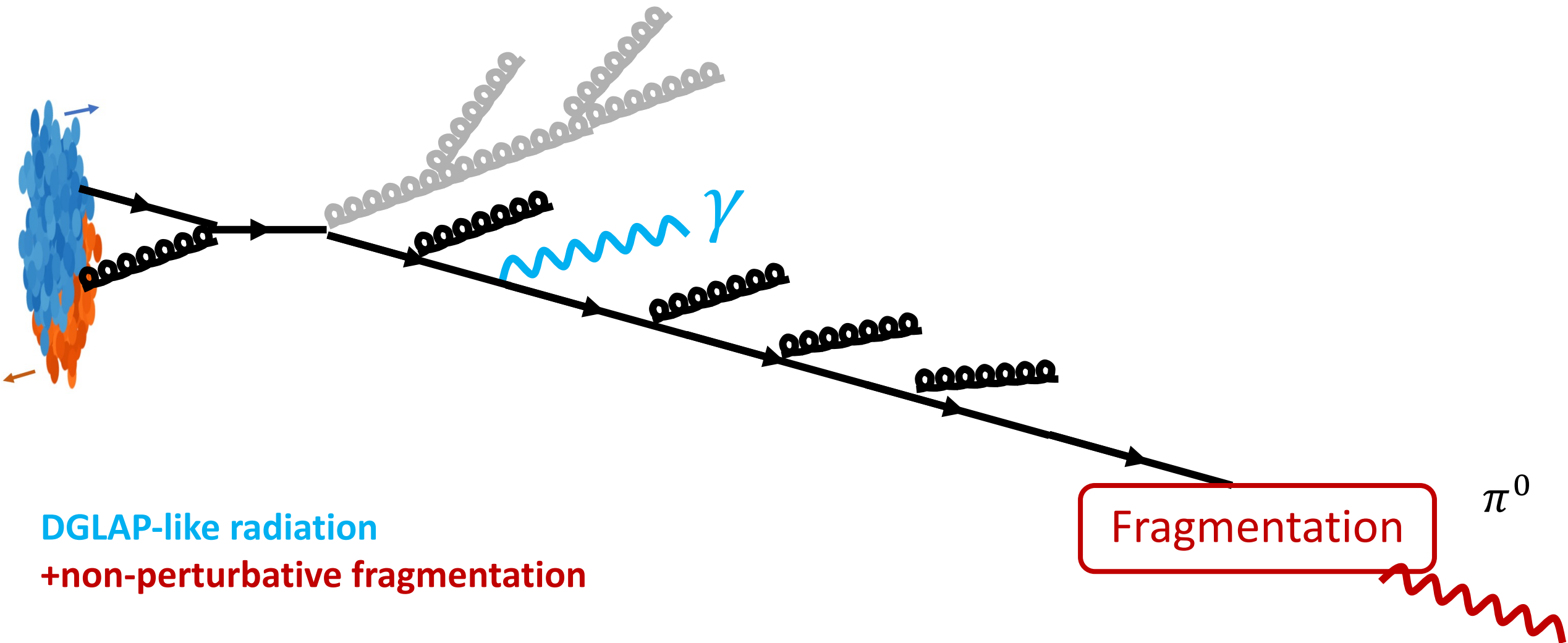
$$q + g \rightarrow q + \gamma$$

$$q + g \rightarrow q + g + \gamma ?$$

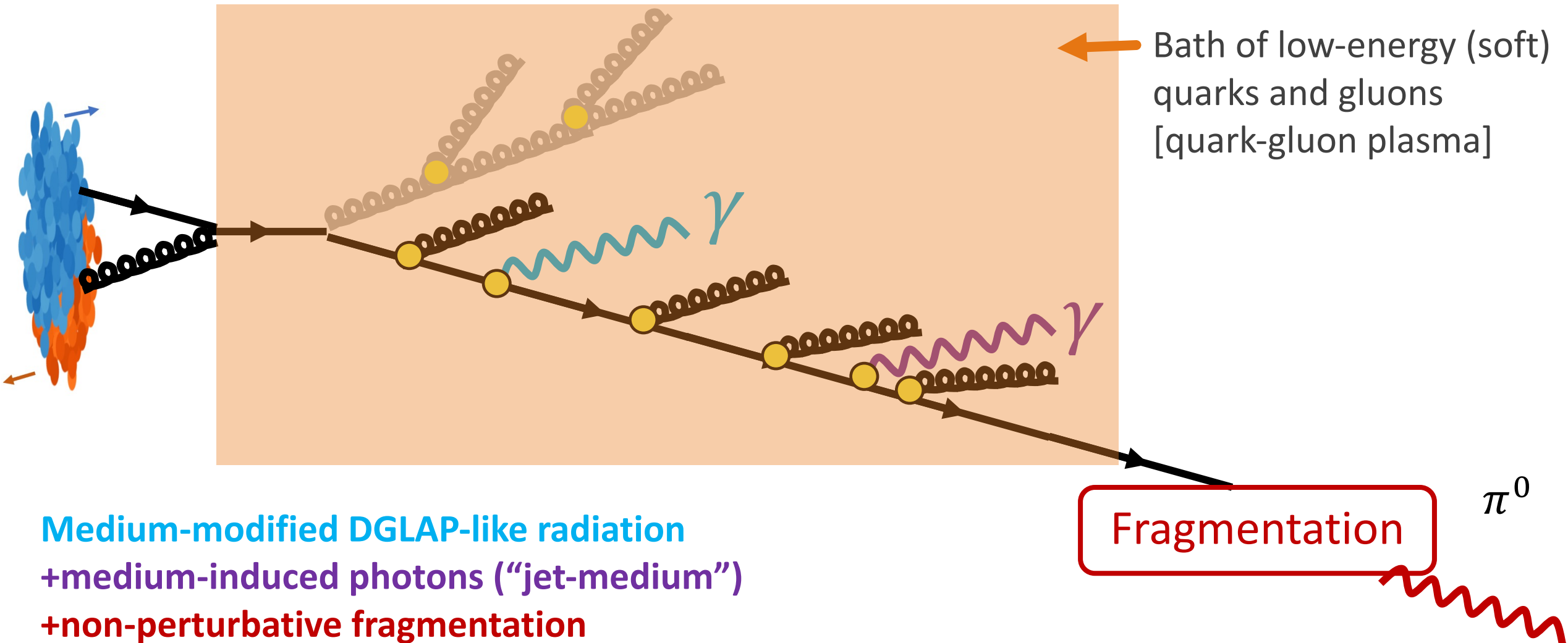


Ref.: PHENIX Collaboration (2012) PRL

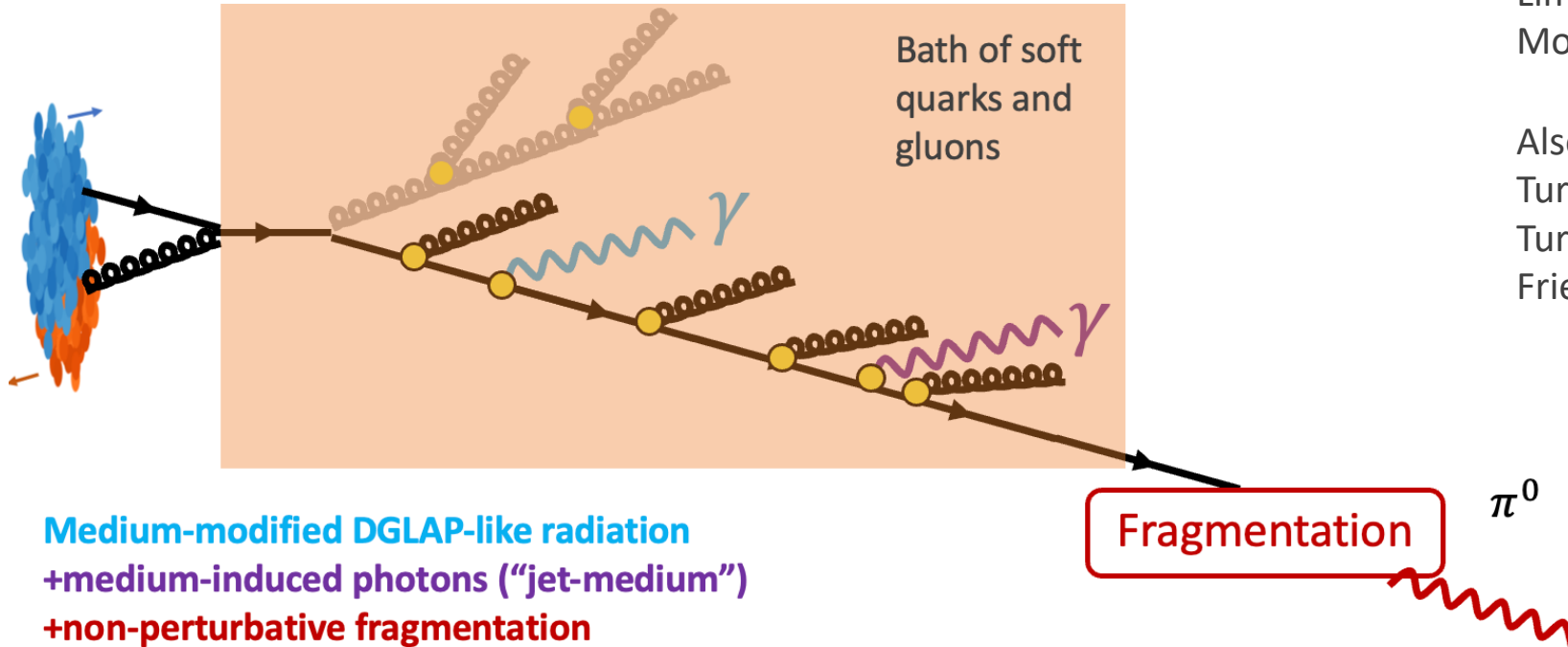
Fragmentation photons and pions in $p+p$ collisions



Fragmentation photons and pions in heavy-ion collisions



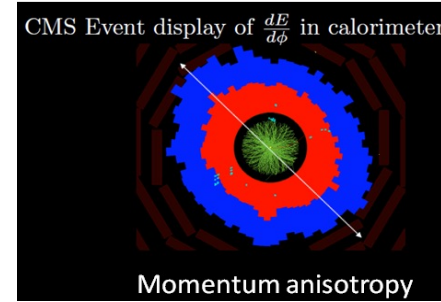
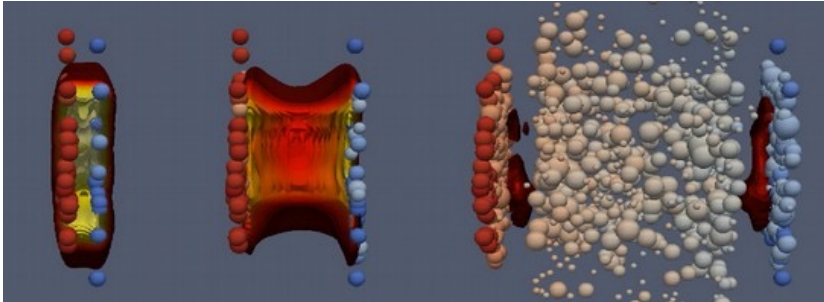
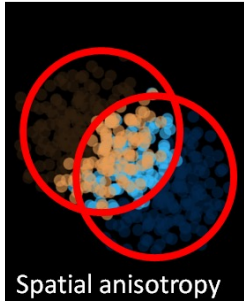
Fragmentation photons and pions in heavy-ion collisions



Limited number of recent studies;
 Modarresi Yazdi, Shi, Gale, Jeon [arXiv:2207.12513]

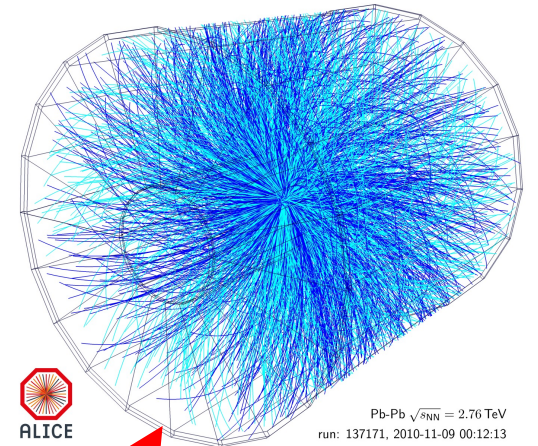
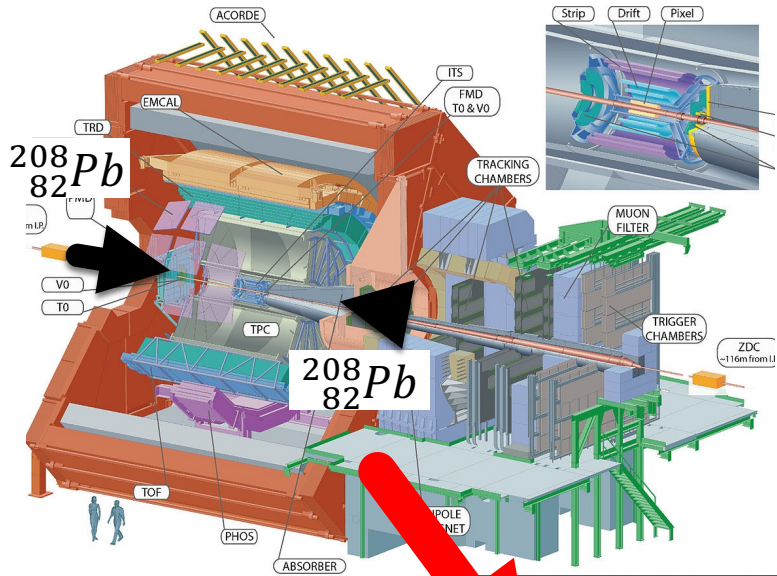
Also:
 Turbide, Gale, Frodermann & Heinz (2008) PRC;
 Turbide, Gale, Srivastava, Fries (2008) PRC;
 Fries, Müller, Srivastava (PRL) 2003

- Challenging because mixes:
 - Spacetime macroscopic description of the soft bath (plasma)
 - Momentum-space description of showers and fragmentation

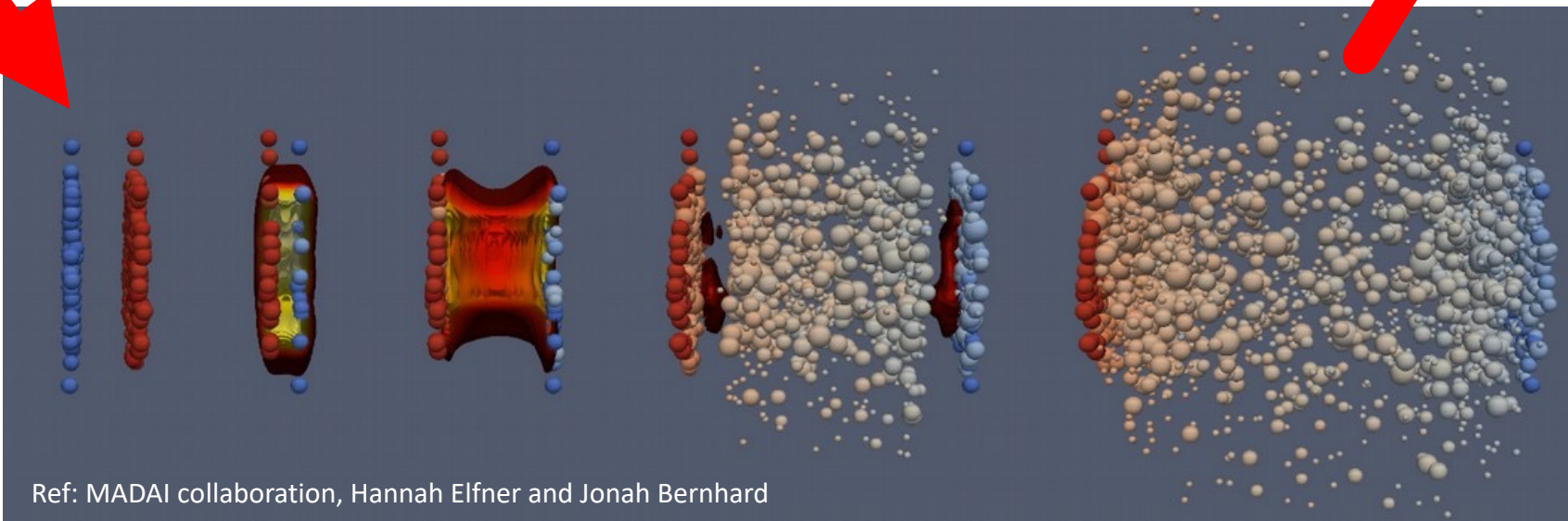


HEAVY-ION COLLISIONS: PHOTONS FROM THE SOFT BATH [QUARK-GLUON PLASMA]

