## Flow in the FXT region

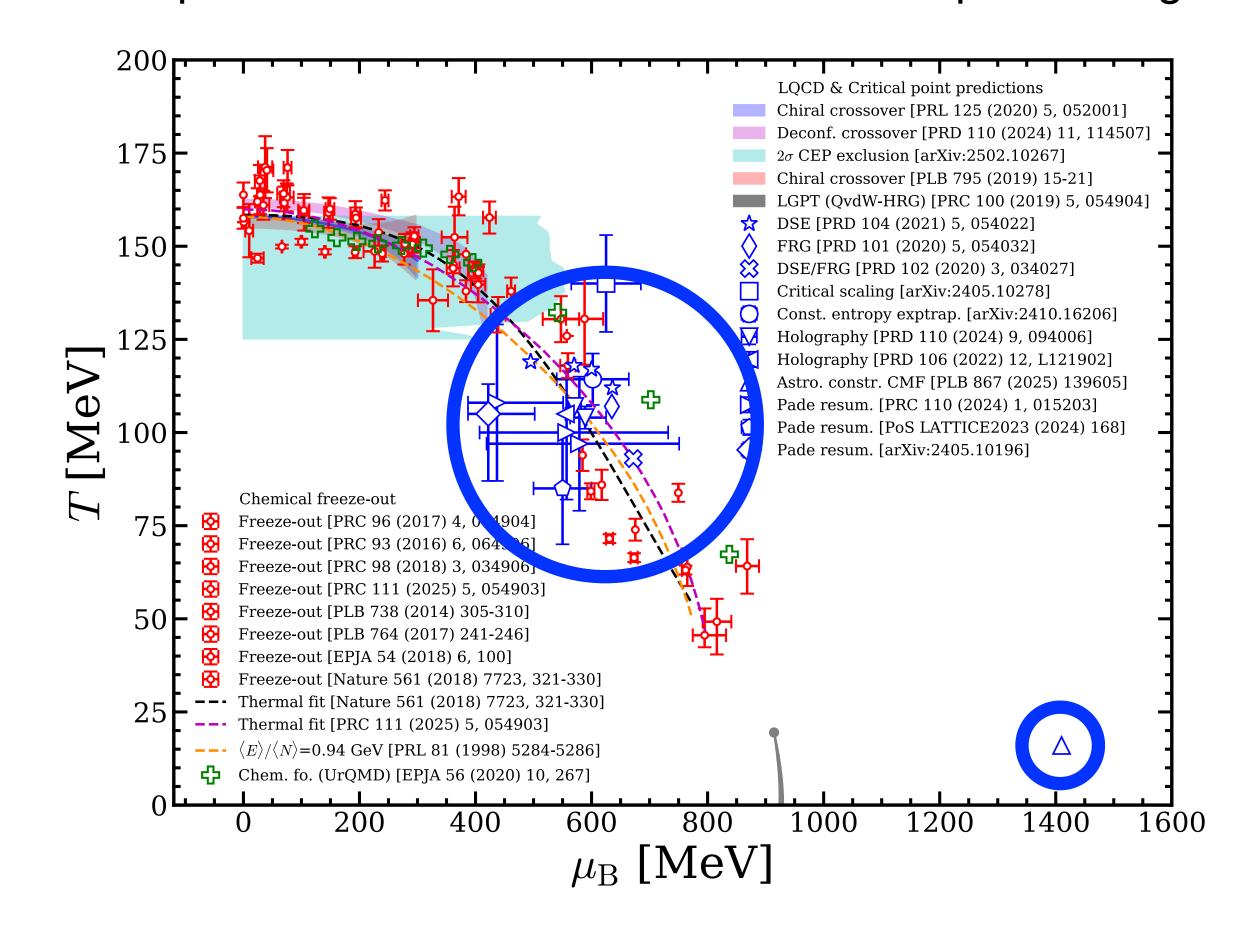
The QCD Critical Point: Are We There Yet?

T. Reichert, R. Seto, and A. Sorensen

Institute for Nuclear Theory (INT), University of Washington, Seattle

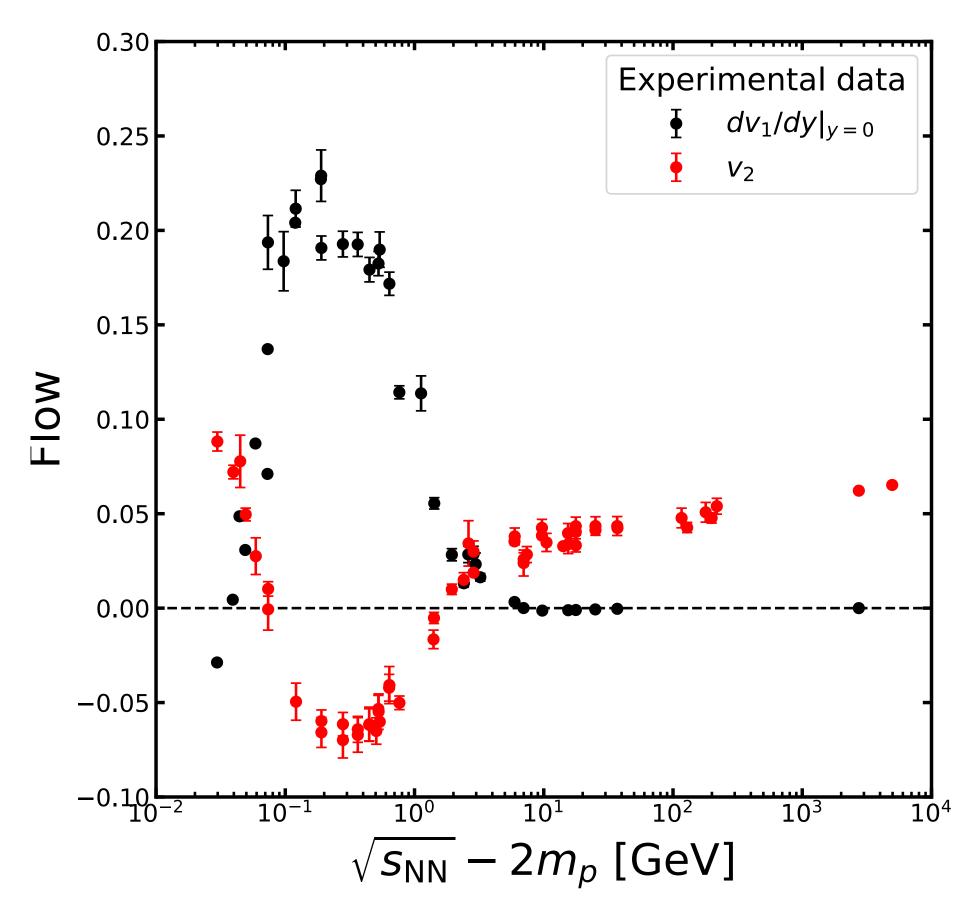
## QCD critical point: Are we there yet?

#### Compilation of statements about the QCD phase diagram



 $\mu_{B,c} \in$  (400, 800) MeV,  $T_c \in$  (75, 125) MeV corresponds to freeze-out parameters for  $\sqrt{s} \in$  (4, 7) GeV

#### Compilation of $p_T$ -integrated flow measurements



Flow varies a lot in this region:

- reflects rich (complex) physics
- opportunity to extract the EOS IF model systematics are understood

## STAR data: what is available now

Published or "official" preliminary: differential measurements

```
FXT: \sqrt{s} = 3.0 - 4.5 \text{ GeV}
Collider: \sqrt{s} = 7.7 - 200 \text{ GeV}
```

Flow Harmonics: using reaction plane method (mostly to  $\Psi_1$ )

- $v_2$ :  $p_T$  distributions for -0.5 < y < 0, rapidity distribution at 4 GeV, all particles, 10-40%
- $v_1$ : rapidity distributions
- $dv_1/dy$ : mostly protons + some  $\pi, K, \Lambda$ , antiparticles, mostly 10-40%
- $v_3$ : rapidity distributions and slopes,  $p_T$  distributions (3 GeV only), protons, mostly 10-40% (at 3 GeV: peripheral, central), centrality distribution

## Comparing models to data

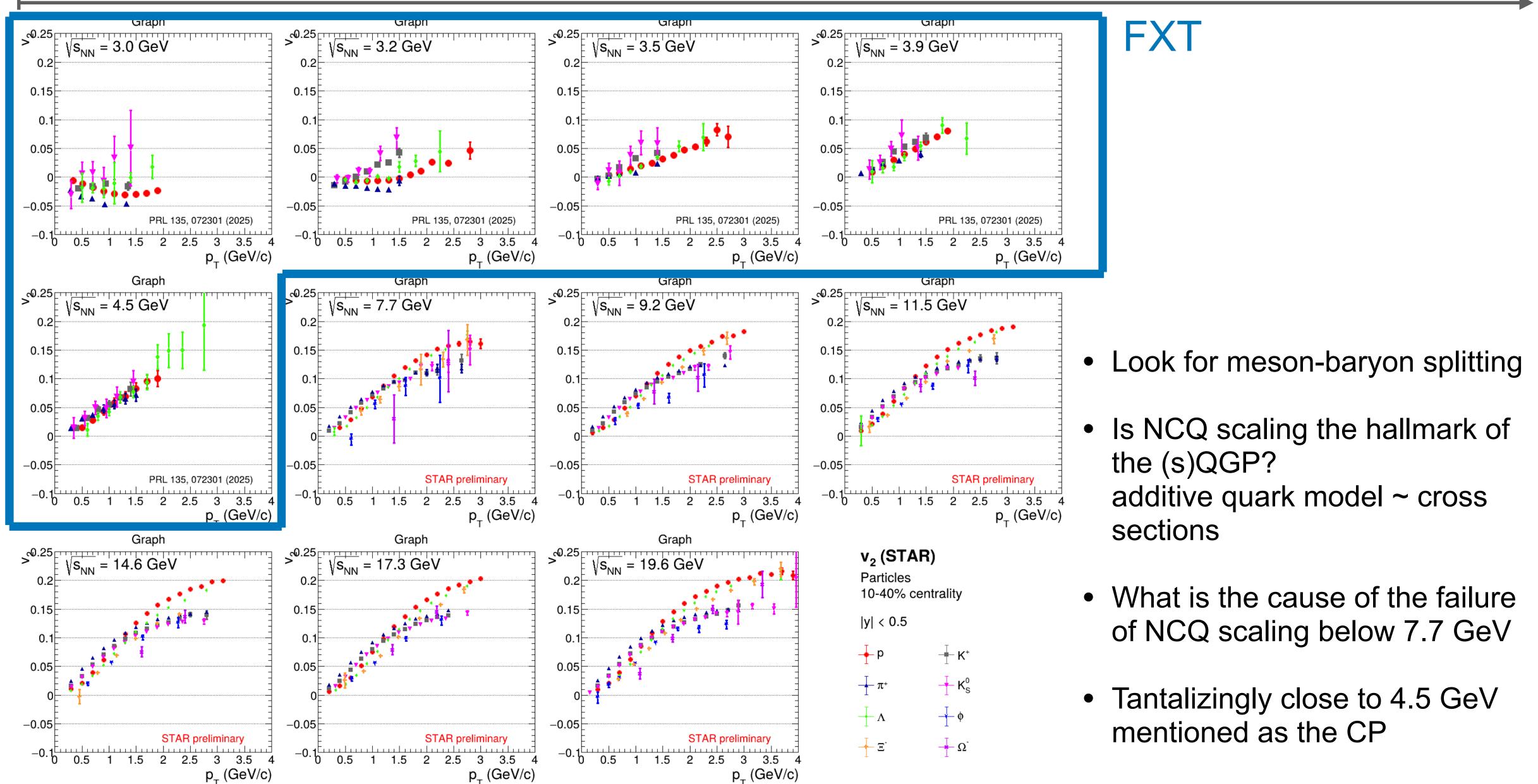
- Energies above  $\sqrt(s) = 10 \text{ GeV}$ : successful paradigm of using hydrodynamics; no such paradigm in the low energy region
- Transport models: UrQMD, SMASH, JAM:
   need: flexible potentials, momentum dependence, maybe in-medium cross sections?
- Can we learn any physics by systematically comparing to models?
   need: reliability
   Experimentalists can help provide benchmarks
   Request for STAR from theoretical community: dN/dy of protons we heard you!
- Wish from the experimentalists:
   Models with same ingredients give same results (e.g. cascade modes)
   Models available to extract physics: e.g., various forms of the potential

