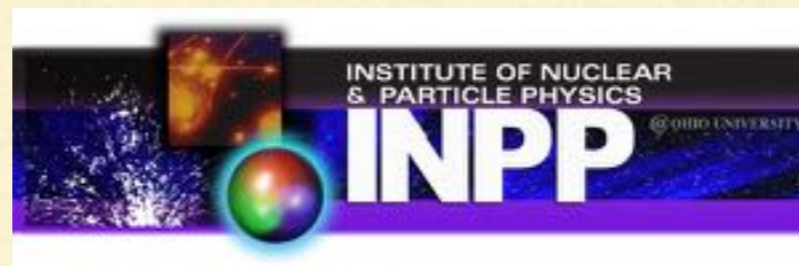

Universal or not?

EFT insights into two-neutron halos and ${}^6\text{Li}$



OHIO
UNIVERSITY



Daniel Phillips

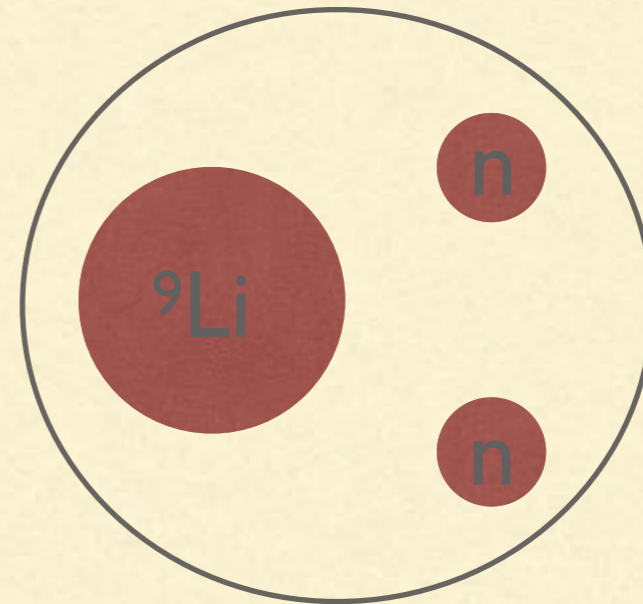
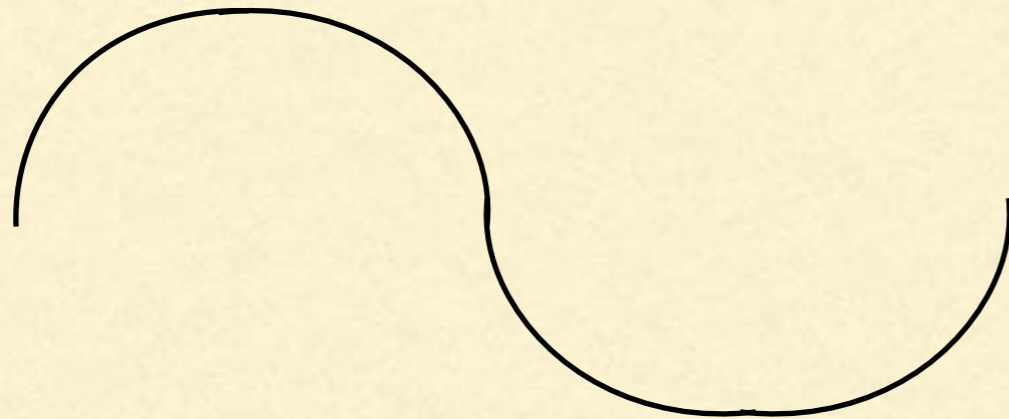
with Matthias Göbel, Hans-Werner Hammer, Chloë Hebborn, and
Carl Brune

**SUPPORTED BY THE US DOE, THE NSF OFFICE OF ADVANCED
CYBERINFRASTRUCTURE, AND THE SWEDISH RESEARCH COUNCIL**

Halo EFT

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Reviews: Hammer, Ji, DP, J. Phys. G 44, 103002 (2017);

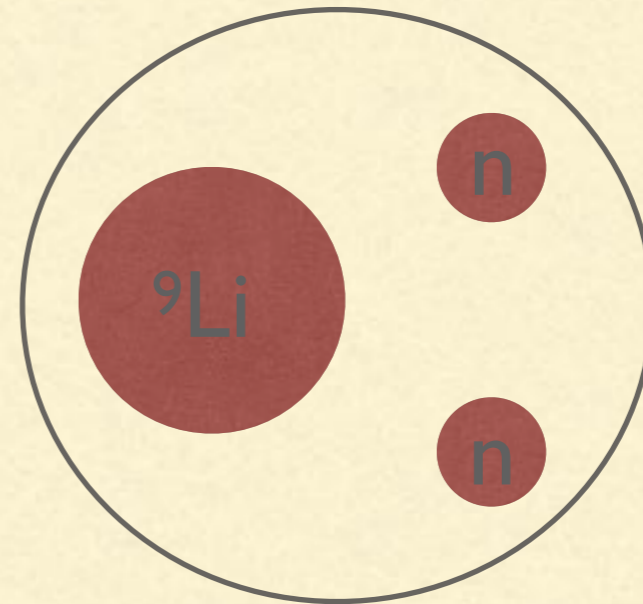
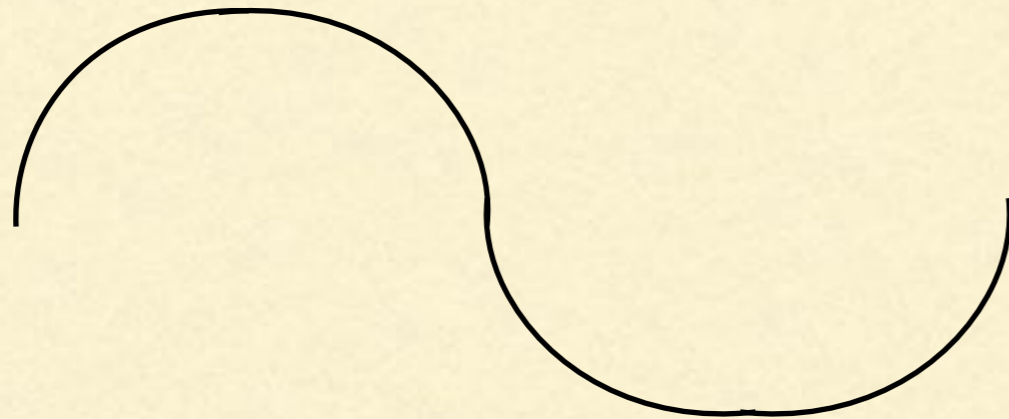
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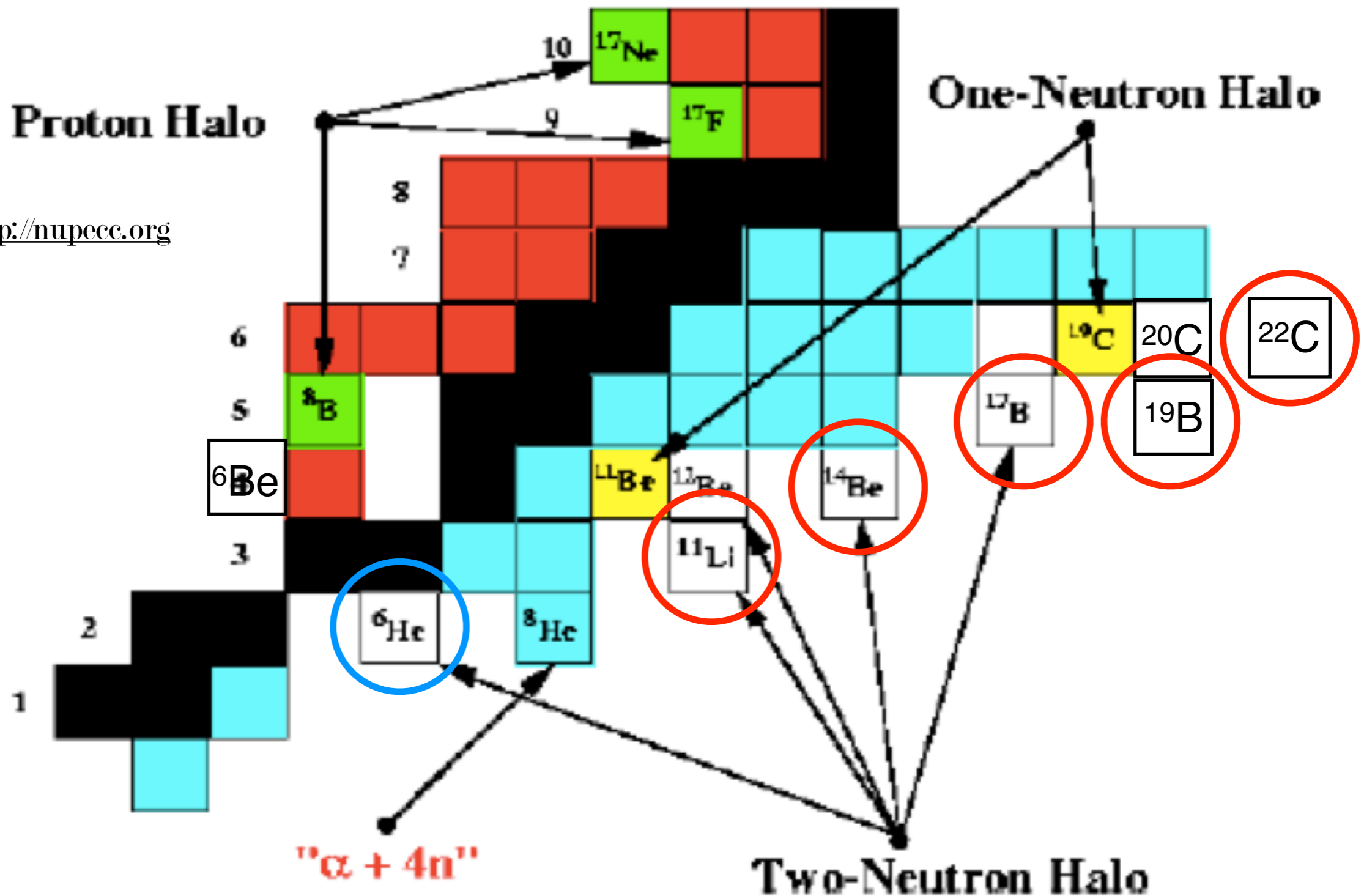
$$\lambda \gg R_{\text{core}}; \lambda \lesssim R_{\text{halo}}$$



- Define $R_{\text{halo}} = \langle r^2 \rangle^{1/2}$. Seek EFT expansion in $R_{\text{core}}/R_{\text{halo}}$. Valid for $\lambda \lesssim R_{\text{halo}}$
- Typically $R \equiv R_{\text{core}} \sim 2$ fm. Since $\langle r^2 \rangle$ is related to the neutron separation energy we seek systems with neutron separation energies less than 1 MeV
- By this definition the deuteron is the lightest halo nucleus, and the pionless EFT for few-nucleon systems is a specific case of halo EFT
- ^{22}C , ^{11}Li , ^{12}Be , ^{19}B , ^{62}Ca (hypothesized), and ^3H : all s-wave 2n halos

Halo nuclei: examples

<http://nupecc.org>



Outline

- What is Halo EFT and what does it do for us?
 - Halo EFT for Borromean s-wave $2n$ halos
 - Measuring nn relative-momentum distributions using fast breakup
 - The unitary limit in momentum distributions of $2n$ halos
 - The surprisingly small uncertainty of $d + {}^4\text{He} \rightarrow {}^6\text{Li} + \gamma$ at low energies
 - What it teaches us about ${}^6\text{Li}$
-

Two-body scattering amplitude in Halo EFT

$$t_0^{2B}(E) = -\frac{2\pi}{m_R} \frac{1}{k \cot \delta(E) - ik}; \quad k = \sqrt{2m_R E}$$

$$k \cot \delta(E) = -\frac{1}{a} + \frac{1}{2}rk^2 + O(k^4 R^3)$$

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Elastic scattering: this is effective-range theory with built-in UQ

Lagrangian: shallow S- and P-states

$$\begin{aligned}\mathcal{L} = & c^\dagger \left(i\partial_t + \frac{\nabla^2}{2M} \right) c + n^\dagger \left(i\partial_t + \frac{\nabla^2}{2m} \right) n \\ & + \sigma^\dagger \left[\eta_0 \left(i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_0 \right] \sigma + \pi_j^\dagger \left[\eta_1 \left(i\partial_t + \frac{\nabla^2}{2M_{nc}} \right) + \Delta_1 \right] \pi_j \\ & - g_0 [\sigma n^\dagger c^\dagger + \sigma^\dagger n c] - \frac{g_1}{2} \left[\pi_j^\dagger (n i\overleftrightarrow{\nabla}_j c) + (c^\dagger i\overleftrightarrow{\nabla}_j n^\dagger) \pi_j \right] \\ & - \frac{g_1}{2} \frac{M - m}{M_{nc}} \left[\pi_j^\dagger i\overrightarrow{\nabla}_j (nc) - i\overleftrightarrow{\nabla}_j (n^\dagger c^\dagger) \pi_j \right] + \dots ,\end{aligned}$$

- c, n: “core”, “neutron” fields. c: boson, n: fermion.
 - σ, π_j : S-wave and P-wave fields
 - Minimal substitution generates leading EM couplings
 - Additional EM couplings at sub-leading order
-

But it's more than just s-wave nn & nc scattering

- So not just two-body scattering: also EM processes
Chen, Rupak, Savage (1999);
Hammer, DP (2011)
 - And other partial waves
Bertulani, Hammer, van Kolck (2003); Bedaque, Hammer, van Kolck (2003); Brown & Hale (2005); Braun et al. (2018); Ando (2016-present)
 - Extension to pp, p-core, and cluster-cluster scattering
Kong & Ravndal (1999); Higa, Hammer, van Kolck (2008);
Ryberg, Forssén, Hammer, Platter (2014, 2016)
 - Expansion around limit of a bound or unbound state near threshold. Include higher-order effects in ERE in proportion to their importance. Expansion in kR_{core} , where R_{core} is scale of unresolved core physics
 - Extends to three-body states at cost of one additional parameter (S_{2n})
Bedaque, Hammer, van Kolck (1999); Hammer & Mehen (2001); Bedaque et al. (2002); Ji, Platter, DP (2009)
 - Then predictive for four-body states (bosons or distinguishable particles) at LO accuracy
Platter, Hammer, Meißner (2005); Bazak, Kirscher, König, Pavon Valderrama, Barnea, van Kolck (2018)
-

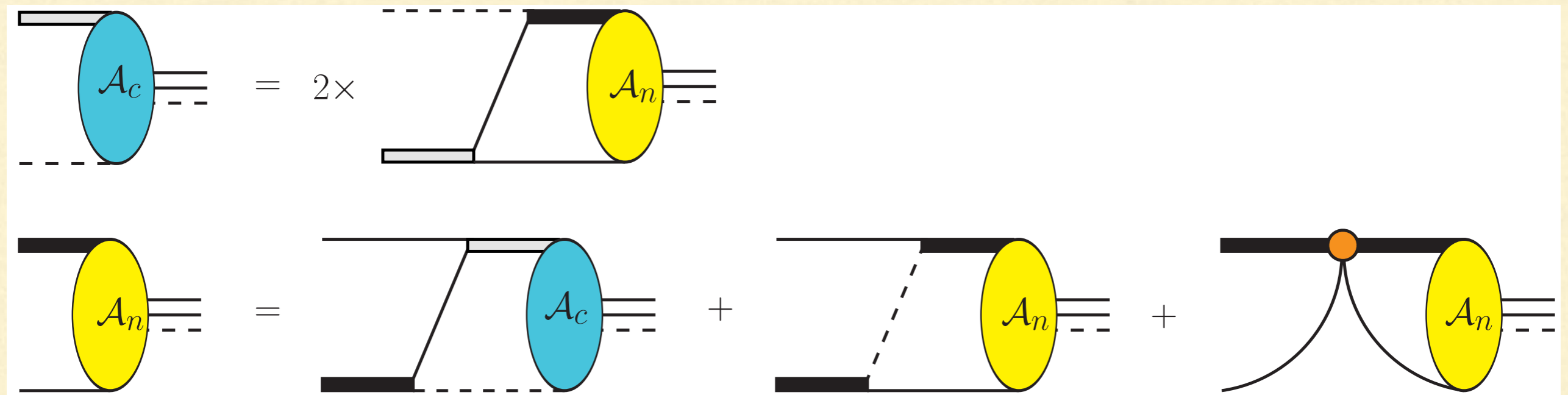
Equations for s-wave $2n$ halo

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- Core- n and n - n contact interactions at leading order: solve 3B problem

