

# Quarkonium production & modification: disentangling cold & hot nuclear effects in small collisions at LHC & RHIC

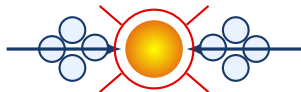
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# Motivation



**QGP fireball**

## Heavy-ion collisions

Pb+Pb/Au+Au create a hot, expanding QCD medium.

LHC: TeV scale; RHIC: 200 GeV.



cold QCD effects + possible QGP

## Small collision systems

p+Pb and O+O help isolate cold nuclear matter from possible hot-medium effects.

O+O: 5.36 TeV and 200 GeV.



## Quarkonia

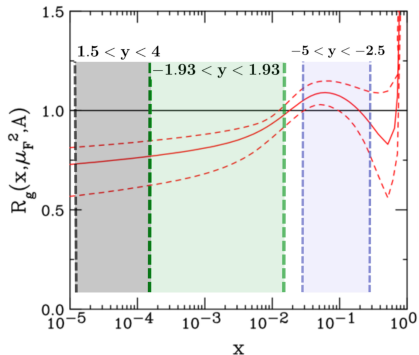
Produced early; survival encodes cold and hot QCD interactions, with clean dimuon decay channels.

## Observables:

$$R_{AA}(p_T, y) = \frac{1}{\langle N_{\text{coll}} \rangle} \frac{d^2 N_{AA}/dp_T dy}{d^2 N_{pp}/dp_T dy}$$

$$\mathcal{D}_{nS/1S}^{AA} = \frac{[N_{\Upsilon(nS)}/N_{\Upsilon(1S)}]_{AA}}{[N_{\Upsilon(nS)}/N_{\Upsilon(1S)}]_{pp}} \simeq \frac{R_{AA}^{\Upsilon(nS)}}{R_{AA}^{\Upsilon(1S)}}$$

# Cold Nuclear Matter Effects



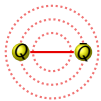
- ▶ **Nuclear PDFs:** parton densities inside bound nucleons differ from free protons.

$$f_i^A(x, Q^2) = R_i^A(x, Q^2) f_i^p(x, Q^2)$$

- ▶ **Momentum broadening (Cronin effect):** multiple scatterings increase transverse momentum,  $p_T^2 \rightarrow p_T^2 + \Delta p_T^2$ .
- ▶ **Energy loss:** pre-resonant  $Q\bar{Q}$  pairs lose energy (coherently) while crossing cold nuclear matter via medium-induced gluon radiation.
- ▶ **Nuclear absorption:** pre-resonant  $c\bar{c}$  may break up before exiting the nucleus. Mainly charmonia at RHIC; LHC formation times are longer, bottomonia compact.

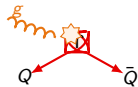
**CNM must be constrained; by itself it does not naturally explain strong  $\Upsilon(2S), \Upsilon(3S)$  suppression.**

# QGP Effects on Quarkonia



## Debye / static screening

Color force is screened; weak states melt first.



## Dynamical dissociation

Thermal particles break bound states, giving large widths.



## Recombination

Unbound heavy quarks can meet again and regenerate quarkonium.

### KSU–Munich: QQS approach

Open Quantum System + pNRQCD at NLO  
density matrix evolution

$$\dot{\rho} = -i[H, \rho] + \sum_n \left( C_n \rho C_n^\dagger - \frac{1}{2} \{ C_n^\dagger C_n, \rho \} \right).$$

Density matrix evolves in the QGP background.

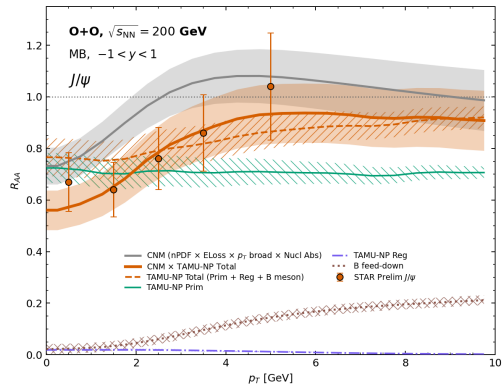
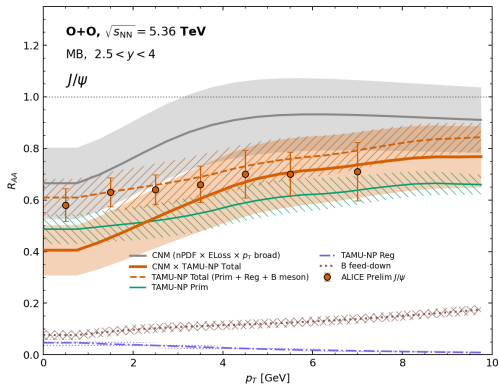
### TAMU: semiclassical transport

Rate equation with loss and gain:

$$\frac{dN_\Upsilon}{d\tau} = -\Gamma_\Upsilon(T) [N_\Upsilon - N_\Upsilon^{\text{eq}}(T)].$$

Gain term accounts for recombination/regeneration.

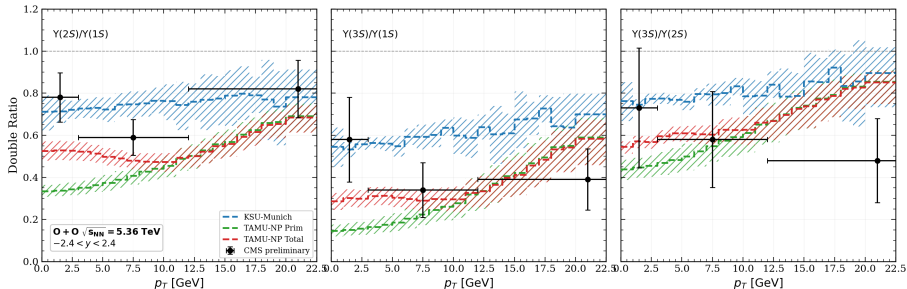
# O+O Charmonia: LHC and RHIC



Across LHC and RHIC energies, CNM constrains the cold-QCD contribution; hot-medium effects modify the final  $J/\psi$  yield.

Thapa, Wu, Fu, Vogt & Rapp: Charmonia production in O+O collisions at LHC & RHIC (in prep)

# O+O Bottomonia: Sequential Suppression



- ▶ Excited states are more fragile than  $\Upsilon(1S)$ .
- ▶ CNM alone does not naturally produce the hierarchy.

- ▶ Hot-medium suppression in addition to CNM is needed in the calculation.
- ▶ KSU-Munich and TAMU descriptions support a hot, short-lived QGP medium in small collisions.

Thapa, Wu, Vogt & Rapp:  $\Upsilon(nS)$  in 5.36 TeV O+O collisions (in prep)

# Thank you!