Goal: pinning down the location of the QCD CP / phase transition. *Are there* specific observables? Inferences from various sources should agree.

# NCQ scaling of $v_2$ : unscaled

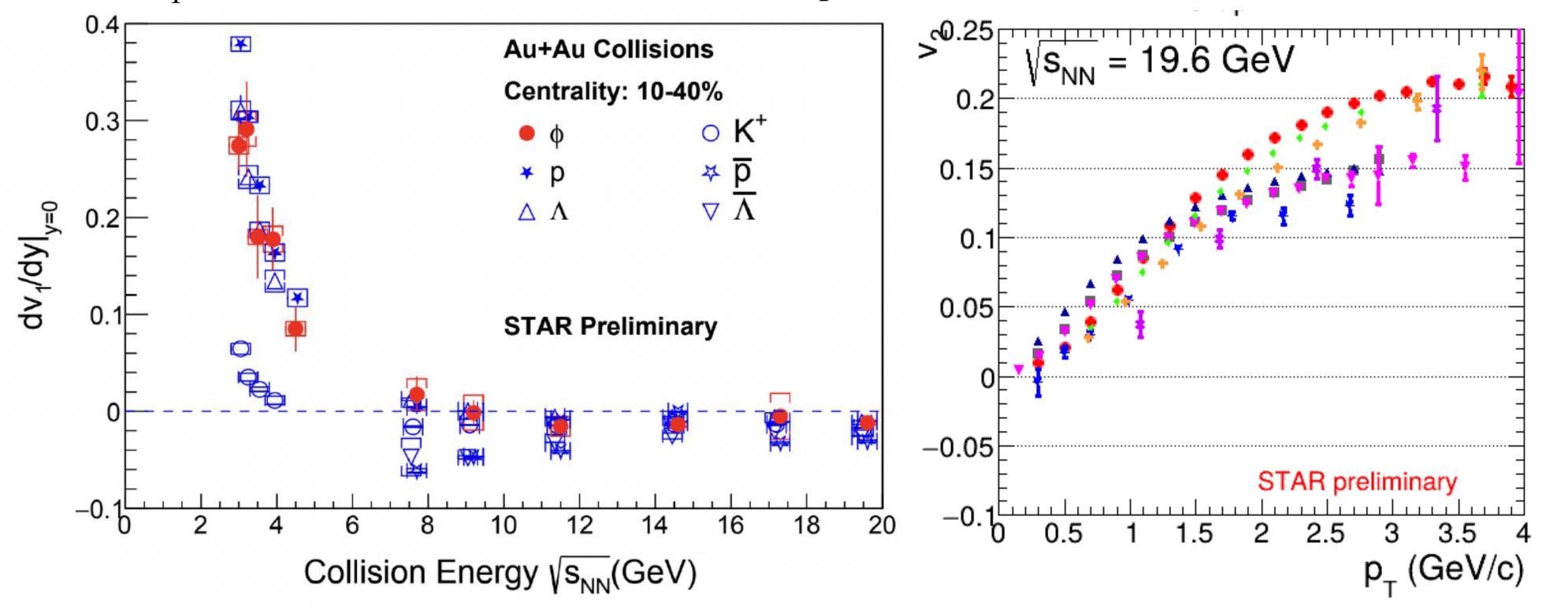
p<sub>T</sub> (GeV/c)

p<sub>T</sub> (GeV/c)



## $v_1$ of mesons and baryons

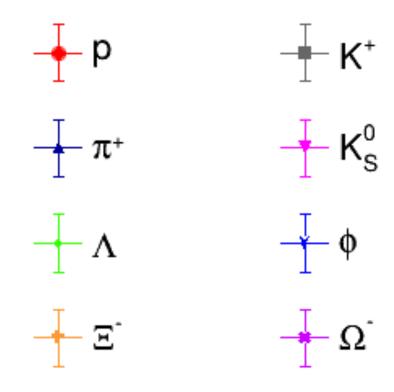
Strong  $v_1$  for protons; different crossing point for p and K



v<sub>2</sub> (STAR)

Particles 10-40% centrality

|y| < 0.5



S. R. Sharma for the STAR Collaboration, QM 2025, <a href="https://indico.cern.ch/event/1334113/contributions/6369544">https://indico.cern.ch/event/1334113/contributions/6369544</a>

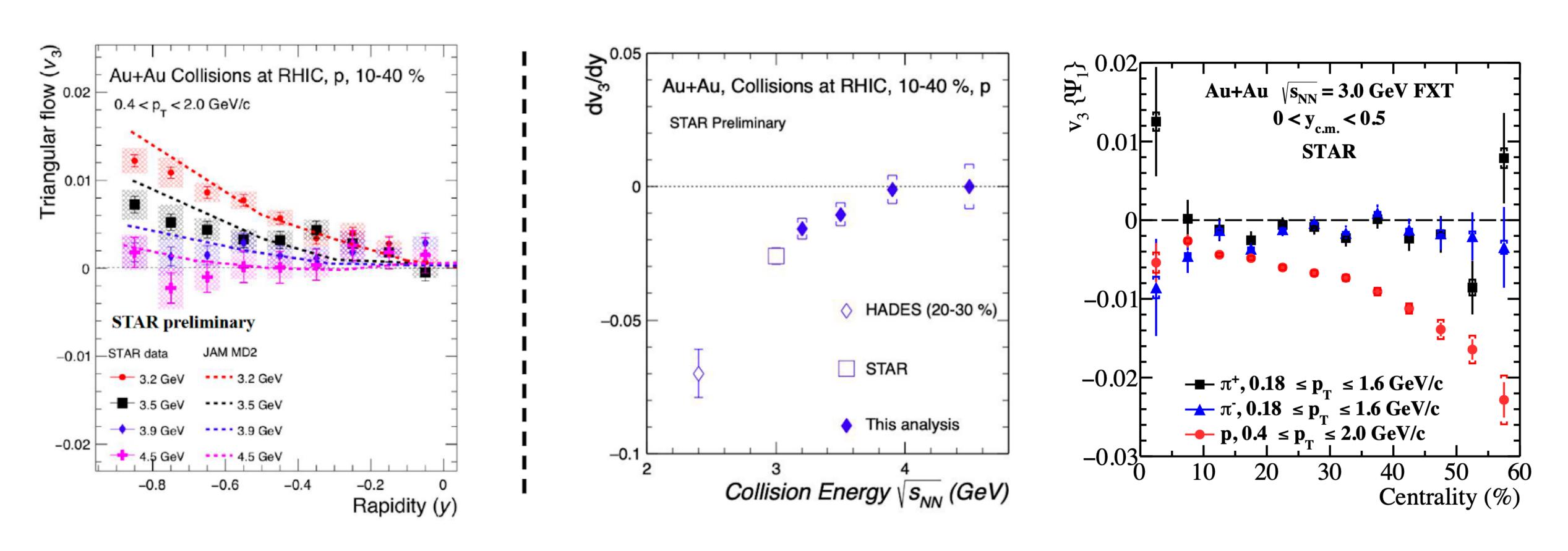
 $v_1$  of  $\phi$ : follows baryons

 $v_2$  of  $\phi$ : follows mesons

Can that be understood?