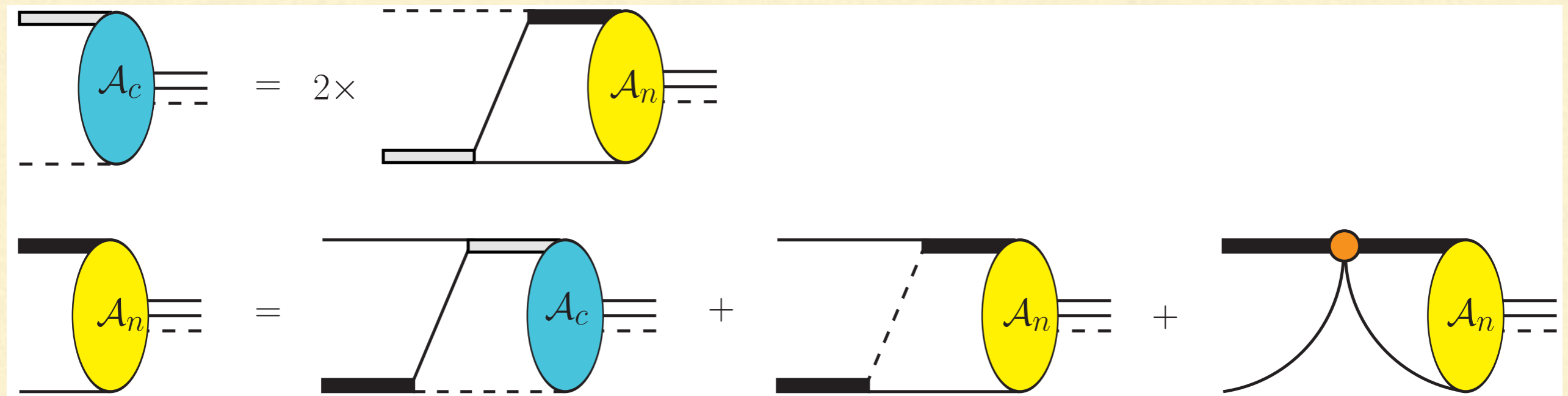


- (cn)- n contact interaction to stabilize three-body system

Equations for s-wave 2n halo

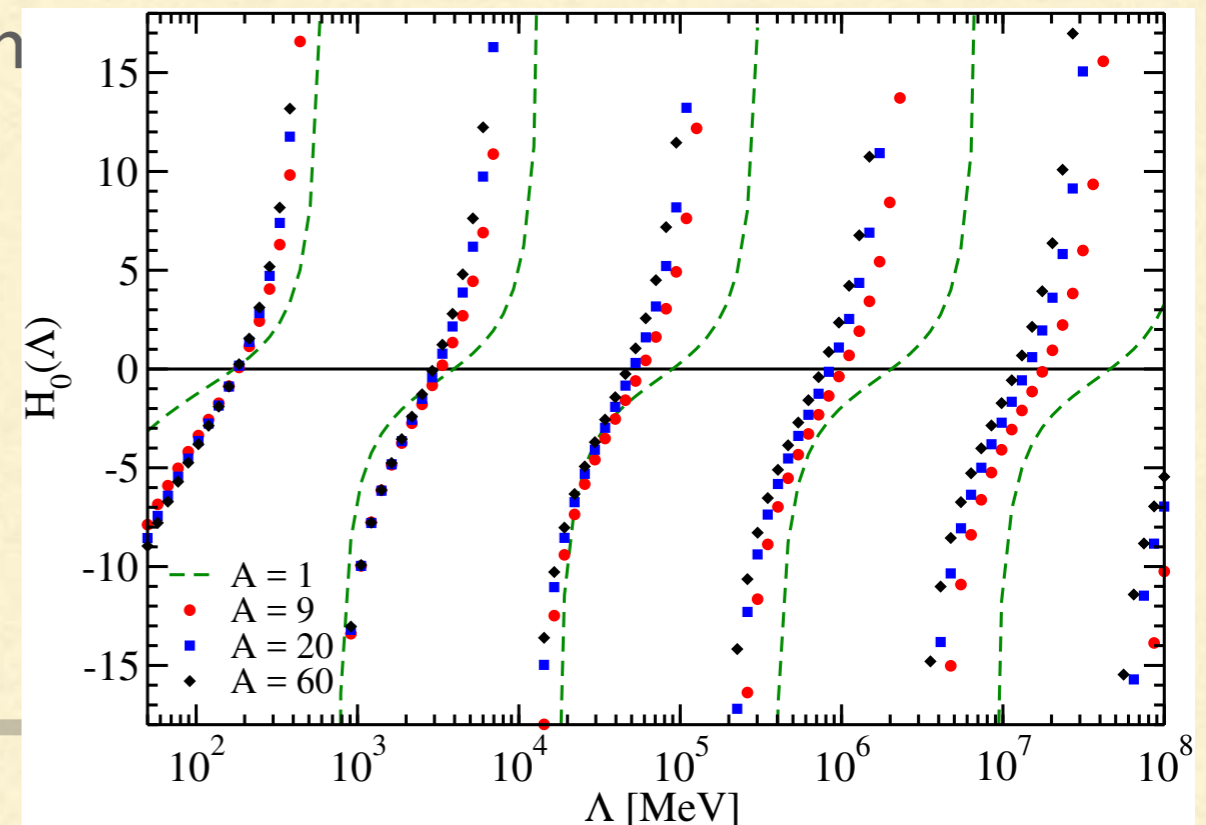
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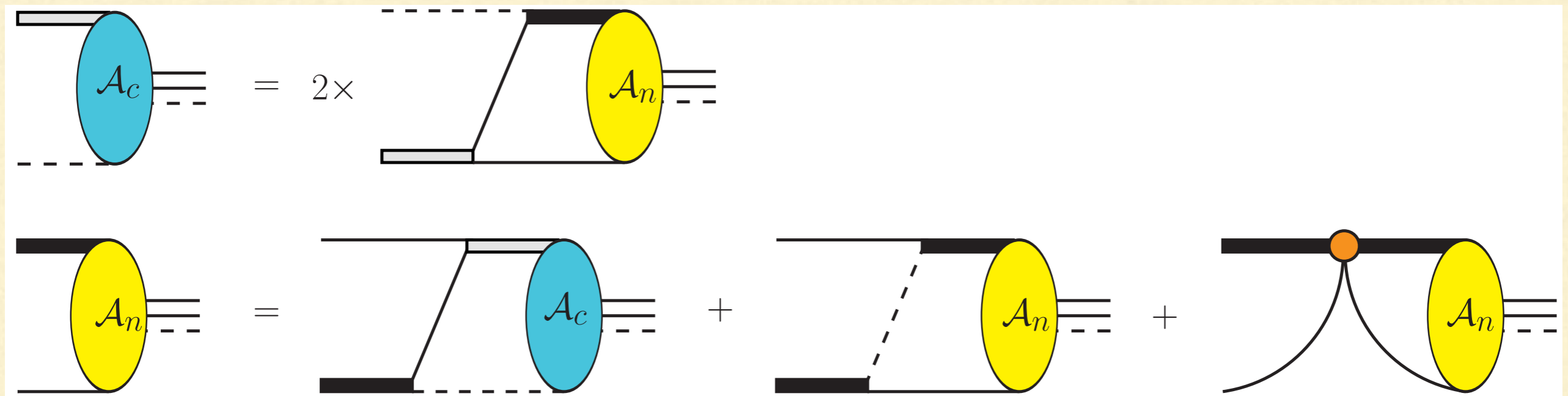
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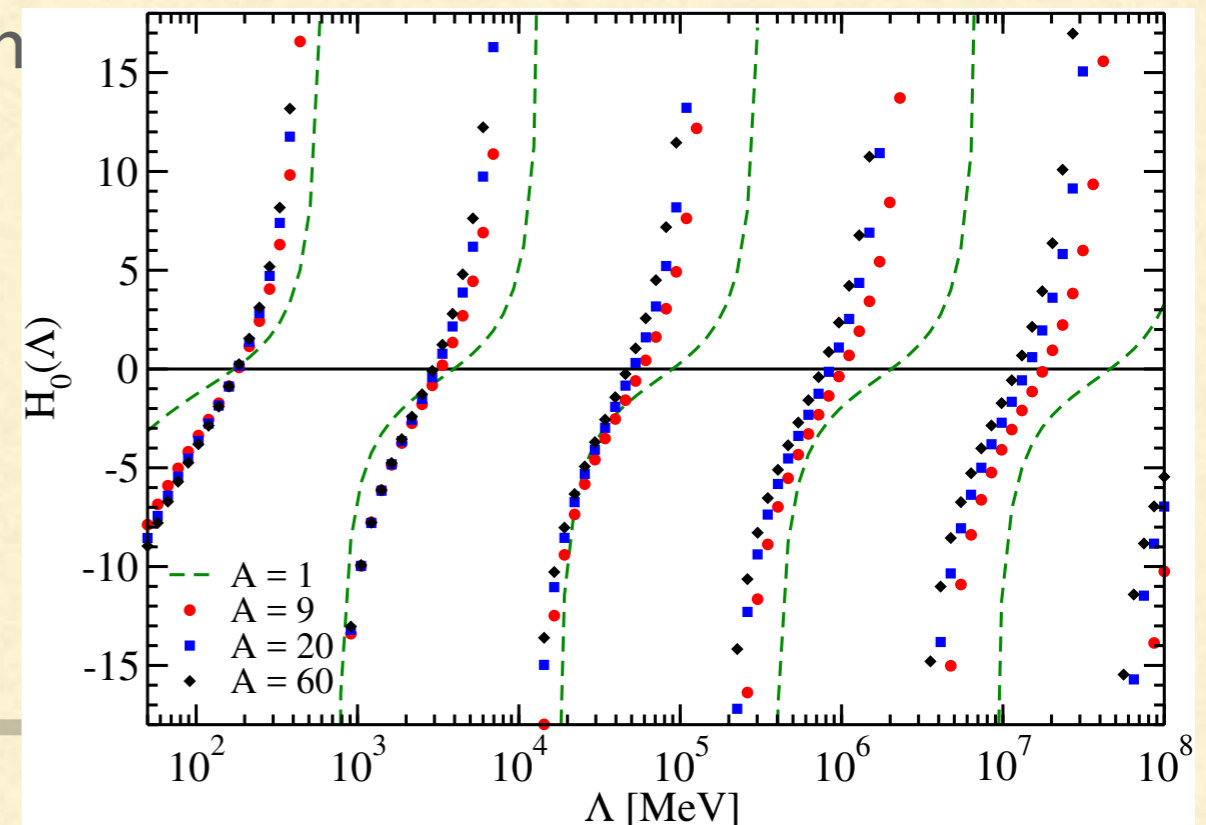


- (cn)-n contact interaction to stabilize the halo

- Efimov-Thomas effects

- Inputs: $E_{nn} = 1/(m a_{nn}^2)$, E_{nc} , $S_{2n} (=B)$

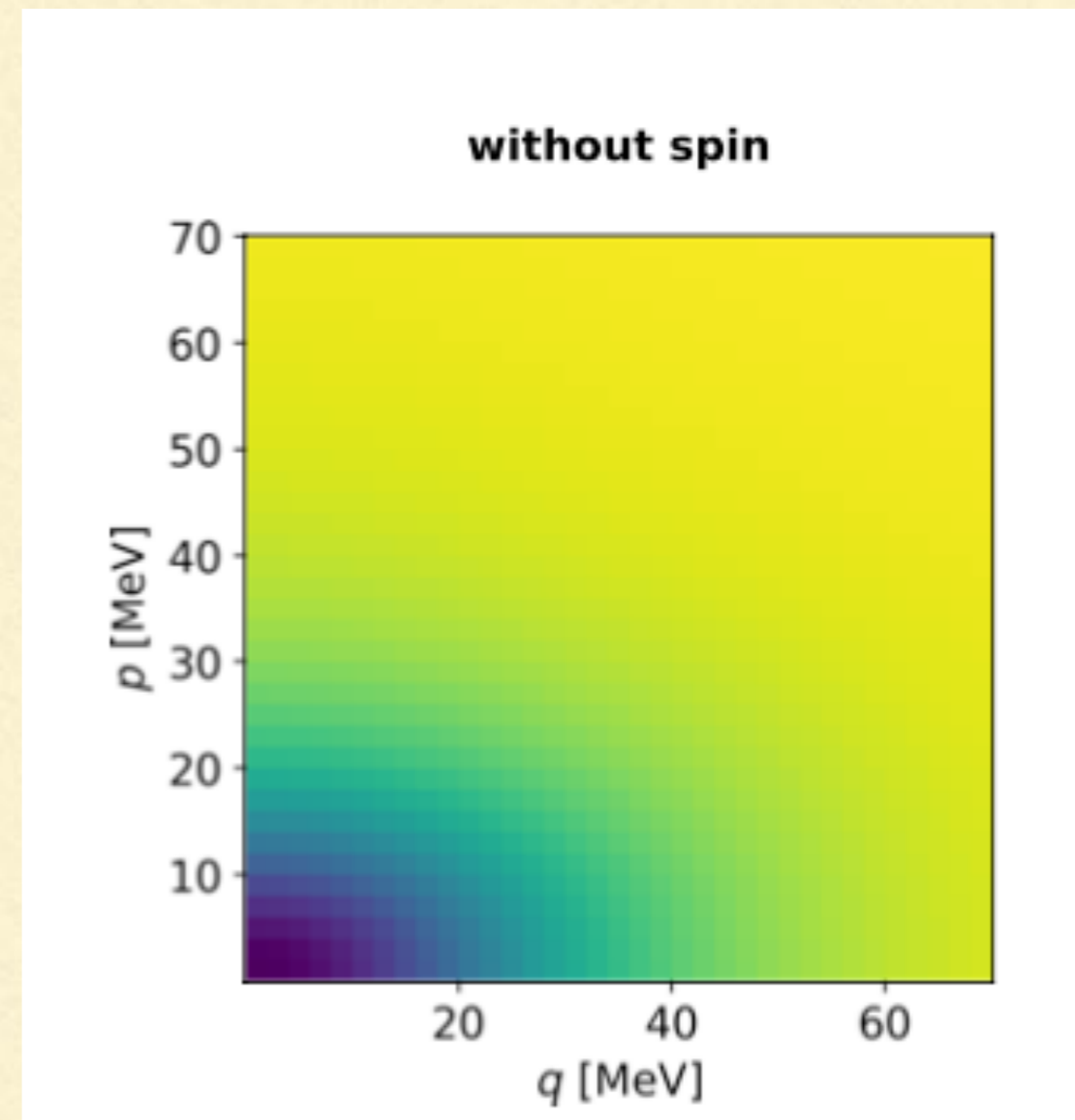
- Output: everything; up to $O(R_{core}/R_{halo})$



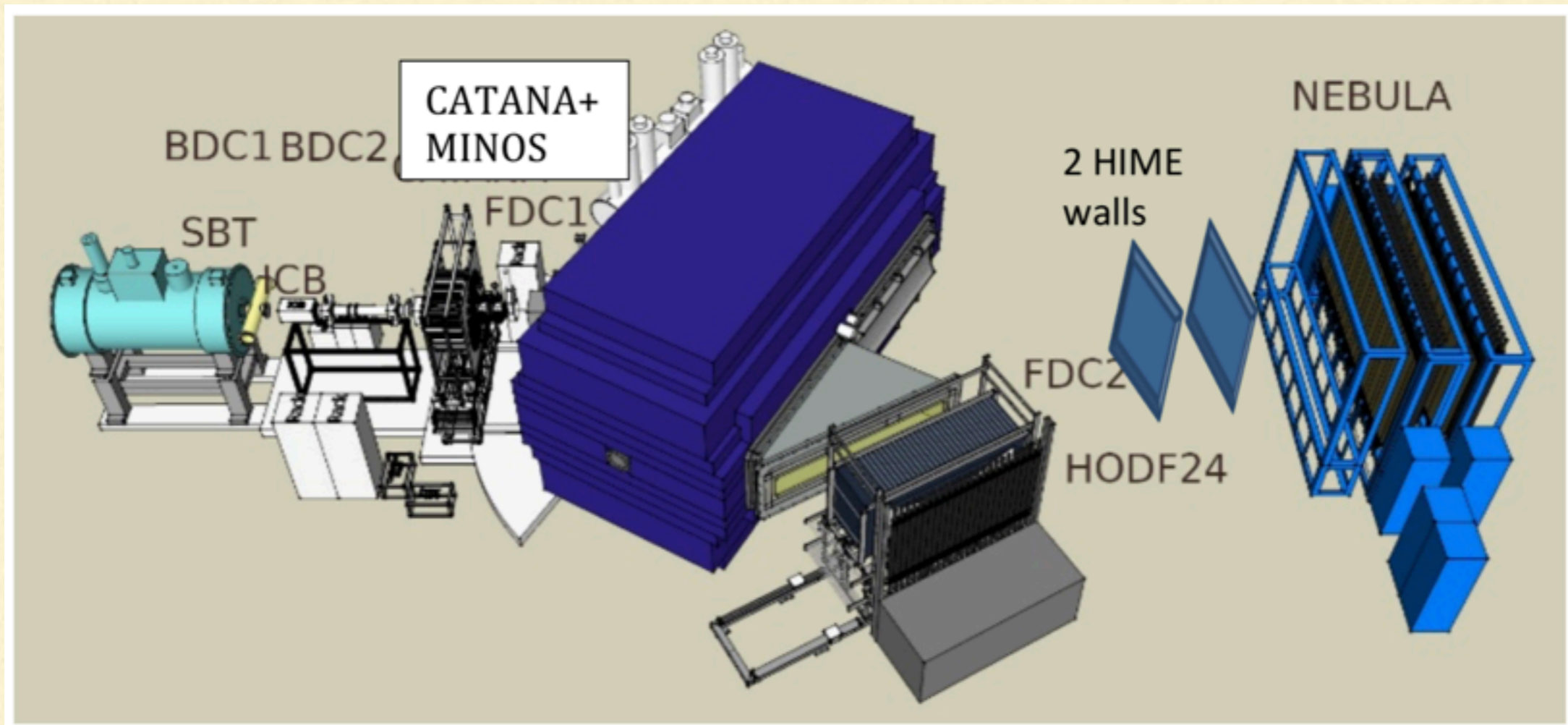
^{11}Li as a $2n$ halo

- $a_{nn} = -18.7$ fm, $E_{nc} = 0.026$ MeV
- $S_{2n} = 369$ keV
- Calculations done with a cutoff of 470 MeV, but results checked for a cutoff of 700 MeV
- Here results with a spin-0 core, but we also examined case of spin-3/2 core
- Results identical if spin-1 and spin-2 nc interactions have equal strength

^{11}Li wave function



RIKEN experiment with ${}^6\text{He}$ beam

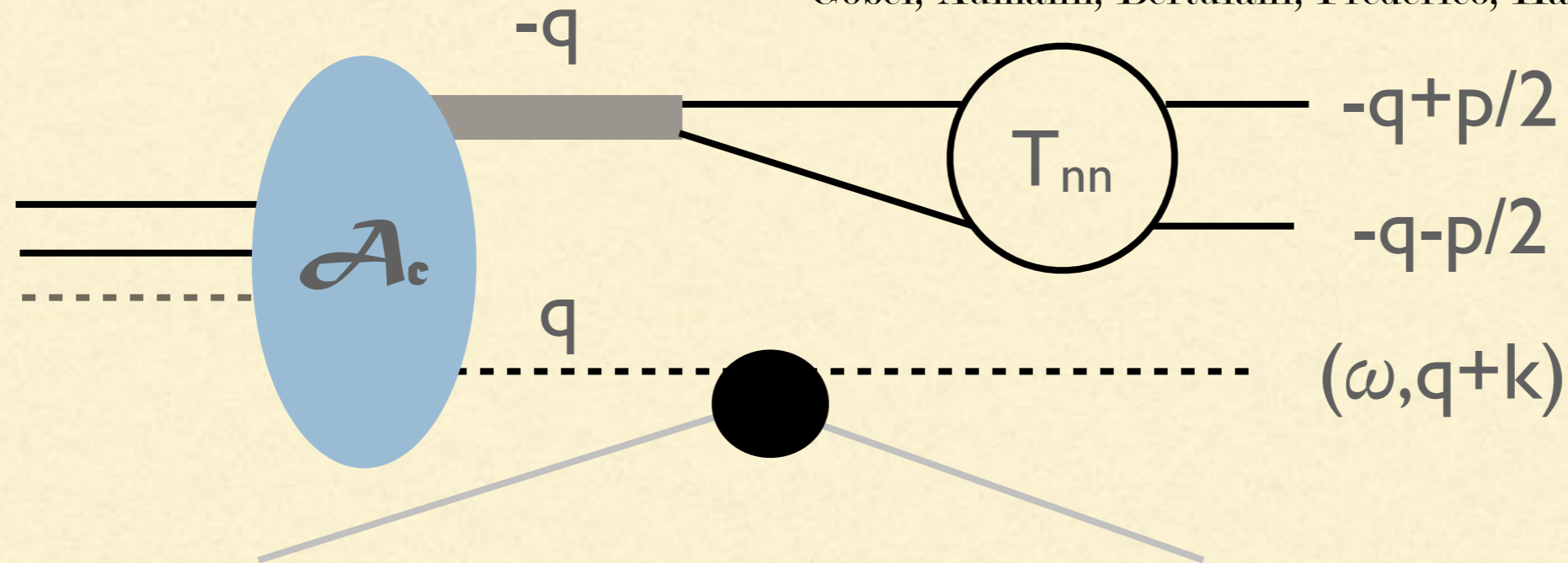


Tom Aumann spokesperson

- Detect proton and alpha in TPC
- Detect neutrons in HIME + NEBULA: excellent energy resolution

