

SMEFT projections using EIC PVDIS asymmetries

Kağan Şimşek

Northwestern
University

in collaboration with

Radja Boughezal, Alexander Emmert, Tyler Kutz, Sonny Mantry,
Michael Nycz, Frank Petriello, Daniel Wiegand, and Xiaochao Zheng

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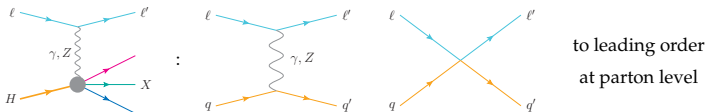
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- We study NC DIS cross-section asymmetries at EIC.
- We adopt SMEFT framework to parameterize BSM effects.
- Higher-dimensional operators are built of existing SM particles:

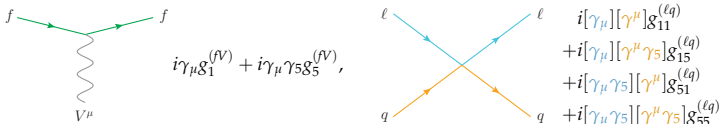
$$\mathcal{L}_{\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \sum_{n>4} \frac{1}{\Lambda^{n-4}} \sum_k C_k^{(n)} O_k^{(n)}$$

- All new physics is assumed to be heavier than SM states and accessible collider energy.
- We focus on $n = 6$ and semi-leptonic 4-fermion $O_k^{(n)}$.
- We find that the EIC can
 - probe complementarily and competitively to LHC DY
 - resolve degeneracies observed in LHC NC DY data fits

We study the NC DIS in the process $\ell + H \rightarrow \ell' + X$, where $\ell = e^-, e^+$ and $H = p, D$:



We parameterize the vertex factors in terms of vector and axial couplings:



We don't consider Yukawa or dipole interactions because they are suppressed by fermion masses, which we assume to vanish.

SMEFT operators shift the usual vector and axial SM couplings in a gauge-invariant way: e.g.

$$g_1^{(fZ)} = g_V^f + \mathcal{O}(C_k), \quad g_5^{(fZ)} = g_A^f + \mathcal{O}(C_k)$$

Operators that contribute to the ffV and $\ell\ell qq$ vertices at dimension 6 are (Grzadkowski *et al.* [[1008.4884](#)]):

ffV	$\ell\ell qq$
$O_{\phi\ell}^{(1)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi)(\bar{\ell}\gamma^\mu\ell)$	$O_{\ell q}^{(1)} = (\bar{\ell}\gamma_\mu\ell)(\bar{q}\gamma^\mu q)$
$O_{\phi\ell}^{(3)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \tau^I \varphi)(\bar{\ell}\gamma^\mu\tau^I\ell)$	$O_{\ell q}^{(3)} = (\bar{\ell}\gamma_\mu\tau^I\ell)(\bar{q}\gamma^\mu\tau^I q)$
$O_{\phi e} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi)(\bar{e}\gamma^\mu e)$	$O_{eu} = (\bar{e}\gamma_\mu e)(\bar{u}\gamma^\mu u)$
$O_{\phi q}^{(1)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi)(\bar{q}\gamma^\mu q)$	$O_{ed} = (\bar{e}\gamma_\mu e)(\bar{d}\gamma^\mu d)$
$O_{\phi q}^{(3)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \tau^I \varphi)(\bar{q}\gamma^\mu\tau^I q)$	$O_{\ell u} = (\bar{\ell}\gamma_\mu\ell)(\bar{u}\gamma^\mu u)$
$O_{\phi u} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi)(\bar{u}\gamma^\mu u)$	$O_{\ell d} = (\bar{\ell}\gamma_\mu\ell)(\bar{d}\gamma^\mu d)$
$O_{\phi d} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi)(\bar{d}\gamma^\mu d)$	$O_{qe} = (\bar{q}\gamma_\mu q)(\bar{e}\gamma^\mu e)$

There is one more:

$$O_{\phi WB} = (\varphi^\dagger \tau^I \varphi) W_{\mu\nu}^I B^{\mu\nu} \Rightarrow \text{causes kinetic mixing of } W^3 \text{ and } B$$