## $v_3$ in the FXT range

 $\Psi_1$ -correlated proton  $v_3$ ; signal gone at 4.5 GeV; no signal for pions (and kaons)

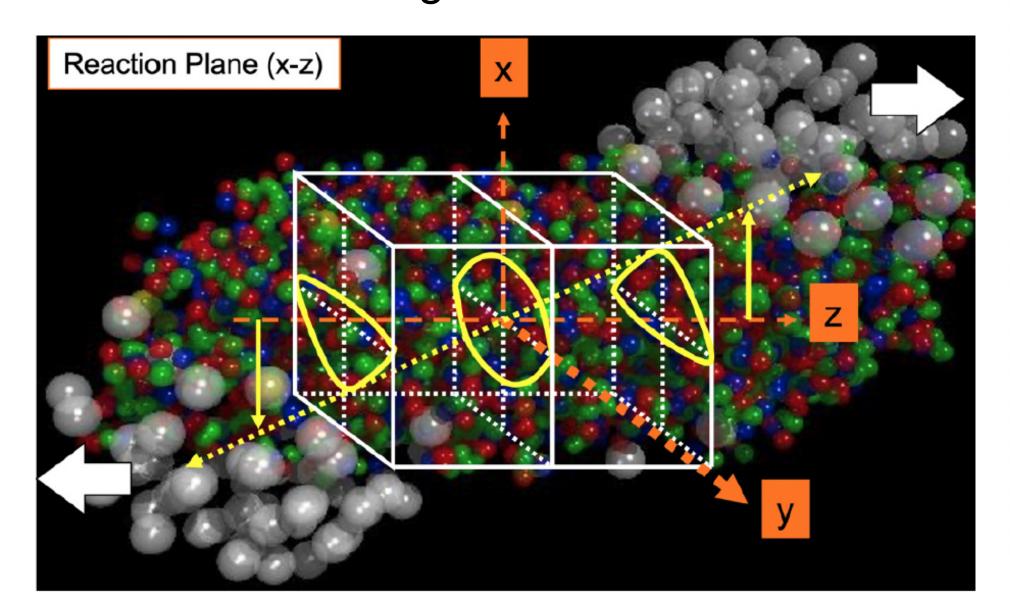


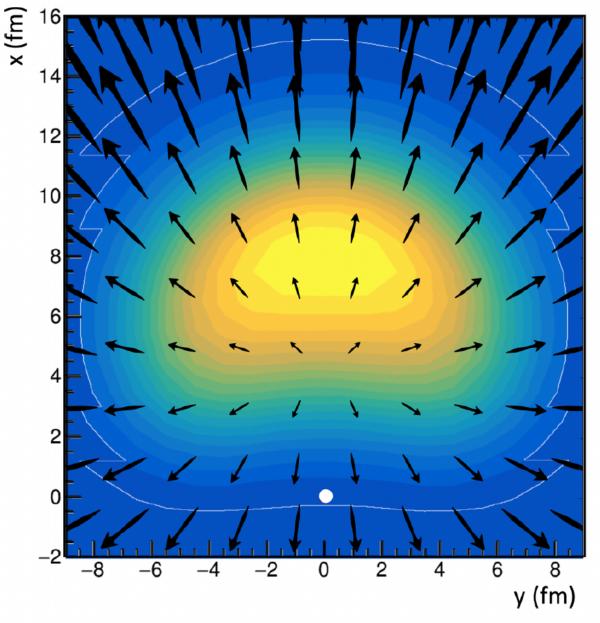
S. R. Sharma for the STAR Collaboration, QM 2025, <a href="https://indico.cern.ch/event/1334113/contributions/6369544">https://indico.cern.ch/event/1334113/contributions/6369544</a>

STAR, *Phys.Rev.C* 109 (2024) 4, 044914, arXiv: 2309.12610

## $v_3$ in the FXT range

Where does the signal come from?





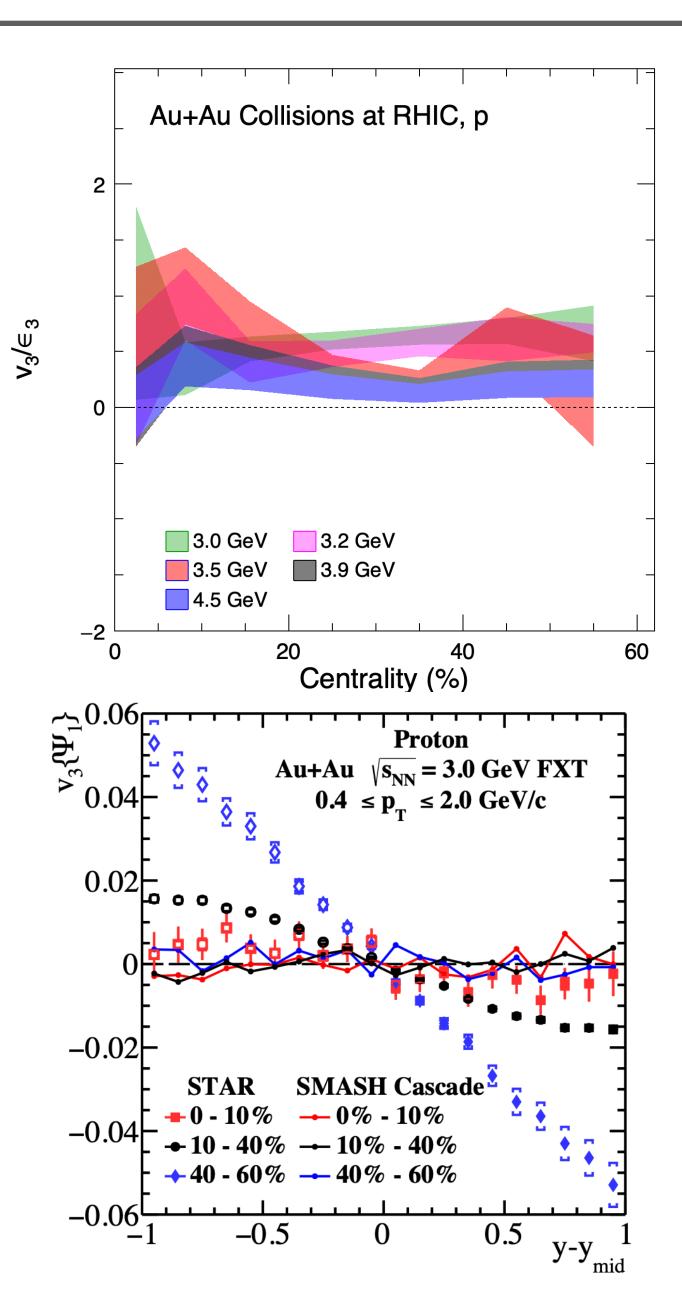
STAR, *Phys.Rev.C* 109 (2024) 4, 044914, arXiv: 2309.12610

It follows triangularity:  $\sim 0$  at  $4.5 \rightarrow v_3 \sim 0$  at 4.5

JAM Md2 potential form doesn't change

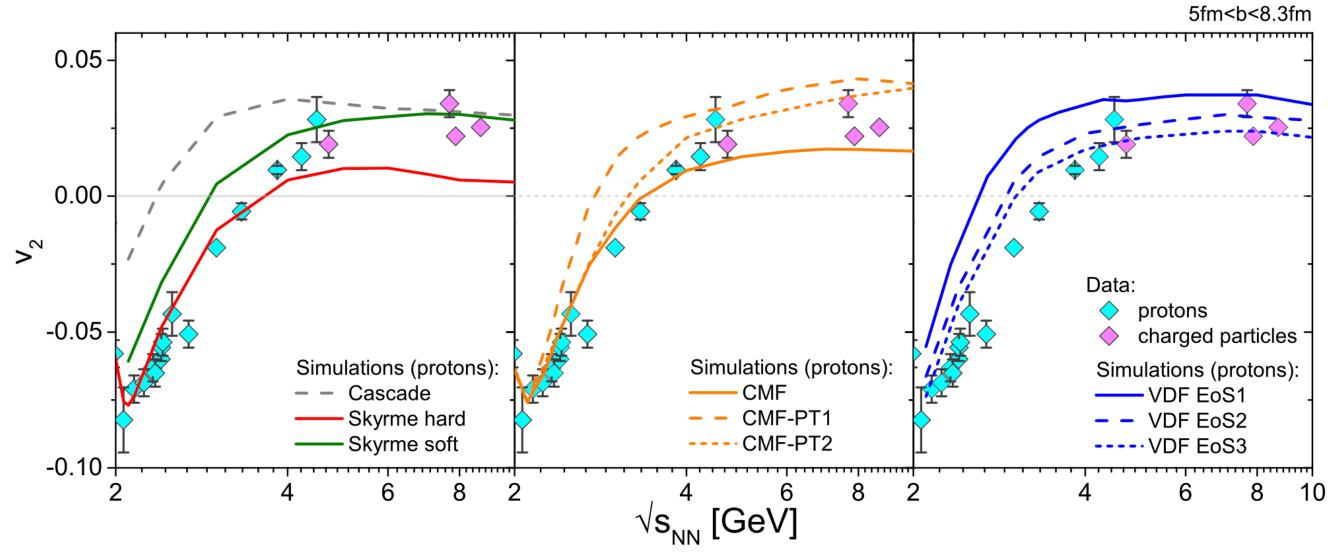
Very sensitive to the potential:

in contrast to  $v_1$  and  $v_3$ , cascade does not generate  $v_3$  (see later)



## Experimental data vs. model predictions

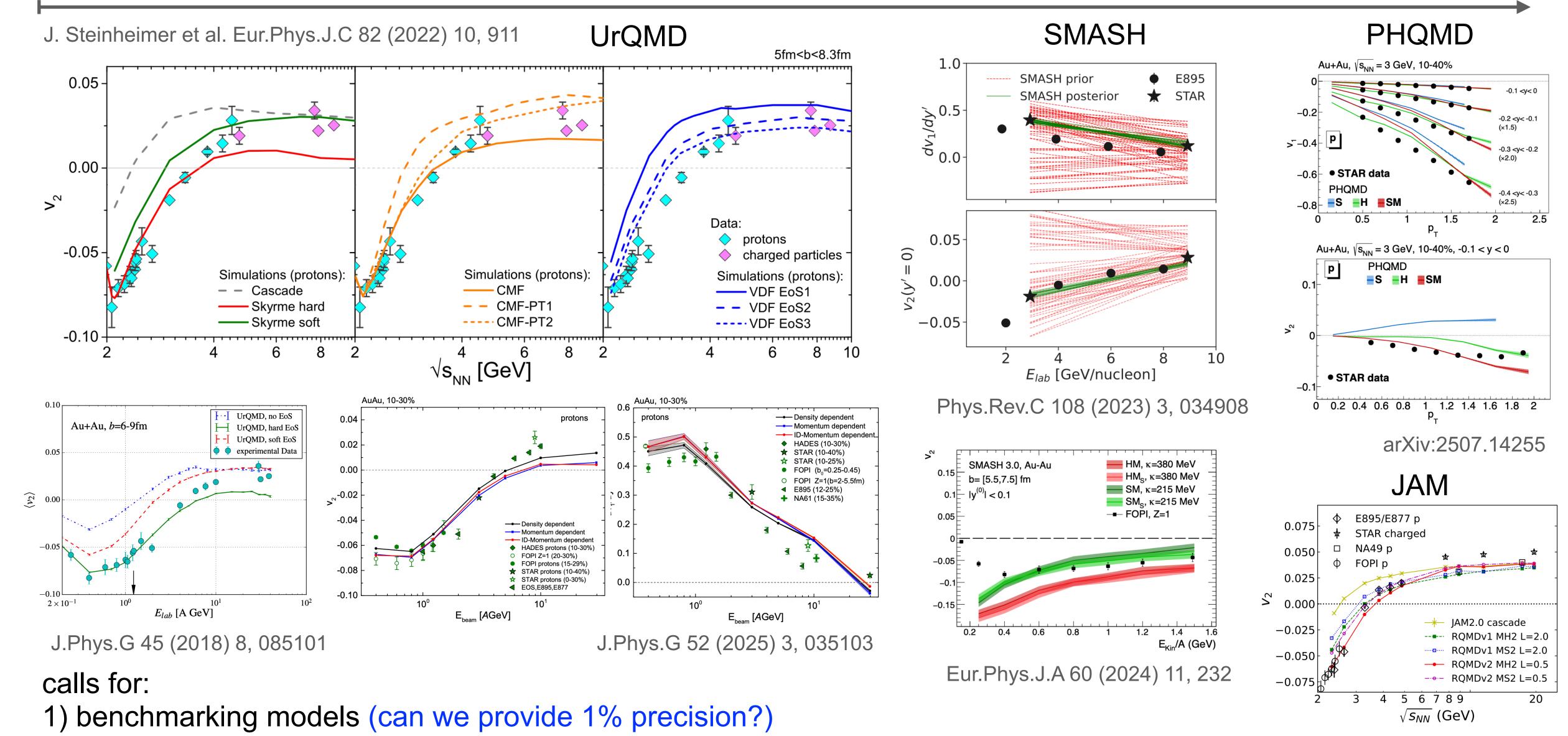
J. Steinheimer et al. Eur. Phys. J. C 82 (2022) 10, 911