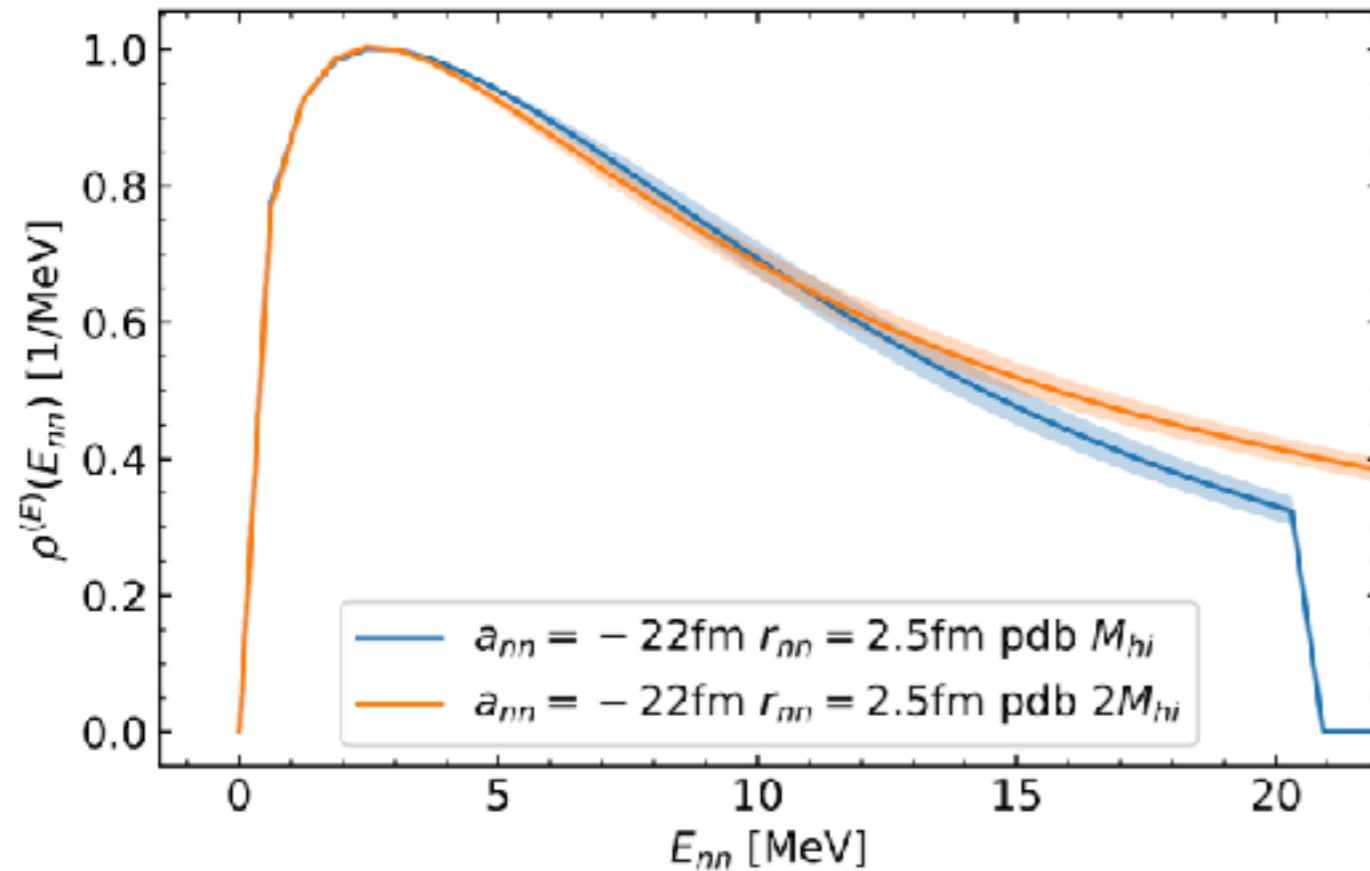
${}^6\text{He}(p,p'\alpha)$ and the nn scattering length

Göbel, Aumann, Bertulani, Frederico, Hammer, Phillips, PRC (2021)



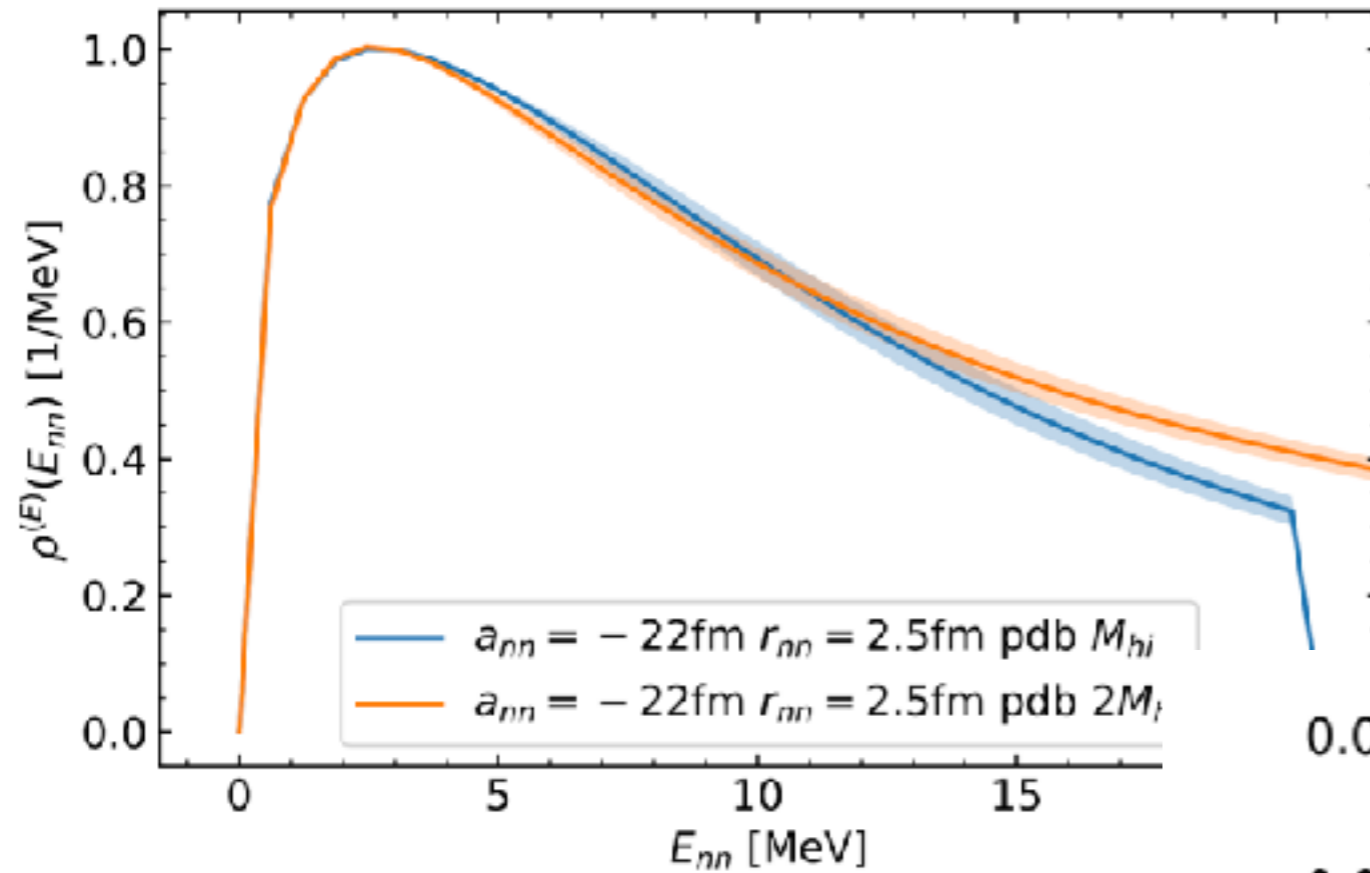
- Quasi-free alpha-particle knockout can leave nn pair almost at rest
- Final-state interaction then generates significant dependence of neutron relative-energy spectrum $f(p^2/m_n)$ on a_{nn}
- ${}^6\text{He}$ acts as a “holder” for low-momentum neutrons
- Neutrons actually move fast in lab. frame: inverse kinematics

Neutron energy distribution in ${}^6\text{He}$

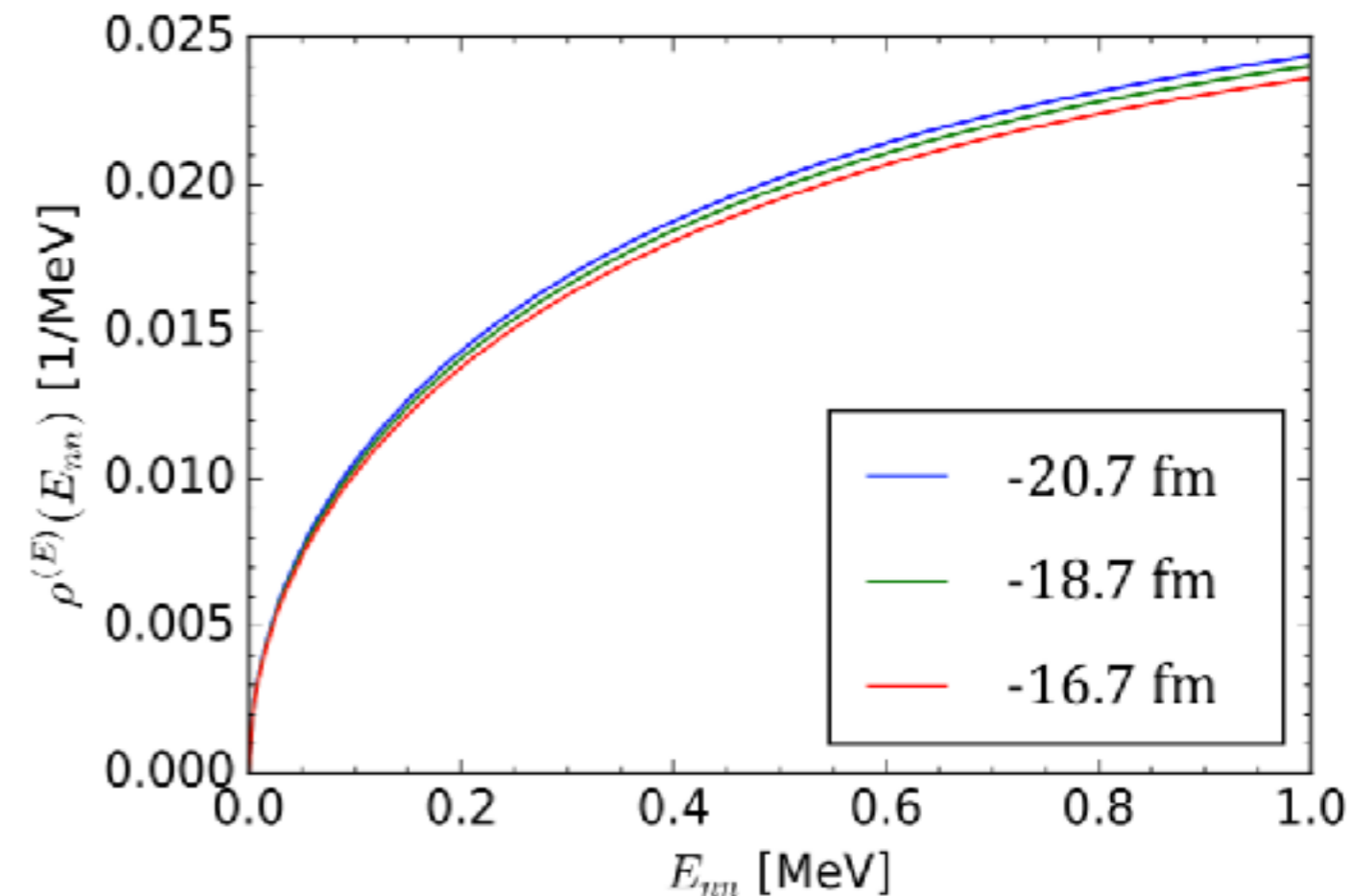


No FSI
included
at first

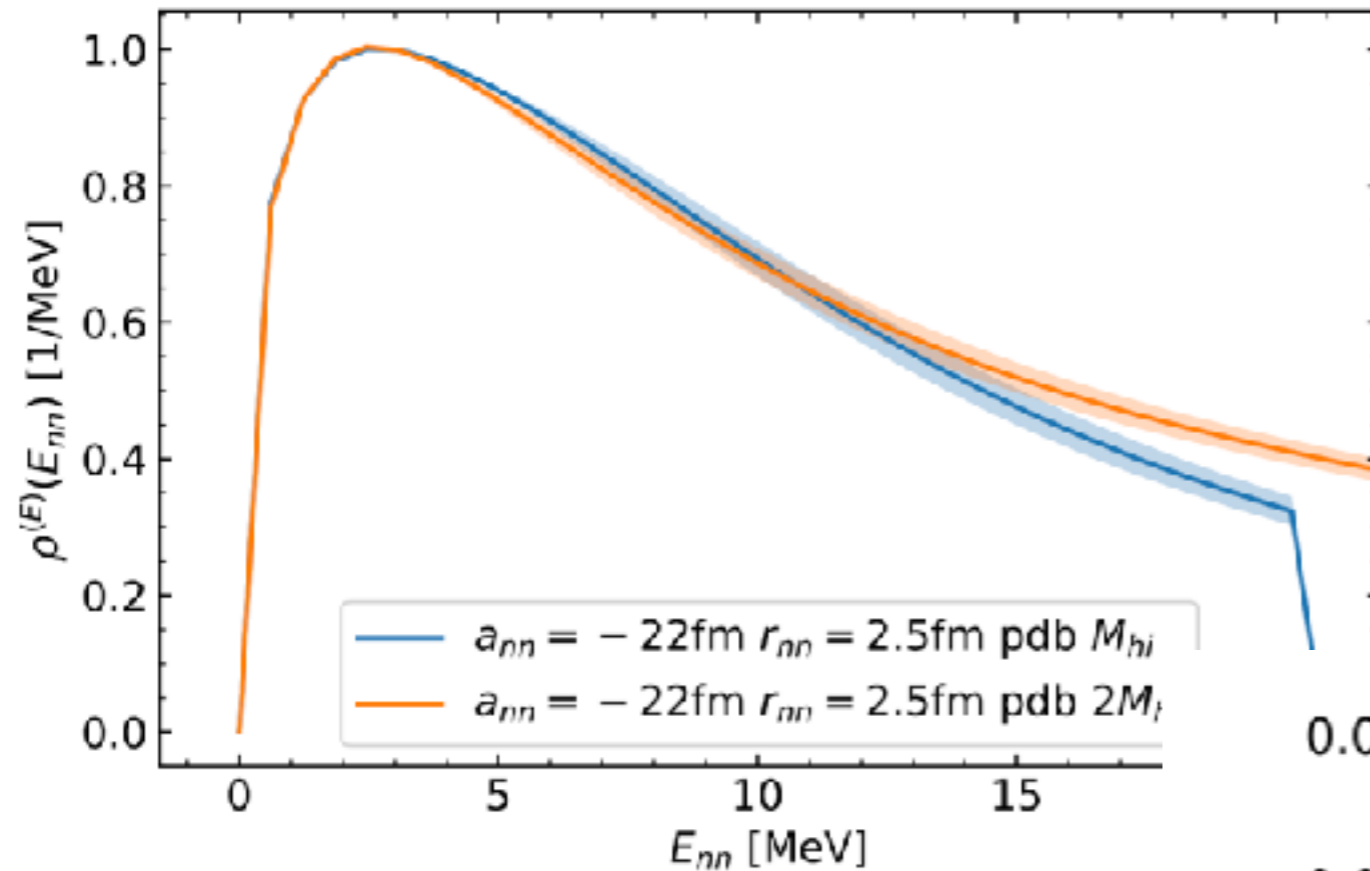
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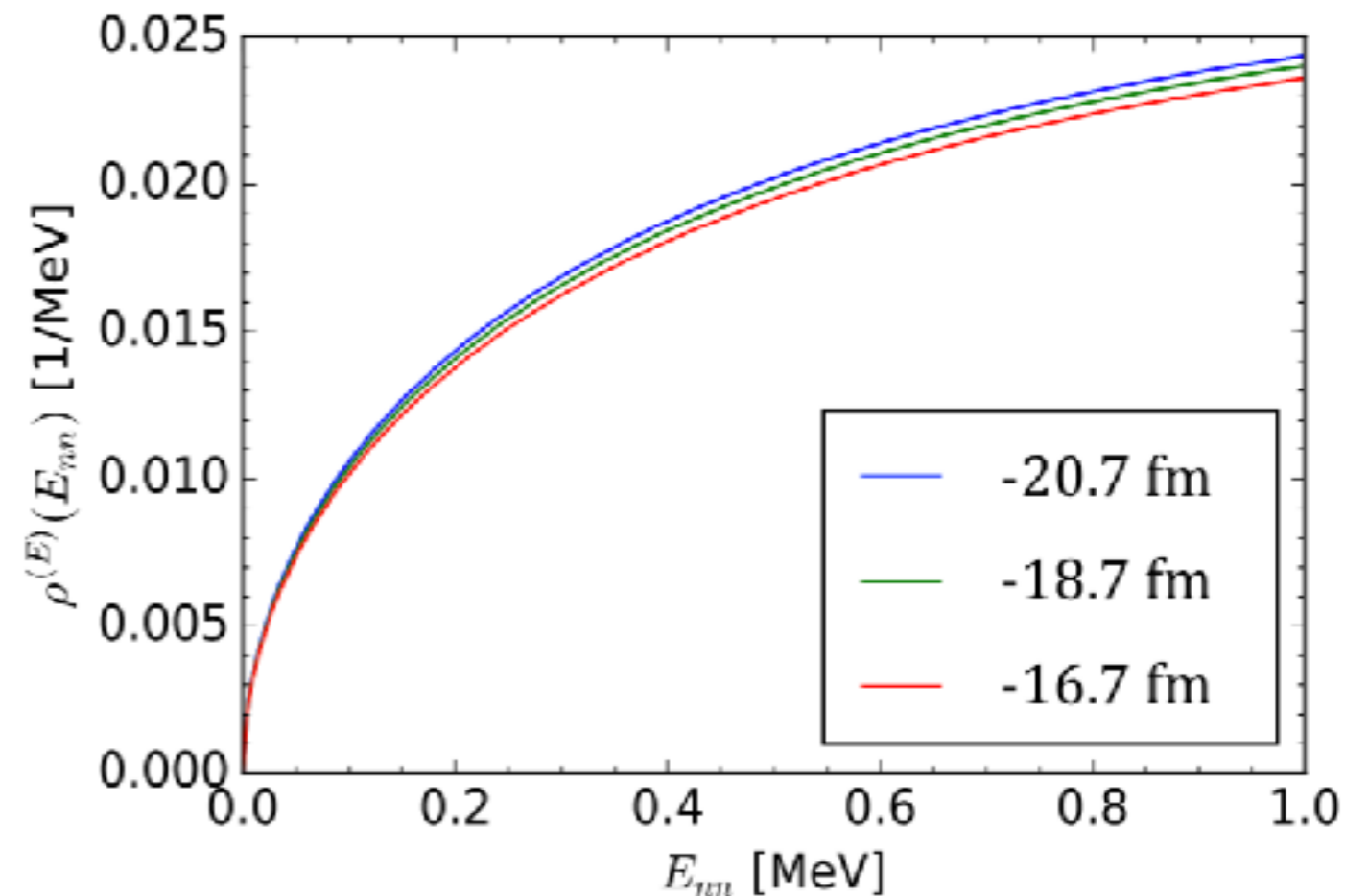