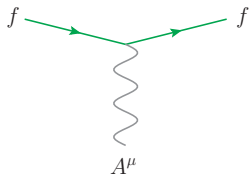
$$\Rightarrow \text{universally shifts the } ffV \text{ vertices after diagonalization that gives physical photon and Z boson states}$$

The ffV operators are already strongly bounded by Z and W pole observables (Dawson & Giardino [1909.02000]):

ffV	C_k	95% CL, $\Lambda = 1$ TeV
$O_{\phi\ell}^{(1)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi) (\bar{\ell} \gamma^\mu \ell)$	$C_{\phi\ell}^{(1)}$	$[-0.043, 0.012]$
$O_{\phi\ell}^{(3)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \tau^I \varphi) (\bar{\ell} \gamma^\mu \tau^I \ell)$	$C_{\phi\ell}^{(3)}$	$[-0.012, 0.0029]$
$O_{\phi e} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi) (\bar{e} \gamma^\mu e)$	$C_{\phi e}$	$[-0.013, 0.0094]$
$O_{\phi q}^{(1)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi) (\bar{q} \gamma^\mu q)$	$C_{\phi q}^{(1)}$	$[-0.027, 0.043]$
$O_{\phi q}^{(3)} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \tau^I \varphi) (\bar{q} \gamma^\mu \tau^I q)$	$C_{\phi q}^{(3)}$	$[-0.011, 0.014]$
$O_{\phi u} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi) (\bar{u} \gamma^\mu u)$	$C_{\phi u}$	$[-0.072, 0.091]$
$O_{\phi d} = (\varphi^\dagger i \overleftrightarrow{D}_\mu \varphi) (\bar{d} \gamma^\mu d)$	$C_{\phi d}$	$[-0.16, 0.060]$
$O_{\phi WB} = (\varphi^\dagger \tau^I \varphi) W_{\mu\nu}^I B^{\mu\nu}$	$C_{\phi WB}$	$[-0.0088, 0.0013]$

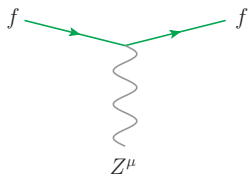
Thus, we restrict our attention only to the operators contributing to the $\ell\ell q\bar{q}$ vertex, which leaves us with seven Wilson coefficients of interest: $C_{eu}, C_{ed}, C_{\ell q}^{(1)}, C_{\ell q}^{(3)}, C_{\ell u}, C_{\ell d},$ and C_{qe} .

Since we consider contributions only to the $\ell\ell qq$ interaction, we assume the usual SM ffV vertices in our analysis:



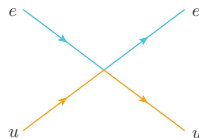
$$g_1^{(fA)} = -eQ_f$$

$$g_5^{(fA)} = 0$$



$$g_1^{(fZ)} = g_V^f$$

$$g_5^{(fZ)} = g_A^f$$

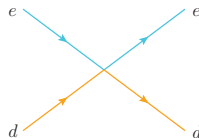


$$g_{11}^{(eu)} = \frac{1}{4} [C_{eu} + (C_{\ell q}^{(1)} - C_{\ell q}^{(3)}) + C_{\ell u} + C_{qe}]$$

$$g_{15}^{(eu)} = \frac{1}{4} [C_{eu} - (C_{\ell q}^{(1)} - C_{\ell q}^{(3)}) + C_{\ell u} - C_{qe}]$$

$$g_{51}^{(eu)} = \frac{1}{4} [C_{eu} - (C_{\ell q}^{(1)} - C_{\ell q}^{(3)}) - C_{\ell u} + C_{qe}]$$

$$g_{55}^{(eu)} = \frac{1}{4} [C_{eu} + (C_{\ell q}^{(1)} - C_{\ell q}^{(3)}) - C_{\ell u} - C_{qe}]$$



the same as for $eeuu$ but with $u \rightarrow d$ and

$$C_{\ell q}^{(1)} - C_{\ell q}^{(3)} \rightarrow C_{\ell q}^{(1)} + C_{\ell q}^{(3)}$$