Heavy-ion collisions

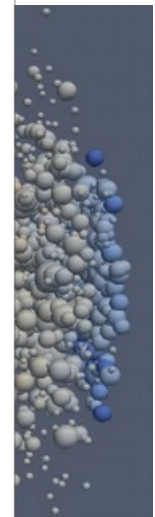
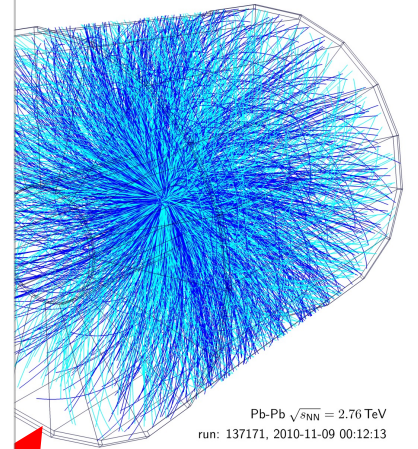
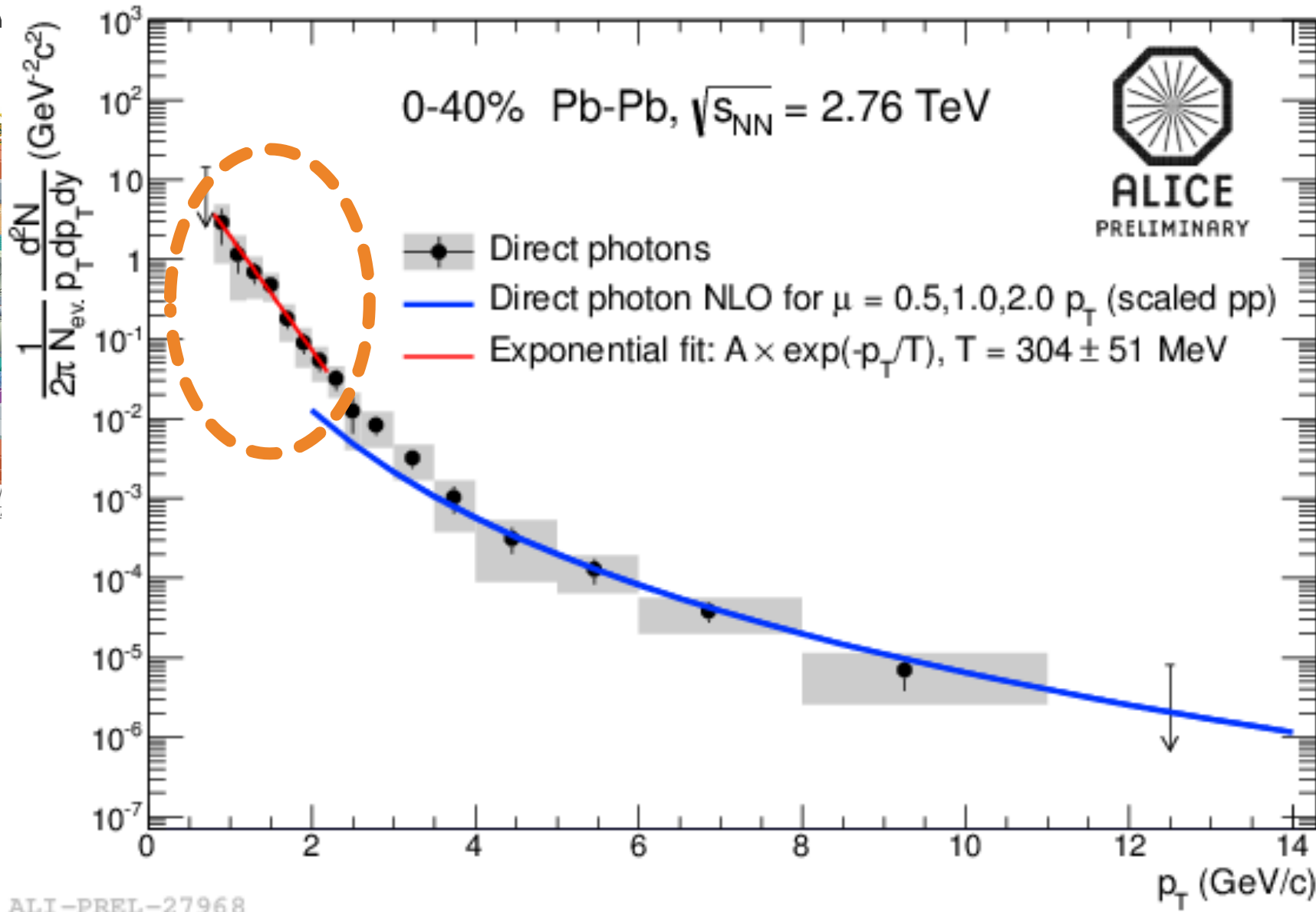
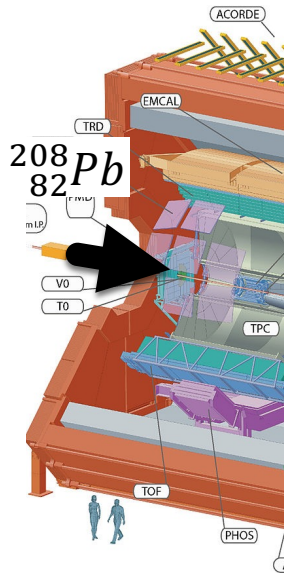


Ref.: ALICE, CERN



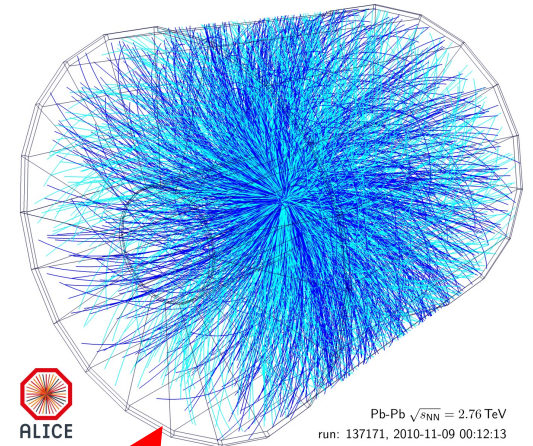
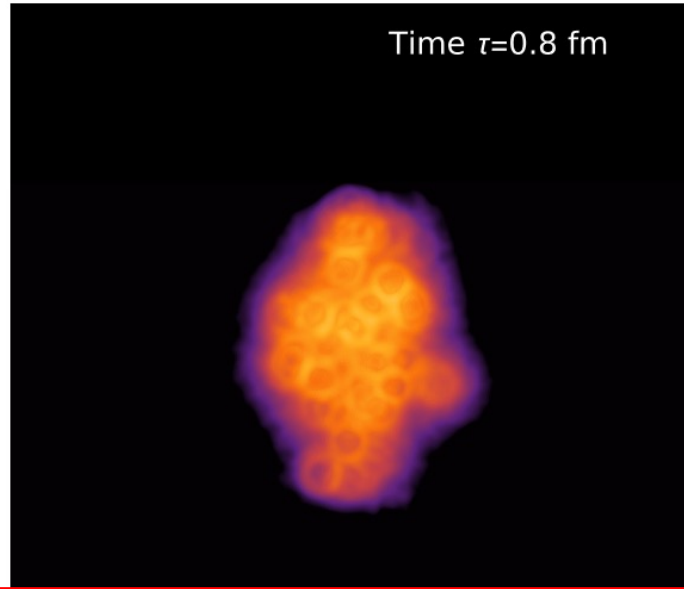
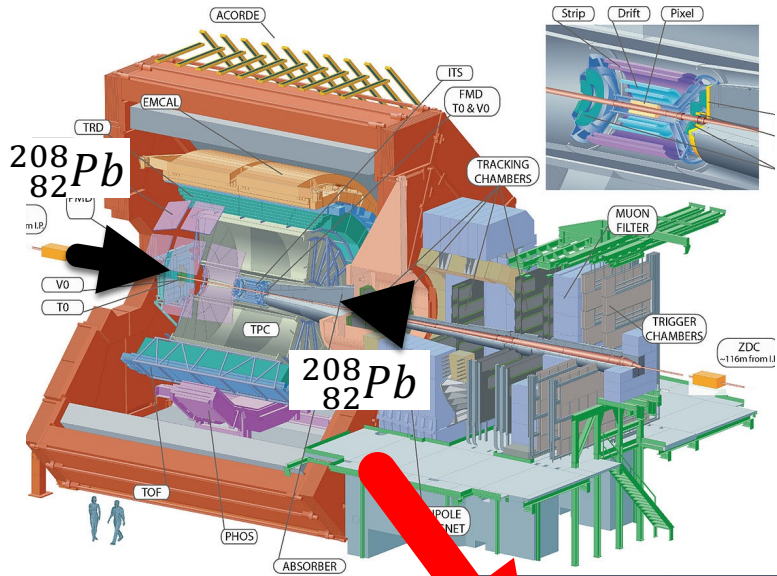
Ref: MADAI collaboration, Hannah Elfner and Jonah Bernhard

Heavy-ion collisions

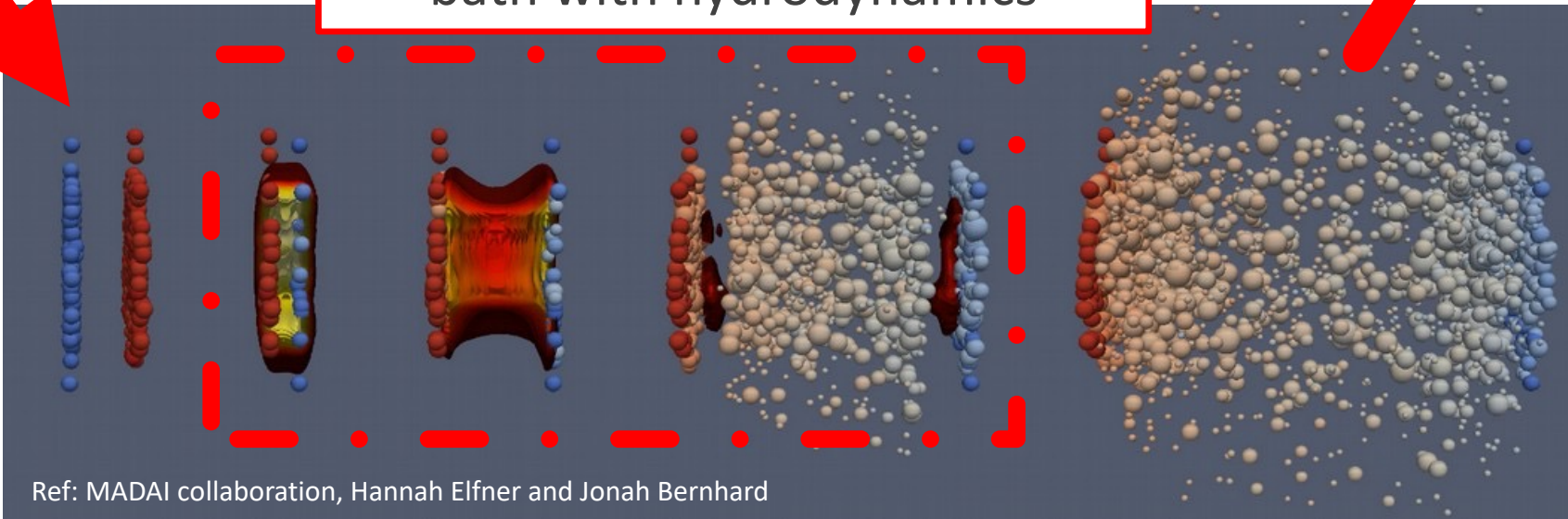


ALI-PREL-27968

Heavy-ion collisions



Macroscopic description of soft bath with hydrodynamics



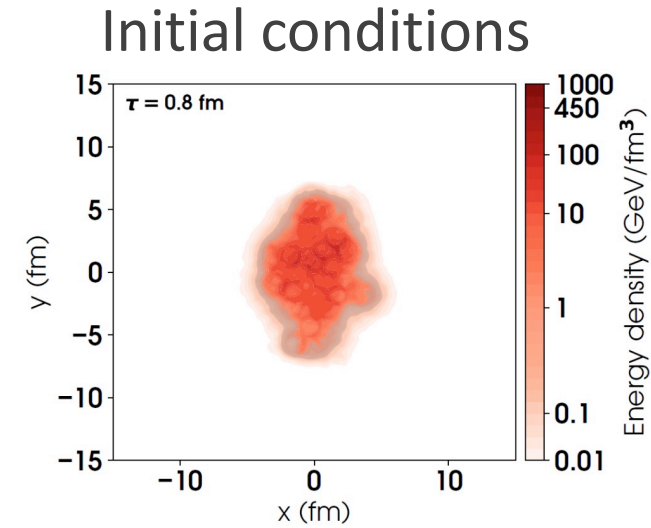
Ref.: ALICE, CERN

Relativistic hydrodynamics

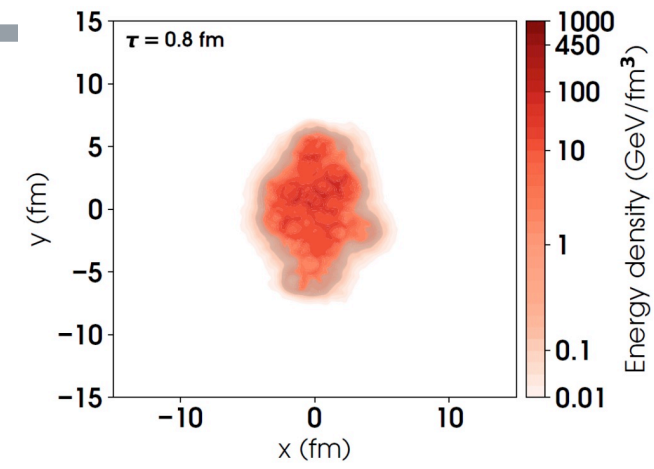
- Evolution of the energy-momentum tensor in space&time

$$T^{\mu\nu} = \epsilon u^\mu u^\nu - (P + \Pi)(g^{\mu\nu} - u^\mu u^\nu) + \pi^{\mu\nu}$$

- ϵ is the energy density
- u^μ is the flow velocity (Landau frame: $T^{\mu\nu} u_\nu = \epsilon u^\mu$)
- Π and $\pi^{\mu\nu}$ are viscous components



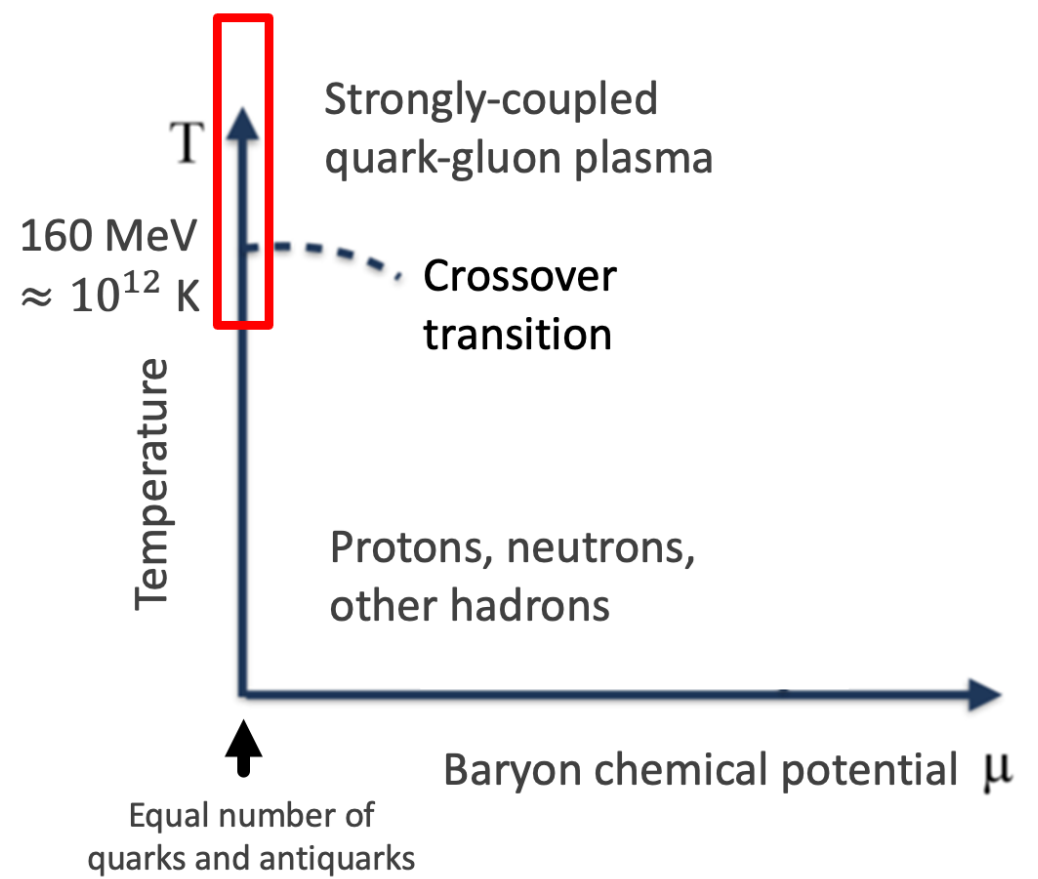
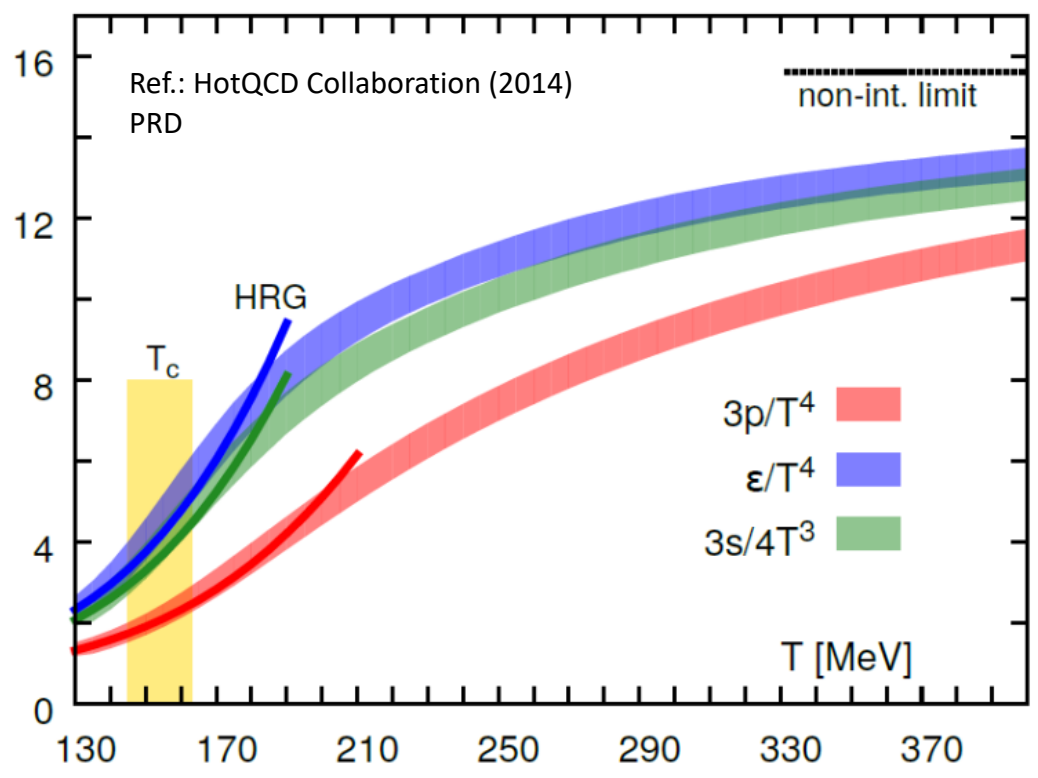
Initial conditions



- Evolution of the energy-momentum tensor in space&time

$$T^{\mu\nu} = \epsilon u^\mu u^\nu - (P + \Pi)(g^{\mu\nu} - u^\mu u^\nu) + \pi^{\mu\nu}$$

- First-principle equation of state



Viscous relativistic hydrodynamics

- Evolution of the energy-momentum tensor in space&time

$$T^{\mu\nu} = \epsilon u^\mu u^\nu - (P + \Pi)(g^{\mu\nu} - u^\mu u^\nu) + \pi^{\mu\nu}$$

- Conservation of energy and momentum:

$$\partial_\nu T^{\mu\nu} = 0$$

Mueller (1967) Zeit. fur Phys;
Israel&Stewart (1979) Ann. Phys.