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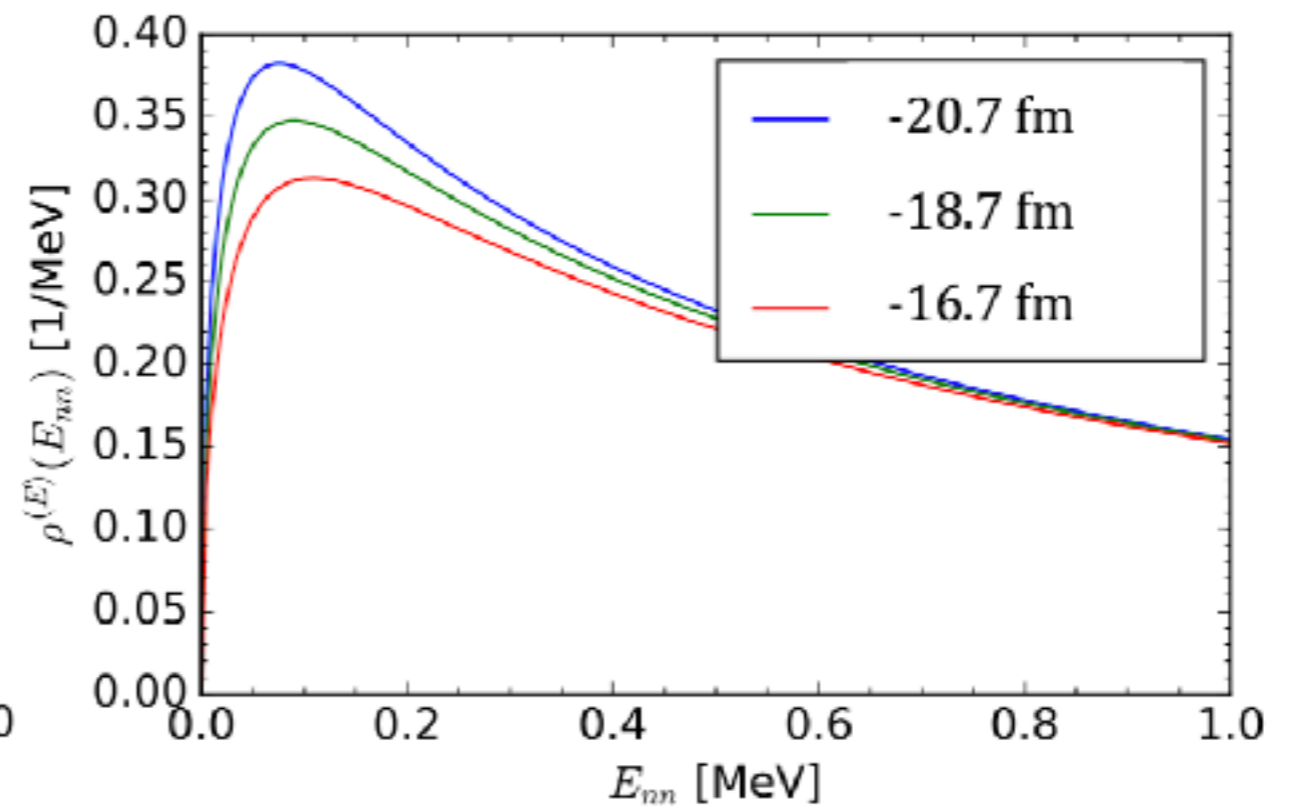
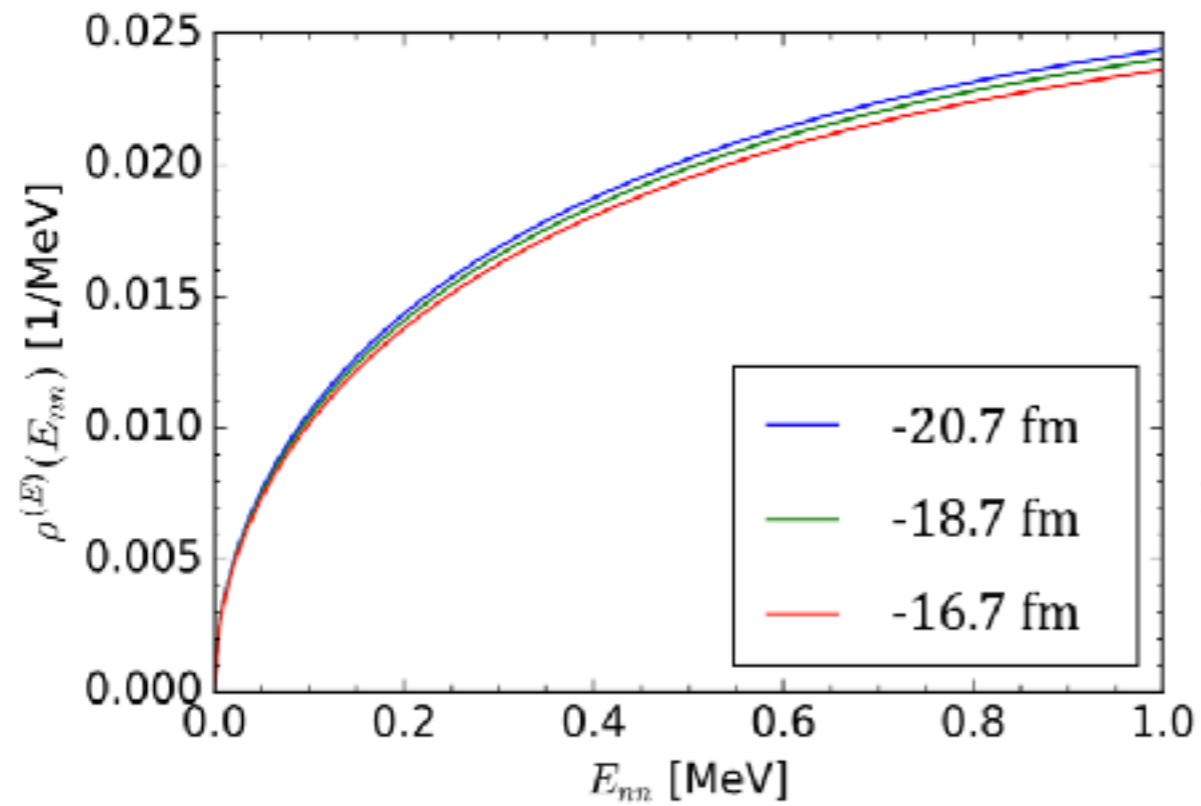


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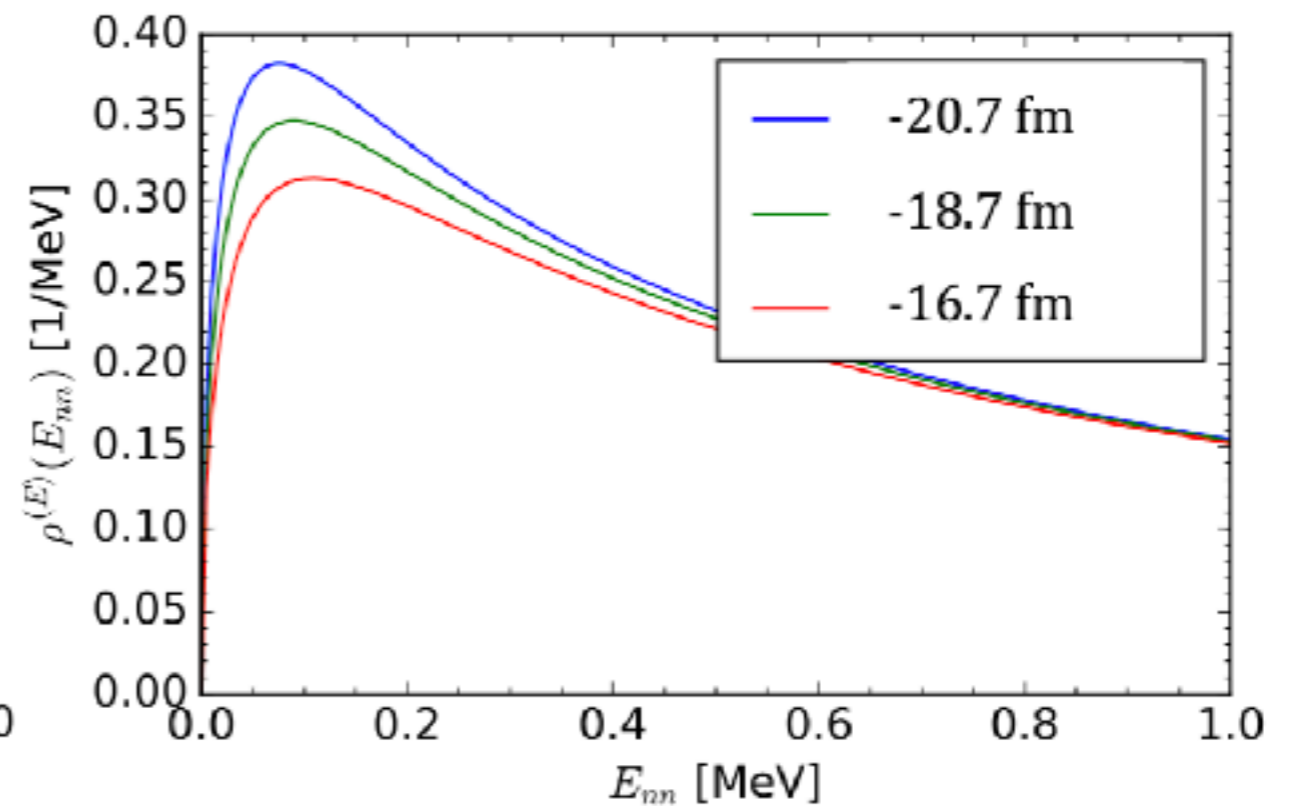
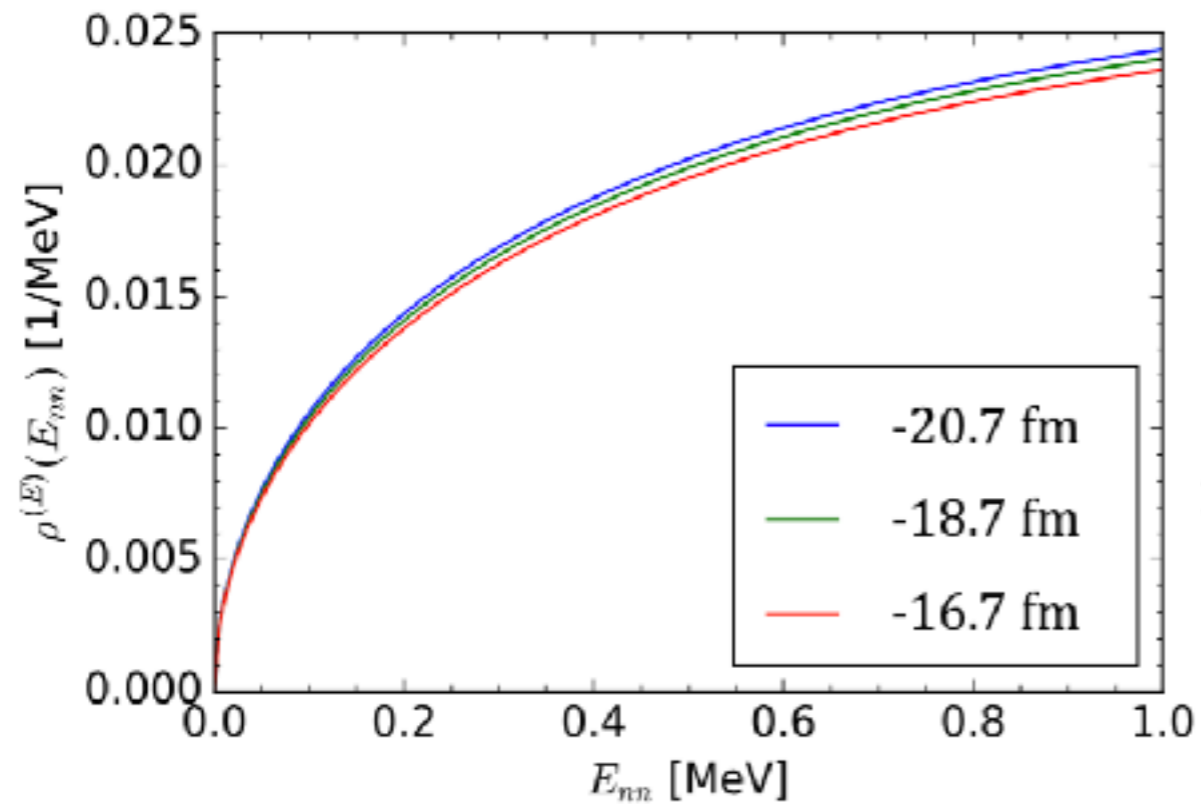
${}^6\text{He}$ structure at low momentum not significantly affected by cutoff or a_{nn} (or r_{nn})



Sensitivity to a_{nn} and (not) r_{nn}

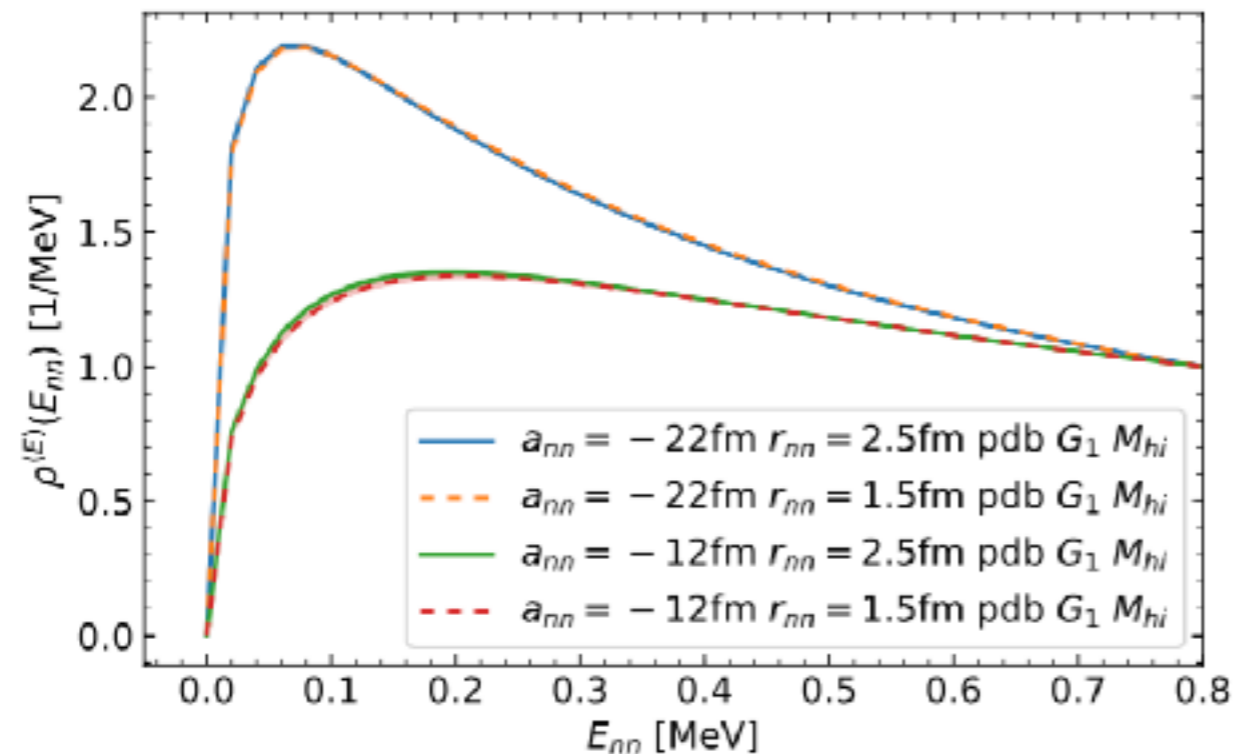
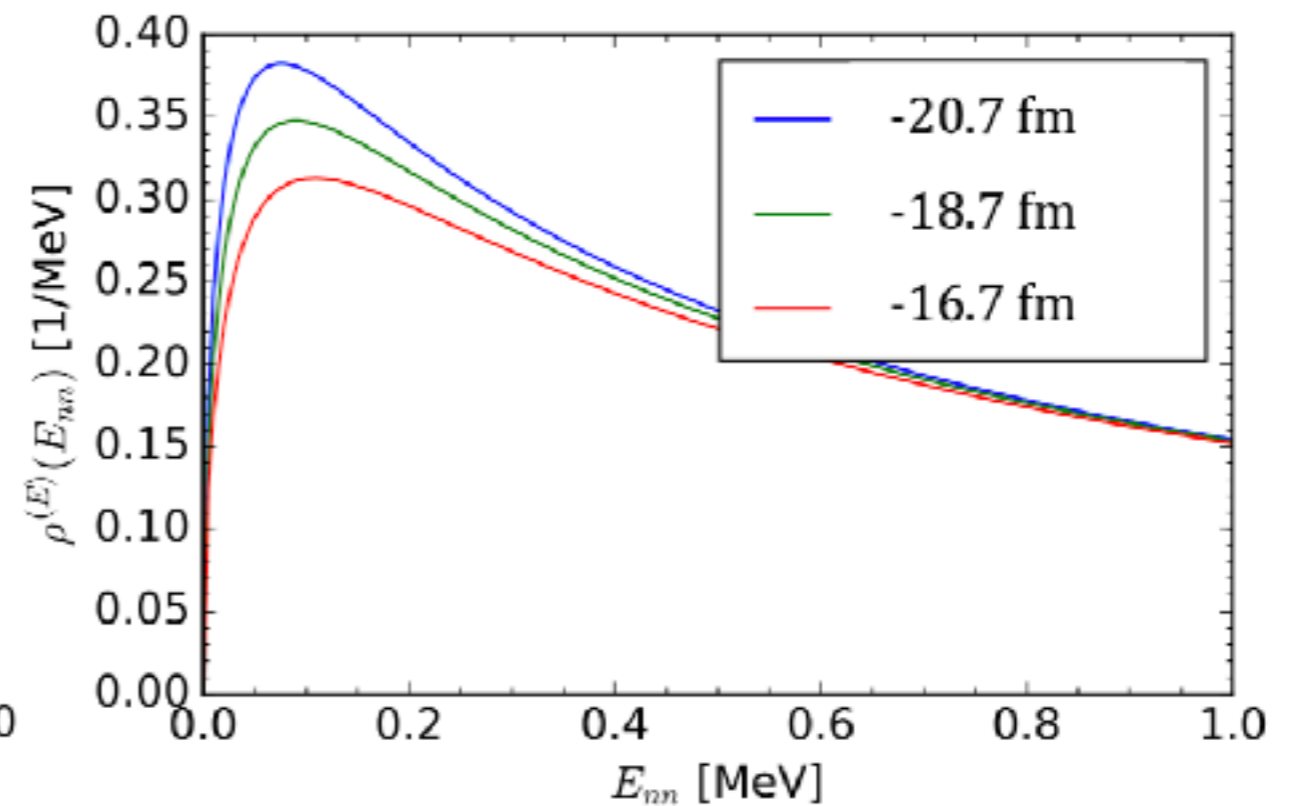
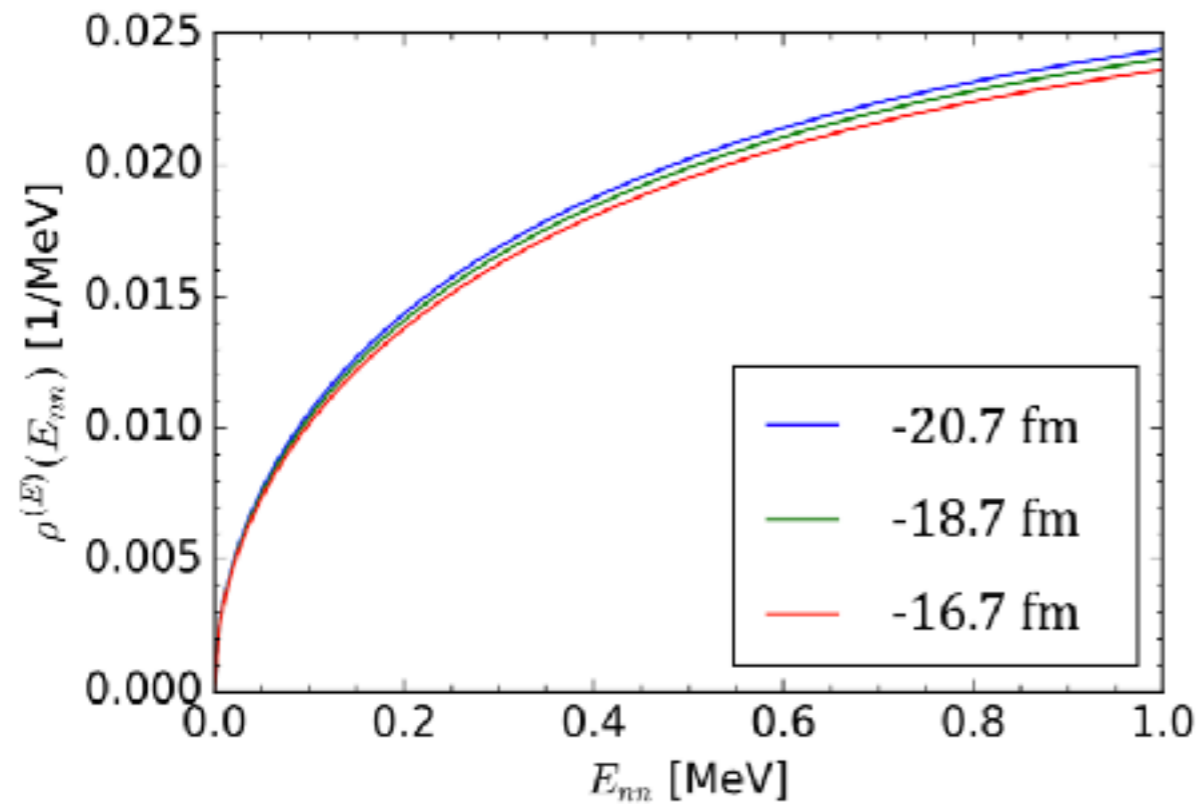


