

# Quantum “Dark” Matter

**Surjeet Rajendran**

# Classical and Quantum Particle Mechanics



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$$i\frac{\partial|\Psi\rangle}{\partial t} = H|\Psi\rangle$$





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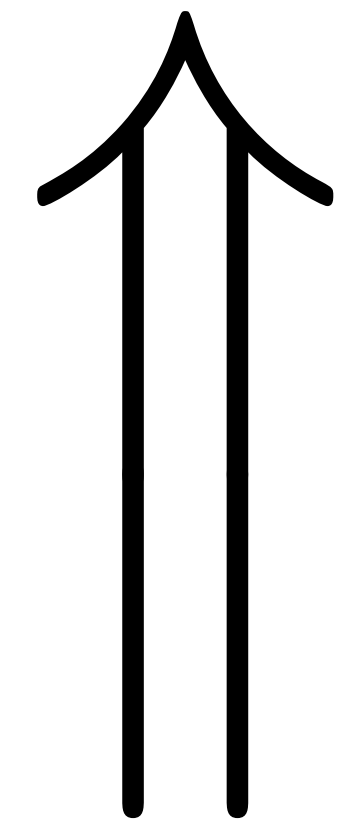
$$\vec{F} = m\vec{a} \quad \frac{d\langle \Psi | \hat{p} | \Psi \rangle}{dt} = -\langle \Psi | \frac{dV}{dx} | \Psi \rangle$$

Ehrenfest's Theorem



$$i\frac{\partial |\Psi\rangle}{\partial t} = H|\Psi\rangle$$

**Quantum Supremacy**



# Classical and Quantum Field Theory

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**(e.g. Klein Gordon)**

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**Quantum State of Fields**  
**(e.g. in Fock states)**

$\phi(x)$

**Time Independent**  
**Operators**

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**Time Evolution**

**Analog of Ehrenfest**  
**Theorem?**

$$i \frac{\partial |\chi(t)\rangle}{\partial t} = H |\chi(t)\rangle$$

**Schwinger Dyson Equations**  
**Quantum Supremacy**



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**Gauge theories?**

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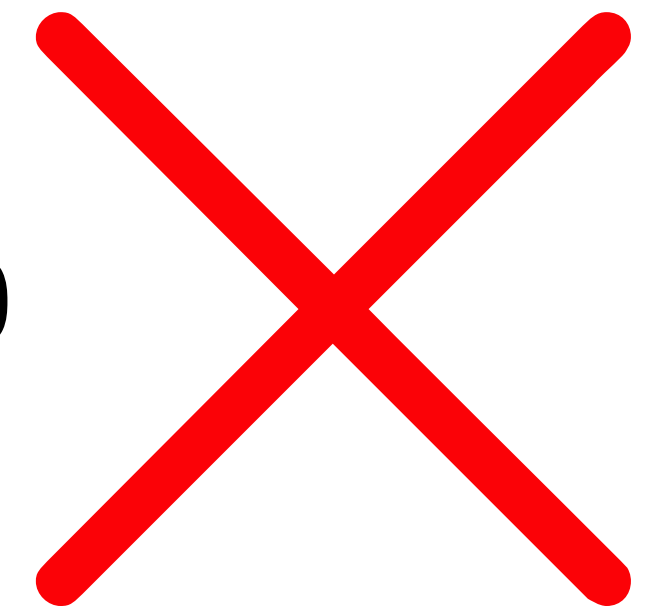
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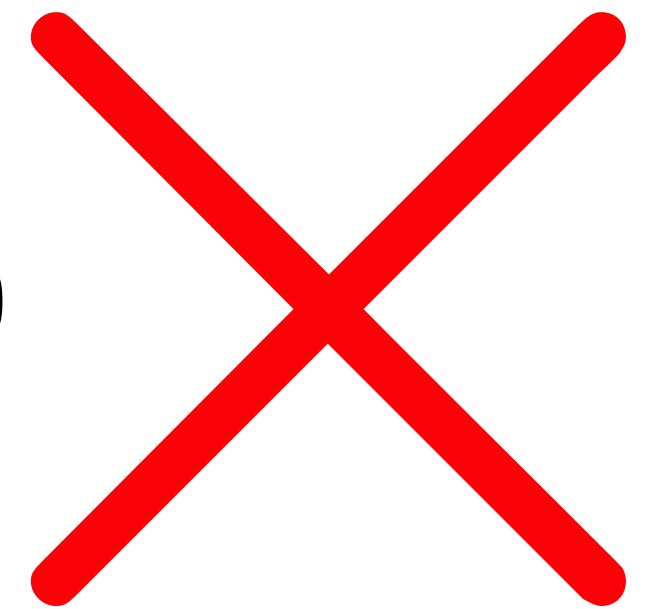
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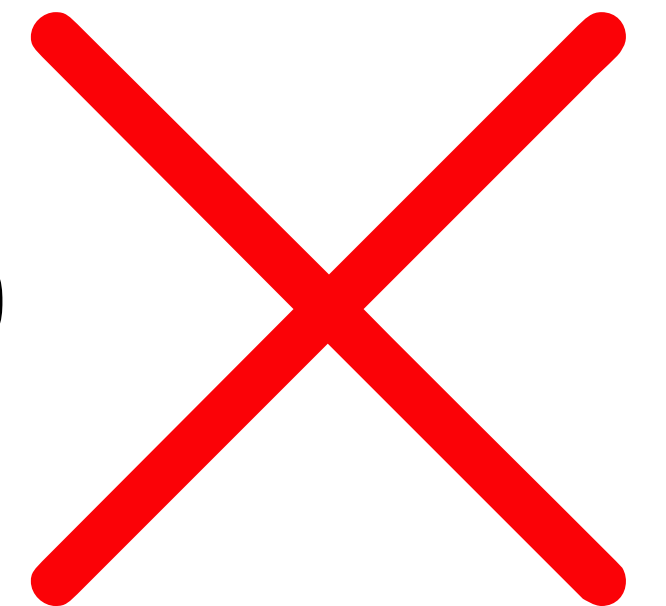
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**Not the Same**  
**???????**

