ТΠ



European Research Council Established by the European Commission

# Constraining parameters of initial condition in isobar collisions

#### Seyed Farid Taghavi

#### Dense & Strange Hadronic Matter Group, Technical University of Munich, Germany

In collaboration with J. Jia

INT program: Intersection of nuclear structure and high - energy nuclear collisions Seattle February 2st, 2023

#### State-of-the-art heavy-ion collision models















# *Initial spatial anisotropy* — *Final momentum anisotropy*



$$\hat{\varepsilon}_{n,m} = -\frac{\{\rho^n e^{im\varphi}\}}{\{\rho^n\}}, \quad r_n = \{\rho^n\} \qquad (\varepsilon_n \equiv |\hat{\varepsilon}_{n,n}|, \quad r_{\mathsf{rms}}^2 \equiv r_2)$$
$$\frac{d^2 N}{p_T dp_T d\varphi} = N(p_T) \left[ 1 + \sum_{n=1}^{\infty} 2\nu_n \cos\left[n(\varphi - \psi_n)\right] \right]$$

**Flow harmonics**,  $(v_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s, ...), ...$ 



collective evolution



$$\hat{\boldsymbol{\varepsilon}}_{n,m} = -\frac{\{\boldsymbol{\rho}^n e^{im\boldsymbol{\varphi}}\}}{\{\boldsymbol{\rho}^n\}}, \quad r_n = \{\boldsymbol{\rho}^n\} \qquad (\boldsymbol{\varepsilon}_n \equiv |\hat{\boldsymbol{\varepsilon}}_{n,n}|, \quad r_{\mathsf{rms}}^2 \equiv r_2)$$
$$\frac{d^2N}{p_T dp_T d\boldsymbol{\varphi}} = N(p_T) \left[ 1 + \sum_{n=1}^{\infty} 2\boldsymbol{\nu}_n \cos\left[n(\boldsymbol{\varphi} - \boldsymbol{\psi}_n)\right] \right]$$

**Flow harmonics**,  $(v_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s, ...), ...$ 





$$\hat{\boldsymbol{\varepsilon}}_{n,m} = -\frac{\{\boldsymbol{\rho}^n e^{im\boldsymbol{\varphi}}\}}{\{\boldsymbol{\rho}^n\}}, \quad r_n = \{\boldsymbol{\rho}^n\} \qquad (\boldsymbol{\varepsilon}_n \equiv |\hat{\boldsymbol{\varepsilon}}_{n,n}|, \quad r_{\mathsf{rms}}^2 \equiv r_2)$$
$$\frac{d^2N}{p_T dp_T d\boldsymbol{\varphi}} = N(p_T) \left[ 1 + \sum_{n=1}^{\infty} 2\boldsymbol{\nu}_n \cos\left[n(\boldsymbol{\varphi} - \boldsymbol{\psi}_n)\right] \right]$$

**Flow harmonics**,  $(v_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s, ...), ...$ 



collective evolution



$$\hat{\boldsymbol{\varepsilon}}_{n,m} = -\frac{\{\boldsymbol{\rho}^n e^{im\boldsymbol{\varphi}}\}}{\{\boldsymbol{\rho}^n\}}, \quad r_n = \{\boldsymbol{\rho}^n\} \qquad (\boldsymbol{\varepsilon}_n \equiv |\hat{\boldsymbol{\varepsilon}}_{n,n}|, \quad r_{\mathsf{rms}}^2 \equiv r_2\}$$
$$\frac{d^2 N}{p_T dp_T d\boldsymbol{\varphi}} = N(p_T) \left[ 1 + \sum_{n=1}^{\infty} 2\boldsymbol{\nu}_n \cos\left[n(\boldsymbol{\varphi} - \boldsymbol{\psi}_n)\right] \right]$$

**Flow harmonics**,  $(v_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s, ...), ...$ 



 $(\varepsilon_n^{(3)}, r_{\rm rms}^{(3)}, \ldots)$ 

$$\hat{\varepsilon}_{n,m} = -\frac{\{\rho^n e^{im\varphi}\}}{\{\rho^n\}}, \quad r_n = \{\rho^n\} \qquad (\varepsilon_n \equiv |\hat{\varepsilon}_{n,n}|, \quad r_{\mathsf{rms}}^2 \equiv r_2$$
$$\frac{d^2 N}{p_T dp_T d\varphi} = N(p_T) \left[1 + \sum_{n=1}^{\infty} 2\nu_n \cos\left[n(\varphi - \psi_n)\right]\right]$$