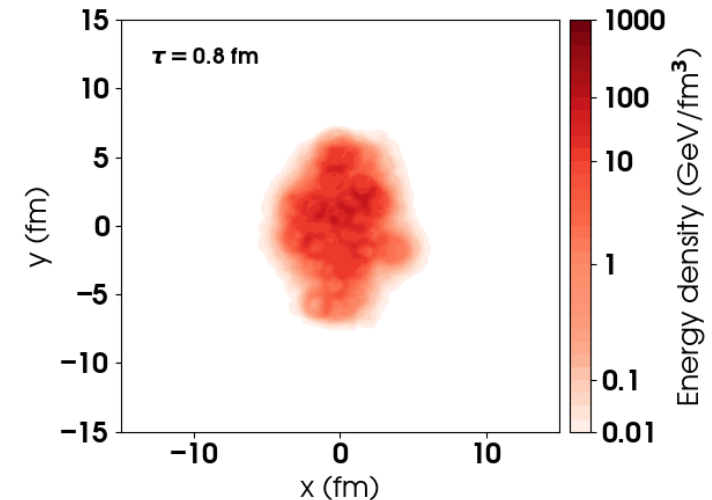
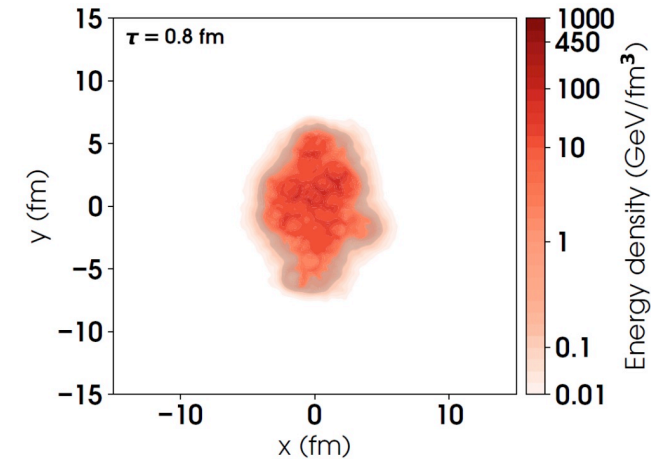
- Mueller-Israel-Stewart relativistic viscous hydrodynamics

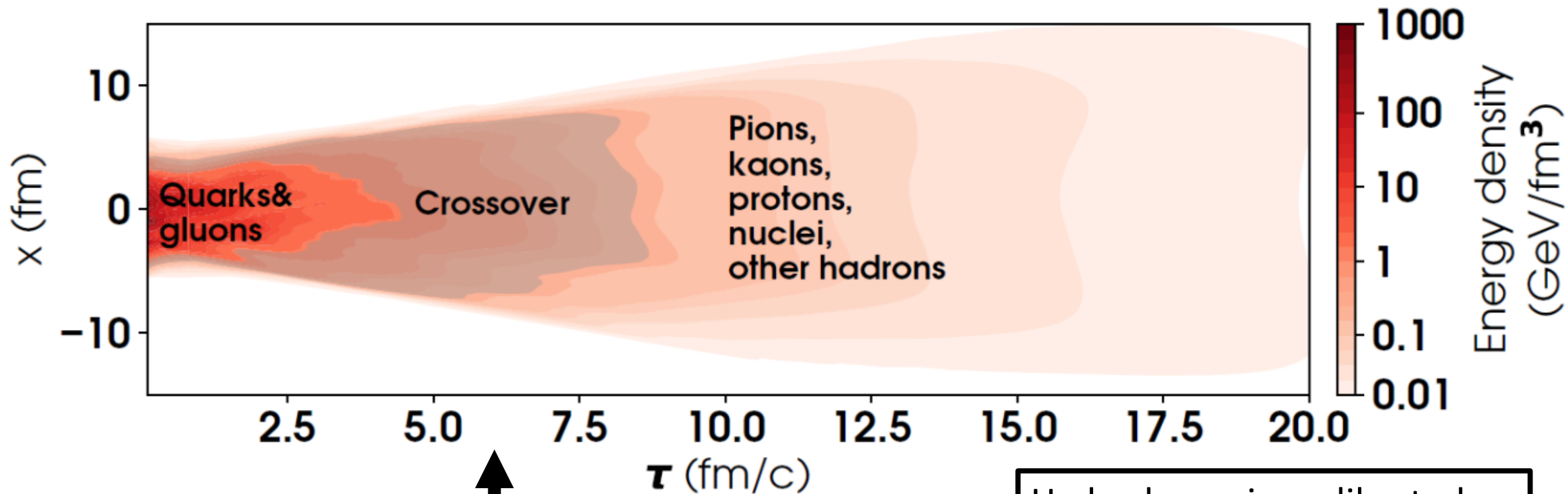
$$\tau_\Pi \dot{\Pi} + \boxed{\Pi = -\zeta \partial_\mu u^\mu} + (2^{\text{nd}} \text{ order terms})$$

$$\tau_\pi \Delta_{\alpha\beta}^{\mu\nu} \pi^{\alpha\beta} + \boxed{\pi^{\mu\nu} = 2\eta(\partial_\mu u^\nu + \dots)} + (2^{\text{nd}} \text{ order terms})$$

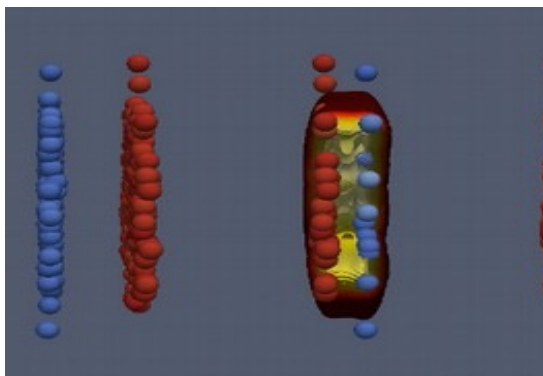
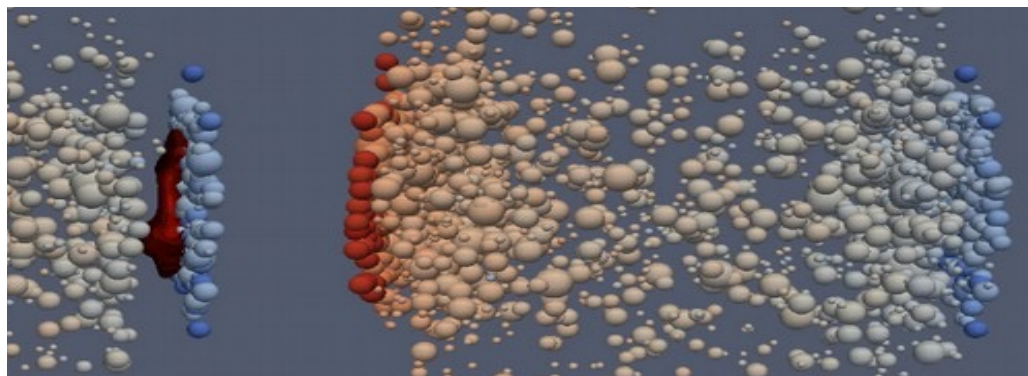
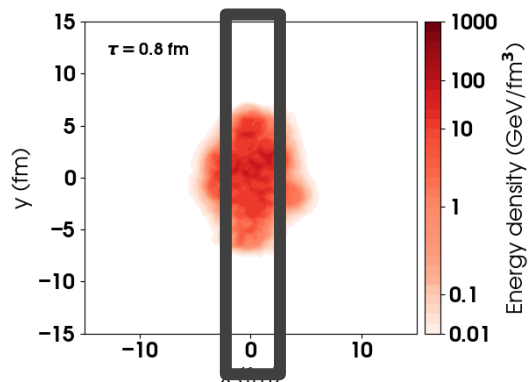
Solve hydrodynamics equations numerically (finite volume)

Initial conditions



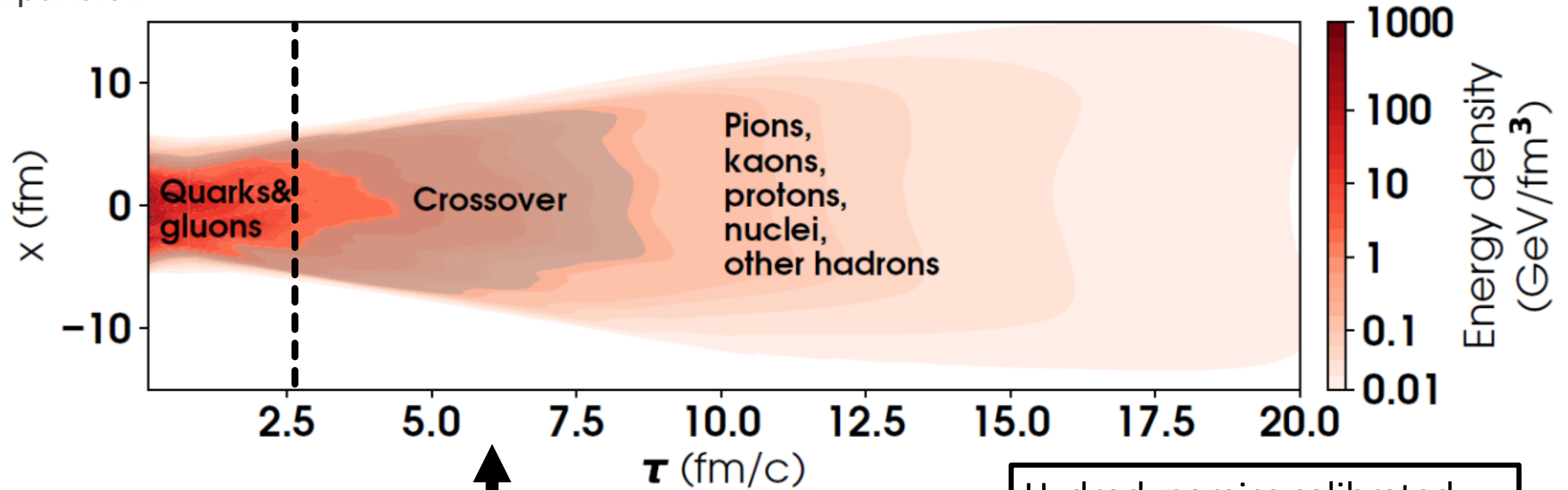


Hydrodynamics calibrated on hadronic measurements

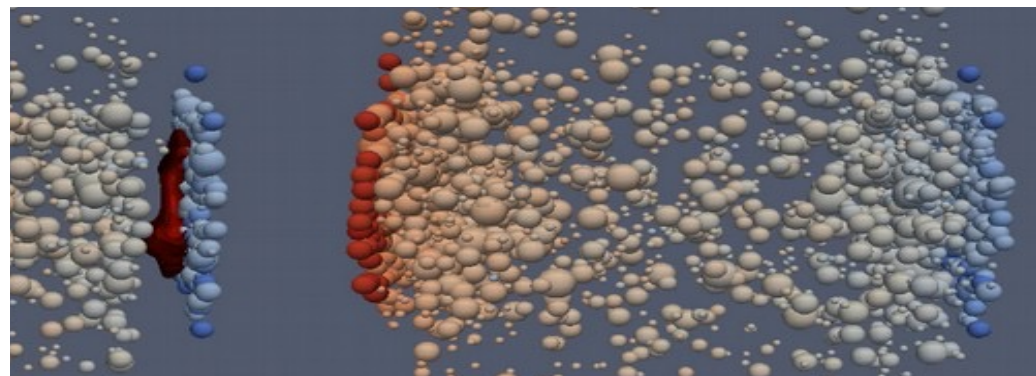
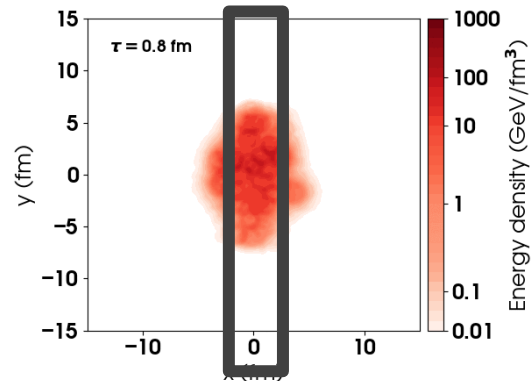
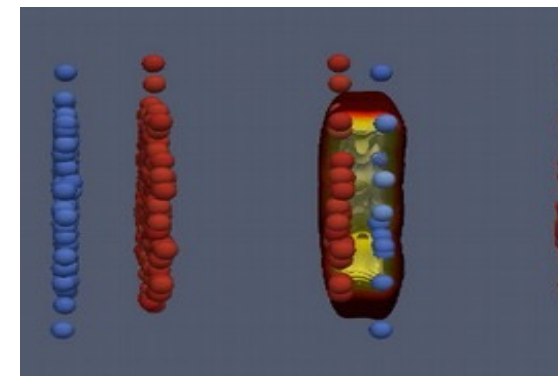


Longitudinal
expansion

Transverse expansion

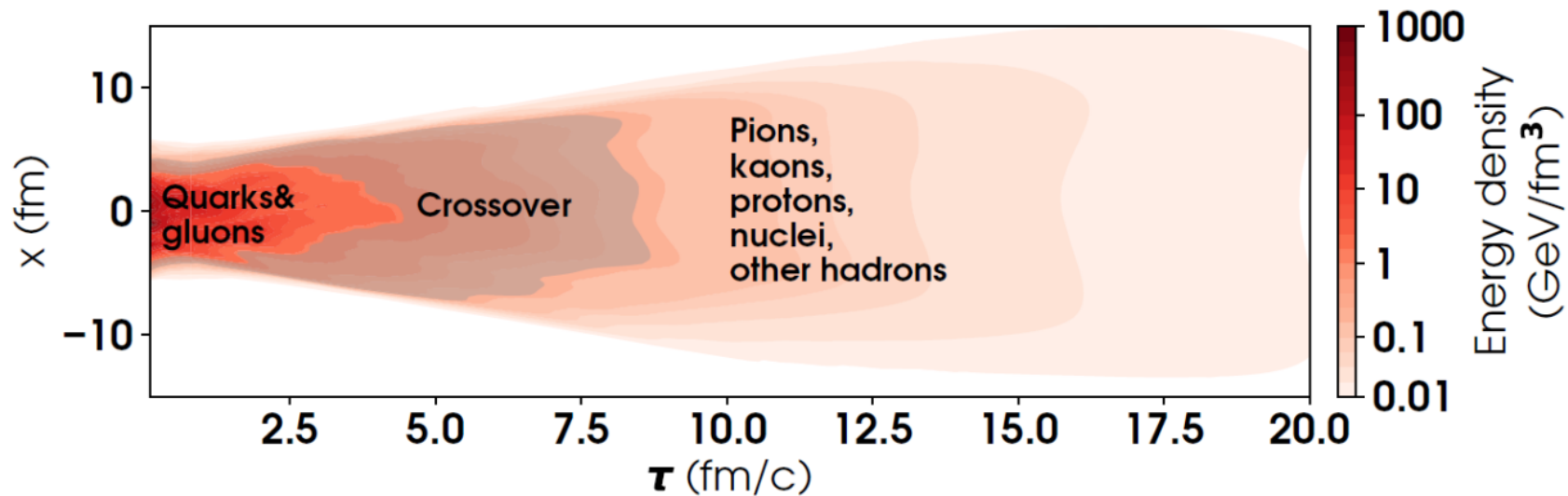


Hydrodynamics calibrated
on hadronic measurements



“Macroscopic” calc. of photons from deconfined plasma

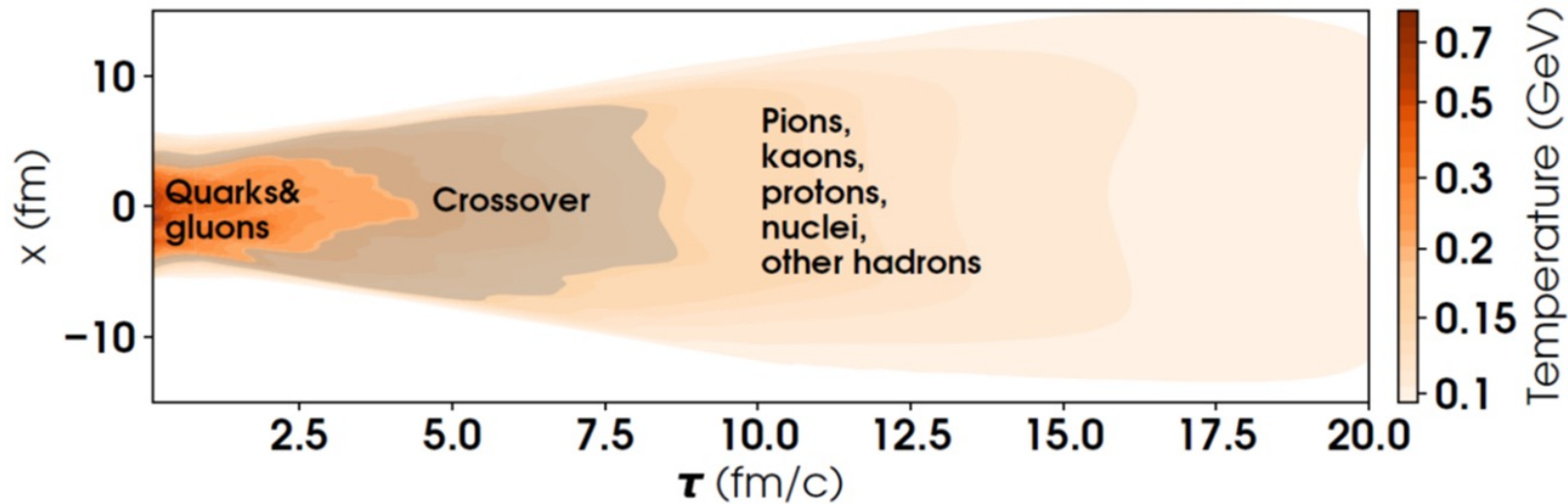
- What is the spacetime and momentum profile of quarks/gluons/hadrons?



- How much radiation is emitted in each region?
- Note: No clear separation between **quark/gluon** phase and **hadronic** phase

“Macroscopic” calc. of photons from deconfined plasma

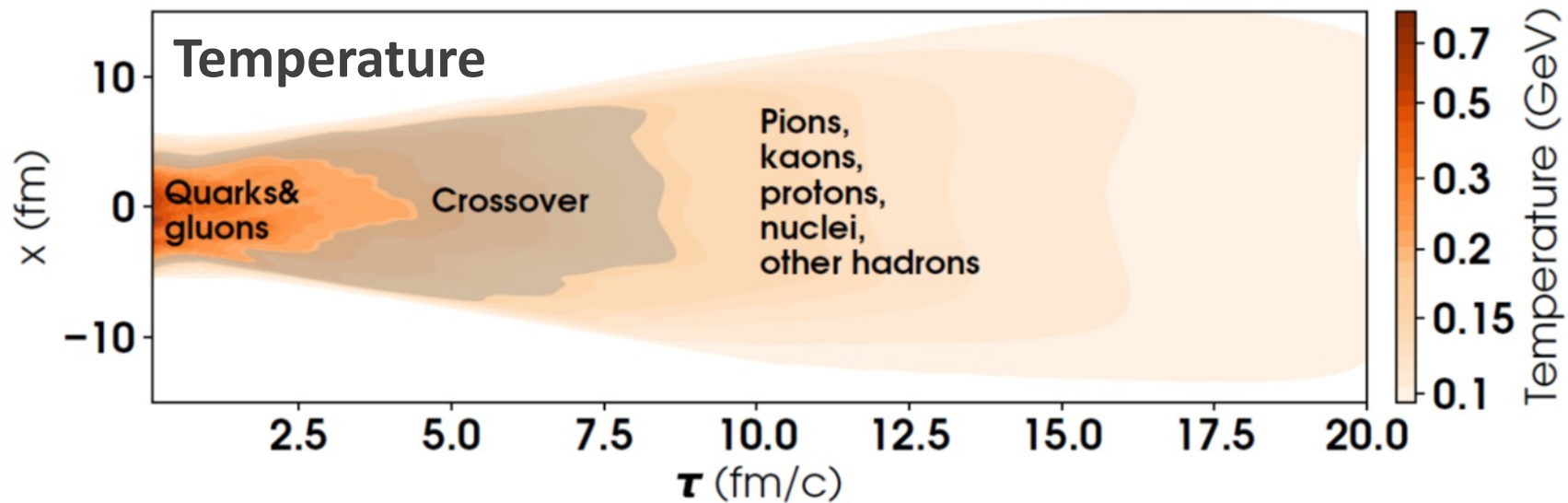
- What is the spacetime and momentum profile of quarks/gluons/hadrons?