- 9 different EOSs in one framework (UrQMD)
- variations on the order of 1% change the extracted EOS

## Experimental data vs. model predictions

2) more differential studies (details, if not integrated out, can provide much needed tests)

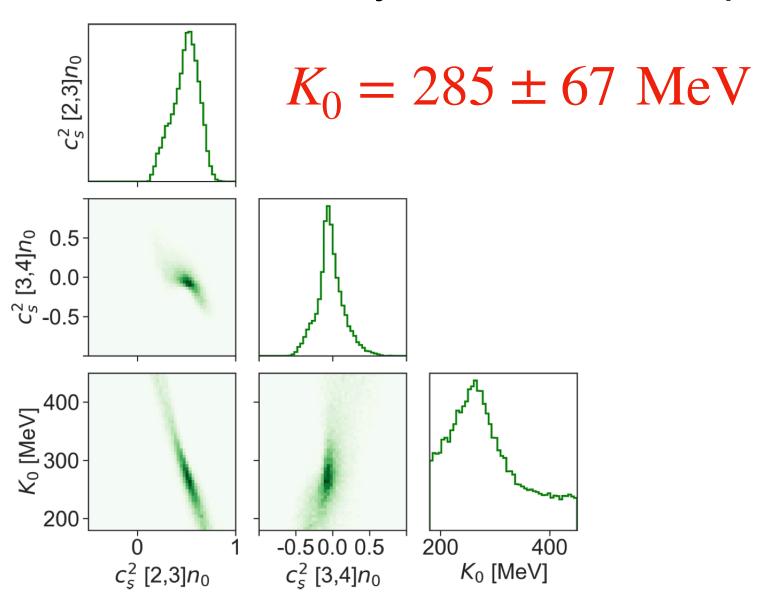


Phys.Rev.C 105 (2022) 1, 014911

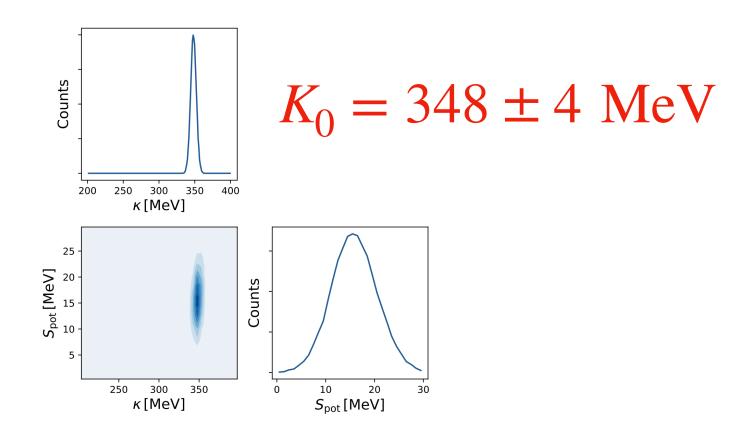
## Bayesian analysis: yes, BUT

Two SMASH analyses extract completely different  $K_0$ 

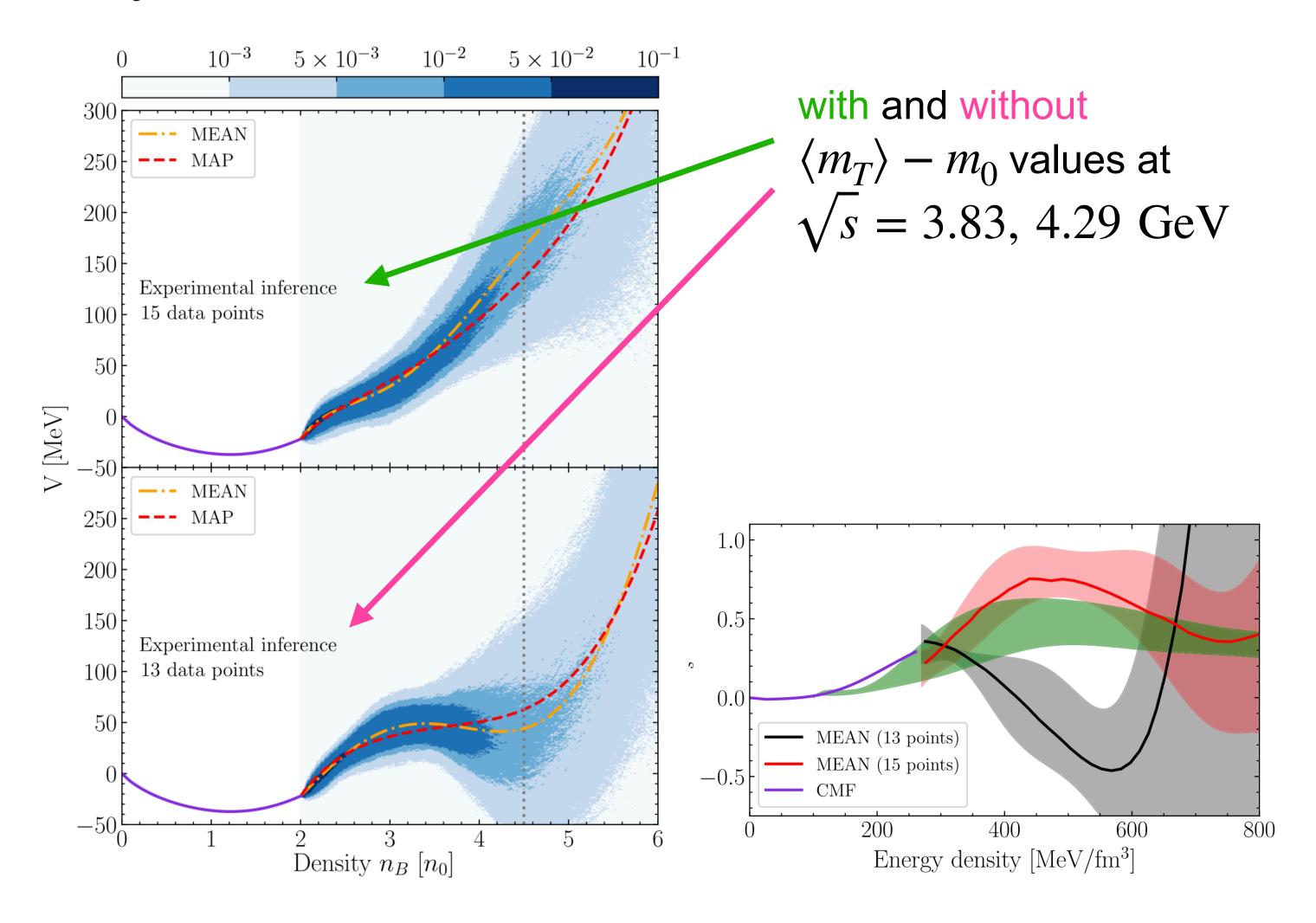
UrQMD results depend strongly on the data set



D. Oliinychenko et al. Phys.Rev.C 108 (2023) 3, 034908



J. Mohs et al. Phys.Rev.C 112 (2025) 4, 044905



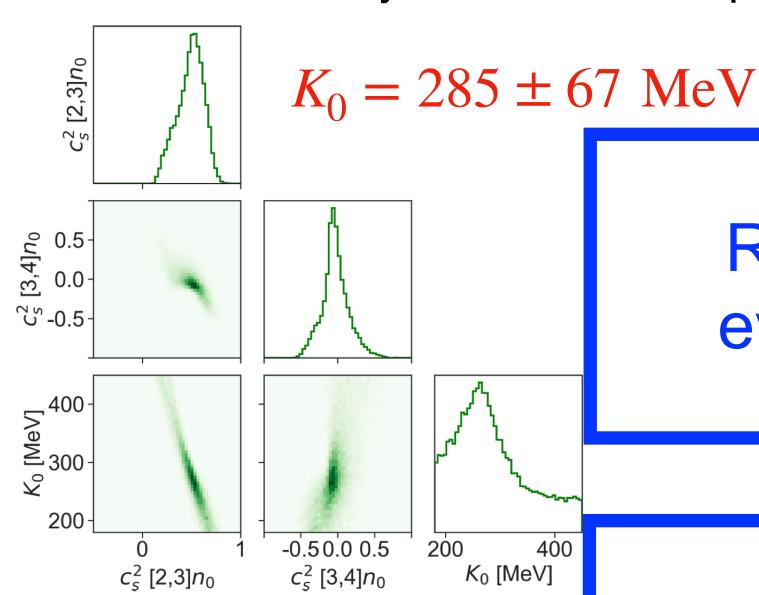
M. Omana Kuttan et al. Phys.Rev.Lett. 131 (2023) 20, 202303