**Flow harmonics**,  $(v_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s, ...), ...$ 

$$p([p_T], v_n, \psi_n, \ldots)$$



$$\hat{\varepsilon}_{n,m} = -\frac{\{\rho^n e^{im\varphi}\}}{\{\rho^n\}}, \quad r_n = \{\rho^n\} \qquad (\varepsilon_n \equiv |\hat{\varepsilon}_{n,n}|, \quad r_{\text{rms}}^2 \equiv r_2]$$
$$\frac{d^2N}{p_T dp_T d\varphi} = N(p_T) \left[ 1 + \sum_{n=1}^{\infty} 2\boldsymbol{\nu}_n \cos\left[n(\varphi - \boldsymbol{\psi}_n)\right] \right]$$

**Flow harmonics**,  $(\nu_n, \psi_n)$ , depend on the initial state parameters, transport coefficients  $(\eta/s, \zeta/s,..), ...$ 

$$p([p_T], v_n, \psi_n, \ldots)$$

$$v_n\{2\} \equiv (\langle v_n^2 \rangle)^{1/2}, \qquad v_n\{4\} \equiv \left(-\langle v_n^4 \rangle + 2\langle v_n^2 \rangle^2\right)^{1/4}, \qquad \cdots$$

[Borghini, Dinh, Ollitrault, PRC, 64, 054901 (2001)]

$$ho_n = rac{\langle [p_T] v_n^2 
angle - \langle [p_T] 
angle \langle v_n^2 
angle}{\sigma_{p_T} \sigma_{v_n^2}}$$

[Piotr Bozek, PRC (2016) 93, 044908 ]





## Theoretical models Vs experimental data

#### Initial state parameters

$N(\sqrt{s_{NN}})$	Overall normalization		
р . р	Entropy deposition parameter		
w	Gaussian nucleon width		

#### **Pre-equilibrium parameters**

#### **QGP** evolution parameters

 $\eta/s(T_c) \ (\eta/s)_{slope} \ (\eta/s)_{curve}$ 

 $\begin{array}{l} \text{Minimum } \eta/s(T) \\ \text{Slope of } \eta/s(T) \text{ above } T_c \\ \text{Curvature of } \eta/s(T) \text{ above } T_c \end{array}$ 

#### Hadronic gas evolution parameters



#### **Experimental observables**

1 1 1	
dN/dy	Particle yields, $\pi^{\pm}$ , $k^{\pm}$ ,
$\langle [p_T] \rangle$	Mean transverse momentum, $\pi^{\pm}$ , $k^{\pm}$ ,
$v_n\{2\}$	Anisotropic flow two-particle correlation
$v_n{4}$	Anisotropic flow four-particle correlation

## Model in the light of experimental data



Bernhard, PhD Thesis, arXiv: 1804.06469; Bernhard, Moreland, Bass, Nature Phys. 15 (2019) 11, 1113-1117
 Auvinen, et al., PRC 102 (2020) 044911, Nijs et al., PRL 126 (2021) 202301, JETSCAPE, PRC 103 (2021) 054904
 JETSCAPE, PRC 104 (2021), 024905
 Mäntysaari, Schenke, Shen, Zhao, arXiv: 2202.01998

## Model in the light of experimental data





#### Theoretical developments: collectivity [2], jet-quenching [3], nucleon substructure [4]

Bernhard, PhD Thesis, arXiv: 1804.06469; Bernhard, Moreland, Bass, Nature Phys. 15 (2019) 11, 1113-1117
 Auvinen, et al., PRC 102 (2020) 044911, Nijs et al., PRL 126 (2021) 202301, JETSCAPE, PRC 103 (2021) 054904
 JETSCAPE, PRC 104 (2021), 024905
 Mäntysaari, Schenke, Shen, Zhao, arXiv: 2202.01998

Initial state, TRENTO model [Moreland, Bernhard, Bass, PRC 92 (2015), 011901; Moreland, Bernhard, Bass, PRC 101 (2020), 024911]

Distribute nucleons based on the Woods-Saxson distribution,

$$WS(r,\theta,\phi) = \frac{n_0}{1 + e^{[r-R(\theta,\phi)]/a_0}}, \qquad R(\theta,\phi) = R_0 \left(1 + \beta_2 Y_2^0 + \beta_3 Y_3^0 + \beta_4 Y_4^0 + \cdots\right)$$