- How much radiation is emitted in each region?
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Photons from deconfined plasma

- What is the spacetime profile of quarks/gluons/hadrons?



Spacetime profile of plasma

- Photon production:
$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, \overbrace{T(X), u^\mu(X), \dots})$$

Photon emission rate

Photon emission rate

Spacetime profile of plasma

■ Photon production:
$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, T(X), u^\mu(X), \dots)$$

Photon emission rate

State of matter/Temperatures

Photon emission rate

Gas of hadrons

Effective hadronic models

Deconfinement

Photon emission from quark gluon plasma: Complete leading order results

Peter Brockway Arnold (Virginia U.), Guy D. Moore (Washington U., Seattle), Laurence G. Yaffe (Washington U., Seattle)
Nov, 2001

Extrapolated rates from low/high temperatures

Strongly-coupled for T

Non-perturbative QCD, holography, effective models

Weakly-coupled QGP at $T \gg 1$ GeV

Perturbative QCD

Photon emission rate

Spacetime profile of plasma

▪ Photon production:
$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, T(X), u^\mu(X), \dots)$$

Photon emission rate

State of matter/Temperatures

Hadronic production of thermal photons

Simon Turbide (McGill U.), Ralf Rapp (Nordita), Charles Gale (McGill U.)

Aug, 2003

Strongly coupled quark/gluons

Thermal photons from heavy ion collisions: A spectral function approach

Kevin Dusling (Brookhaven), Ismail Zahed (SUNY, Stony Brook)

Nov, 2009

Photon emission rate

Effective hadronic models

Extrapolated rates from low/high temperatures

Lattice QCD, holography, effective models

Perturbative QCD

Photon emission rate

Spacetime profile of plasma

▪ Photon production:
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Photon emission rate

State of matter/Temperatures

**Photon emissivity of the quark-gluon plasma:
A lattice QCD analysis of the transverse
channel**

Marco Cè (U. Bern, AEC and Bern U. and CERN), Tim
Harris (Edinburgh U.), Ardit Krasniqi (U. Mainz, PRISMA), Harvey B.
Meyer (U. Mainz, PRISMA and Helmholtz Inst., Mainz and
Darmstadt, GSI), Csaba Török (U. Mainz, PRISMA)

May 5, 2022

26 pages

weakly coupled QCD at $T \gg T_{c,QCD}$

Photon emission rate

Effective hadronic models

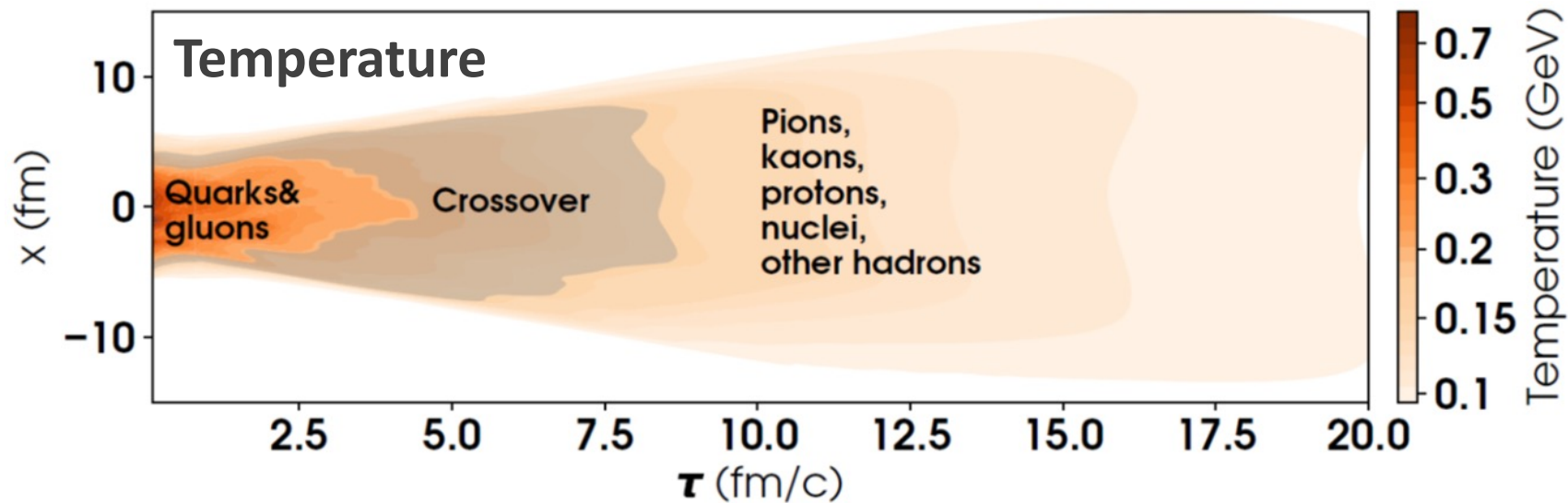
Extrapolated rates from low/high
temperatures

Lattice QCD, holography, effective
models

Perturbative QCD

Photons from deconfined plasma

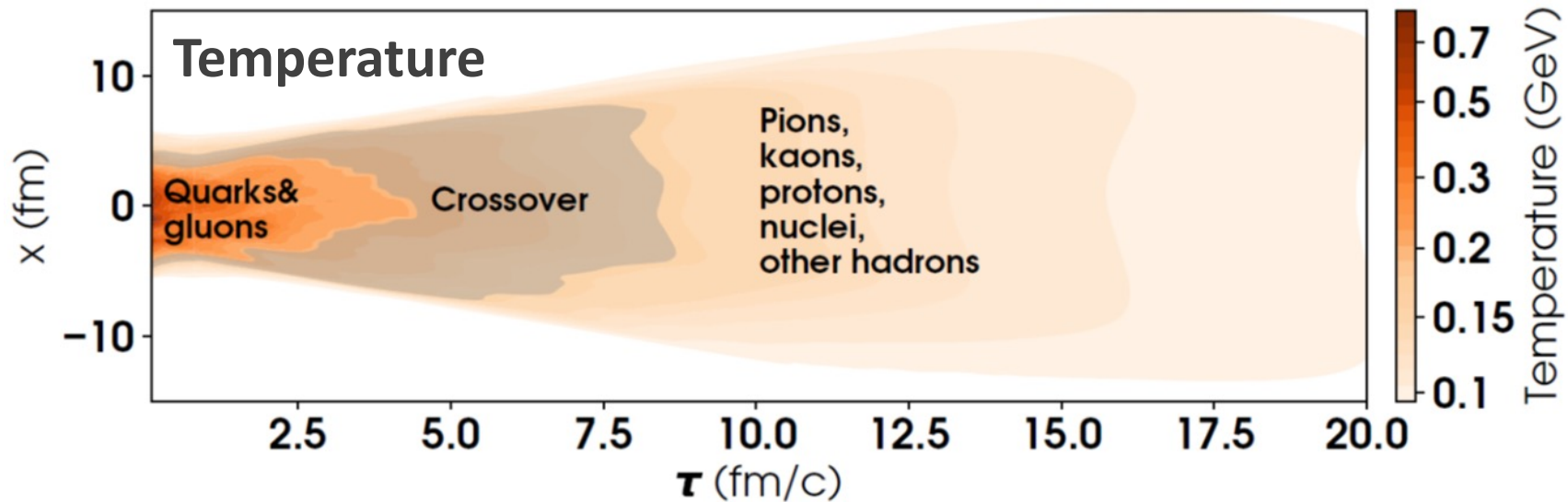
- What is the spacetime profile of quarks/gluons/hadrons?



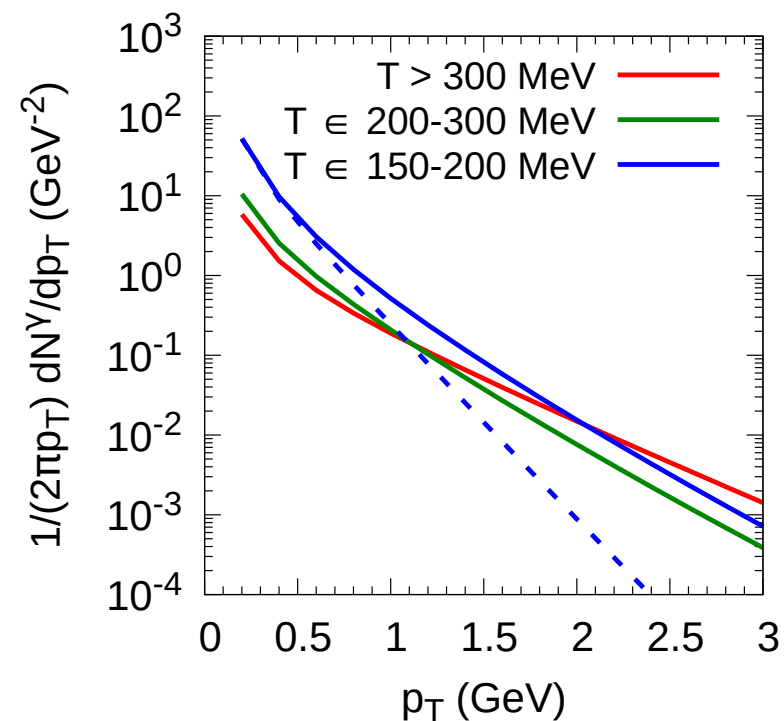
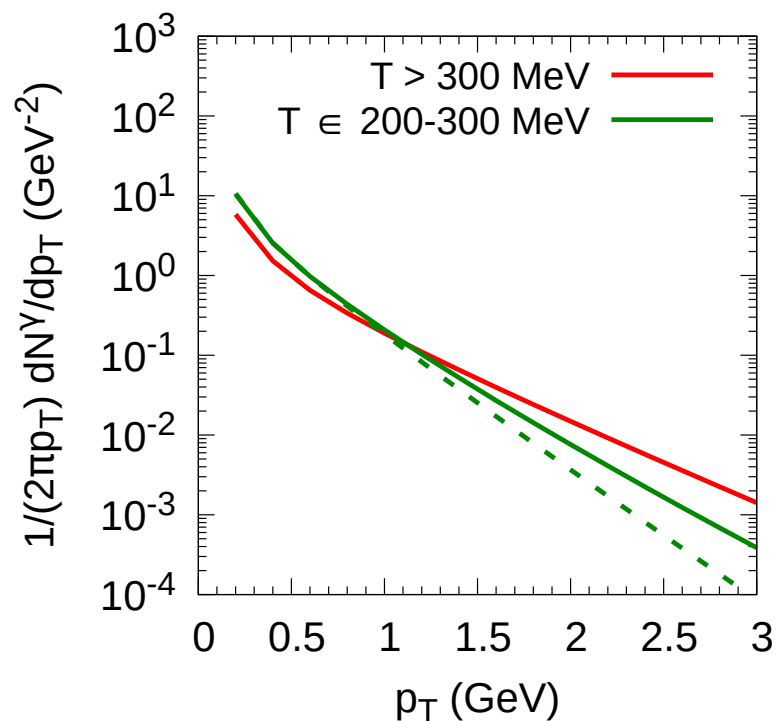
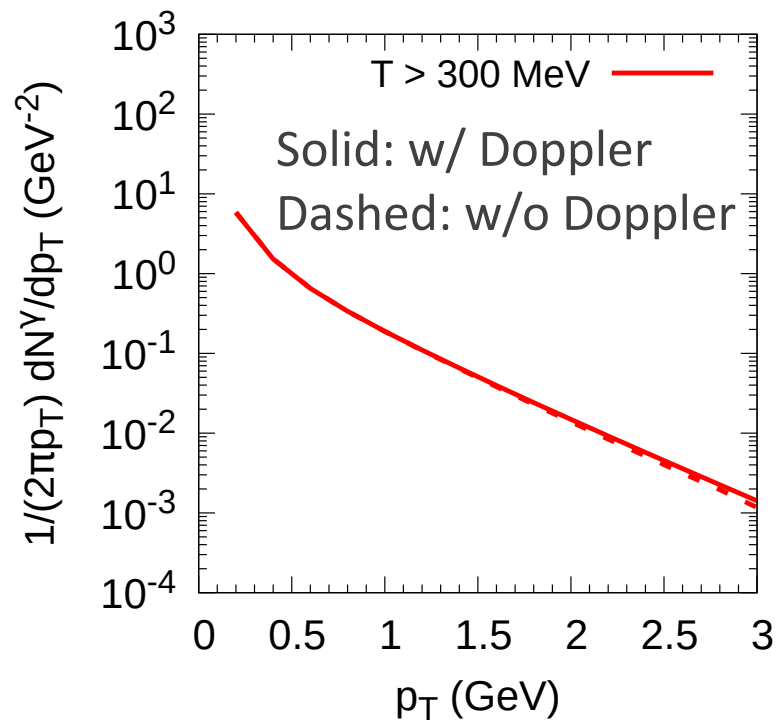
Spacetime profile of plasma

- Photon production:
$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, \overbrace{T(X), u^\mu(X), \dots})$$