## Bayesian analysis: yes, BUT

Two SMASH analyses extract completely different  $K_0$ 

UrQMD results depend strongly on the data set

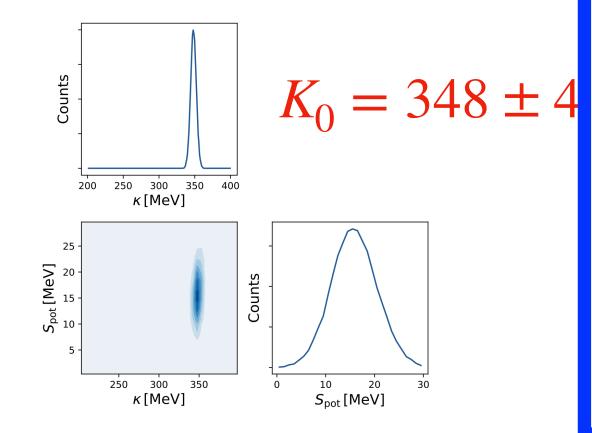
 $5 \times 10^{-2}$ 



Results are meaningless without evaluation of model uncertainties

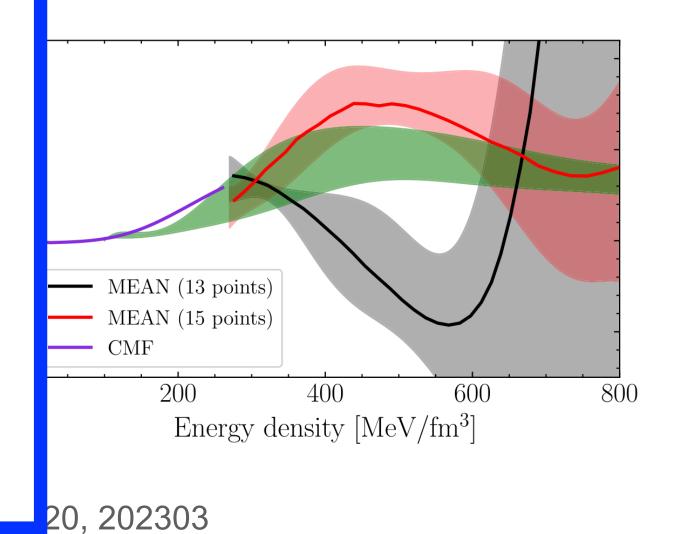
with and without  $\langle m_T \rangle - m_0$  values at  $\sqrt{s} = 3.83, 4.29 \text{ GeV}$ 

D. Oliinychenko et al. Phys.Rev.C 108 (



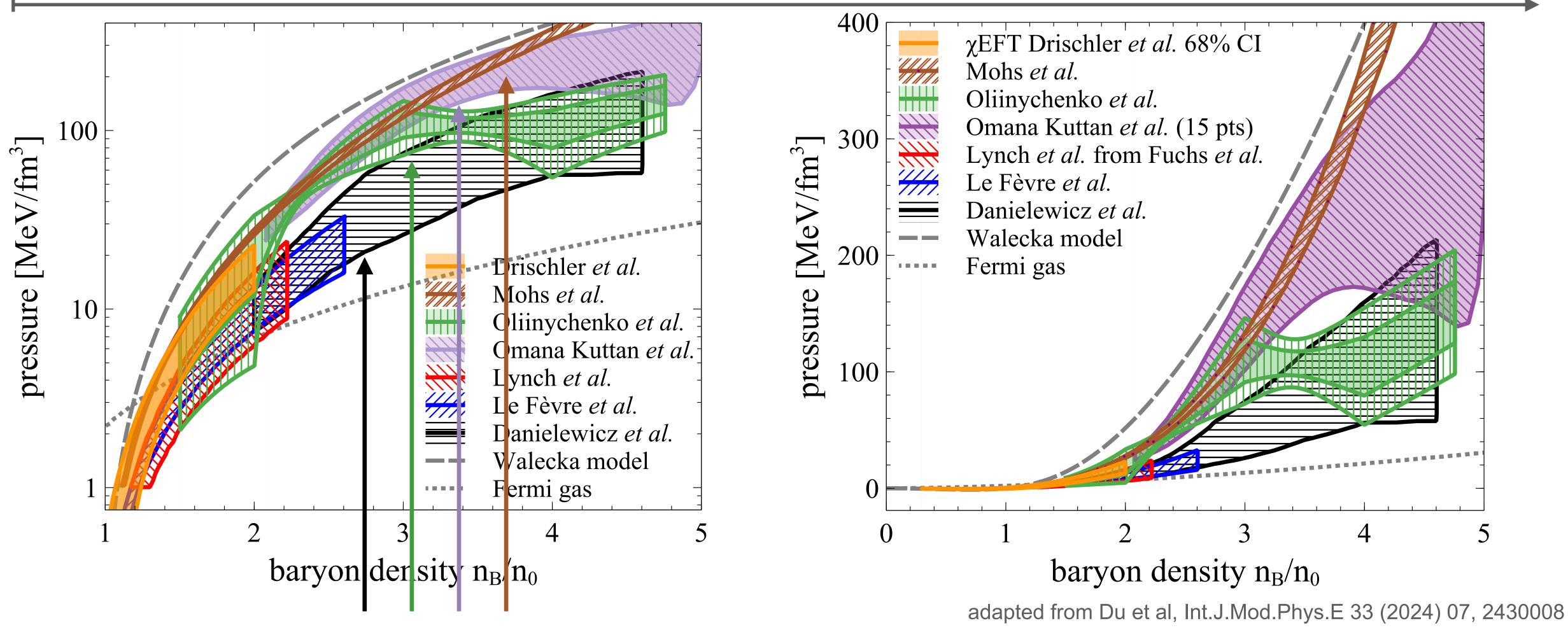
On the other hand, transport:

- Has good control of the initial state
- Handles non-equilibrium evolution
- No model interfaces (another source of uncertainty)



J. Mohs et al. Phys.Rev.C 112 (2025) 4, 044905

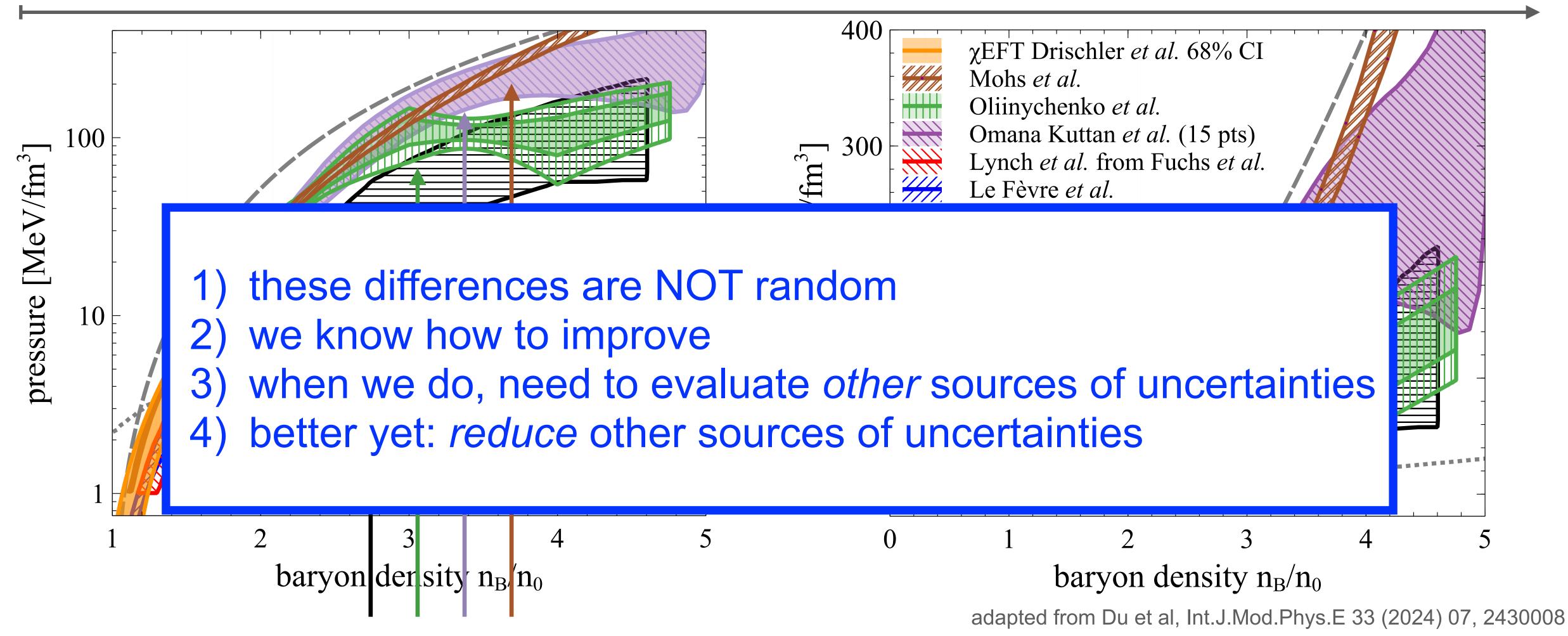
### Good news: we understand the differences



Danielewicz et al: has p-dependence, but can't vary the EOS at high  $n_B$ 

Oliinychenko et al: no p-dependence, but can vary the EOS at high  $n_B$ ; however, did not vary above  $4n_0$  Omana Kuttan et al: no p-dependence, but can vary the EOS at high  $n_B$ ; observable choice not constraining at high  $n_B$  Mohs et al: has p-dependence, but can't vary the EOS at high  $n_B$ ; however, p-dependence likely incorrect