- ▶ We impose a constraint on the minimum distance that two nucleons can have *d*<sub>min</sub>.
- ► It is assumed that nucleons have a Gaussian shape  $\rho_N(\vec{r}) \propto \exp\left[-\frac{r^2}{2w^2}\right]$  with width w.
- Two nucleons *participate* in a collision with probability

$$P_{\rm coll} = 1 - \exp\left[-\sigma_{gg}\int dxdy\int dz \rho_{N_1}(\vec{r})\int dz \rho_{N_1}(\vec{r})\right]. \label{eq:P_coll}$$

 $\sigma_{gg}$  is fixed by the nucleon nucleon total cross section measurements.

- Adding up participants to make  $\rho_{A,B}(\vec{r})$ , participant thickness function is defined as  $T_{A,B}(x,y) = \int dz \rho_{A,B}(\vec{r})$
- The deposited entropy into the collision region is obtained via

$$T_R(p,T_A,T_B) = \left(\frac{T_A^p + T_B^p}{2}\right)^{1/p}, \qquad \langle N_{\mathsf{Ch}} \rangle \propto \int dx dy T_R(x,y).$$

p qualitatively controls the mechanism of entropy production during the collision.

The parameters: w,  $d_{min}$ , p,  $R_0$ ,  $a_0$ ,  $\beta_2$ ,  $\beta_3$ , ...

(and some more:  $\sigma_{fluc}$ : subnucleonic structure, . . . )

#### Linear response hvdrodvnamics



[Niemi, Denicol, Holopainen, Huovinen, PRC 87 (2013) 5, 054901]

#### Transverse momentum linear response [1,2]



For an observable *O*, define event-by-event deviation and variance  $\delta O \equiv O - \langle O \rangle$ ,  $\sigma_O^2 = \langle O^2 \rangle - \langle O \rangle^2$ . Considering the overlap region  $A_{\perp} = \pi r_{rms}^2 \sqrt{1 - \varepsilon_2^2}$ , one can define a predictor  $d_{\perp} = \sqrt{N_{\text{part}}/A_{\perp}}$ . Linear approximation for deviation: Linear approximation for standard deviation: Linear approximation for average?  $\langle [p_T] \rangle \approx k_0 \frac{\sigma_{d_T}}{\langle d_T \rangle}$ 

Broniowski, Chojnacki, Obara PRC, 80 (2009), 051902; Bozek, Broniowski PRC 96 (2017), 014904
 Schenke, Shen, Teaney, PRC, 102 (2020), 034905

#### Isobar Ratios, a good choice for the linear response approximations



The ratio of observables should have very low sensitivity to the collective evolution.

#### If we are lucky enough:

the linear response coefficients, k<sub>2</sub>, k<sub>3</sub>, k<sub>0</sub>, k'<sub>0</sub>, are constants at least in some range of multiplicity

$$\frac{v_n\{2\}|_{Ru}}{v_n\{2\}|_{Zr}}\approx\frac{\varepsilon_n\{2\}|_{Ru}}{\varepsilon_n\{2\}|_{Zr}},$$



10/17

#### Setup

Scanning the parameter space:

w, 
$$d_{min}$$
,  $p$ ,  $R_0$ ,  $a_0$ ,  $\beta_2$ ,  $\beta_3$ 

Nuclei size and shape:

	$R_0$ [fm]	$a_0[fm]$	$\beta_2$	$\beta_3$
Ru96	5.09	0.46	0.162	0
Rull	5.09	0.46	0.06	0
RullI	5.09	0.46	0.06	0.2
RulV	5.09	0.52	0.06	0.2
Zr96	5.02	0.52	0.06	0.2

Initial state internal structure:

$$(w, d_{\min}, p) \in \{0.2, 0.3, \dots, 0.9\} \otimes \{0., 0.2 \dots, 1.0\} \otimes \{-1, 0, 1\}.$$

30M minimum bias events per each point, 240 points in total.

#### **Multiplicity distribution ratio**

 $\langle N_{ch} \rangle = N \int dx dy T_R(x, y)$ Finding overall normalization using experimental measurements.