Photon emission rate

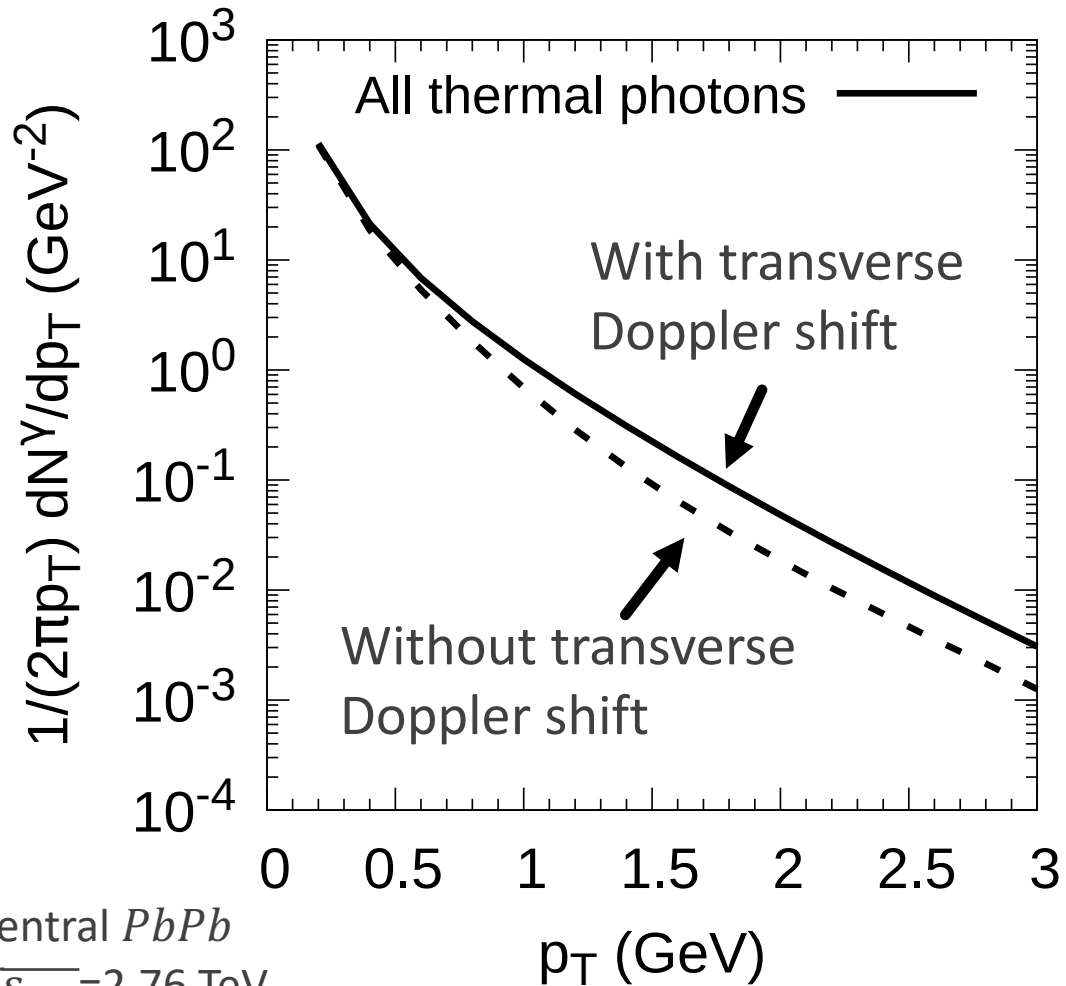


The later (or lower temperature) photons are produced, the more Doppler-shifted

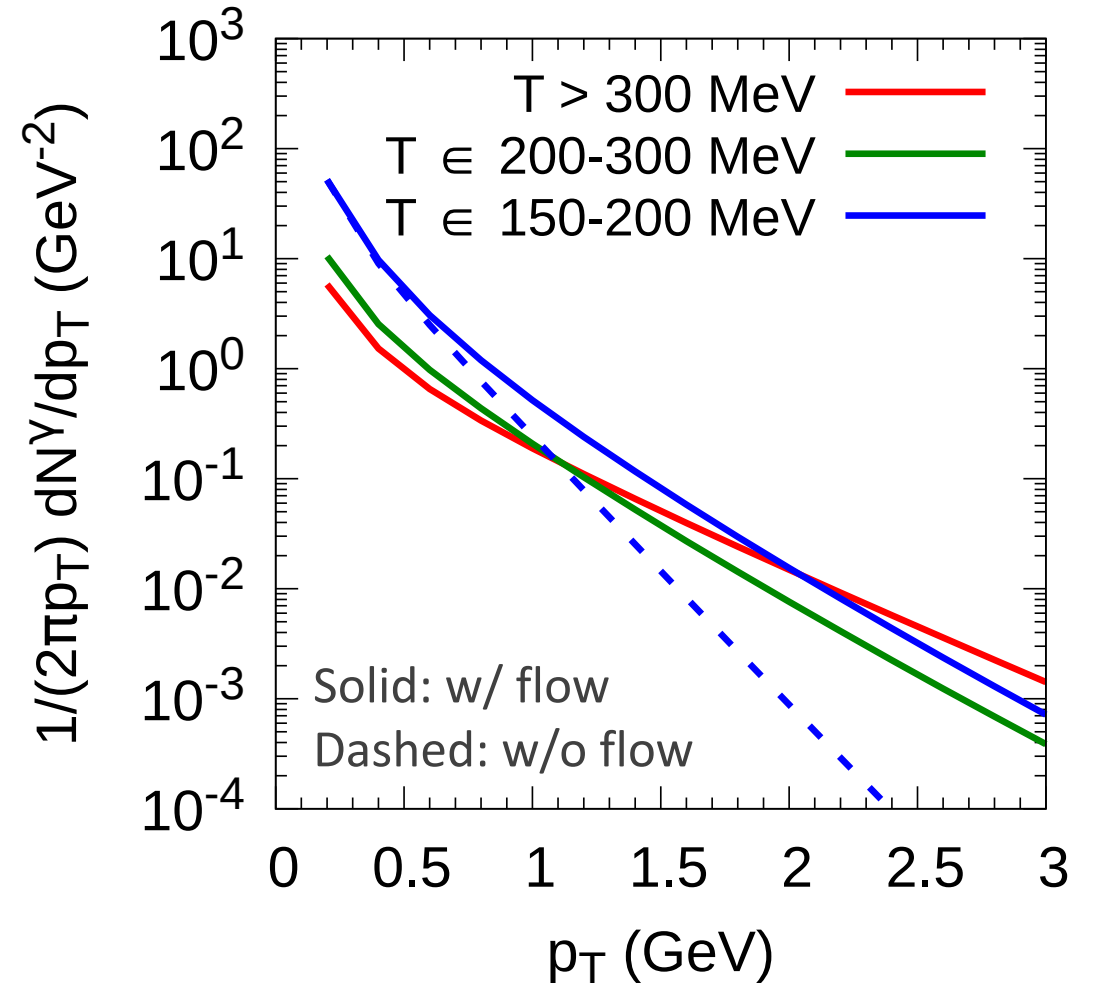


Effect of transverse Doppler shift in realistic calculation

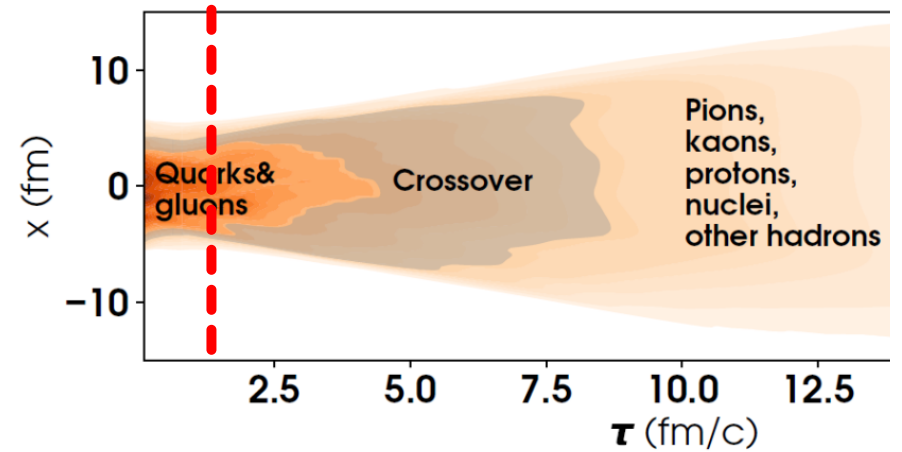
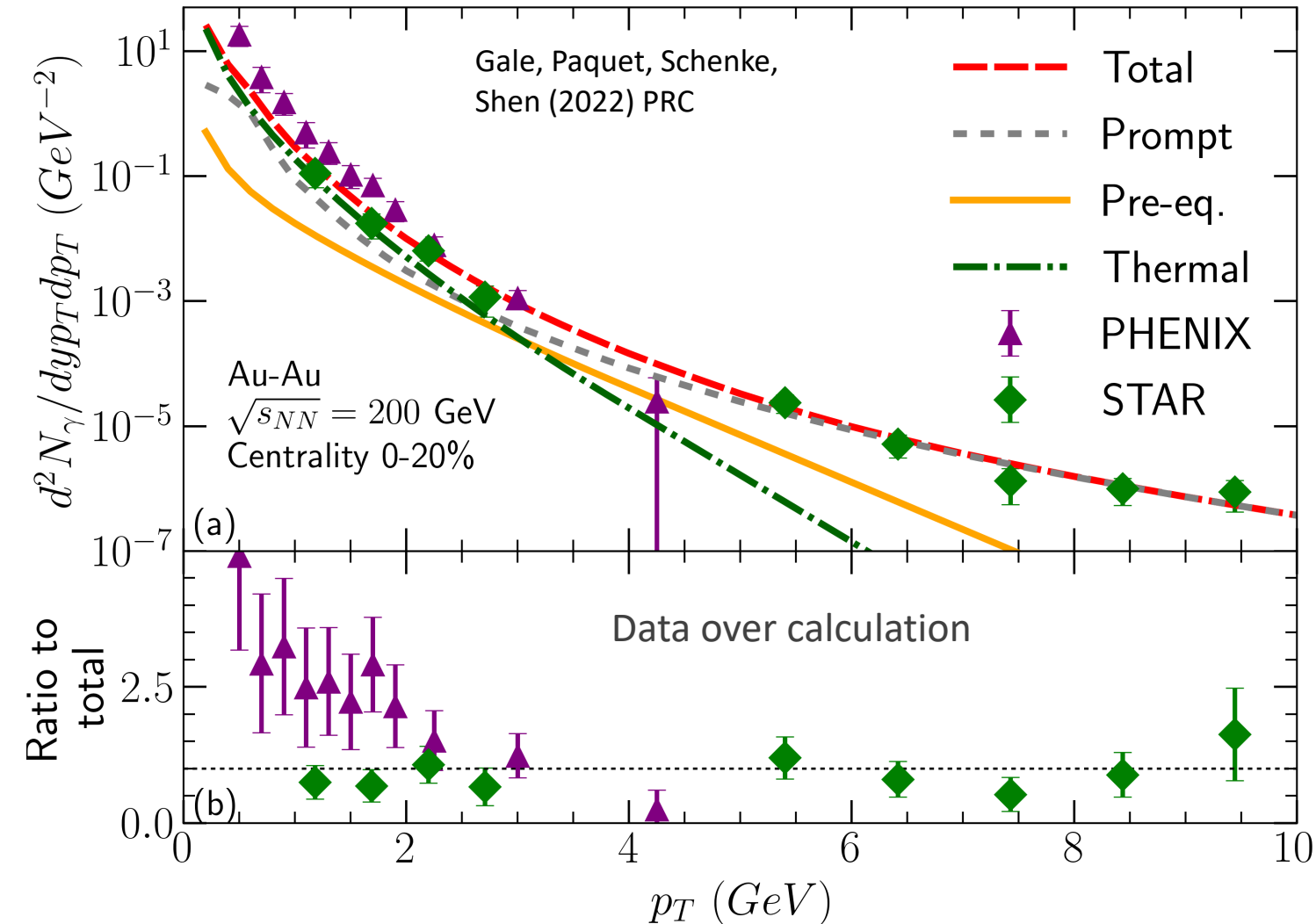
Global effect of Doppler shift



Local effect of Doppler shift

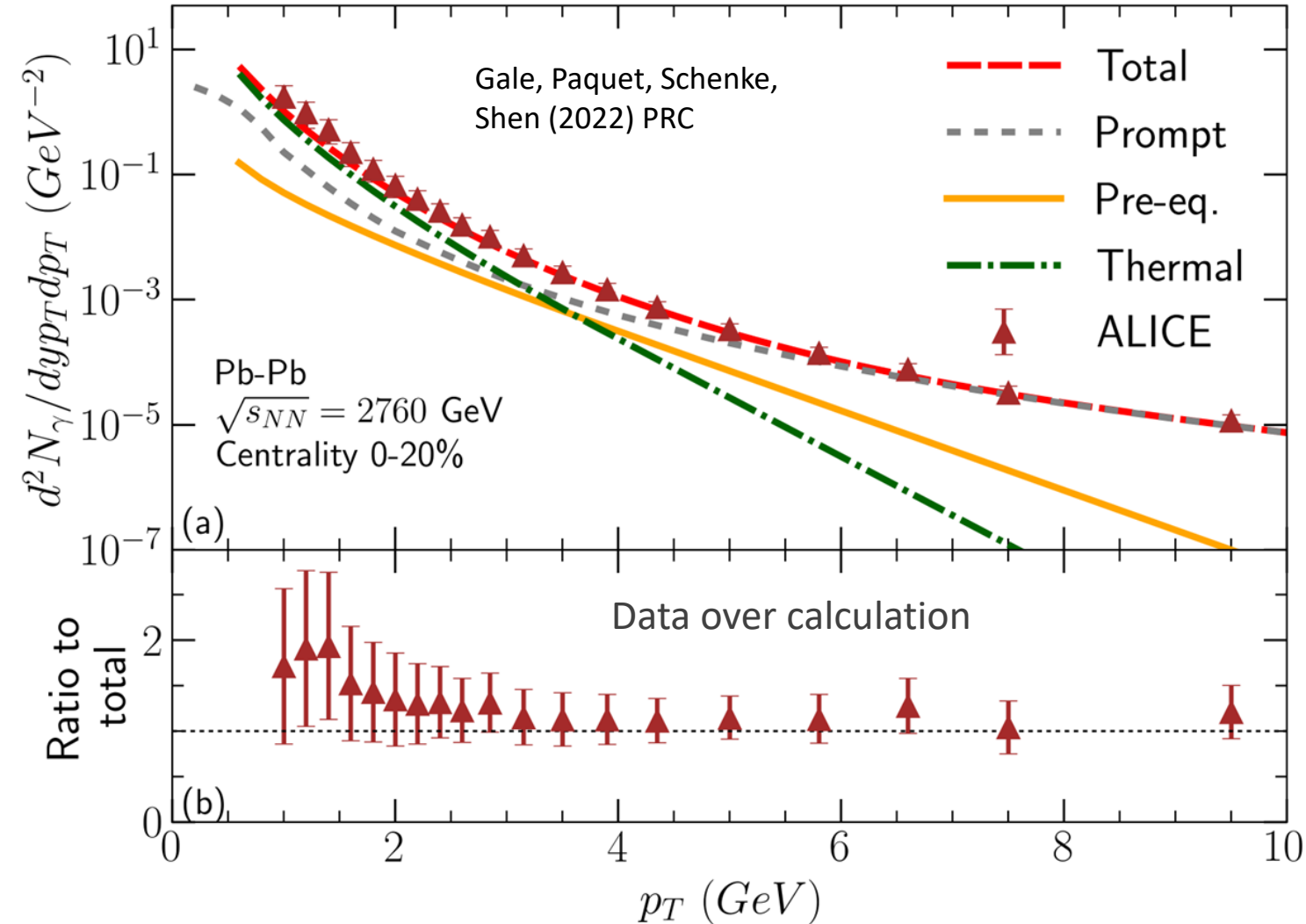


Results: Au-Au $\sqrt{s_{NN}} = 200$ GeV, 0-20%



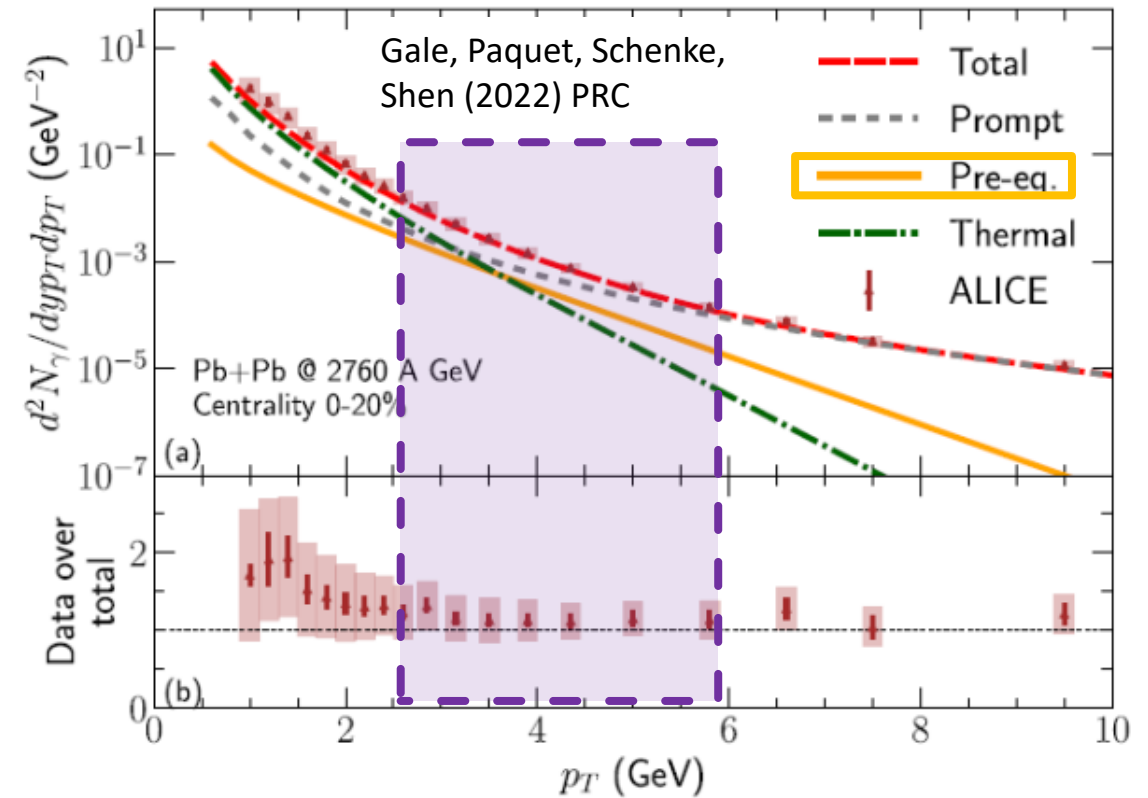
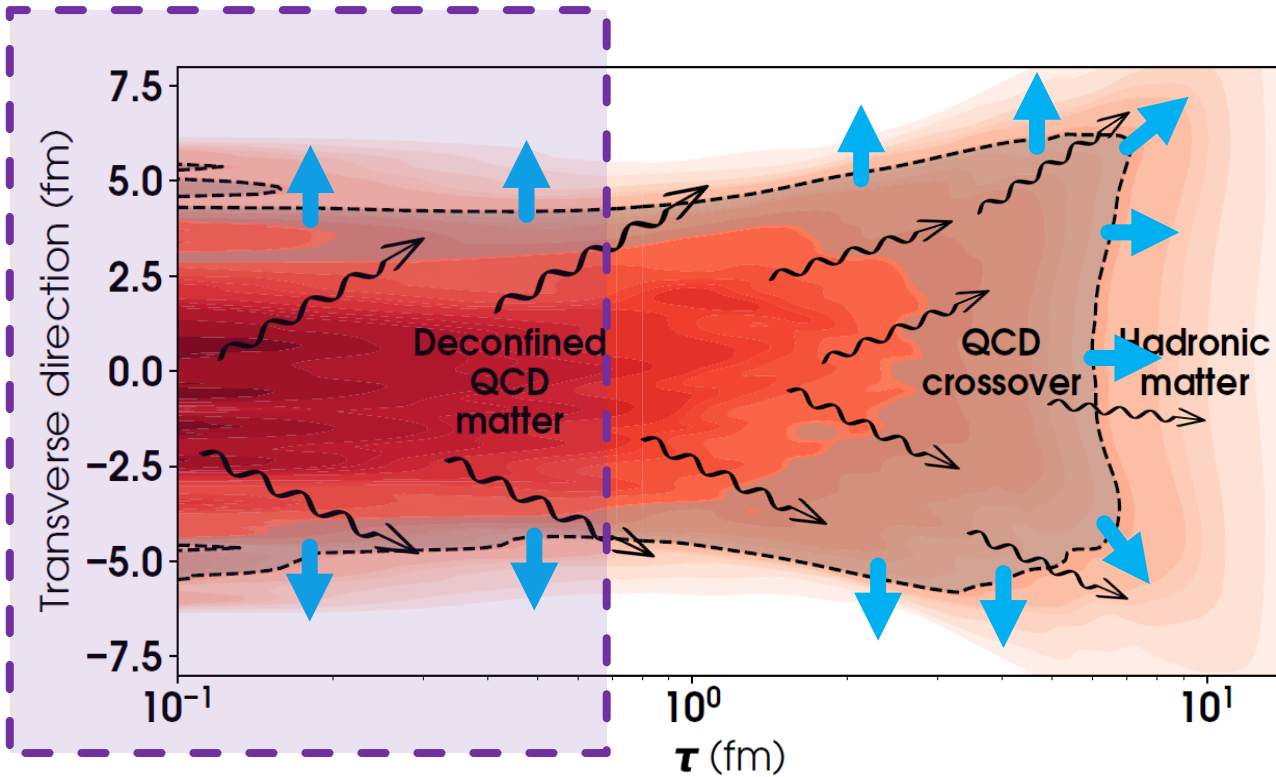
- Soft-bath ("thermal") photons dominate at low energy/ p_T
- Experimental uncertainties large
- \pm Some tension between measurements

Results: Pb-Pb $\sqrt{s_{NN}} = 2760$ GeV, 0-20%



- Soft-bath ("thermal") photons dominate at low energy/ p_T
- Experimental uncertainties large
- \pm Some tension between measurements

Photons from the early stage of the collision



- Use photons to study the earliest stage of the collisions:
 - Approach to chemical equilibrium (from gluon dominated to quark&gluon equilibrium)
 - Understand formation of the soft bath and “emergence of hydrodynamics”