Amplitude and cross section for $\ell + q \rightarrow \ell' + q'$:

$$\mathcal{M} = \mathcal{M}_\gamma + \mathcal{M}_Z + \mathcal{M}_\times \Rightarrow d\sigma^{\lambda_\ell \lambda_q} = \frac{d^2\sigma}{dx dQ^2} = \frac{1}{16\pi x^2 s^2} |\mathcal{M}|^2 + \mathcal{O}(C_k^2)$$

Asymmetry definitions:

- unpolarized PV asymmetries: $A_{PV} = \frac{d\sigma_\ell}{d\sigma_0}$
- polarized PV asymmetries: $\Delta A_{PV} = \frac{d\sigma_H}{d\sigma_0}$
- lepton-charge asymmetries: $A_{LC} = \frac{d\sigma_0(e^+H) - d\sigma_0(e^-H)}{d\sigma_0(e^+H) + d\sigma_0(e^-H)}$

where

$$d\sigma_0 = \frac{1}{4} \sum_q f_{q/H} [d\sigma^{++} + d\sigma^{+-} + d\sigma^{-+} + d\sigma^{--}] : \text{unpol. } \ell + \text{unpol. } H$$

$$d\sigma_\ell = \frac{1}{4} \sum_q f_{q/H} [d\sigma^{++} + d\sigma^{+-} - d\sigma^{-+} - d\sigma^{--}] : \text{pol. } \ell + \text{unpol. } H$$

$$d\sigma_H = \frac{1}{4} \sum_q \Delta f_{q/H} [d\sigma^{++} - d\sigma^{+-} + d\sigma^{-+} - d\sigma^{--}] : \text{unpol. } \ell + \text{pol. } H$$

Data sets, shown with beam energies and nominal annual luminosities:

D1	5 GeV × 41 GeV eD , 4.4 fb ⁻¹
D2	5 GeV × 100 GeV eD , 36.8 fb ⁻¹
D3	10 GeV × 100 GeV eD , 44.8 fb ⁻¹
D4	10 GeV × 137 GeV eD , 100 fb ⁻¹
D5	18 GeV × 137 GeV eD , 15.4 fb ⁻¹
P1	5 GeV × 41 GeV ep , 4.4 fb ⁻¹
P2	5 GeV × 100 GeV ep , 36.8 fb ⁻¹
P3	10 GeV × 100 GeV ep , 44.8 fb ⁻¹
P4	10 GeV × 275 GeV ep , 100 fb ⁻¹
P5	18 GeV × 275 GeV ep , 15.4 fb ⁻¹
P6	18 GeV × 275 GeV ep , 100 fb ⁻¹

P6: Yellow Report reference setting [[2103.05419](#)]

Data set labels:

D, P: unpolarized PV asymmetry

ΔD, ΔP: polarized PV asymmetry

LD, LP: lepton-charge asymmetry

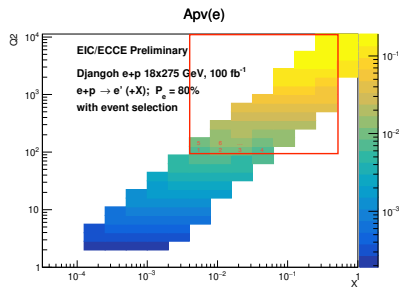
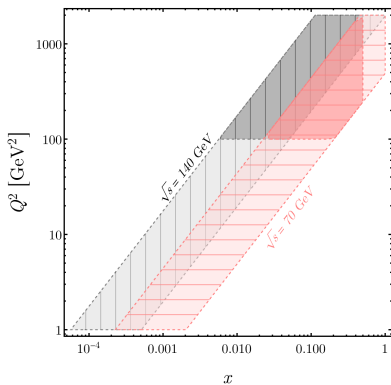
Cuts on projected data:

$Q > 1 \text{ GeV}$	to avoid nonperturbative QCD
$y > 0.1$	to avoid bin migration and unfolding uncertainty
$y < 0.9$	to avoid high photoproduction background due to final-state hadron
$ \eta < 3.5$	to restrict events in main acceptance of ECCE detector
$E' > 2 \text{ GeV}$	to have high-purity e^- samples

Additional cuts in SMEFT analysis:

$x < 0.5$	to avoid <i>large</i> uncertainties from nonperturbative QCD and nuclear dynamics
$Q > 10 \text{ GeV}$	

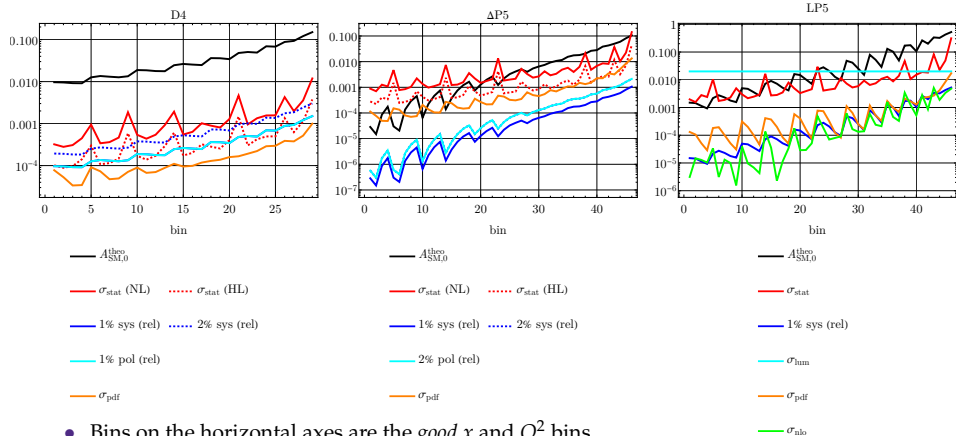
Phase space of the data sets ($\sqrt{s} = 70\text{-}140\text{ GeV}$, $0.1 \leq y \leq 0.9$):



The red box indicates the *good* bins used in our SMEFT analysis.

Error type	$A_{PV}(D, P)$	$\Delta A_{PV}(\Delta D, \Delta P)$	$A_{LC}(LD, LP)$
statistical (NL)	$\sigma_{\text{stat}} = \frac{1}{P_\ell \sqrt{N}}$	$\frac{P_\ell}{P_H} \sigma_{\text{stat}}$	$\sqrt{10} P_\ell \sigma_{\text{stat}}$
statistical (HL)	$\frac{1}{\sqrt{10}} \sigma_{\text{stat}}$	$\frac{1}{\sqrt{10}} \frac{P_\ell}{P_H} \sigma_{\text{stat}}$	x
uncorrelated systematic	1% rel.	1% rel.	1% rel.
fully correlated beam polarization	1% rel.	2% rel.	x
fully correlated luminosity	x	x	2% abs.
uncorrelated QED NLO	x	x	$5\% \times (A_{LC}^{\text{NLO}} - A_{LC}^{\text{Born}})$
fully correlated PDF	✓	✓	✓

PDF sets used: NNPDF3.1 NLO and NNPDFpo11.1



- Bins on the horizontal axes are the *good* x and Q^2 bins.
- Stat error dominates in PV asymmetries in NL case.
- Systematic and **beam-polarization** errors become comparable to stat error in HL case.
- **Luminosity** error dominates in LC asymmetries.
- Stat error competes with **luminosity** error at high- x high- Q^2 bins.
- PDF errors are the least dominant in **unpolarized** PV asymmetries but become significant in the **polarized** case.