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Note that since this is not an absolute measurement we need to decide how to normalize the spectra

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What we learn from ${}^6\text{He}$

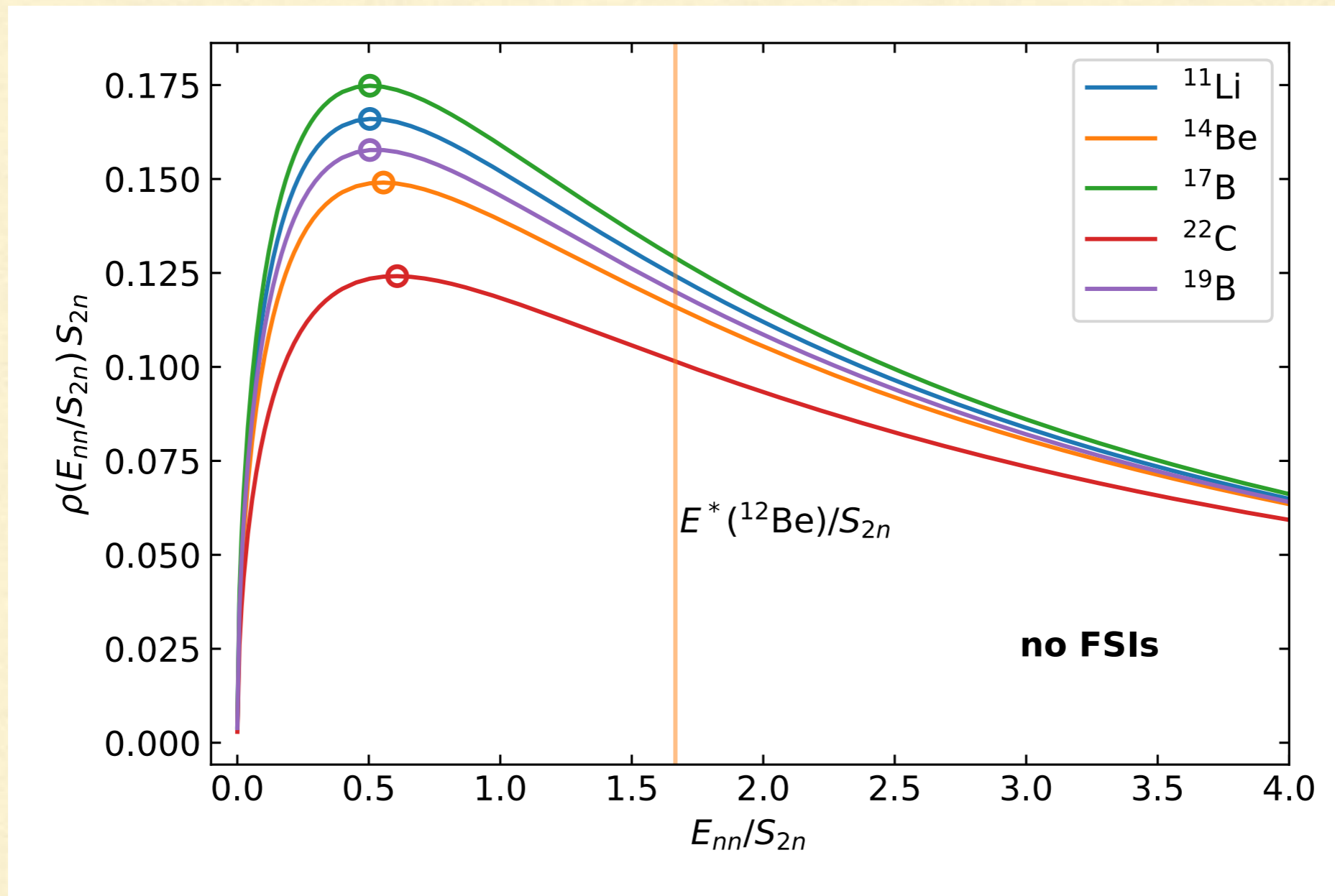
So ${}^6\text{He}$ relative-momentum distribution work shows:

- Very little sensitivity to a_{nn} in “structure part”
- NLO corrections to structure part should be small (not this talk)
- Even less sensitivity to r_{nn}
- Strong a_{nn} modification from FSI
- This modification can be well described by an enhancement factor

$$\rho^{\text{full}}(E_{nn}) \approx G(E_{nn}, a_{nn}, r_{nn})\rho^{\text{gs}}(E_{nn})$$

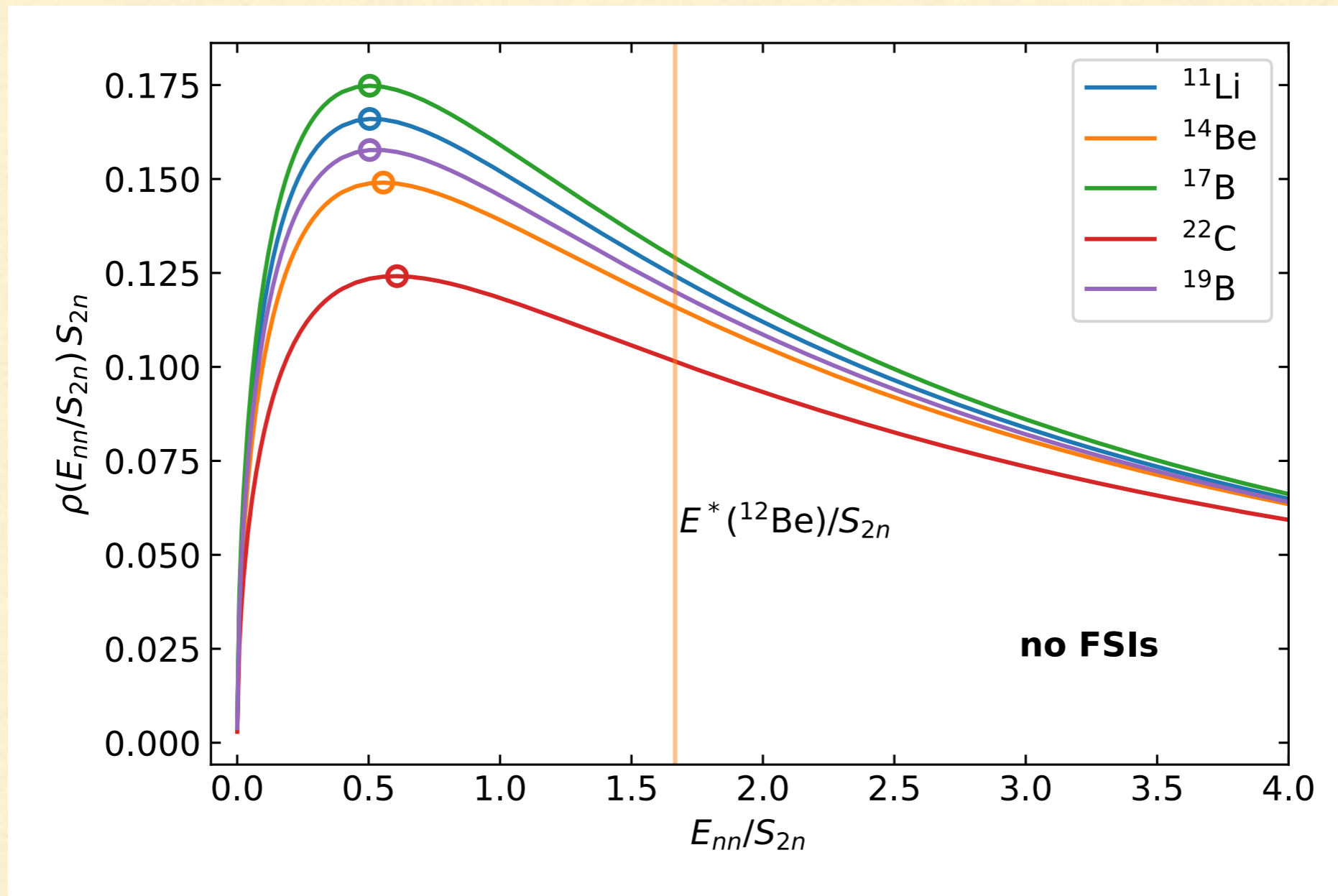
nn momentum distributions for s-wave 2n halos

Göbel, Hammer, DP, PRC (2024)



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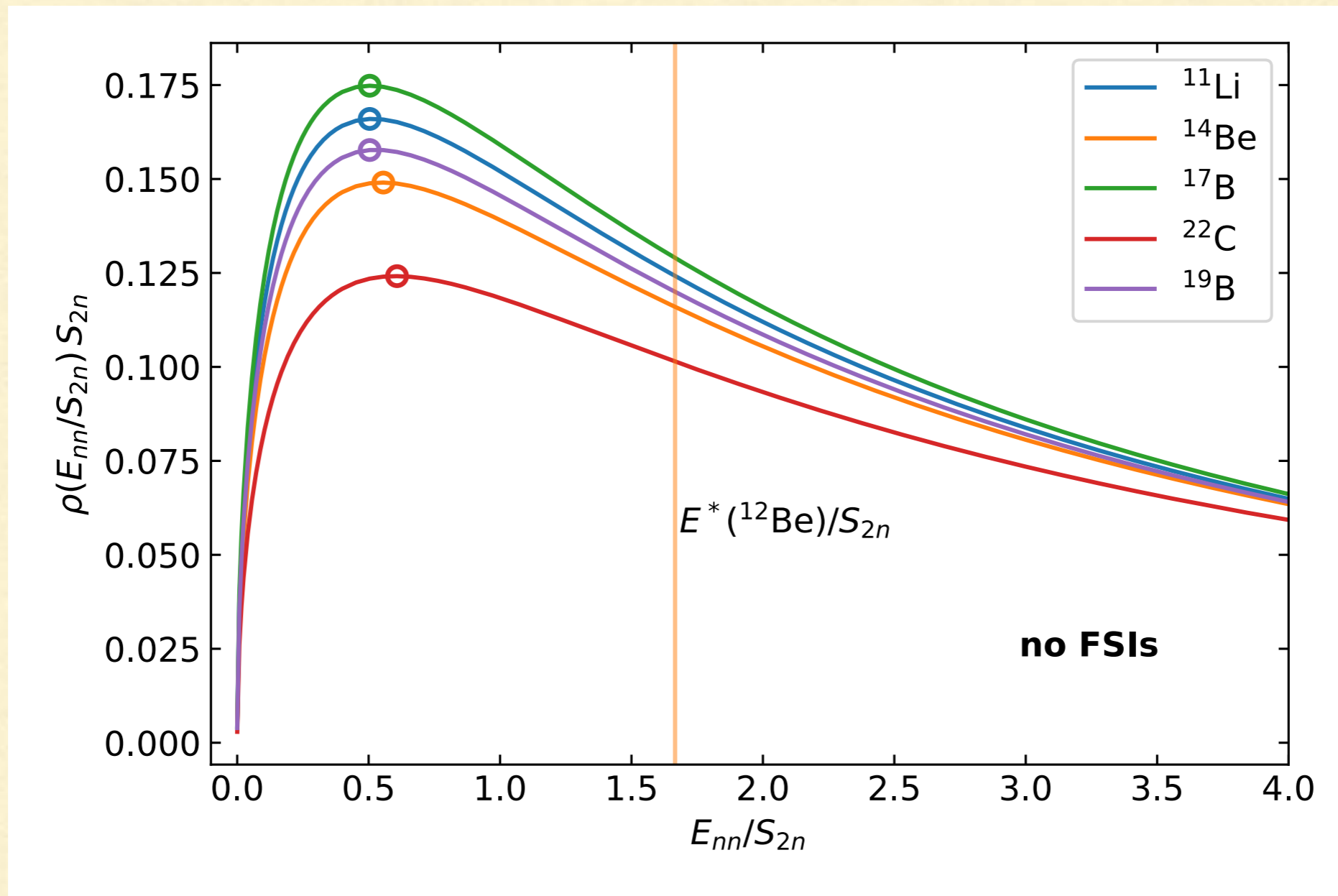
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- Want to plot in dimensionless units so all 2n halos fall on same plot

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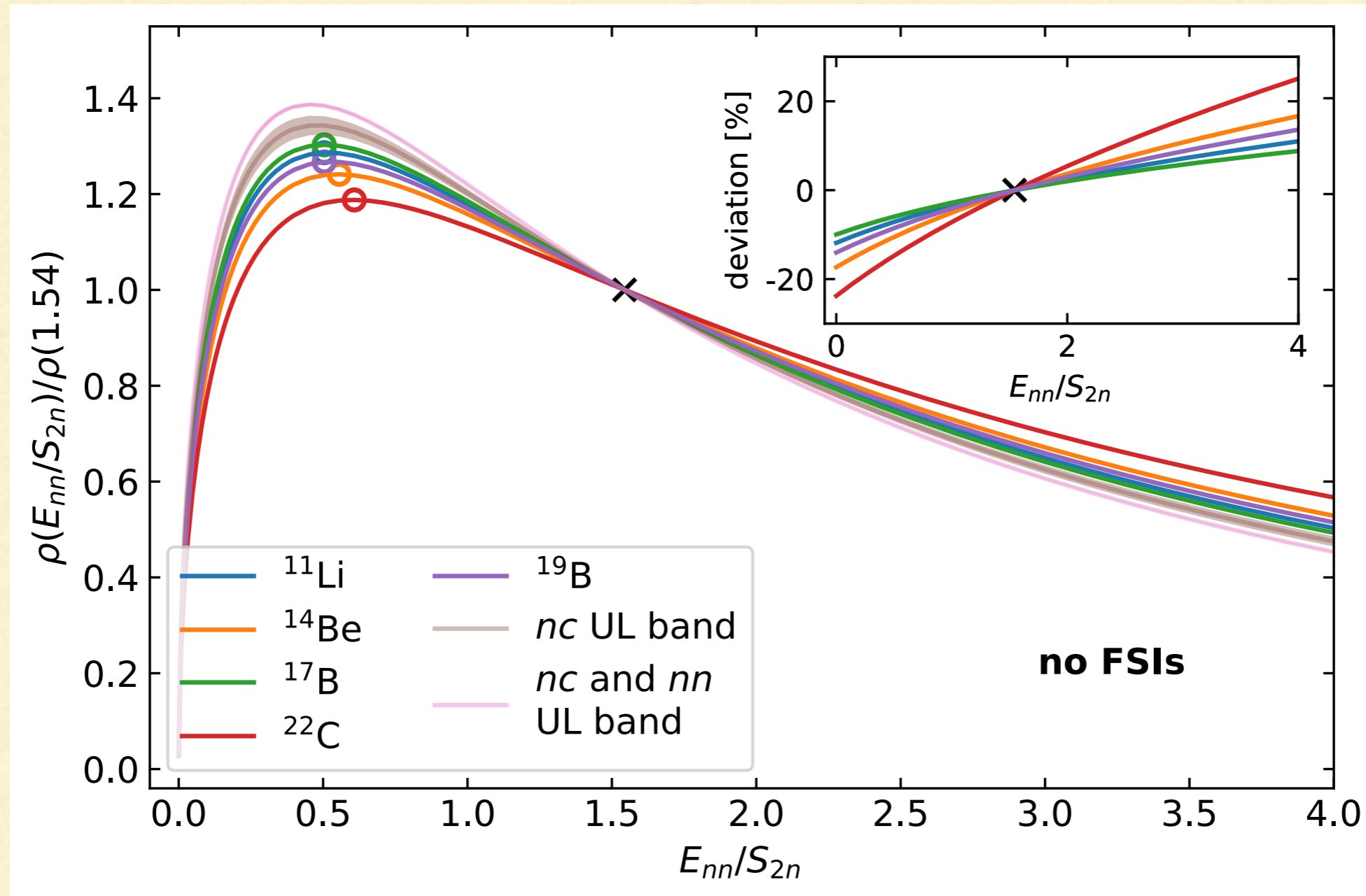
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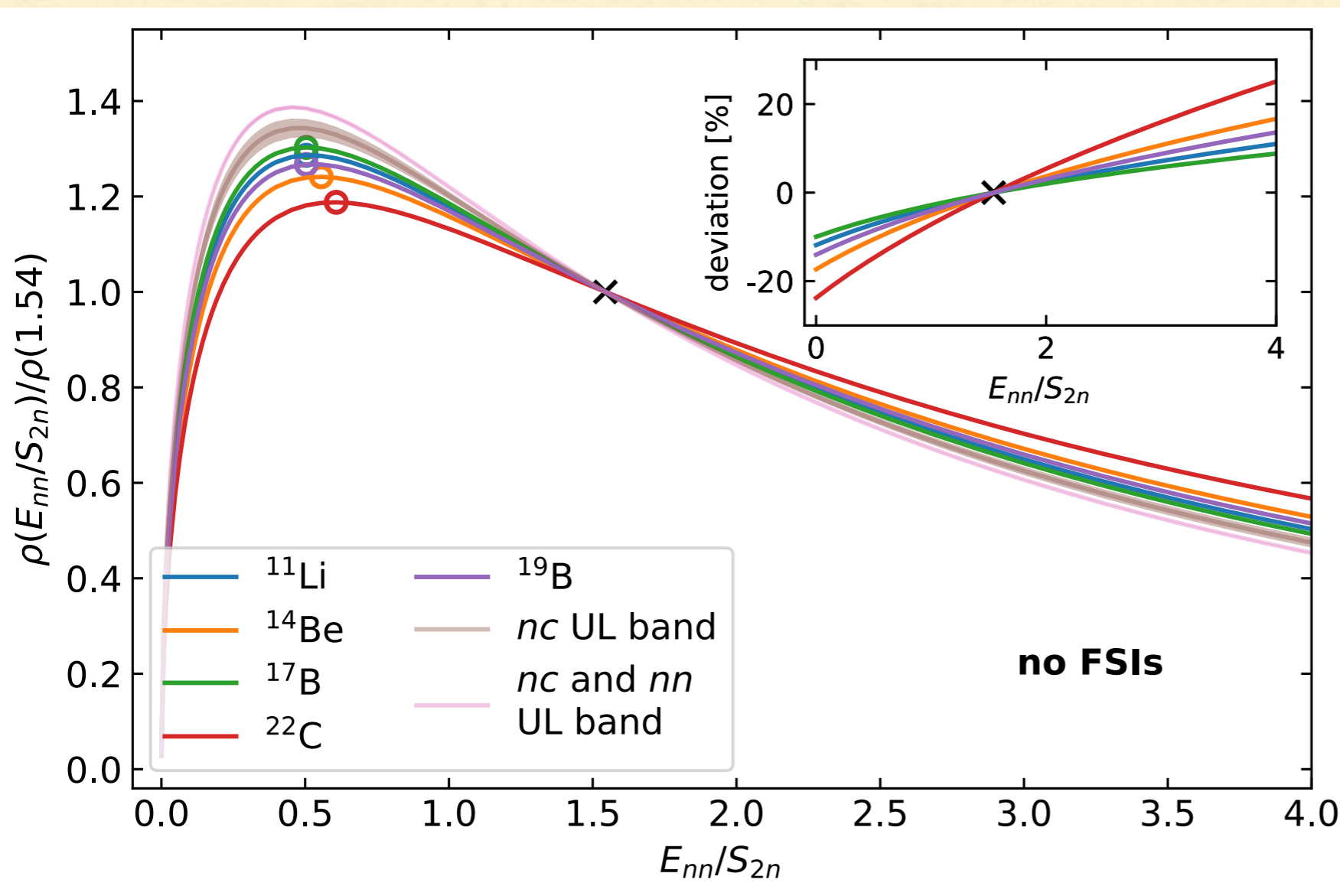
Going to the unitary limit