Results: momentum anisotropy

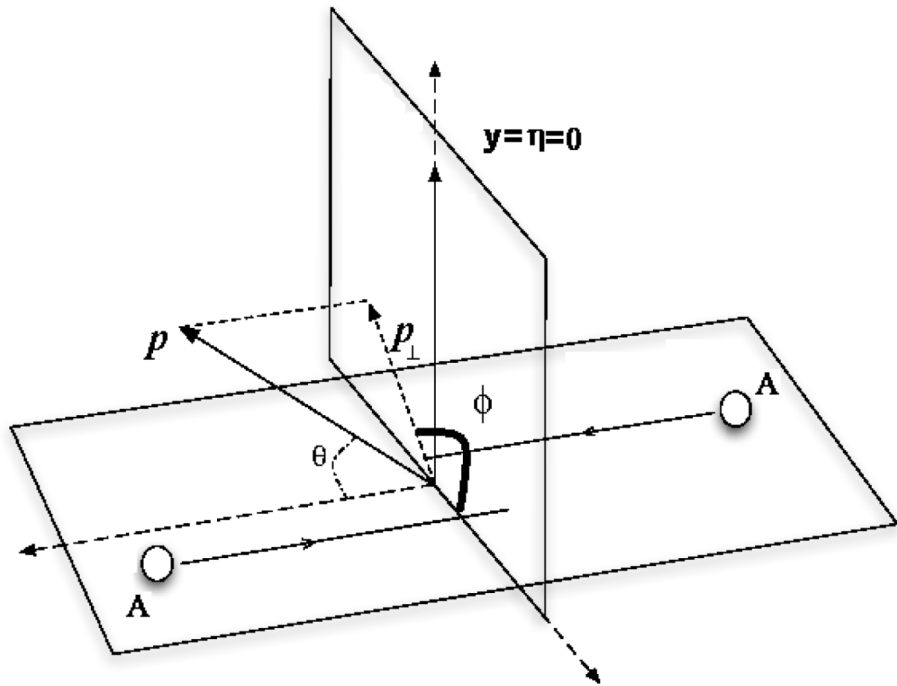


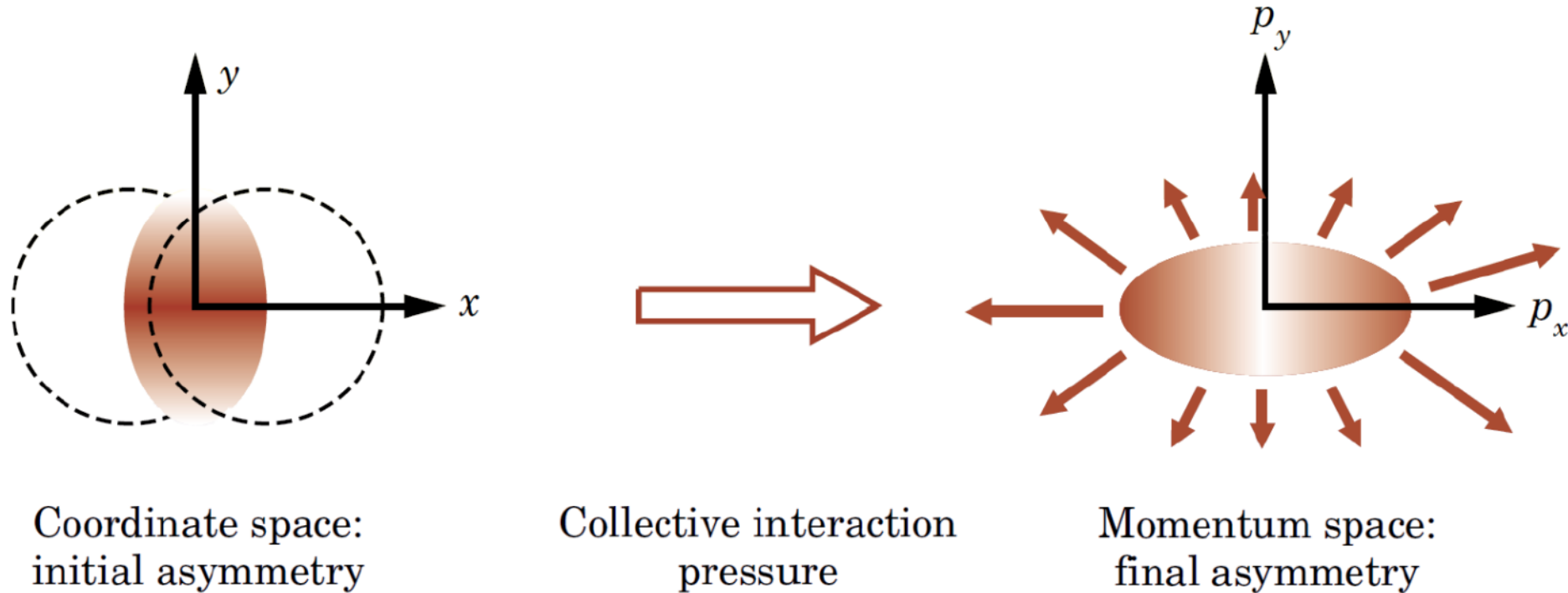
Figure adapted from K. Tuchin (2013) AHEP

$$\frac{1}{2\pi p_T} \frac{dN}{dp_T d\phi} = \left(\frac{1}{2\pi p_T} \frac{dN}{dp_T} \right) \left[1 + 2 \sum_{n=1}^{\infty} v_n \cos(n(\phi - \Psi_n)) \right]$$

- More precisely:
momentum anisotropy through photon-hadron correlation

$$v_n\{SP\}(p_T) = \frac{\left\langle v_n^y(p_T) v_n^h \cos \left(n(\Psi_n^y(p_T) - \Psi_n^h) \right) \right\rangle}{\sqrt{\left\langle (v_n^h)^2 \right\rangle}}$$

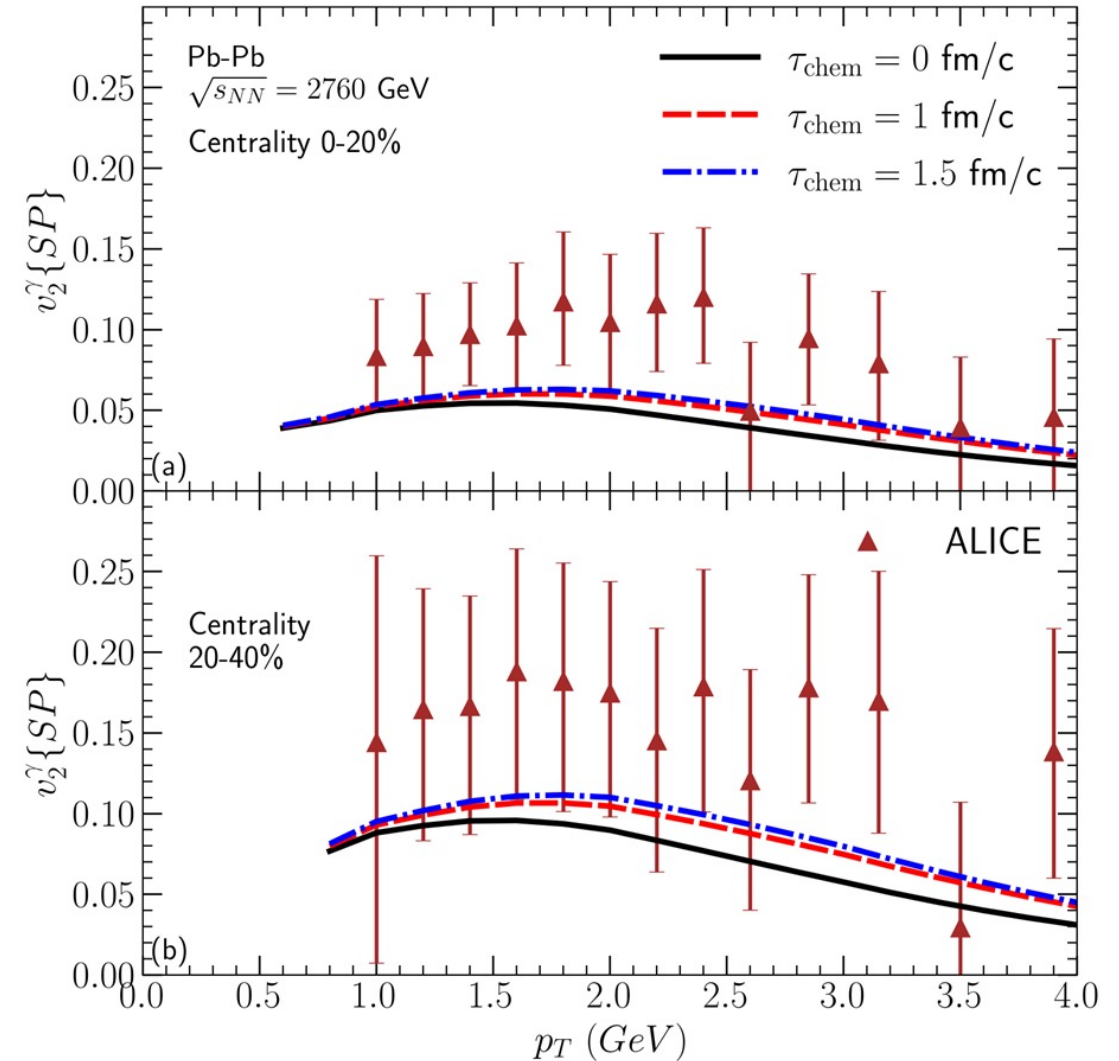
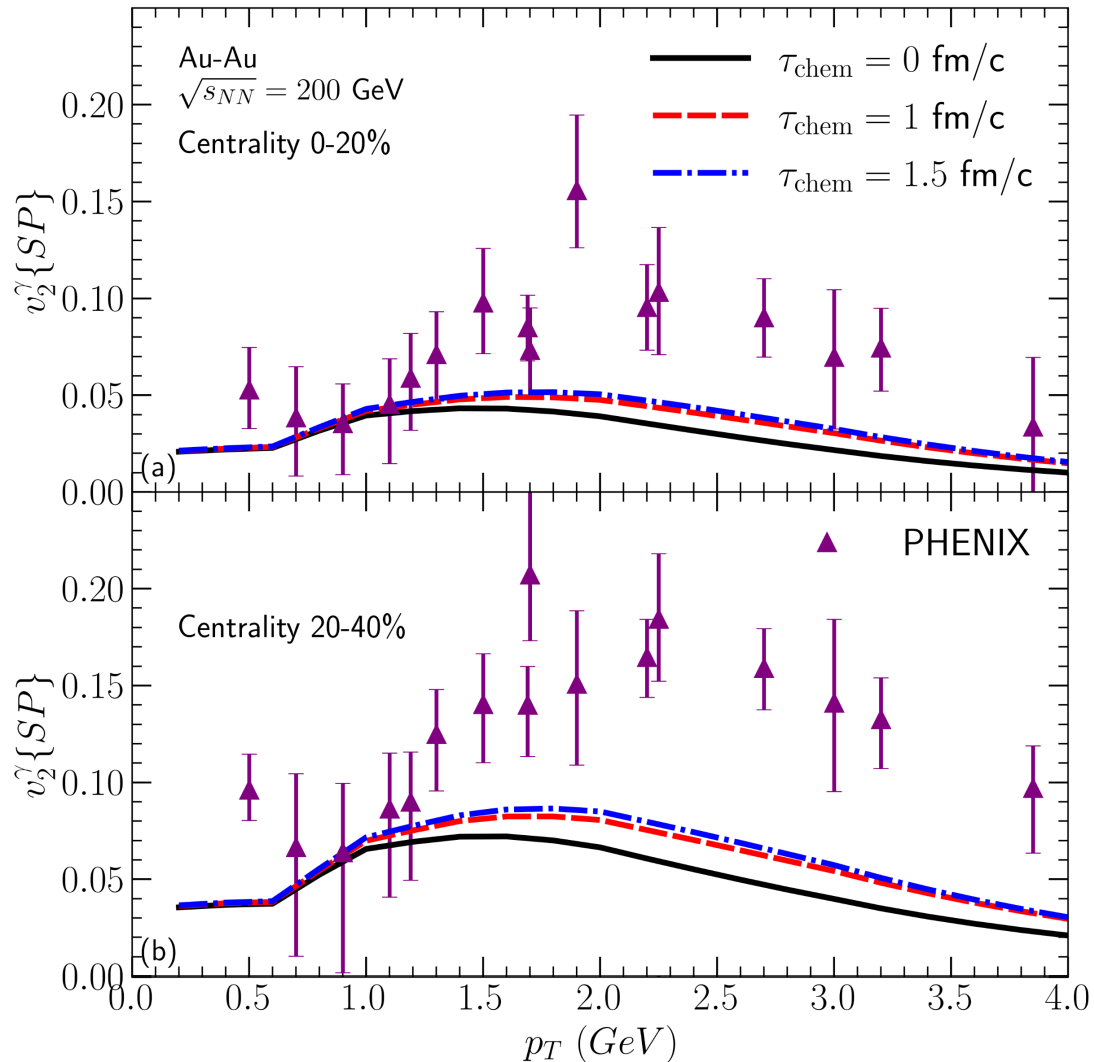
Momentum anisotropy from geometrical anisotropy

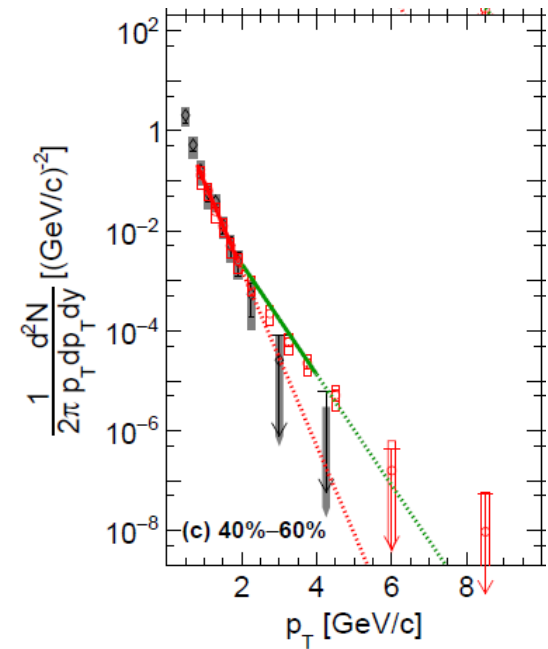
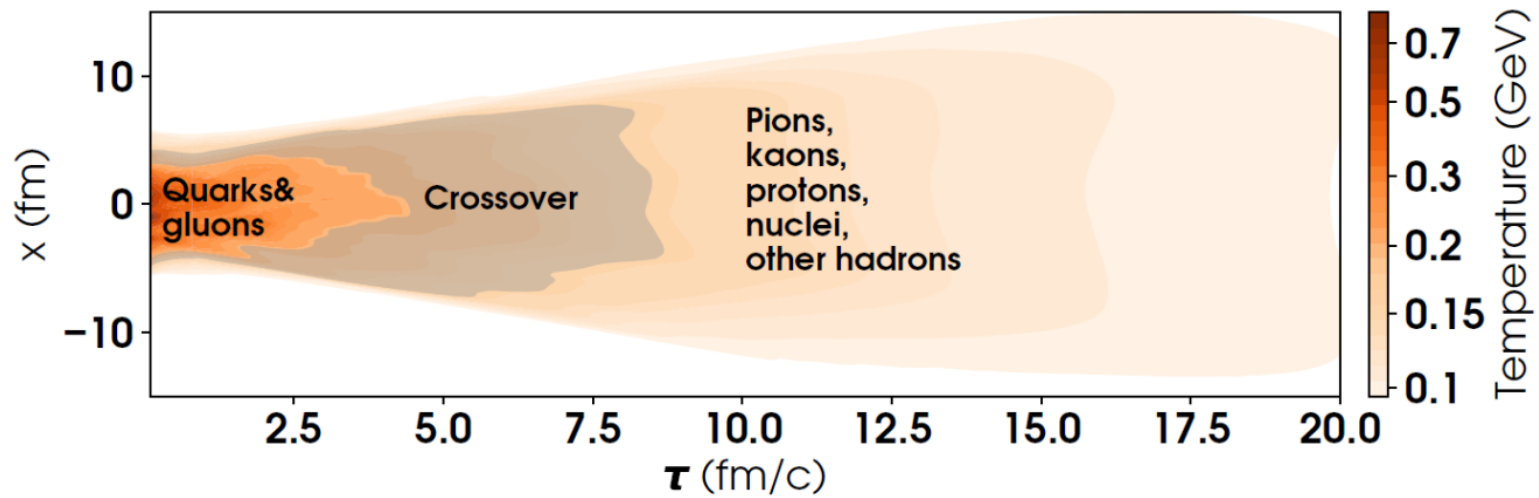


$$\frac{1}{2\pi p_T} \frac{dN}{dp_T d\phi} = \left(\frac{1}{2\pi p_T} \frac{dN}{dp_T} \right) \left[1 + 2 \sum_{n=1}^{\infty} v_n \cos(n(\phi - \Psi_n)) \right]$$

Photon momentum anisotropy

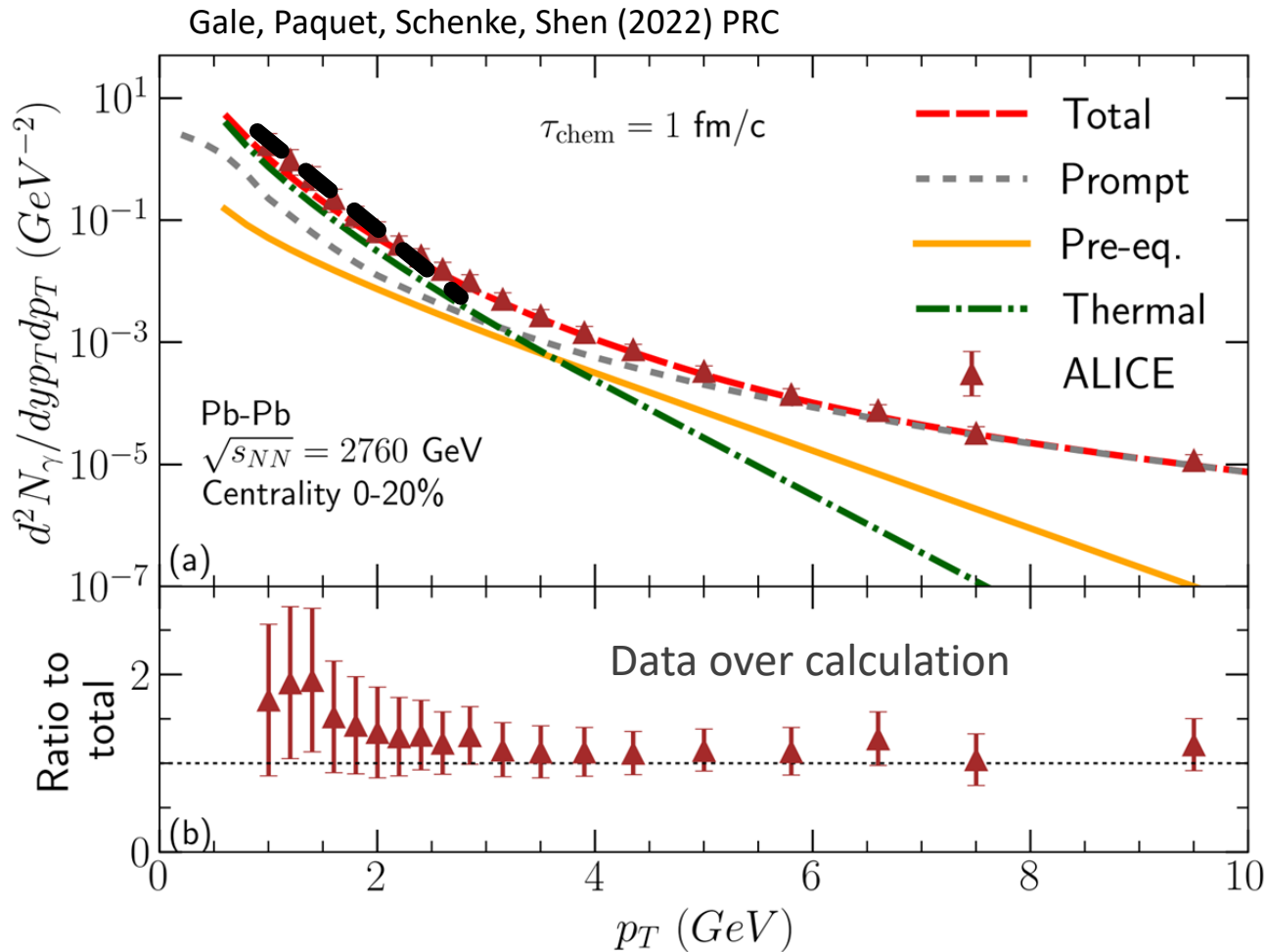
(Compare different chemical equilibration scenarios)