Pseudodata generation:

$$A_{\text{pseudo},b}^{(e)} = A_{\text{SM},b} + r_b^{(e)} \sigma_b^{\text{unc}} + r'^{(e)} \sigma_b^{\text{cor}}$$

$$b \in \text{Range}(N_{\text{bin}}), \quad e \in \text{Range}(N_{\text{exp}}), \quad N_{\text{exp}} = 10^3, \quad r_b^{(e)}, r'^{(e)} \sim \mathcal{N}(0, 1)$$

$$\sigma_b^{\text{unc}} = \sigma_{\text{stat},b} \oplus \sigma_{\text{sys},b}$$

$$\sigma_b^{\text{unc}} = \sigma_{\text{stat},b} \oplus \sigma_{\text{sys},b} \oplus \sigma_{\text{nlo},b}$$

$$\sigma_b^{\text{cor}} = \sigma_{\text{pol},b}$$

$$\sigma_b^{\text{cor}} = \sigma_{\text{lum},b}$$

SMEFT asymmetry expressions:

$$A_{\text{SMEFT},b} = A_{\text{SM},b} + \sum_{k=1}^{N_{\text{fit}}} C_k \delta A_{k,b} + \mathcal{O}(C_k^2), \quad N_{\text{fit}} \in \text{Range}(7)$$

χ^2 function for each pseudoexperiment:

$$\chi^2{}^{(e)} = \sum_{b,b'=1}^{N_{\text{bin}}} [A_{\text{SMEFT},b} - A_{\text{pseudo},b}^{(e)}] H_{bb'} [A_{\text{SMEFT},b'} - A_{\text{pseudo},b'}^{(e)}]$$

Polarimetry and **luminosity difference** can be limiting factors.

- ⇒ use data itself to constrain these systematic effects
- ⇒ simultaneous fits of C_k with **beam polarization**, P , and **luminosity difference**, A_{lum} , in an attempt to obtain stronger bounds for C_k

Fits of C_k with P :

$$\chi^2(e) = \sum_{b,b'=1}^{N_{\text{bin}}} [PA_{\text{SMEFT},b} - A_{\text{pseudo},b}^{(e)}] \left[H_{bb'} \Big|_{\sigma_{\text{pol}} \rightarrow 0} \right] [PA_{\text{SMEFT},b'} - A_{\text{pseudo},b'}^{(e)}] + \frac{(P - \bar{P})^2}{\delta P^2}$$

unpolarized PV asymmetries:

- $|\rho(C_k, P)| \gtrsim 0.7$
- **30-50% stronger bounds**

polarized PV asymmetries:

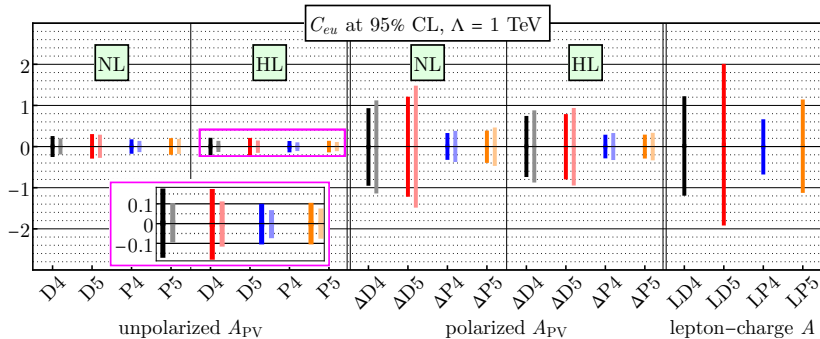
- $|\rho(C_k, P)| \lesssim 0.2$
- **15-20% weaker bounds**

Improvement is more significant than worsening ⇒ include P in fits.

Fits of C_k with A_{lum} :

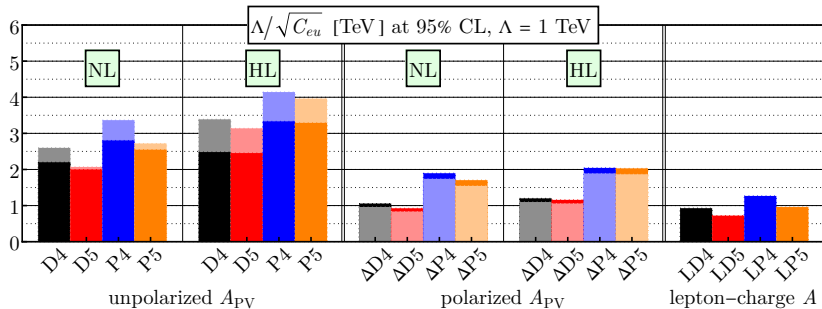
$$\chi^2(e) = \sum_{b,b'=1}^{N_{\text{bin}}} [A_{\text{SMEFT},b} - A_{\text{pseudo},b}^{(e)} - A_{\text{lum}}] \left[H_{bb'} \Big|_{\sigma_{\text{lum}} \rightarrow 0} \right] [A_{\text{SMEFT},b'} - A_{\text{pseudo},b'}^{(e)} - A_{\text{lum}}]$$