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# Outline

**1. Classical Mini Superspace Cosmology**

**2. Path Integral**

**3. Hamiltonian**

**4. Wheeler deWitt**

**5. Conclusions**



# **Mini Superspace Cosmology**

# Homogeneous Universe

**Study homogeneous space-times in General Relativity**

$$g_{\mu\nu} = -N(t)^2 dt^2 + a(t)^2 (dx^2 + dy^2 + dz^2)$$

**Couple this to homogenous sources of matter, for e.g. rolling scalar field  $\phi$**

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**Classical Equations?**

**2 FRW + 1 Scalar Field**

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**2nd FRW Equation**

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**1st FRW Equation  
(Hubble)**



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**1st FRW Equation  
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**Classical Solutions are over-constrained (3 equations for 2 variables)**

**Solve: Klein Gordon + 2nd FRW with boundary condition from 1st FRW**

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$$\dot{N} = 0 \implies N(t_2) = N_0, N(t_1) = N_0$$

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**Is this a problem?**



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**Perfectly Reasonable for the time evolution operator to be gauge dependent**

**We need gauge invariant physics. What does this mean?**

# Gauge Invariant Physics

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$$| \phi_i, a_i \rangle \rightarrow \sum_{\phi_f, a_f} c_{\phi_f, a_f}(t_2, t_1) | \phi_f, a_f \rangle = \sum_{\phi_f, a_f} c_{\phi_f, a_f}(\tilde{t}_2, \tilde{t}_1) | \phi_f, a_f \rangle$$

$$| \Psi(t_1) \rangle = | \tilde{\Psi}(\tilde{t}_1) \rangle = | \Sigma \rangle \rightarrow | \Psi(t_2) \rangle = | \tilde{\Psi}(\tilde{t}_2) \rangle = | \Omega \rangle$$

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**Moral: Gauge invariant physics from gauge dependent path integral**  
**Reasonable since time is gauge choice!**



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**Heisenberg Picture**

$$\langle \Psi | \partial_t \left( \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) - \frac{\partial \mathcal{L}}{\partial \phi} | \Psi \rangle = 0 \quad \text{Klein Gordon}$$

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**Klein Gordon**

$$\langle \Psi | \frac{\partial \mathcal{L}}{\partial \lambda} - \frac{\partial \mathcal{L}}{\partial N} | \Psi \rangle = 0$$

**Tells you how  $\lambda$  evolves in the path integral - not 1st FRW Equation?????**

**Heisenberg Picture**

# Hamiltonian

# Hamiltonian Construction

$$g_{\mu\nu} = -N(t)^2 dt^2 + a(t)^2 (dx^2 + dy^2 + dz^2)$$

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$$\Pi_N = \frac{\partial \mathcal{L}}{\partial \dot{N}} = 0$$

**What to do with  $N$ ?**

# Hamiltonian Construction

$$\Pi_a = \frac{\partial \mathcal{L}}{\partial \dot{a}}, \Pi_\phi = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \quad \text{Quantize These}$$

$$[a, \Pi_a] = [\phi, \Pi_\phi] = i$$

$$a = A + A^\dagger, \Pi_a = i(A - A^\dagger) \quad \phi = B + B^\dagger, \Pi_\phi = i(B - B^\dagger)$$