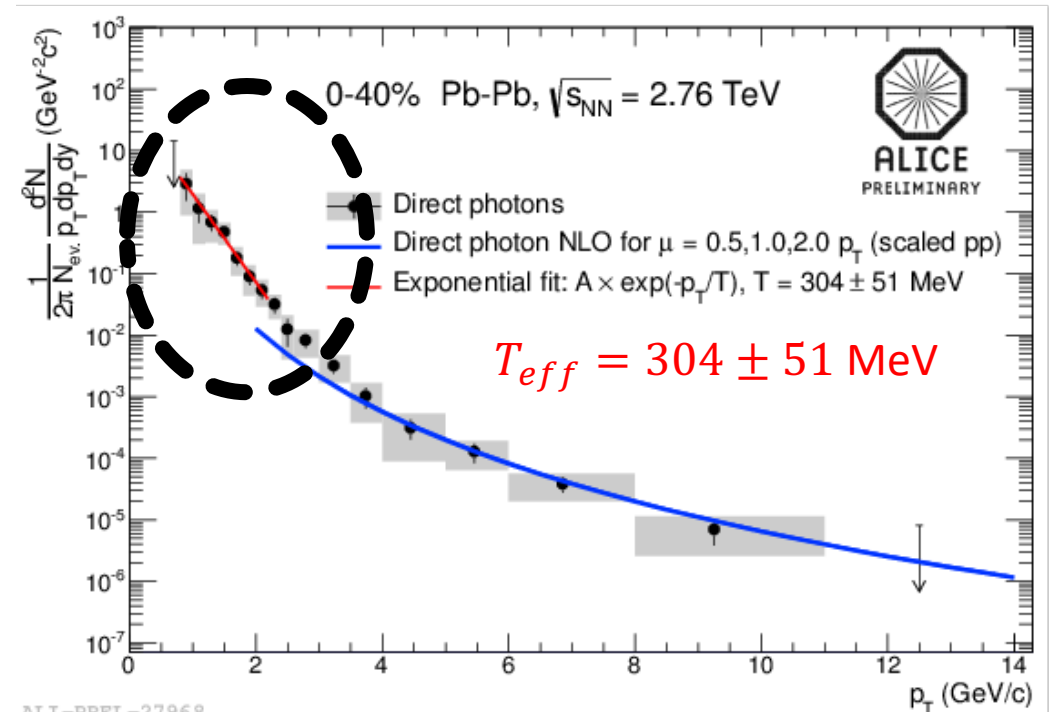


PLASMA TEMPERATURE FROM PHOTON ENERGY SPECTRUM

Results: Pb-Pb $\sqrt{s_{NN}} = 2760$ GeV, 0-20%

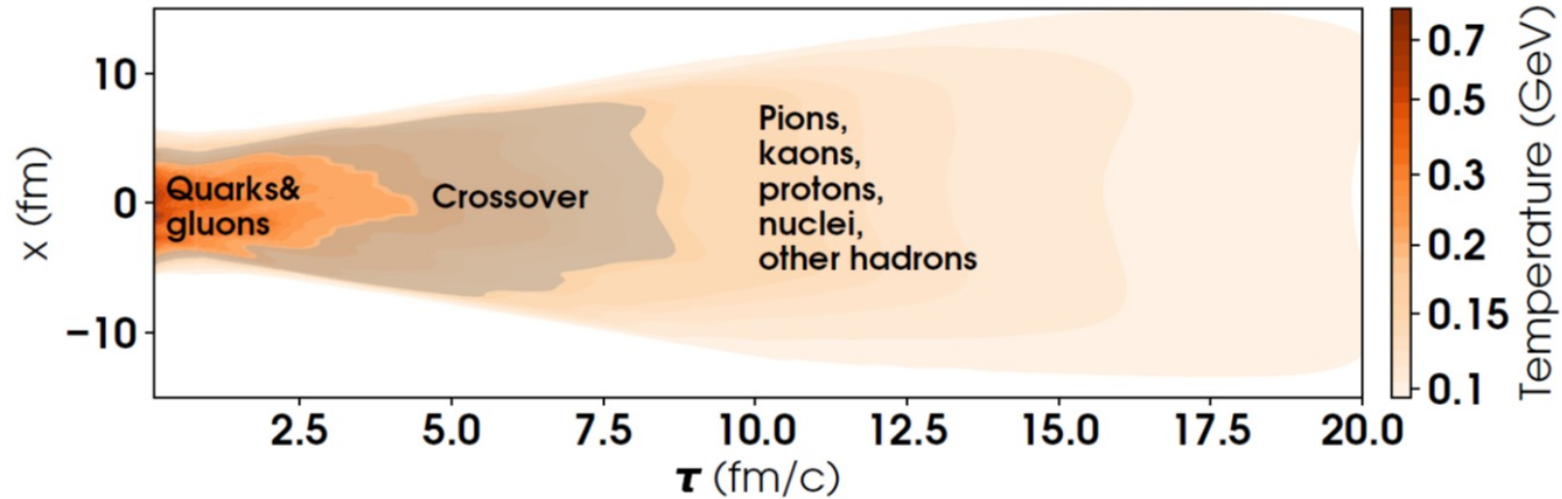


$$\ln \left(\frac{1}{2\pi E} \frac{dN}{dE dy} \right) = cte - \frac{E}{T_{eff}}$$

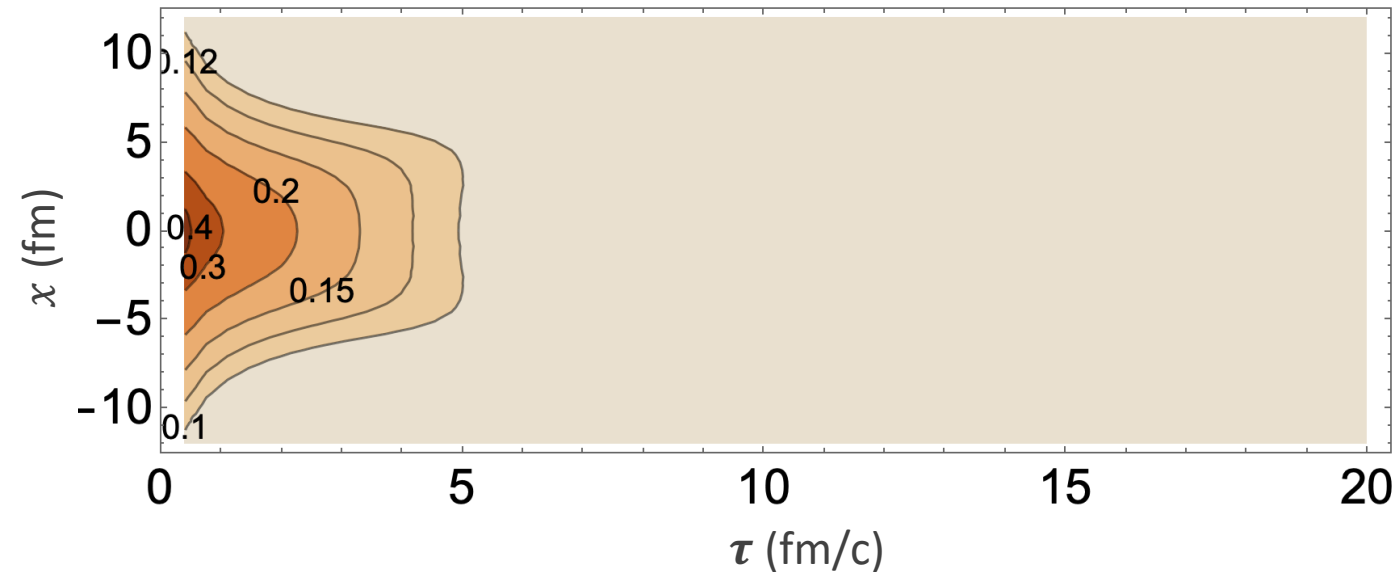


ALI-PREL-27968

- Can we estimate the maximum temperature of the plasma with the photon spectrum?



$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, T, u^\mu, \dots)$$



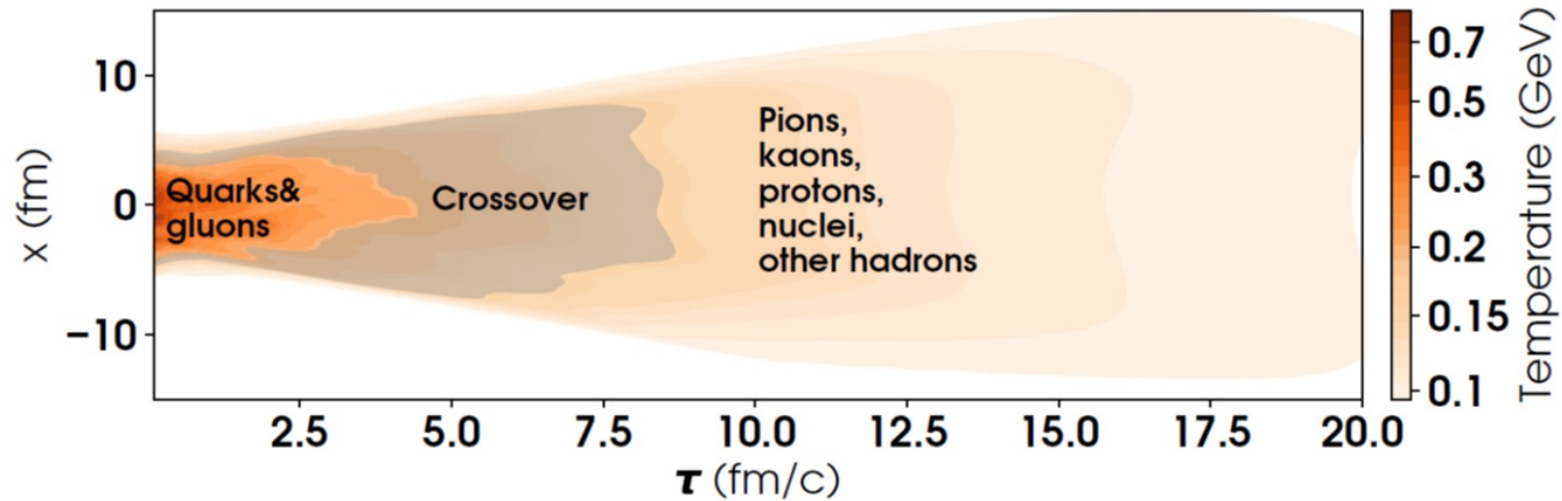
Symmetry constraints on generalizations of Bjorken flow

Steven S. Gubser

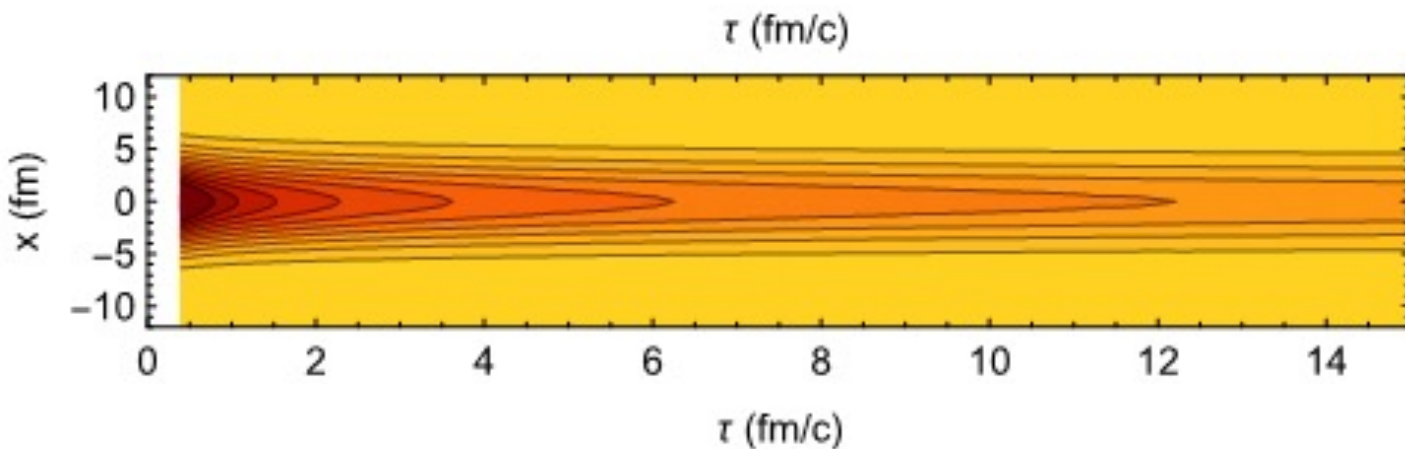
Phys. Rev. D **82**, 085027 – Published 26 October 2010

$$T(\tau, r) \propto \frac{(2q\tau)^{\frac{2}{3}}}{\tau(1 + 2q^2(\tau^2 + r^2) + q^4(\tau^2 - r^2)^2)^{1/3}}$$

- Can we estimate the maximum temperature of the plasma with the photon spectrum?



$$\frac{dN_\gamma}{d^3p} = \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, T, u^\mu, \dots)$$

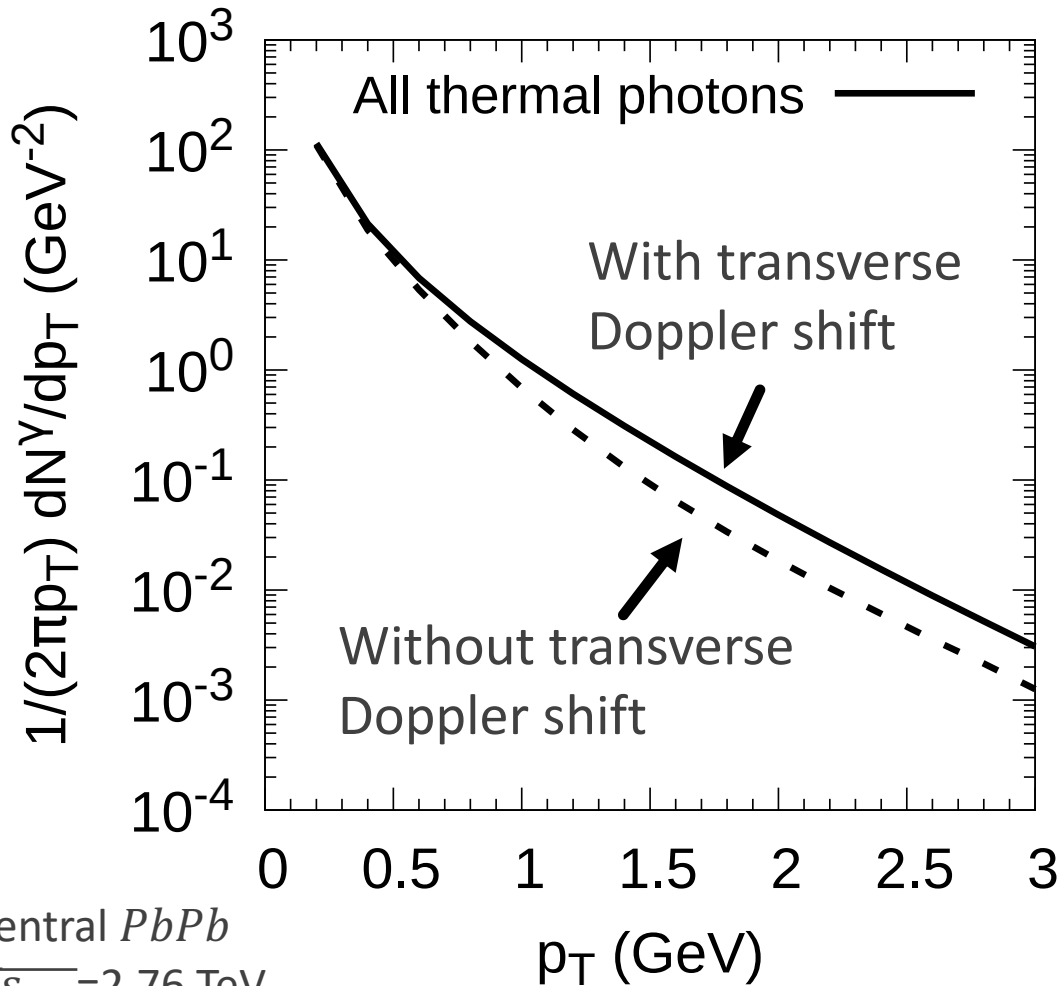


Gaussian transverse profile, but only longitudinal hydrodynamic expansion

$$T(\tau, r) = T_0 e^{-\frac{r^2}{2\sigma^2}} \left(\frac{\tau_0}{\tau}\right)^{c_s^2}$$

Effect of transverse Doppler shift

Global effect of Doppler shift



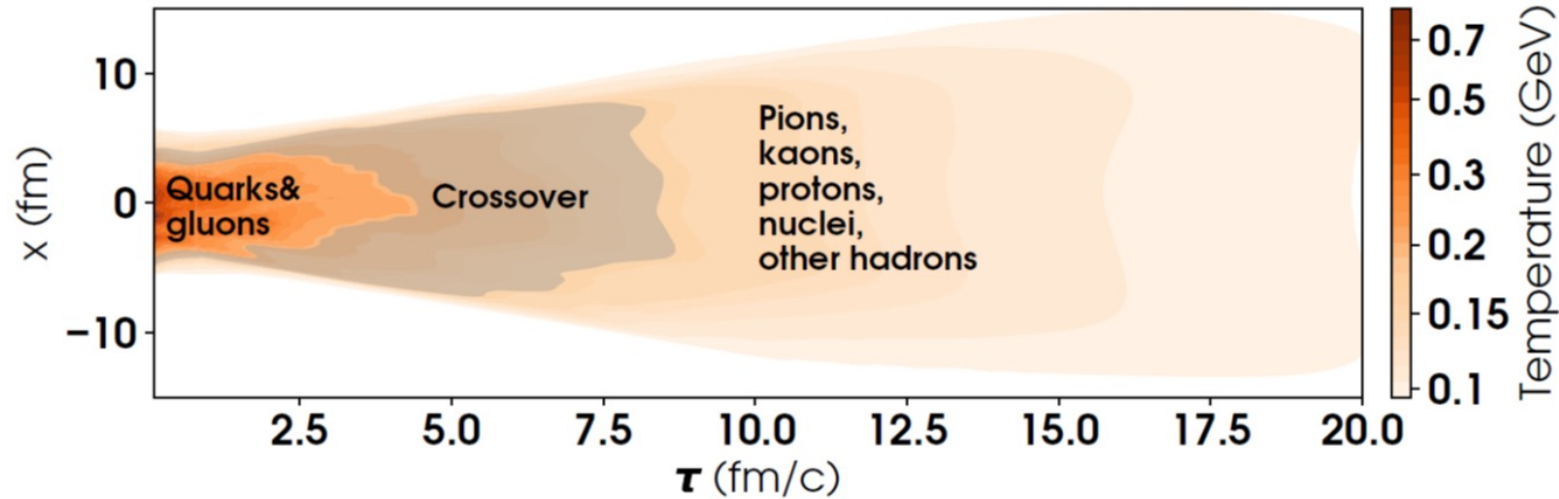
$$\frac{dN_\gamma}{d^3p} \approx \int d^4X \frac{d\Gamma_\gamma}{d^3p} (p, T, u^\mu = (1,0,0,0))$$

$$\frac{d\Gamma_\gamma}{d^3p} \sim e^{-p/T}$$

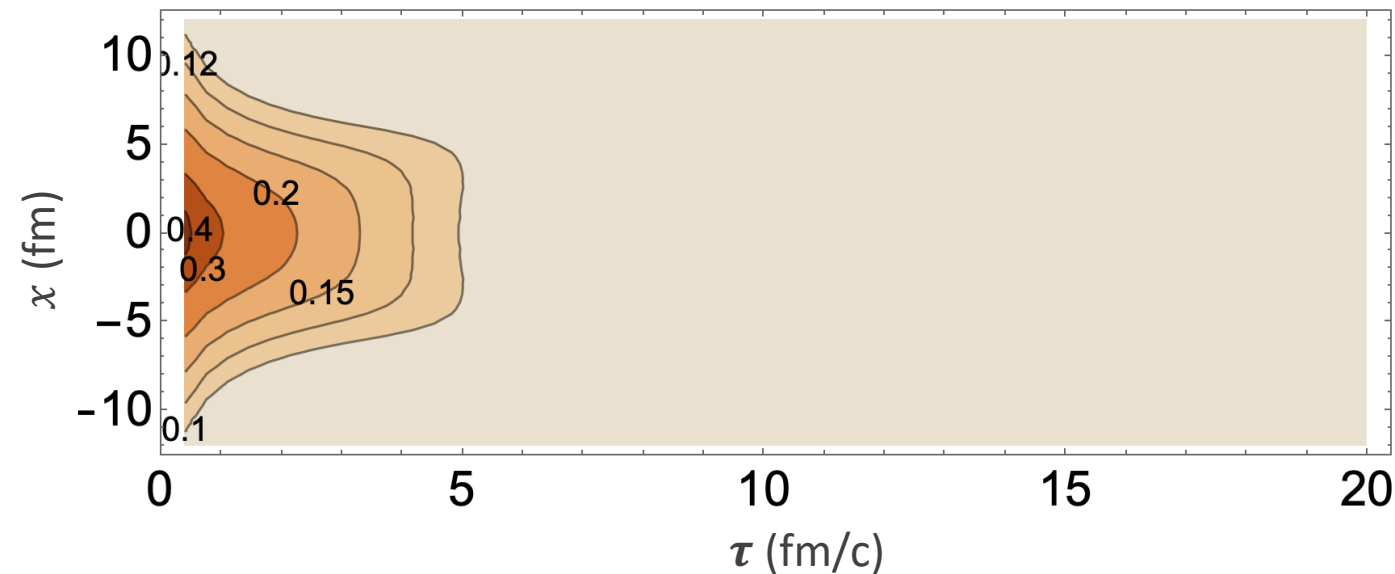
Paquet and Bass [arXiv:2205.12299]