- The “unitary limit” is another limit on top of LO Halo EFT: $|a| \rightarrow \infty$
 - The 2B state is then right at threshold. No scales left: $r \rightarrow 0, |a| \rightarrow \infty$.
 - 2B amplitude: $t^{2B}(E = k^2/m_R) \sim \frac{1}{ik}$, 2B problem has conformal invariance
 - Efimov effect in 3B system: infinite tower of bound states $E^{(n)}/E^{(n-1)} = 515$
 - Ratio of 4B and 3B binding energies $E^{4B,n}/E^{3B,n} = 4.6$ + excited tetramer
Platter & Hammer (2007); Deltuva (2012)
 - Scaling dimension of multi-neutron momentum distributions calculable
Son & Hammer (2022); Chowdry, Mishra, Son (2023)
 - **This talk:** momentum distribution of nn relative-momentum distributions in Borromean s-wave 2n halos
-

The unitary limit can be seen in 2n halos

Cf. for ^{19}B : Hiyama, Lazauskas, Marqués, Carbonell (2019); Hiyama, Lazauskas, Carbonell, Frederico (2023)

$$\rho^{g.s.}(E_{nn}/S_{2n}; V_{nn}, V_{nc}, S_{2n}, A) \approx \rho^{g.s.}(E_{nn}/S_{2n})$$



i.e., ρ is the same function for all halos to better than 20%

- Works because halos are sufficiently bound that precise values of a_{nn} and a_{nc} do not matter.
- A dependence also goes away

But can it be measured?

Results for other 2n halos after FSI modification

- Use Møller operator to include nn FSI:

$$\psi_c^{(\text{wFSI})}(p, q) = \langle p, q; \zeta_c, \xi_c | (1 + t_{nn}(E_p)G_0(E_p)) | \Psi \rangle$$

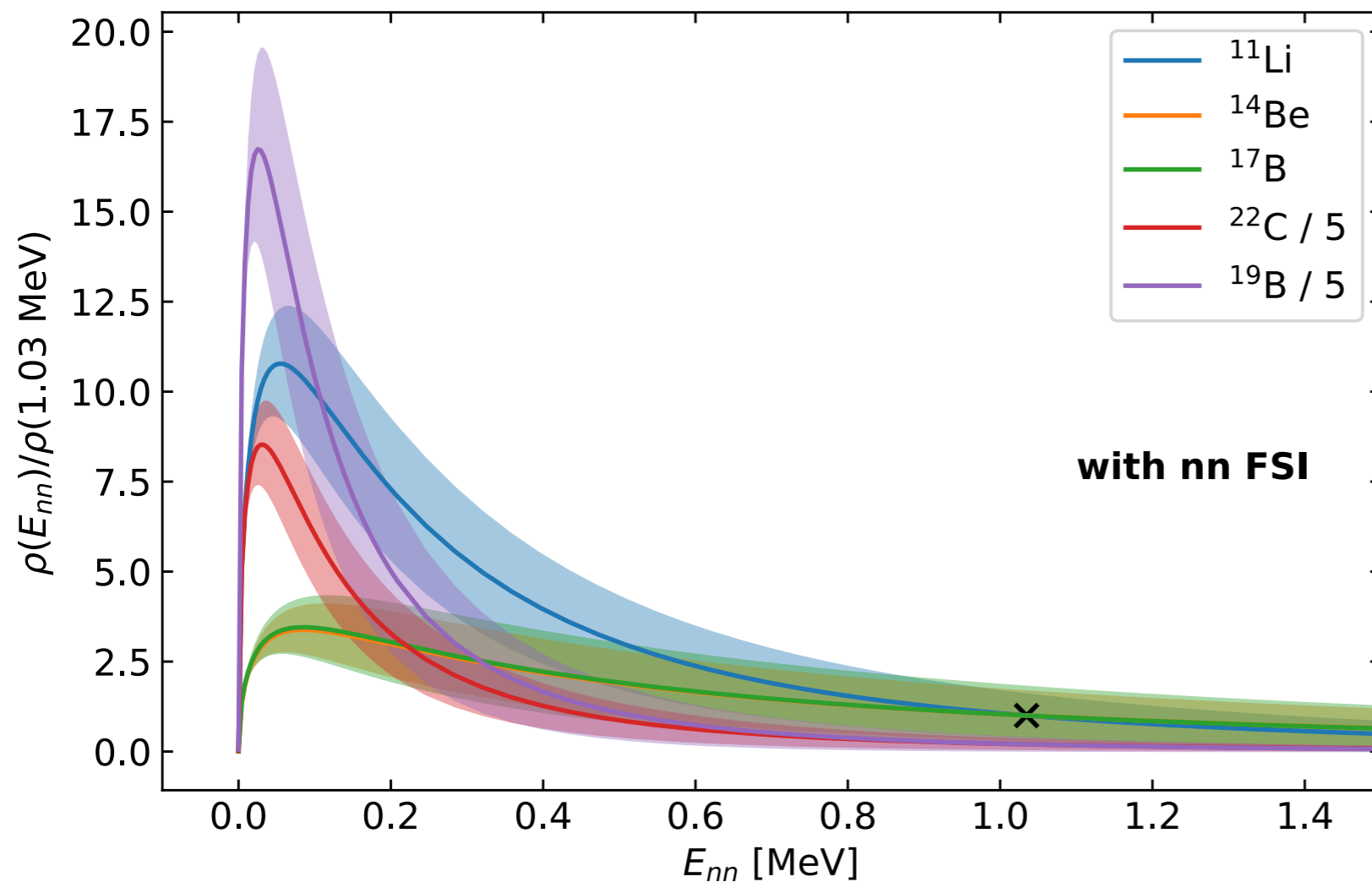
- Relative energy distribution: $\rho(E_{nn}) = \sqrt{\frac{m_n}{4E_{nn}}} \int_0^\Lambda dq q^2 |\Psi_c(p_{nn}, q)|^2 p_{nn}^2$

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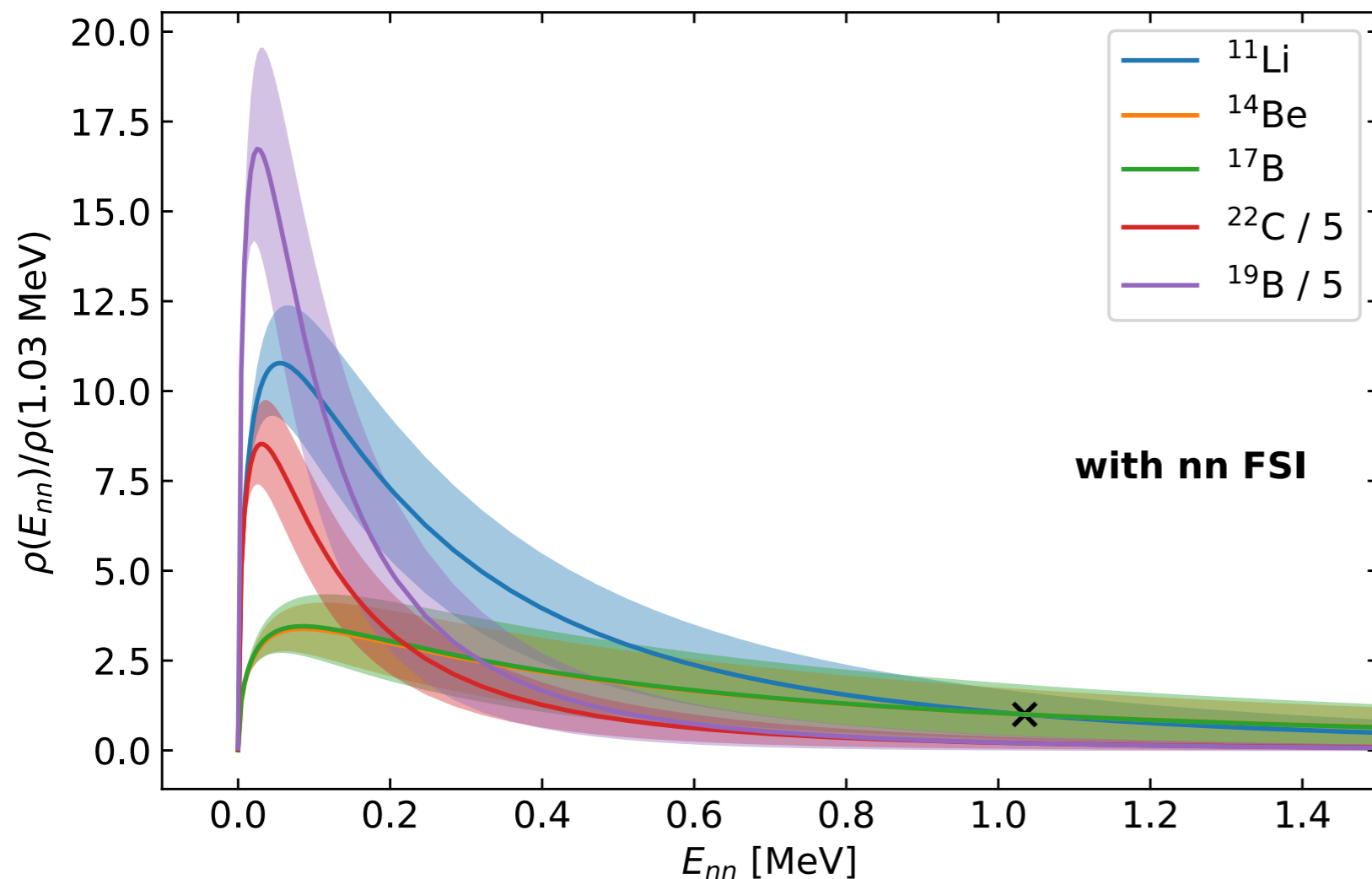


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nn interaction
produces
variation on scale
 $1/(m_n a_{nn}^2)$

Ground-state
distribution varies
on scale S_{2n}

Divide out by FSI factor

Hypothesis: $\rho^{(\text{wFSI})}(E_{nn}/S_{2n}; a_{nn}, r_{nn}) \approx G(E_{nn}; a_{nn}, r_{nn})\rho^{g.s.}(E_{nn}/S_{2n})$

Divide out by FSI factor

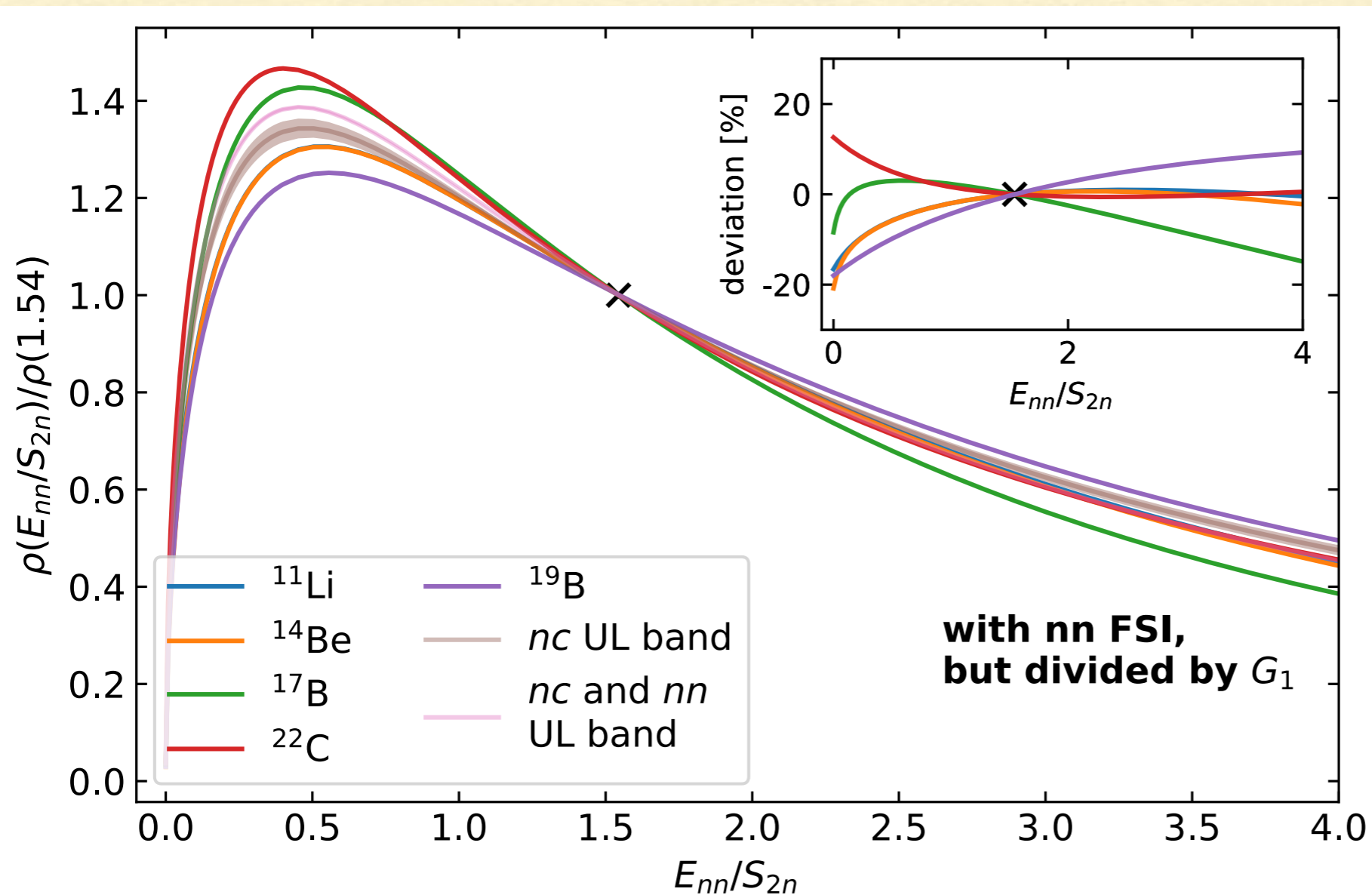
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$$\text{So we plot: } \rho(E_{nn}/S_{2n}) = \frac{\rho^{\text{full LO Halo EFT}}(E_{nn}/S_{2n}; a_{nn})}{G(E_{nn}; a_{nn}, r_{nn})}$$

Divide out by FSI factor

Hypothesis: $\rho^{(\text{wFSI})}(E_{nn}/S_{2n}; a_{nn}, r_{nn}) \approx G(E_{nn}; a_{nn}, r_{nn})\rho^{g.s.}(E_{nn}/S_{2n})$

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Distributed $\pm 20\%$
around UL
result for ρ

Summary and outlook: part I

Summary:

- Halo EFT describes the low-momentum physics of halo nuclei
- There is a unique nn momentum distribution in (s-wave) $2n$ halos
- Approximately the unitary limit momentum distribution: nothing about the nn and nc interactions matters except that they're strong
- This claim can be checked by measuring the nn relative energy distribution on several halos and dividing out the effects of FSI

To do:

- NLO corrections & comparison to ab initio calculations
 - ${}^3\text{H}$ and other non-Borromean halos?
-

Analyzing ab initio calculations of ${}^6\text{Li}$ in Halo EFT

Hebborn, Brune, Phillips, in preparation

- Want to describe $\alpha(d,\gamma){}^6\text{Li}$ (motivated by Big Bang production of ${}^6\text{Li}$). ${}^6\text{Li}$ has a deuteron separation energy of 1.5 MeV: comparable to the deuteron binding energy of 2.2 MeV, cf. proton separation energy in $\alpha \approx 20$ MeV.

