Self Interacting Dark Matter (SIDM)

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Why SIDM?

Aghanim+ Planck Collaboration, Planck 2018

- $\Omega_{DM}h^2 = 0.120 \pm 0.001$
- Small scale structure problems
 - Missing satelline
 - Core cusp problem
 - Too big too fail
- => SIDM (Spergel+ 2000)



What is SIDM?

The cold DM model is not able to explain observations on scales smaller than a few Mpc The key properties of the SIDM dark matter particles

- have a large scattering cross-section
- negligible annihilation or dissipation
- do not interact with ordinary particles in the standard model
- DM scater elastically with each other at astrophysically interesting rates

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$$\Gamma = n\sigma v = \rho \frac{\sigma}{m} v$$

• $\frac{\sigma}{m} \sim 1 cm^2 g^{-1}$

Model SIDM

with RH neutrinos provides two Higgs bosons: one is scalar or CP-even and another is pseudoscalar or CP odd particle having properties of candidates for dark matter (Lan+ 2006)

• Coupling of the interaction of DM to the standard Higgs boson

$$\begin{split} L(\sigma,\zeta_{\eta}) &= \left[\frac{\sigma(x)}{\sqrt{\lambda_{5}^{2}v^{2} + \lambda_{6}^{2}u^{2}}} \left(\lambda_{1}\lambda_{5}v^{2} + \frac{\lambda_{4}\lambda_{6}}{2}u^{2} \right) + \frac{H_{1}(x)\sigma(x)}{(\lambda_{5}^{2}v^{2} + \lambda_{6}^{2}u^{2})} \left(\lambda_{1} - \frac{\lambda_{4}}{2} \right) \lambda_{5}\lambda_{6}uv \right. \\ &+ \left. \frac{\sigma^{2}(x)}{2(\lambda_{5}^{2}v^{2} + \lambda_{6}^{2}u^{2})} \left(\lambda_{5}^{2}v^{2} + \frac{\lambda_{6}^{2}}{2}u^{2} \right) \right] \left(\xi_{\eta}^{'2} + \zeta_{\eta}^{'2} \right). \end{split}$$

• The cross section for hh - > hh (h stands for ζ'_{η} and ξ'_{η}) with quartic interaction is $\sigma = \frac{\lambda_1}{4\pi m_h^2}$

• DM elastic cross section to its mass is:

$$2.05 \times 10^3 \text{ GeV}^{-3} \le \frac{\sigma}{m_h} \le 2.57 \times 10^4 \text{ GeV}^{-3}$$

• SIDM mass. $4.7 MeV \leq m_h \leq 23 MeV$

$$\Omega_h = 2g(T_\gamma)T_\gamma^3 \frac{m_h\beta}{\rho_c g(T)}$$

- Cosmic density of the h scalar given by
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Constraint of the SIDM

- 1. SIDM and small scale structure problems
- 2. SIDM and tidal deformability NS
- 3. SIDM and GW gravitational lensing

SIDM and Small Scale Structure

• SIDM and core cusp problem: constrain with CMB (Lan+ 2021)



SIDM and Small Scale Structure

SIDM and core cusp problem: constrain with galaxy simulation (Lan+ 2021)

For simulations, we set the initial conditions (Komatsu+ 2011).

 $\Omega\Lambda=0.734,\,\Omega_{m}=0.266,\,\Omega_{b}=0.0449,\,n_{s}=0.963,\,h=0.71,\,\sigma_{8}=0.801$



SIDM and Gravitational Wave

• Effect of dark matter on Binary Neutron Merges (Nelson+ 2019)

• SIDM and tidal deformability NS (Lan+ 2022 prep.)

SIDM Experiment

• beryllium-8 nuclei for a new boson at the 10 MeV-scale that can be tested by nuclear and atomic spectroscopy experiments (Battaglieri + 2017)

THANK YOU!