Paquet [arXiv:2305.10669]

- Can we estimate the maximum temperature of the plasma with the photon spectrum?



$$\frac{dN_\gamma}{d^3p} \approx \int d^4X \frac{d\Gamma_\gamma}{d^3p}(p, T)$$



$$\frac{dN_\gamma}{d^3p} \propto \left(\frac{T_0}{p_T}\right)^{5/2} E \frac{d\Gamma_\gamma}{d^3p}(p_T, T_0)$$

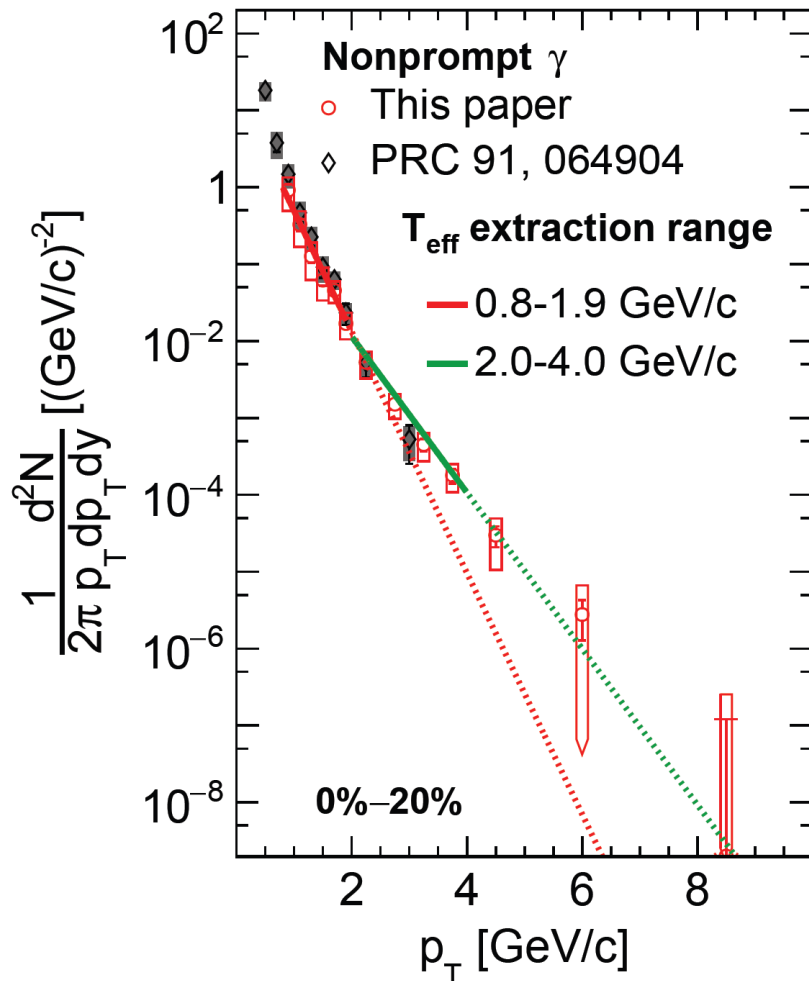
Inverse slope:

$$-\frac{1}{T_{eff}} \approx -\frac{1}{T_0} - \frac{5}{2} \frac{1}{p_T} + O\left(\frac{T_0}{p_T^2}\right)$$

Au-Au $\sqrt{s_{NN}} = 200$ GeV, 0-20%

Caveats: other sources of photons (e.g. pre-equilibrium), viscosity, ...

Ref.: PHENIX Collaboration [arXiv:2203.17187]



p_T cut	T_{eff}	$T_0 = \frac{T_{eff}}{1 - \frac{5 T_{eff}}{2 p_T}}$
$0.8 < p_T < 1.9$ GeV	277 MeV	570 MeV
$2 < p_T < 4$ GeV	428 MeV	670 MeV

From hydro fit to hadronic data: $T_0 \approx 530$ MeV

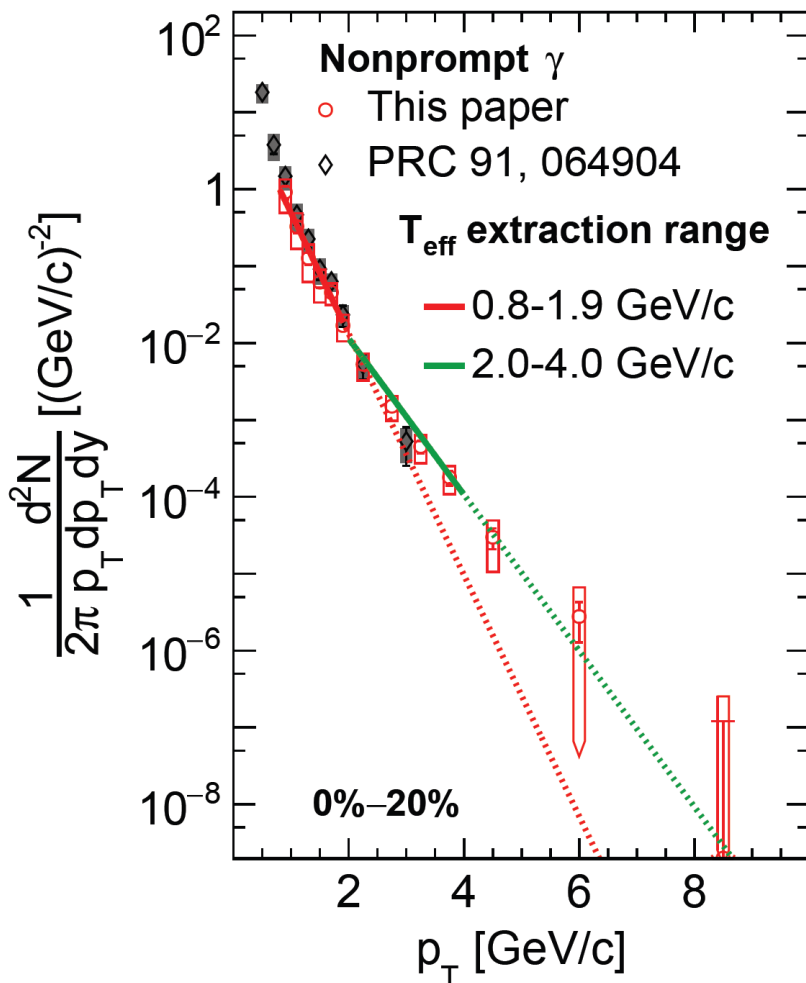
[from Gale, Paquet, Schenke, Shen (2022) PRC]

Partly explains large p_T -cut dependence of T_{eff}

Au-Au $\sqrt{s_{NN}} = 200$ GeV, 0-20%

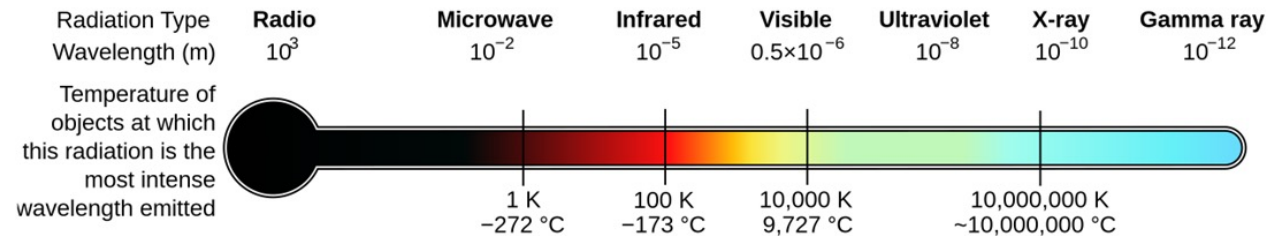
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p_T cut	T_{eff}	$T_0 = \frac{T_{eff}}{1 - \frac{5 T_{eff}}{2 p_T}}$
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$$T = 500 \text{ MeV} = 6 \times 10^{12} \text{ K}$$



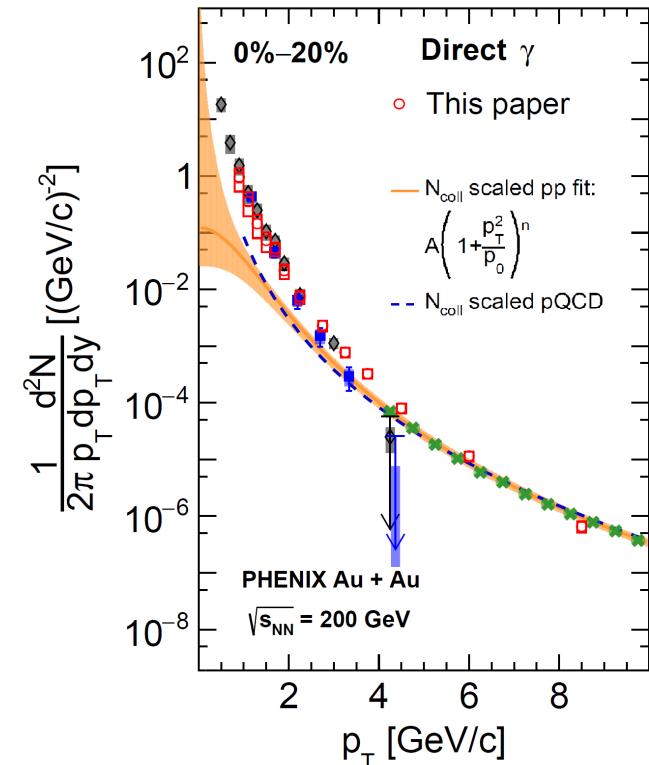
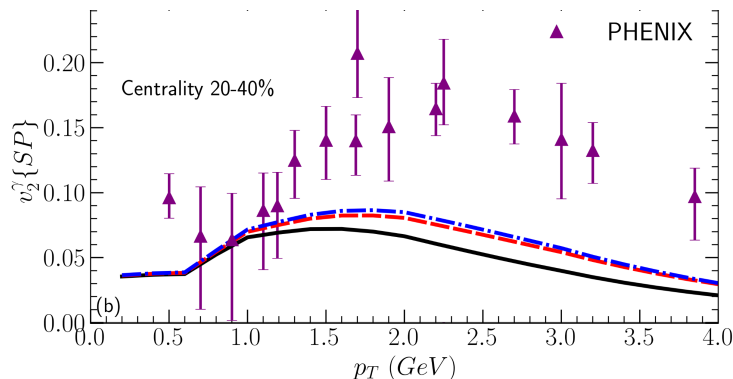
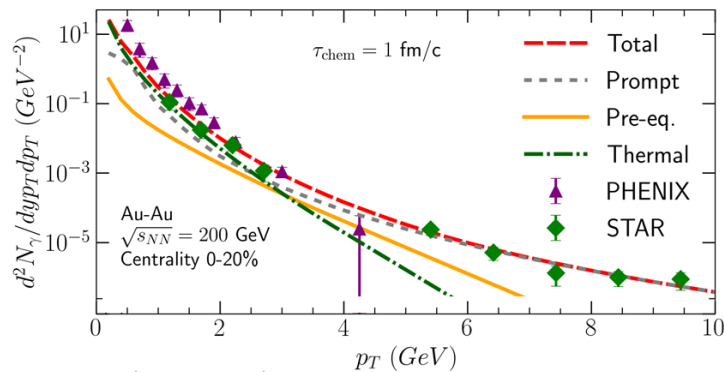
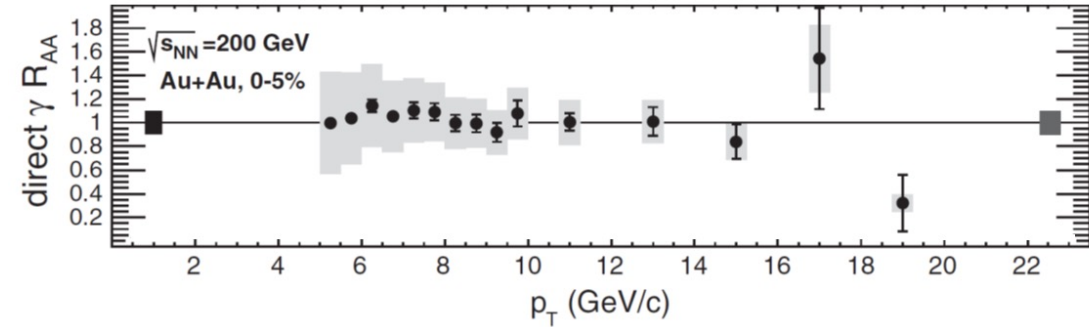
$$E_\gamma \sim 1 \text{ GeV}$$

$$T \sim 10^{12} \text{ K}$$

Summary and outlook

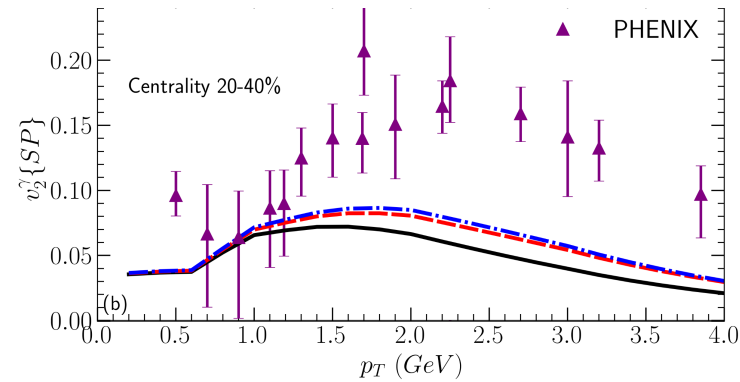
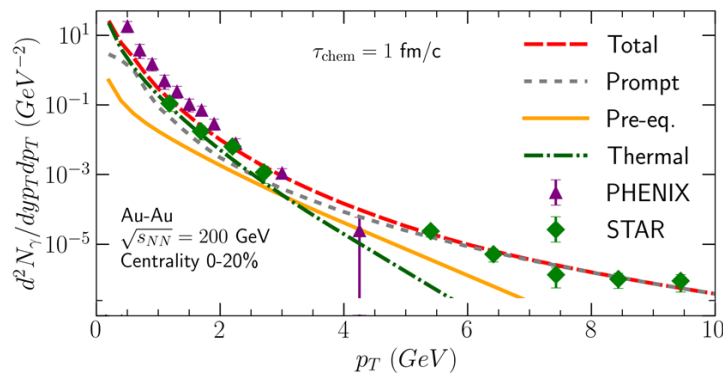
Summary

- High-energy photons: heavy-ion collisions similar to proton-proton case
- Low-energy photons:
 - Enhancement with respect to proton-proton collisions
 - Exponential spectrum \pm consistent with thermal radiation from $T_{\text{max}} \sim 500$ MeV deconfined plasma
 - Azimuthal anisotropy: important complementary information



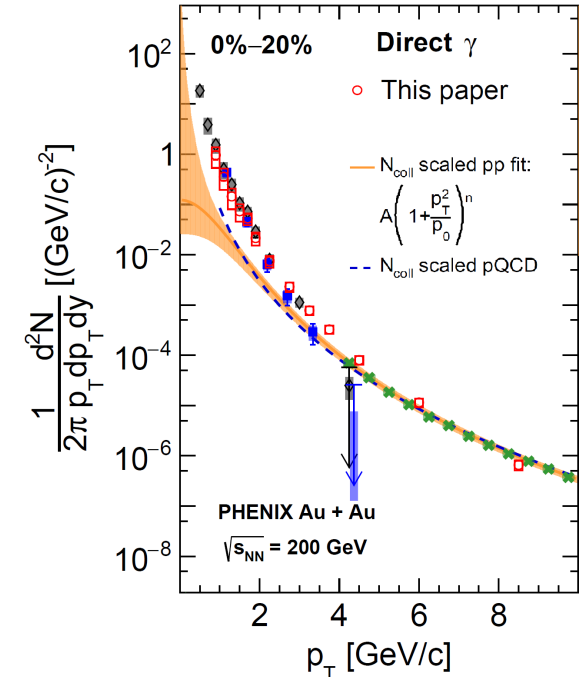
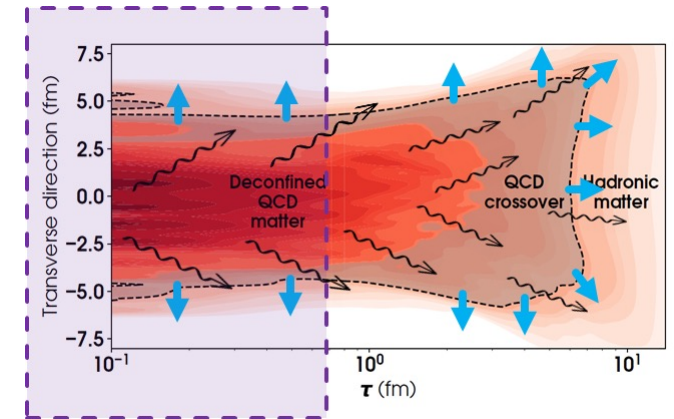
Outlook

- Studying the early stage of heavy-ion collisions with photons
- “Multi-messenger” study of heavy-ion collisions
- Understanding low p_T photons in proton-proton collisions?



- Many opportunities with dileptons as well

Gale, Paquet, Schenke, Shen (2022) PRC



Questions