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- NCSMC calculation: diagonalization of nuclear Hamiltonian using an over-complete basis

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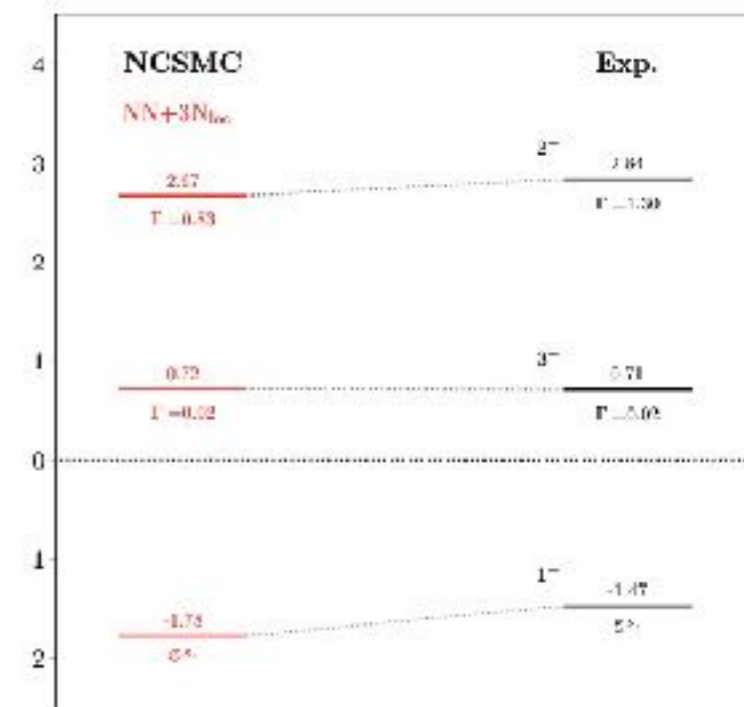
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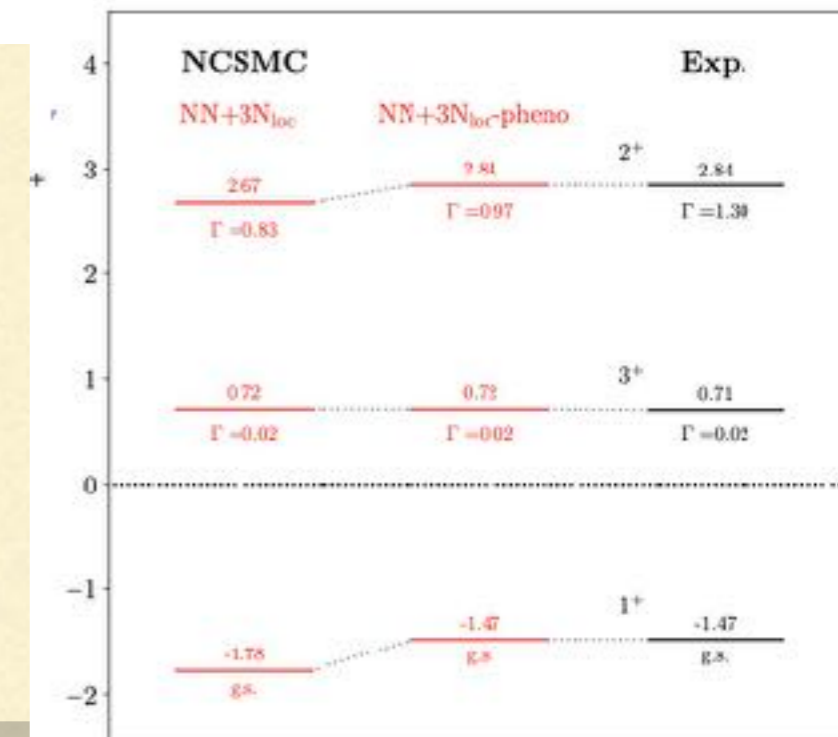
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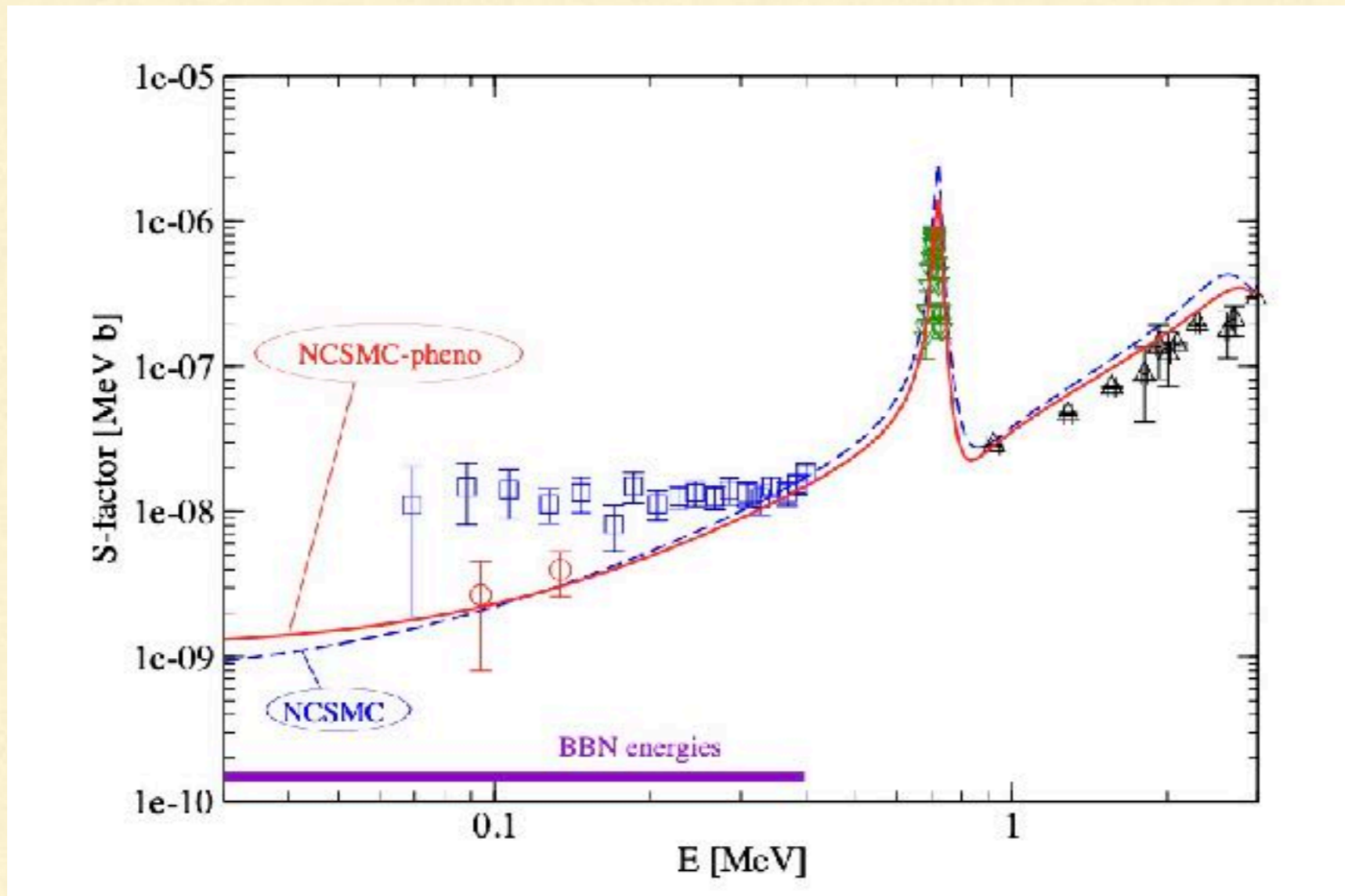
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Hebborn et al., PRL (2022)

- Convergence with 10 positive parity and 5 negative parity ${}^6\text{Li}$ states, and deuteron ground state + 8 pseudo states for continuum at $N_{\text{max}}=11$
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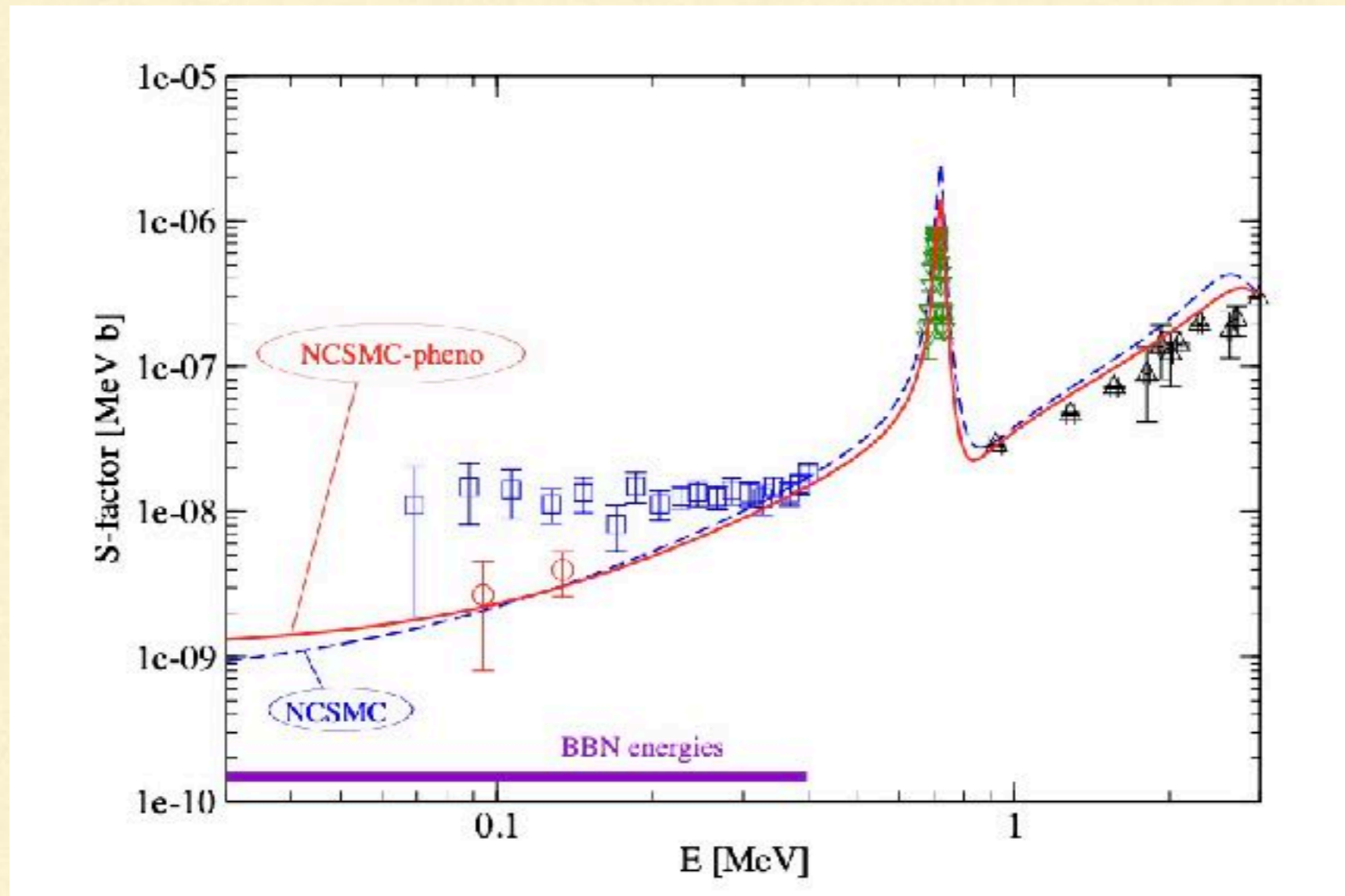
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Hebborn et al., PRL (2022)



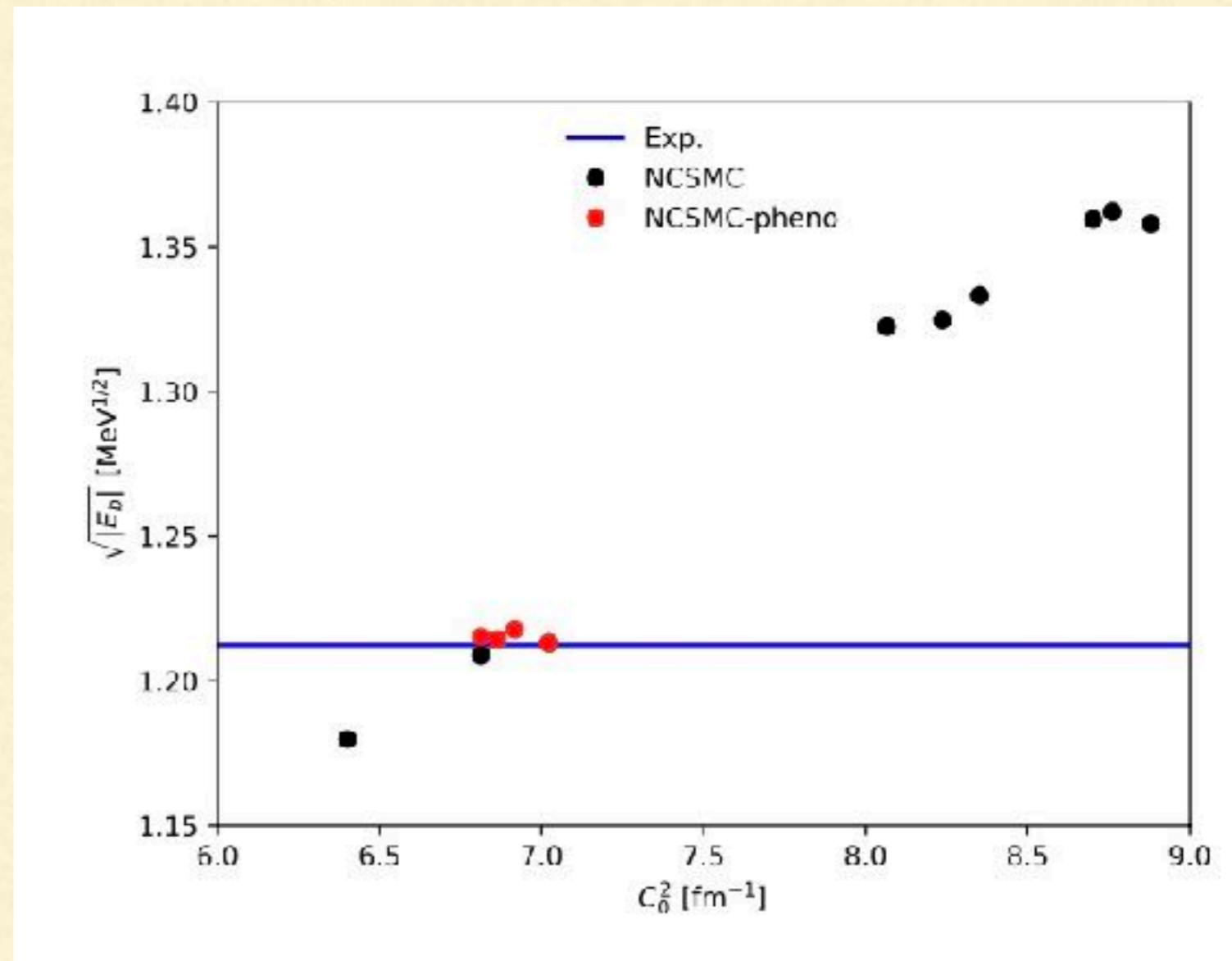
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ANC?

Is ANC of bound state (or width for resonance) stable against variations of force, etc., once deuteron separation energy, S_d , has been adjusted?

- Results indicate a one-parameter correlation between C_0^2 and the binding momentum $\sqrt{S_d}$

- $C_0^2 = \frac{2\gamma_0}{1 - \gamma_0 r_0}$ so C_0^2 should scale approximately linearly with $\sqrt{S_d}$



Is this correlation indicative of universality?

Fitting *ab initio* phase shifts to CMERE

Bethe (1949)
Sparenberg, Capel, Baye (2010)

$$\frac{2\pi\eta}{e^{2\pi\eta} - 1} k \cot(\delta) + 2k_C \operatorname{Re}[H(\eta)] \equiv K(k^2)$$

$$k_C = m_R Z_1 Z_2 \alpha_{\text{em}}$$

$$\eta = k_C / k$$

$$K(k^2) = -\frac{1}{a_0} + \frac{1}{2} r_0 k^2 + \frac{1}{4} P_0 k^4 + \frac{1}{8} Q_0 k^6 + \frac{1}{16} R_0 k^8 + \dots$$

- CMERE=Coulomb Modified Effective Range Expansion
 - $K(k^2)$ is real and analytic in k^2 within a radius of convergence defined by the first (non-Coulomb) analytic structure
 - Extrapolate Coulomb Modified Effective Range Theory amplitude to $k^2 < 0$ to find zero of inverse amplitude
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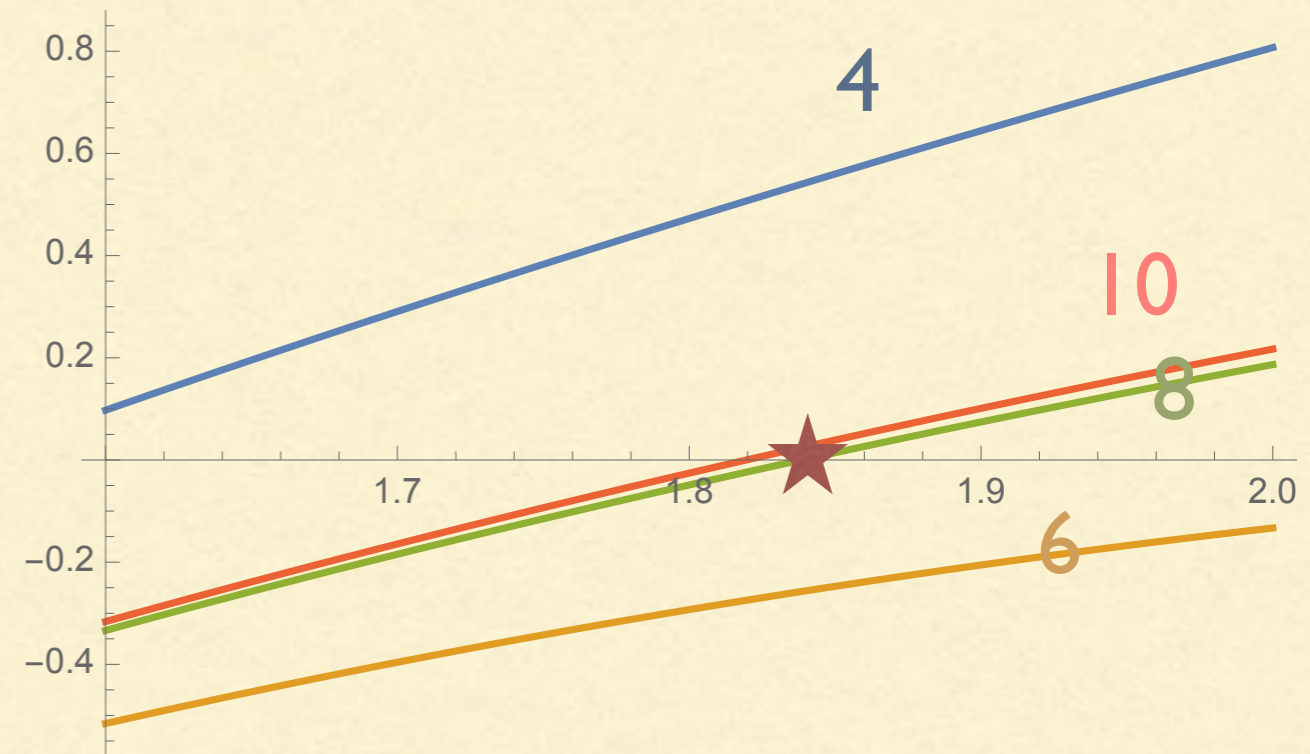
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Constrained CM-ERE

$$K(k^2) = 2k_C H(\eta(-S_d)) + \frac{1}{2}\rho_0(k^2 + \gamma_0^2) + \frac{1}{4}P_0(k^2 + \gamma_0^2)^2 + \dots \quad \gamma_0^2 = 2m_R S_d$$

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Fit to $E_{\max}=3$ MeV of NN-only phase shifts

- Then $C_0^2 = 6k_C \frac{\Gamma(1 + |\eta(S_d)|)^2}{\tilde{H}(-\eta(S_d)) - 3\rho_0 k_C}$
- $C_0^2=9.38$ fm⁻¹ at sixth order; 8.46 fm⁻¹ at eight order; 8.53 fm⁻¹ at tenth order.
- Not bad, but high order required
- Cf. $6k_C=0.56$ fm⁻¹ and $2\gamma_0=0.69$ fm⁻¹.
Fine tuned?
- Ab initio ANC=8.70 fm⁻¹