## Good news: we understand the differences

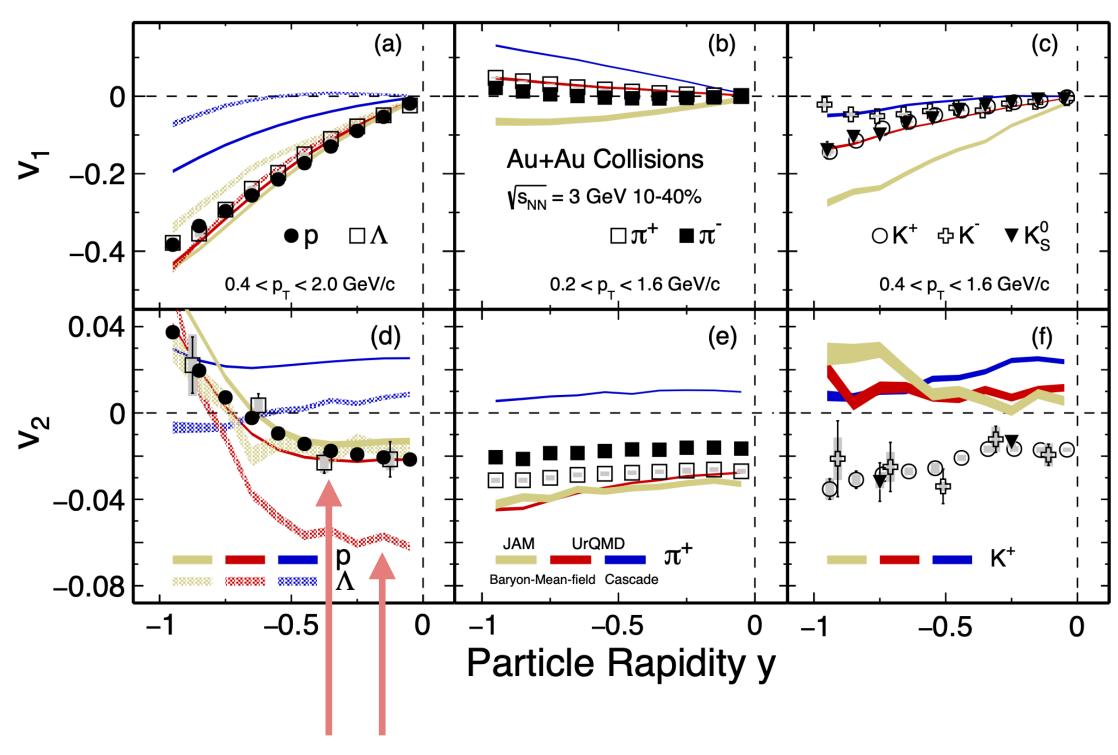


Danielewicz et al: has p-dependence, but can't vary the EOS at high  $n_B$ 

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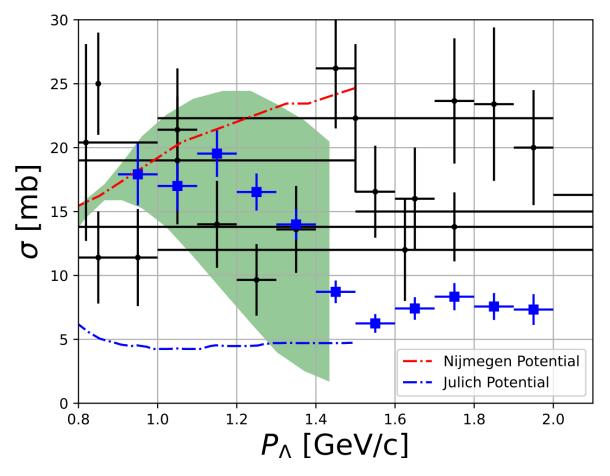
## Importance of fundamental input

STAR, Phys.Lett.B 827 (2022) 137003

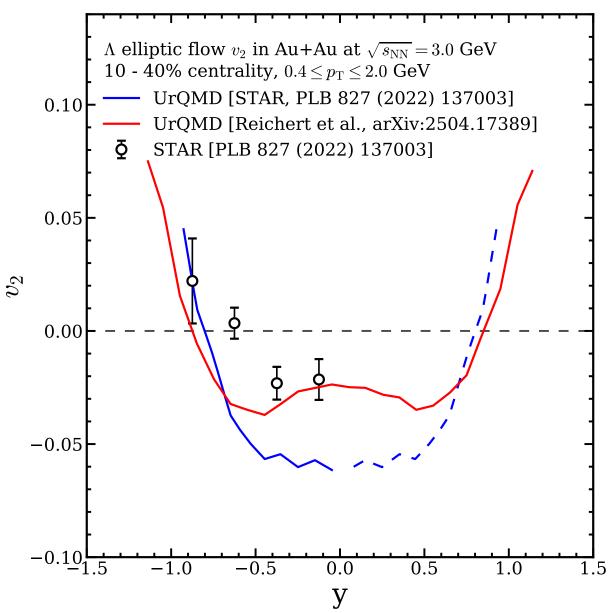


STAR data and model results differ by a factor of 2

#### Input to the model: $\Lambda N$ cross section



CLAS, Phys.Rev.Lett. 127 (2021) 27, 272303



 $\Lambda$  flow is reproduced when employing more precise measurements of the  $\Lambda N$  cross section by CLAS

T. Reichert et al, arXiv:2504.17389

## Importance of comparing apples to apples

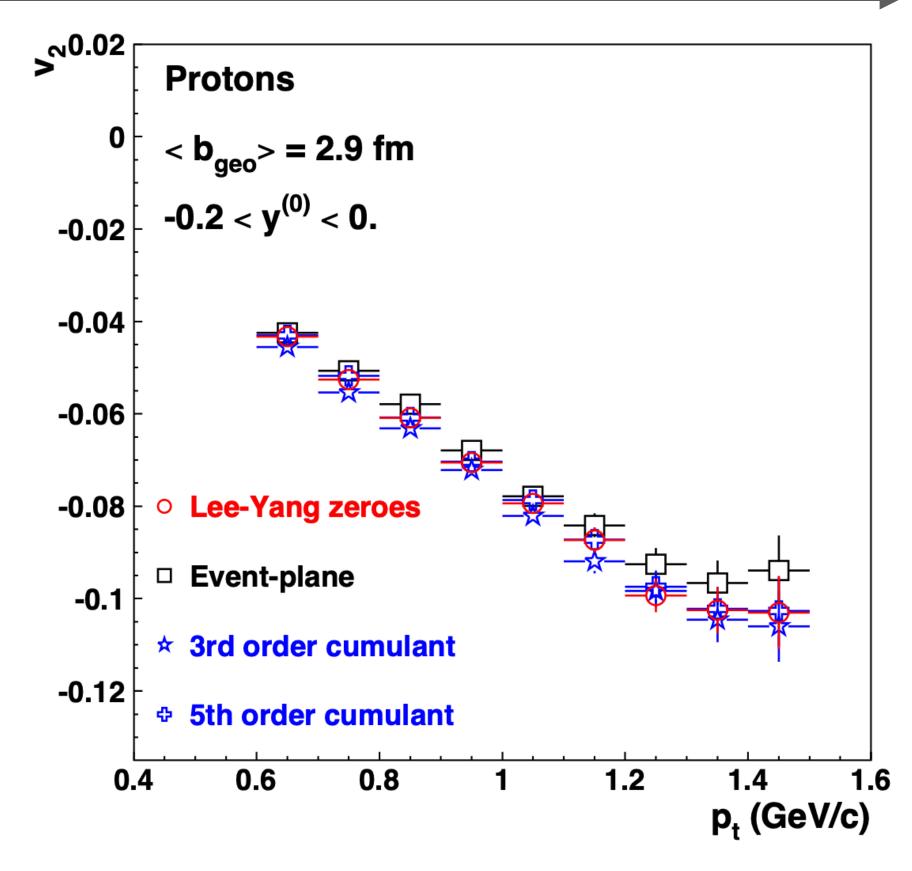
How to measure flow?

- The theory method:  $\Psi_{\rm RP} \equiv 0$
- 1st order forward-spectator event plane
- *n*-th order event plane
- Q-vectors
- 2- and 4-particle cumulants
- Lee-Yang zeros

Use the same method in the measurement and in model analysis

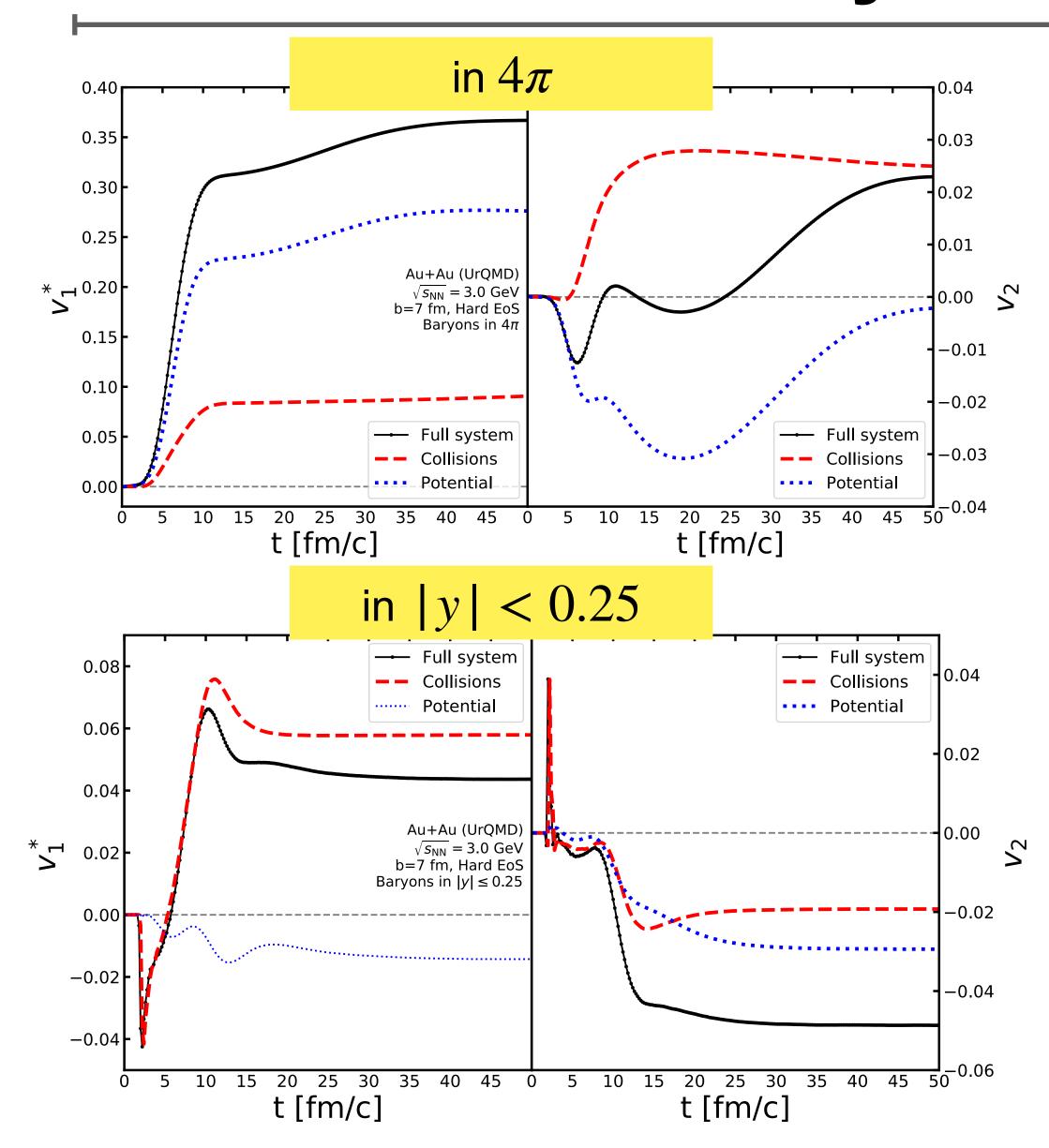
Should we compare to fits or data points?