Mild correlations, $|\rho(C_k, A_{\text{lum}})| \lesssim 0.4$, leading to **15-20% weaker bounds**
 ⇒ do not include A_{lum} in fits.



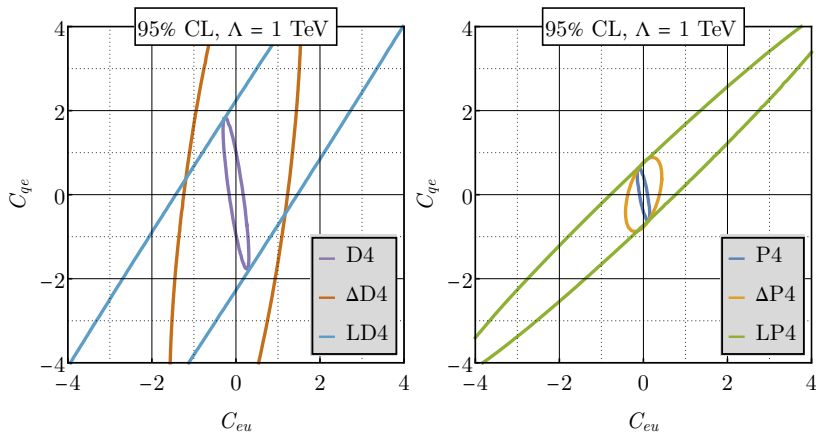
In terms of the strength of bounds:

- proton > deuteron
- low-lum. high-energy (4th sets) > high-lum. low-energy (5th sets)
- unpolarized PV > polarized PV > lepton-charge
- improvement: unpolarized PV > polarized PV if NL \rightarrow HL



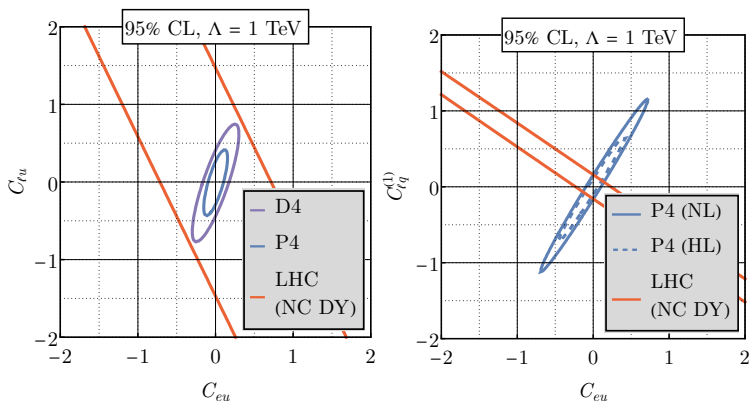
- UV scales ~ 3 TeV in NL case
- UV scales ~ 4 TeV in HL case

Compare the bounds from deuteron vs. proton data in the nominal-luminosity case for all the three types of asymmetries:



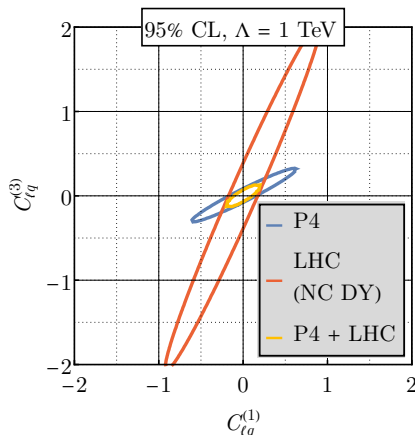
- Unpolarized PV asymmetries lead to strongest bounds.
- Proton data imposes stronger bounds.