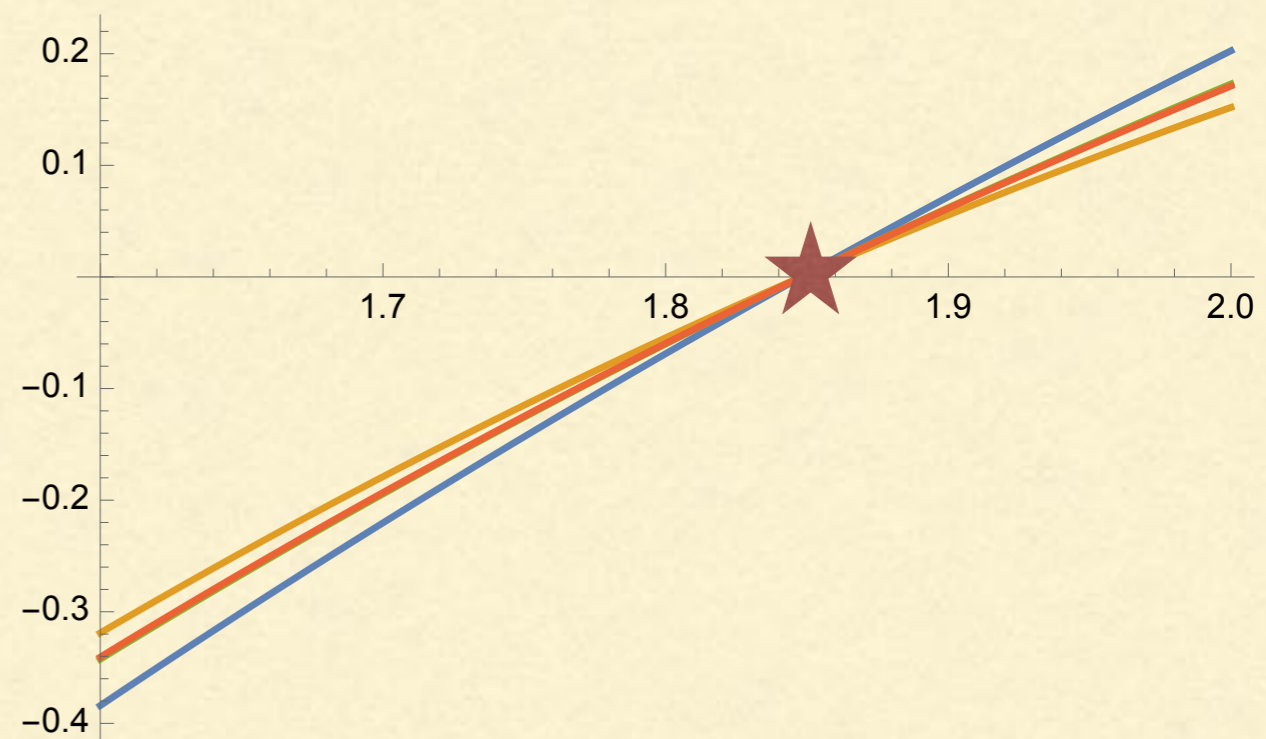
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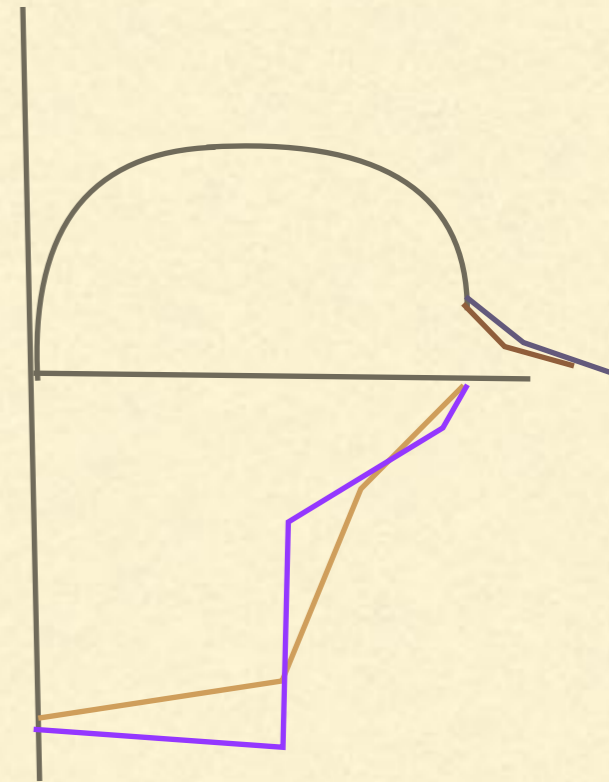
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Fit to $E_{\text{max}} = 3 \text{ MeV}$ of NN-only phase shifts



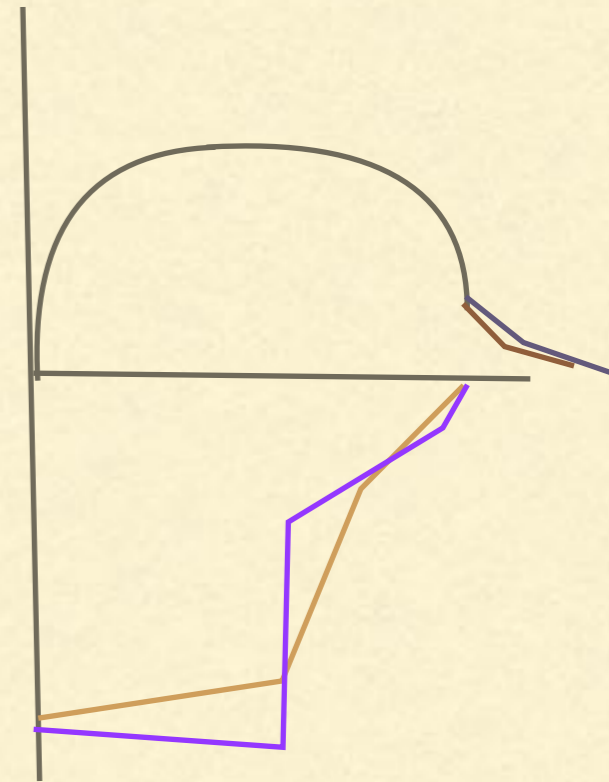
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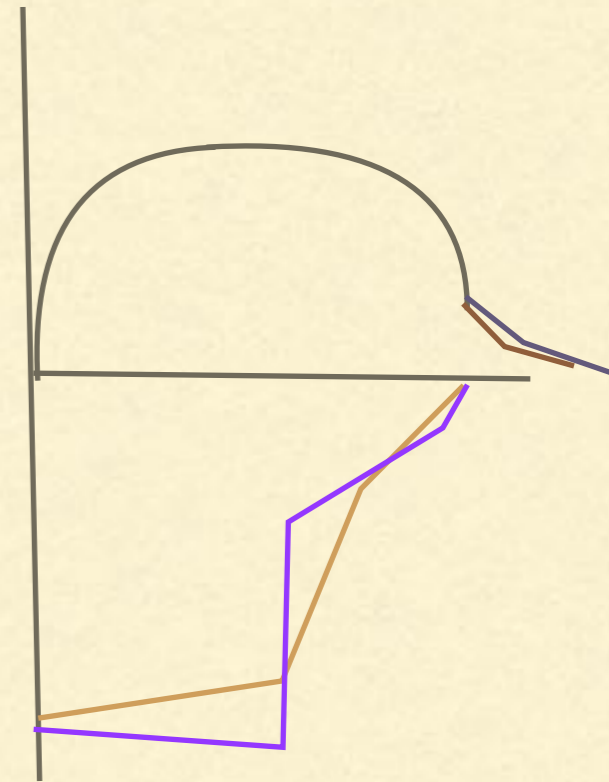
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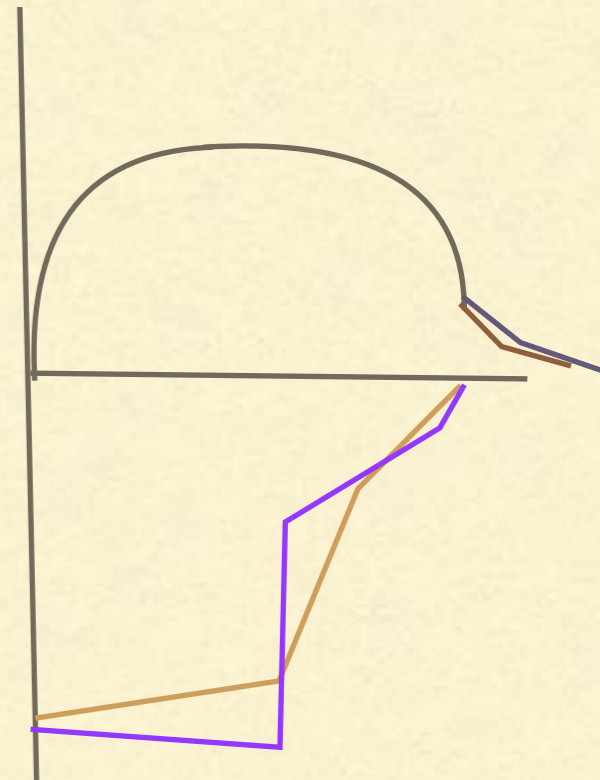
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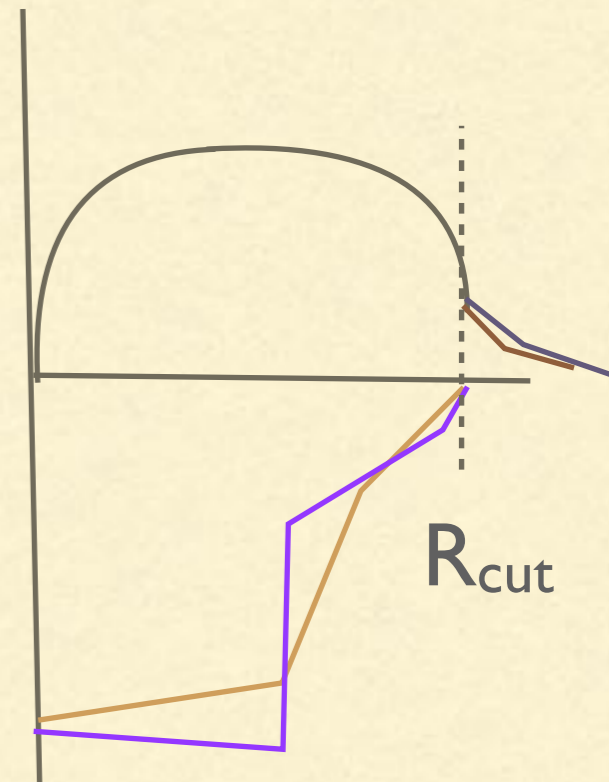
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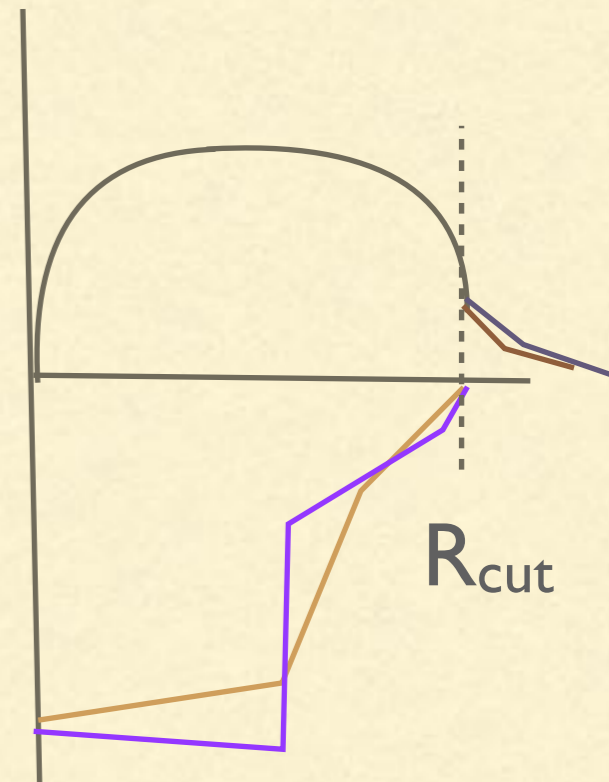
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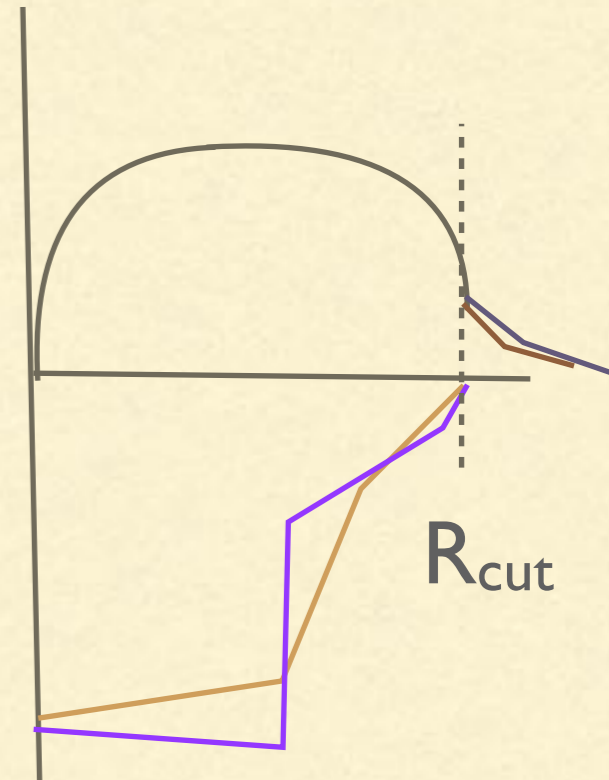
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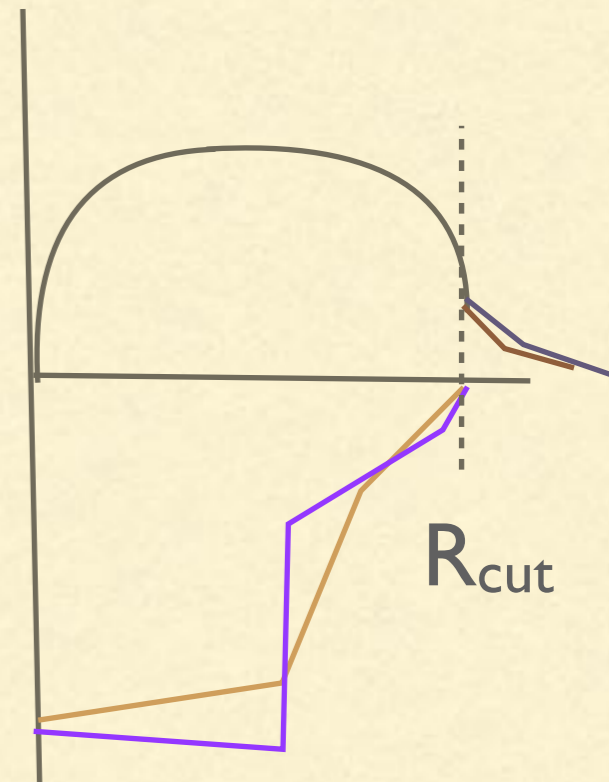
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- So if $f(x)$ becomes independent of x then $C_{0,j}^2$ will grow linearly with $\gamma_{0,j}$ in that region

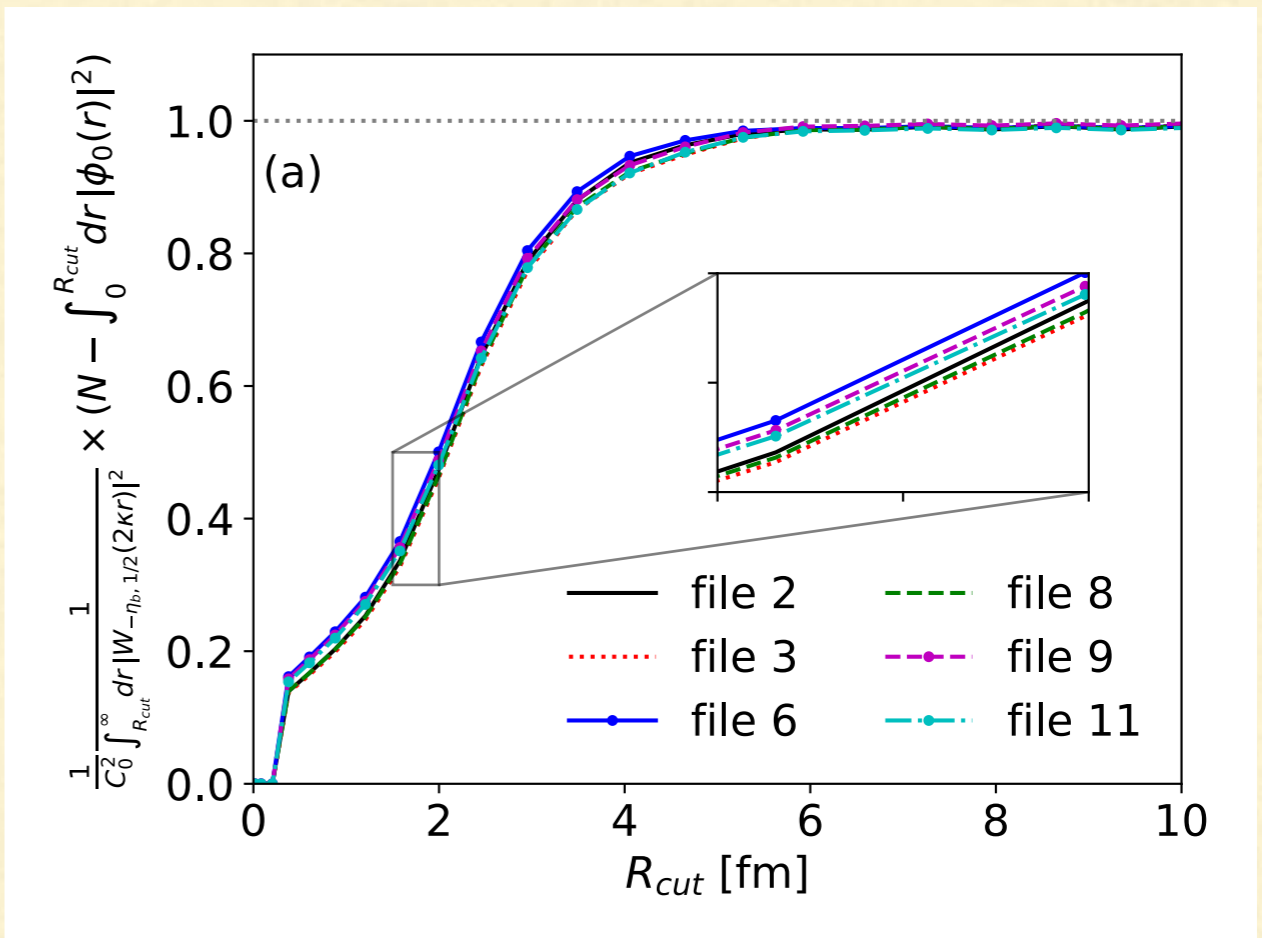


Getting on the scaling curve

- How far do we have to go in R_{cut} for $f(\gamma_0 R_{\text{cut}})$ to become R_{cut} independent?
- Calculate exterior probability via:

1. $N-I(R_{\text{cut}})$

2.
$$\int_{R_{\text{cut}}}^{\infty} dr C_0^2 W_{-\eta, 1/2}(2\gamma r)$$



- R_{cut} in asymptotic region once they're equal
- By this measure asymptotic wave function reached already at 5 fm
- Scaling region! $f(\gamma_0 R_{\text{cut}})$ independent of R_{cut}

Summary and outlook: part 2

- NCSMC calculations yield consistent scattering and bound-state results
 - Offer the possibility to compute ANCs and separation energies *ab initio*
 - But getting the proton (or deuteron or neutron or ...) separation energy of halos accurately is hard: NNLO ChiEFT only predicts it to ~a few hundred keV
 - NCSMC-pheno adjusts last few hundred keV of separation energy
 - So can be expected to also get ANC right if correlation is one parameter
 - ${}^6\text{Li}$ is a halo (fine tuned) nucleus: one-parameter ANC- S_d correlation expected from independence of short-distance part of wave function to fine tuning
 - Offers possibility to perform halo calculations in smaller model spaces or without three-body forces and then pheno adjust to reproduce separation energy and ANC
-