Fits are finicky! Let's move beyond 
$$dv_1/dy \Big|_{y=0}$$
 and  $v_2(y=0)$ 

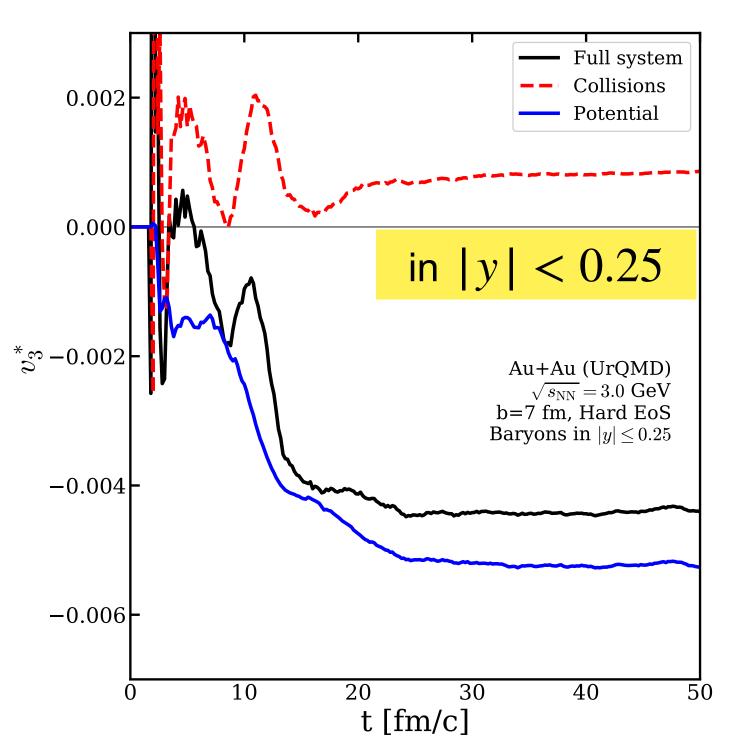


FOPI, Phys.Rev.C 72 (2005) 011901

## Understand the dynamics behind observables



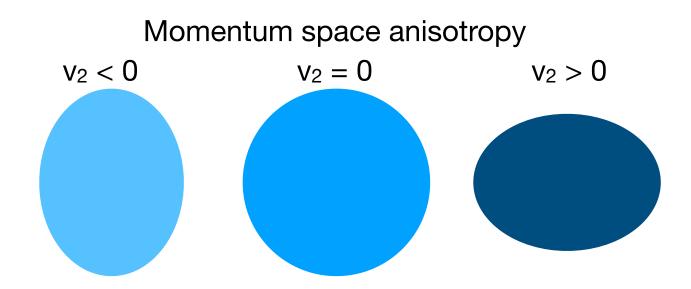
- Do we understand how both v1 and v2 are generated in transport models? We think we do!
- However, reality is more complicated than naive pictures suggest
- Midrapidity  $v_1$  is mostly collisions,  $v_2$  is 50:50 collisions and potential

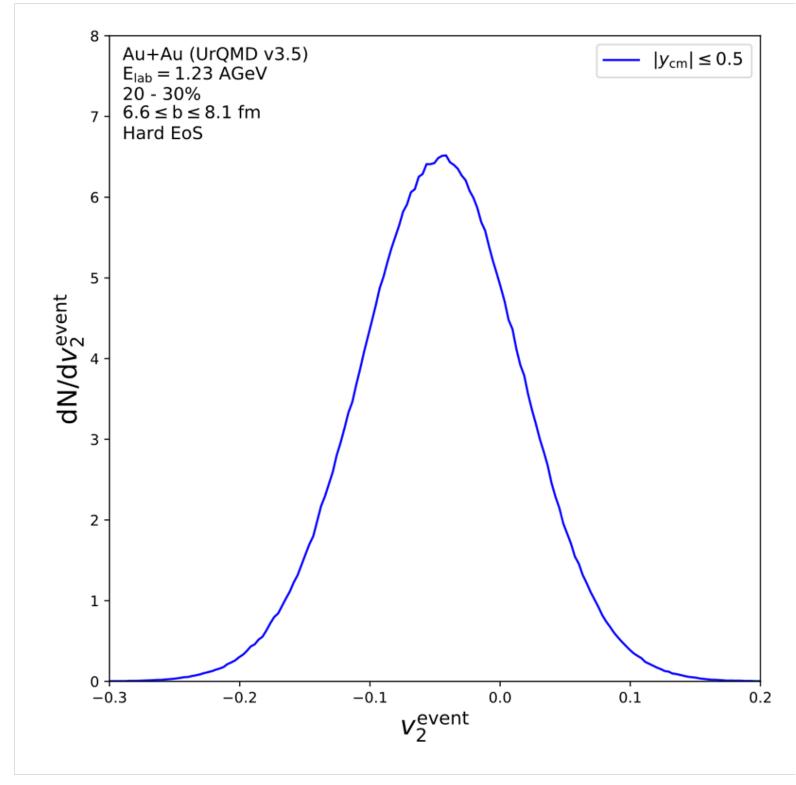


- Triangular flow data starts to become available
- Can we understand its generation?
- Interestingly,  $v_3$  seems to be drastically driven by the EoS
- Go-to place to study the EoS?

T. Reichert and J. Aichelin, Phys.Rev.C 111 (2025) 5, 054916

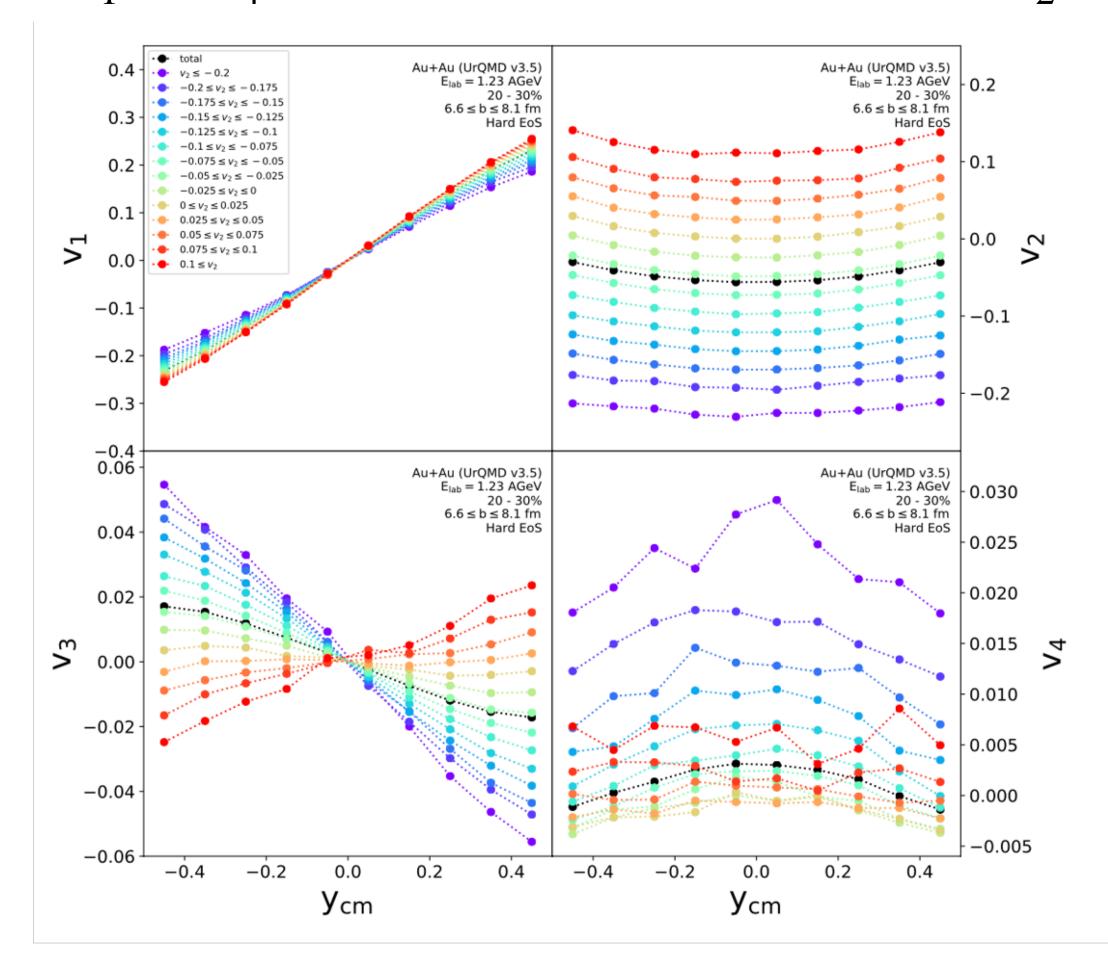
#### Differential observables



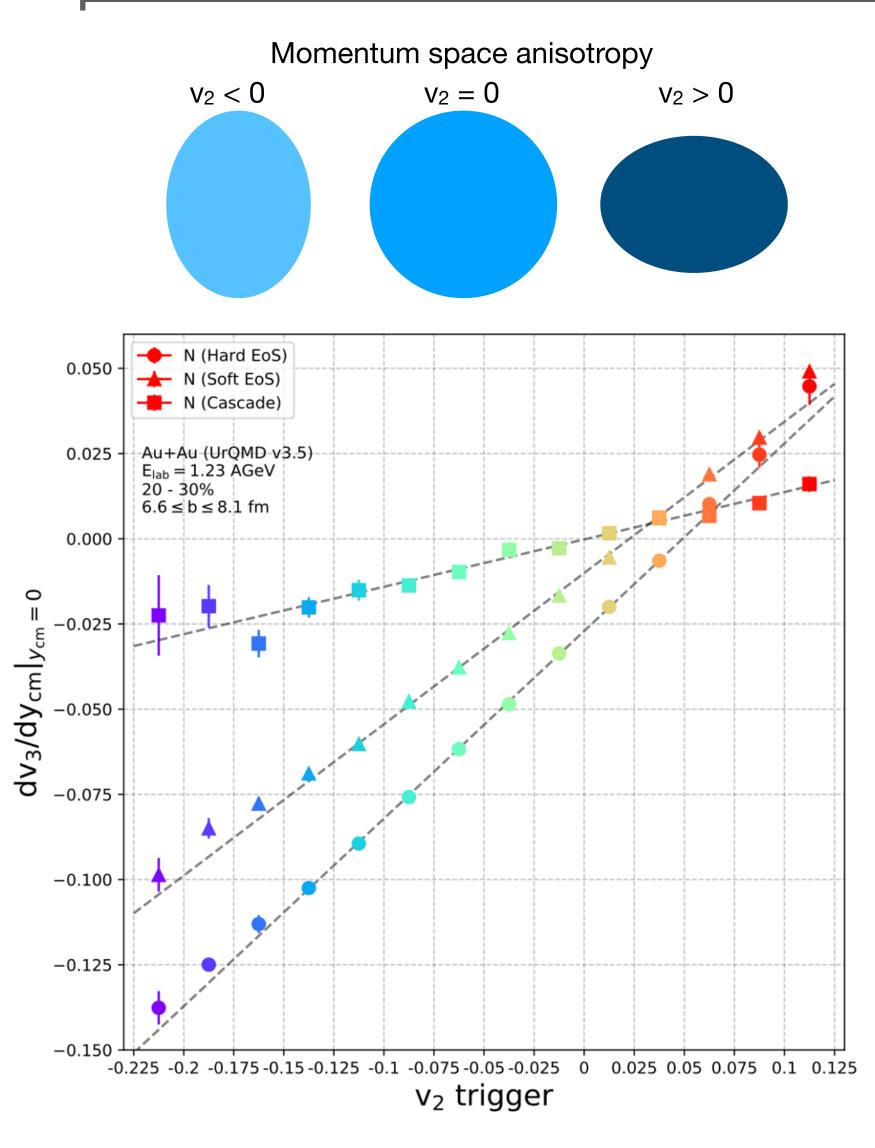


T. Reichert et al. Eur. Phys. J. C 82 (2022) 6, 510

- Event-by-event elliptic flow at midrapidity fluctuates strongly
- Define event classes basen on  $\langle v_2 \rangle_{|y| \le 0.5}^{\text{event}}$
- Analyze  $v_1, ..., v_4$  as a function of rapidity in different " $v_2$  classes"