Compare the bounds from deuteron and proton data of unpolarized PV asymmetries to the 8-TeV 20-fb⁻¹ LHC NC DY data (Boughezal, Petriello, & Wiegand [2004.00748, 2104.03979]):



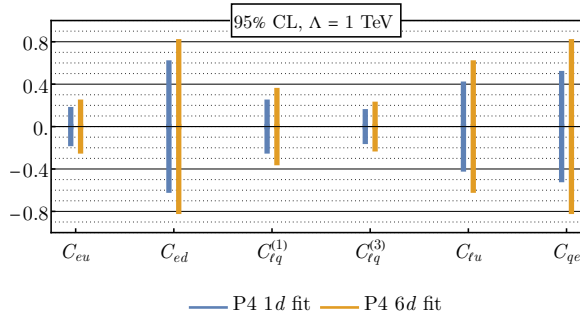
The LHC fits are highly degenerate and exhibit a flat direction, which remain even in the high-luminosity case. The EIC can resolve and constrain this parameter space strongly.

Compare proton data of **unpolarized** PV asymmetries to the 8-TeV 20-fb⁻¹ LHC NC DY data (Boughezal, Petriello, & Wiegand [2004.00748]) when the LHC fit doesn't have a flat direction:



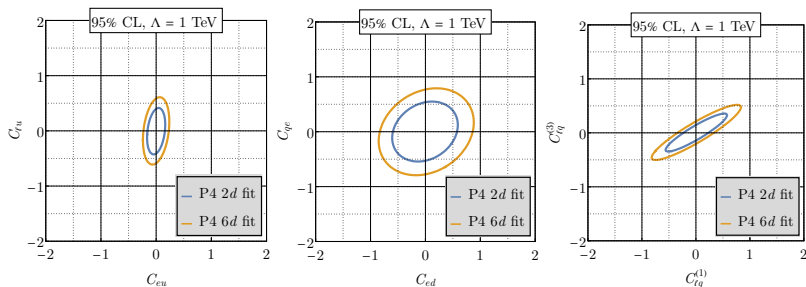
When the LHC fit gives a strong bound without showing a flat direction, the EIC can constrain the same parameter space even more strongly.

N_{fit}	N_{exp}
2	10^3
3	10^4
4	10^5
5	10^6
6	10^7
7	10^8 (?!)



- Number of pseudoexperiments increases to reflect the required statistics.
- **Beam polarization** parameter, P , is not included here.
- Bounds become 25 to 40% weaker due to increased number of fitted parameters and correlations among them.

Compare the two-parameter fits of Wilson coefficients to the projections from a six-parameter fit:



- The $eeuu$ vertex contains the combination $C_{\ell q}^{(1)} - C_{\ell q}^{(3)}$ and the $eedd$ vertex has $C_{\ell q}^{(1)} + C_{\ell q}^{(3)}$.
- These may lead to degeneracies and flat directions in a multi-parameter fits of Wilson coefficients.
- The EIC can resolve this part of the parameter space, imposing strong bounds.

- We investigate the BSM potential of EIC in the model-independent SMEFT framework by focusing on semi-leptonic four-fermion operators at dimension 6 by giving a detailed accounting of uncertainties.
- We obtain bounds on Wilson coefficients from single-, double-, and even multiple-parameter fits by using techniques to simultaneously fit P and A_{lum} together with SMEFT parameters.

- We find that UV scales up to 3 TeV (or 4 TeV) can be probed with nominal (or $10\times$ high) annual luminosity.
- We observe that the strongest bounds come from **unpolarized** PV asymmetries of proton.
- EIC is shown to be complementary and competitive to LHC NC DY by
 - equally or more strongly confining the Wilson coefficients with distinct correlations;
 - resolving the degeneracies observed in the LHC data.

EIC was designed as a QCD machine and it shows strong potential for BSM physics.