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**Fully non-linear General Relativity - no “free” theory with “free” kinetic term**

**But, can still construct Hilbert space with Fock states of A, B - these are operator level statements independent of kinetic terms of the theory**

$$A|0\rangle = 0, A^\dagger|0\rangle = |1\rangle \text{ etc.}$$

# Hamiltonian Construction

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**Construct Canonical Hamiltonian from this Lagrangian**

$$H_N = N H_0(a, \Pi_a, \phi, \Pi_\phi)$$

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**What is N? Different values of N yield different Hamiltonians  
Different physics? Gauge Invariance?**

**If N is a non-trivial operator on the Fock space, no way to make physics gauge invariant  
Possible choice: N is a c-number  
But still, different choices of N yield different Hamiltonians!**

# Schrodinger Equation

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**Different choices of N correspond to different choices of time co-ordinate.**

**Gauge invariant physics!**

$$\langle \phi_f, a_f | T(t_2; t_1) | \phi_i, a_i \rangle = \int_{\phi(t_1)=\phi_i, a(t_1)=a_i}^{\phi(t_2)=\phi_f, a(t_2)=a_f} D\phi D a e^{i \int_{t=t_1}^{t=t_2} (\tilde{\mathcal{L}}(N_0))}$$

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**Consistent with Path Integral - you just pick N**

# Consequence of Schrodinger Equation

$$i\frac{\partial|\chi(t)\rangle}{\partial t} = N H_0|\chi(t)\rangle \implies \frac{d\langle\chi(t)|H_0|\chi(t)\rangle}{dt} = 0$$

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Thus:  $\frac{d \langle \chi | a^3 \frac{\partial \mathcal{L}}{\partial N} | \chi \rangle}{dt} = 0$

**This is almost the 1st FRW equation - but not quite.  
1st FRW equation only satisfied up to overall constant**

# Initial State

$$a = A + A^\dagger, \Pi_a = i(A - A^\dagger) \quad \phi = B + B^\dagger, \Pi_\phi = i(B - B^\dagger)$$

**Create quantum states of  $a, \phi$**

$$|\chi\rangle = f(A, A^\dagger, B, B^\dagger) |0\rangle$$

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**First order ODE - no issue with time evolving**

$$\langle \chi | a^3 \frac{\partial \mathcal{L}}{\partial N} | \chi \rangle \neq 0$$



# Violating 1st FRW

$$\langle \chi | a^3 \frac{\partial \mathcal{L}}{\partial N} | \chi \rangle = c$$

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## Implied Classical Dynamics

$$\langle \chi | \partial_t \left( \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) - \frac{\partial \mathcal{L}}{\partial \phi} | \chi \rangle = 0$$

**2nd FRW Equation**

$$\langle \chi | \partial_t \left( \frac{\partial \mathcal{L}}{\partial \dot{a}} \right) - \frac{\partial \mathcal{L}}{\partial a} | \chi \rangle = 0$$

**Klein Gordon**

$$\langle \chi | \frac{\partial \mathcal{L}}{\partial N} | \chi \rangle = \langle \chi | \frac{c}{a^3} | \chi \rangle$$

**1st FRW but with “Dark”  
Matter**

# Quantum “Dark” Matter

$$\langle \chi | a^3 \frac{\partial \mathcal{L}}{\partial N} | \chi \rangle = c$$

$$a = A + A^\dagger, \Pi_a = i(A - A^\dagger) \quad \phi = B + B^\dagger, \Pi_\phi = i(B - B^\dagger)$$

**Create quantum states of  $a, \phi$**

**Quantum Dynamics: Just 1 first order ODE (Schrodinger)**

**No reason to constrain initial state!**

**Failure manifests classically as “dark” matter - though no real particle excitation there. Conservation implies super-selection sector.**

**Can be positive or negative!**

**Wheeler deWitt**

# Can we get all of Einstein's Equations?

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**Universe is in an energy eigenstate**

**But... Time???**

# Path Integral Version

$$\langle \phi_f, a_f | T (t_2; t_1) | \phi_i, a_i \rangle = e^{iP(\phi_f, a_f, t_2; \phi_i, a_i, t_1; N_0)}$$

**Gauge fixed path integral implies we lose 1st FRW equation**

**Path Integral explicitly depends upon random gauge choice  $N_0$**

**Path integral obviously not gauge invariant**



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**Integrate over all  $N_0$**

**Unsurprisingly, this yields infinity**

**Demanding gauge invariant path integral implies only possible states are static (Wheeler deWitt)**

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**Quantum Supremacy**

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With Gauge Invariant Physics**

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Evolution, Infinite path integral**

**All Equations Hold...but...**

**Only static states...no time!**

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**Full General Relativity, similar issues in Electromagnetism**

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