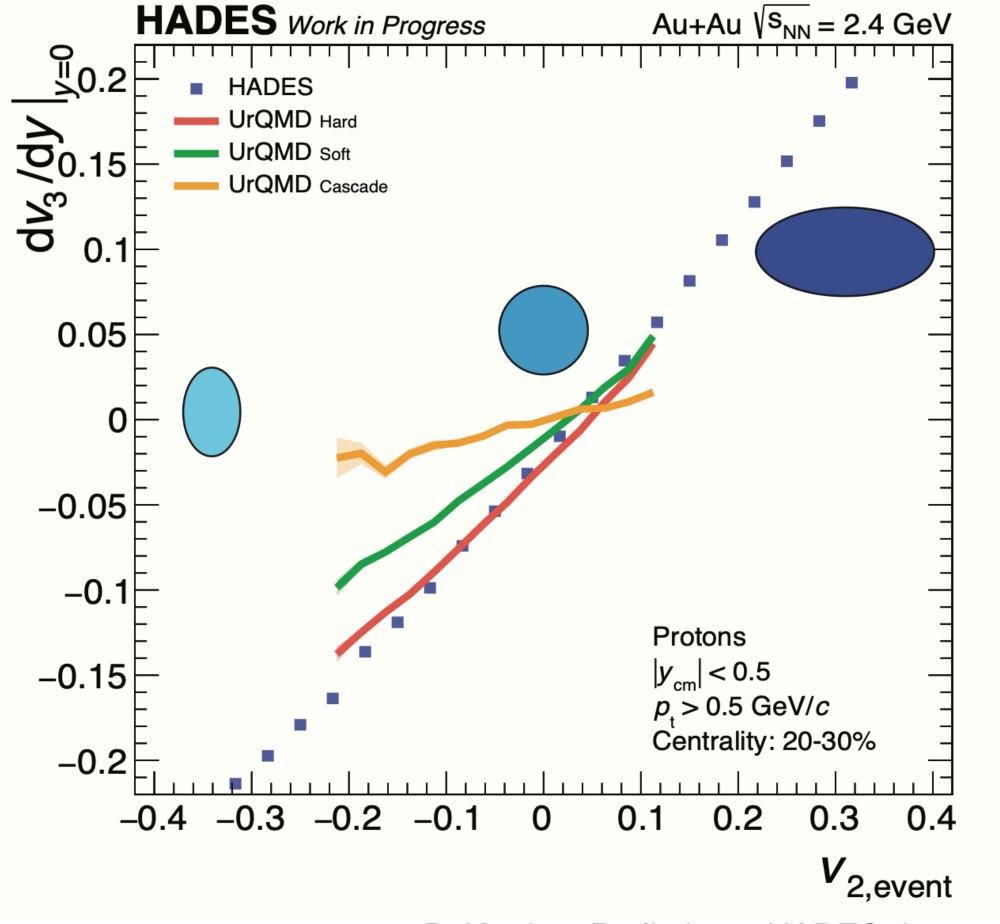


# Differential observables: $v_3$



T. Reichert et al. Eur. Phys. J. C 82 (2022) 6, 510

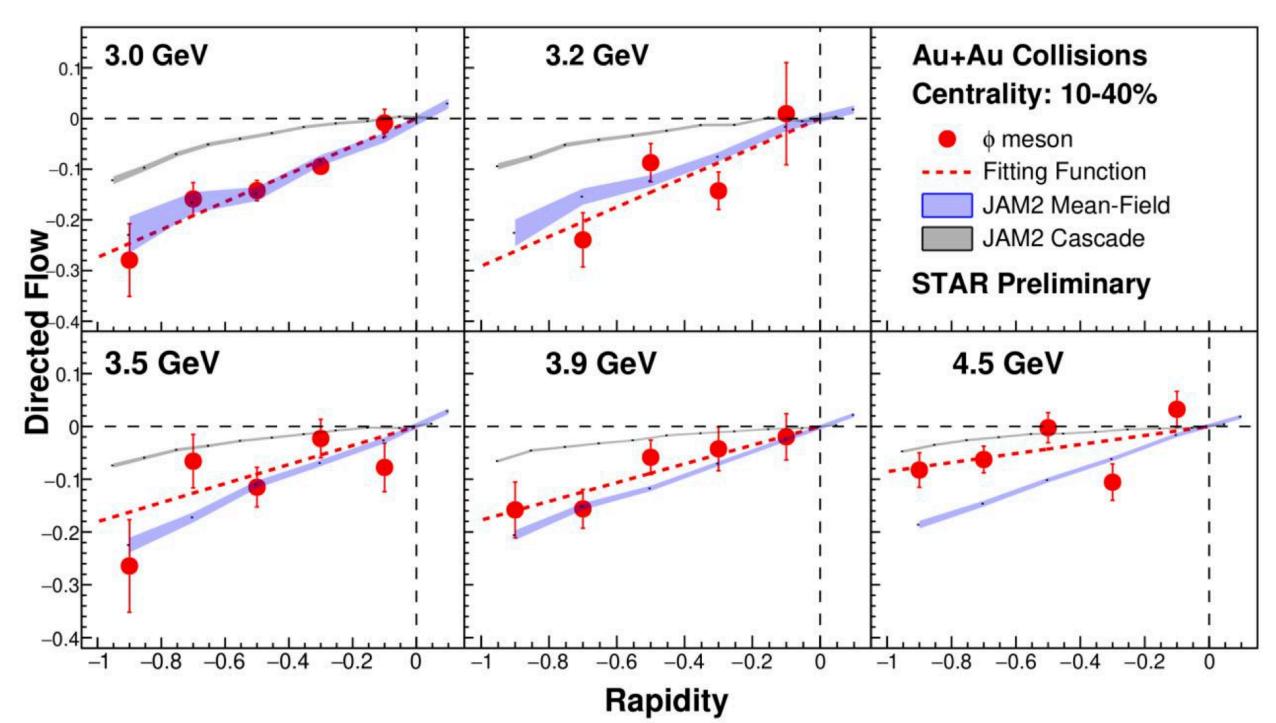
- Strong sensitivity to the employed EoS due to flow correlations
- Under investigation by HADES, insights from STAR would be valuable



B. Kardan. Preliminary HADES data. https://indico.cern.ch/event/1139644/contributions/5456429/

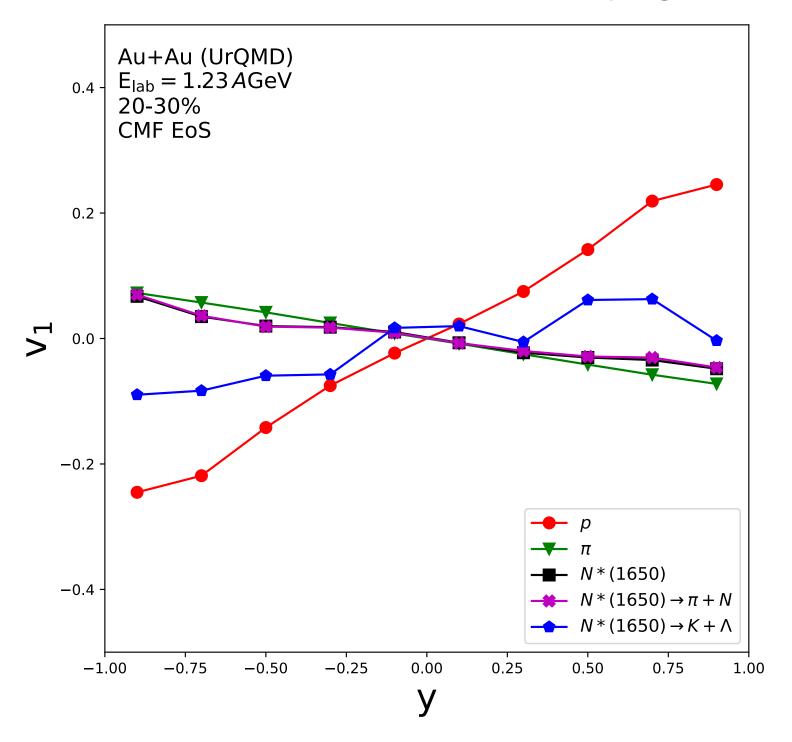
### Flow of resonances

G. Zheng for STAR, QM25 poster, Preliminary



- $\phi$  directed flow follows baryons
- Quark coalescence?
- Production from heavy N\*?

#### T. Reichert, work in progress



- Reconstruction effect
- $N^*$ (1650) provides a handle on the in-medium  $K\!N$  and  $\Lambda N$  cross section

## Wishlist

#### 1. @Transportlers:

Please, do your best to scrutinize and eradicate model systematics. Discuss results in light of the capabilities of your model.

#### 2. @Everybody:

Please, in the appendix of your papers, put as much and as detailed information about the exact procedure of how you have a) measured and/or b) simulated your results. (@Experimentalists: please make sure you know what the models do.)

- 3. @Experimentalists: Please, triple differential  $d^3N/d^3p$  data is highly valuable and needed to shrink transport models' parameter freedom.
- 4. Bayesianing provides parameters that describe data. Bayesianing does not tell you whether your physics is correct.