12/17

#### **Results:**



# Best fit with $\chi^2$ , only $v_2$ {2}





14/17









































15/17

















## Conclusion / Summary / Outlook

# $v_2$ {2} isobar ratio is sensitivity to *w*, $d_{min}$ initial state parameters.

The best fit point for high multiplicity events:

$$w = 0.6$$
 [fm],  $d_{\min} = 0.6$  [fm]

Nucleon minimum distance:

Nucleon width:

- $0.4 \lesssim d_{\min} \lesssim 0.6 \,[\text{fm}]$  $0.4 \leq w \leq 0.8 \,[\text{fm}]$
- Nucleon size and nucleon minimum distance from other studies:

	w [fm]	d <sub>min</sub> [fm]	Ref(s).
Bayesian analysis	$\sim 0.8$ to $\sim 1.0$	$\sim 0.5$ to $\sim 1.5$	[1]
$\rho_n([p_T], v_n^2)$	$\sim 0.5$	-	[2]
nucleus-nucleus cross-section	$\sim 0.7$	-	[3]

- The effect of other parameters need to be explored.
- Hydrodynamic calculation increase the prediction power by including lower multiplicity events.

- [2] Giacalone, Schenke, Shen, PRL 128 (2022) 4, 042301
- [3] Nijs, van der Schee, PRL, 129 (2022) 23, 232301

<sup>[1]</sup> Bernhard, Moreland, Bass, Nature Phys. 15 (2019) 11, 1113-1117; Nijs, van der Schee, Gürsoy, Snellings, PRL, 126, 2010.15130; JETSCAPE, PRC, 103 (2021), 054904; Parkkila, Onnerstad, Kim, PRC 104 (2021) 5, 054904; Parkkila, Onnerstad, Taghavi, Mordasini, Bilandzic, Kim, Virta, PLB 835 (2022) 137485.

# Thank You!

# **Backup Slides**

#### References of slide in page 2

[IP-Glasma] Bjoern Schenke, Prithwish Tribedy, and Raju Venugopalan, Phys. Rev. Lett. 108, 252301 (2012); Bjoern Schenke, Prithwish Tribedy, and Raju Venugopalan, Phys. Rev. C 86, 034908 (2012)

[T<sub>R</sub>ENTo] J. Scott Moreland, Jonah E. Bernhard, and Steffen A. Bass, Phys. Rev. C 92, 011901 (2015)

[MC-Glauber] Wojciech Broniowski, Maciej Rybczynski, and Piotr Bozek, Comput. Phys. Commun. 180, 69?83 (2009)

[MC-KLN] H. J. Drescher and Y. Nara, Phys. Rev. C 75, 034905 (2007); Hans-Joachim Drescher and Yasushi Nara, Phys. Rev. C 76, 041903 (2007) [Free-streaming] Jonah E. Bernhard. PhD Thesis. arXiv: 1804.06469:

[KøMPøST] Aleksi Kurkela, Aleksas Mazeliauskas, Jean-François Paquet, Sören Schlichting, and Derek Teaney, Phys. Rev. Lett. 122, 122302 (2019)

[Gauge/gravity] W. van der Schee, P. Romatschke, S. Pratt, Phys. Rev. Lett. 111, 222302 (2013)

[VISH2+1] Huichao Song, Steffen A. Bass, and Ulrich Heinz, Phys. Rev. C 83, 024912 (2011)

[MUSIC] Bjoern Schenke, Sangyong Jeon, and Charles Gale, Phys. Rev. C 82, 014903 (2010)

[Trajectum] G. Nijs, W. van der Schee, U. Gürsoy, R. Snellings, Phys.Rev.C 103, 054909 (2021)

[VH2+1] Matthew Luzum and Paul Romatschke, Phys. Rev. C 78, 034915 (2008). [Erratum: Phys.Rev.C 79, 039903 (2009)]; Paul Romatschke and Ulrike Romatschke, Phys. Rev. Lett. 99, 172301 (2007)

[UrQMD] M. Bleicher et al., J. Phys. G 25, 1859?1896 (1999)

[SMASH] J. Weil et al., Phys. Rev. C 94, 054905 (2016)

[B3D] John Novak, Kevin Novak, Scott Pratt, Joshua Vredevoogd, Chris Coleman-Smith, and Robert Wolpert, Phys. Rev. C 89, 034917 (2014)

#### Add *p* dependent



#### Add *p* dependent

