

# On the detectability of nucleonic to quark matter transition from future gravitational wave detection using a null hypothesis

Chiranjib Mondal  
Caen-Meudon Virgo Group



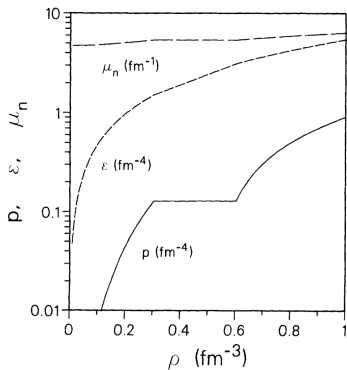
Neutron Rich Matter on  
Heaven and Earth,  
INT 22-2a  
July 13, 2022

# Nuclenic to quark matter

Equations of state with Phase transition

# Nuclenic to quark matter

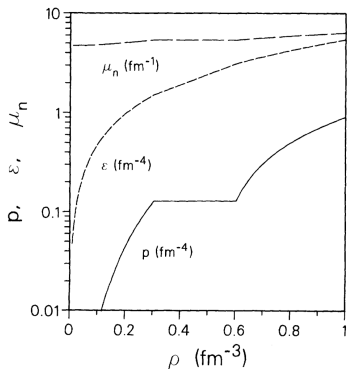
Equations of state with Phase transition



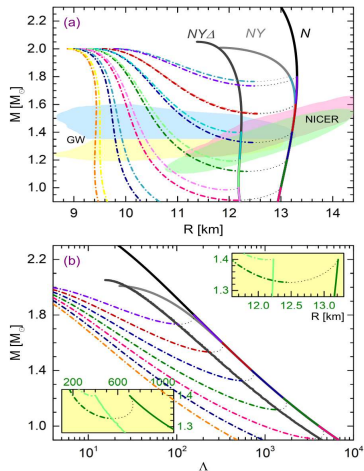
Glendenning PRD 46, 1274 (1992)

# Nuclenic to quark matter

Equations of state with Phase transition



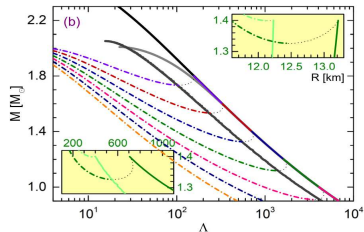
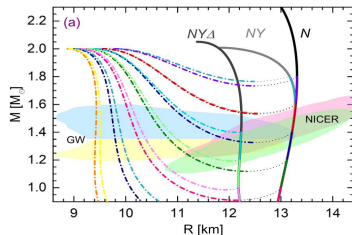
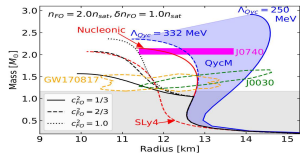
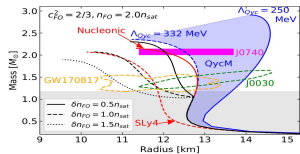
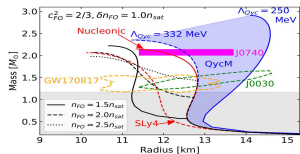
Glendenning PRD 46, 1274 (1992)



Li, Sedrakian and Alford, PRD 101, 063022 (2020)

# Nuclenic to quark matter

Equations of state with Phase transition



Somasundaram and Margueron,  
2104.13612 (2021)

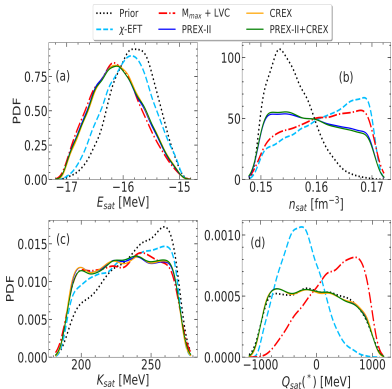
Li, Sedrakian and Alford, PRD 101, 063022  
(2020)

# The Null hypothesis

Nucleonic metamodel

# The Null hypothesis

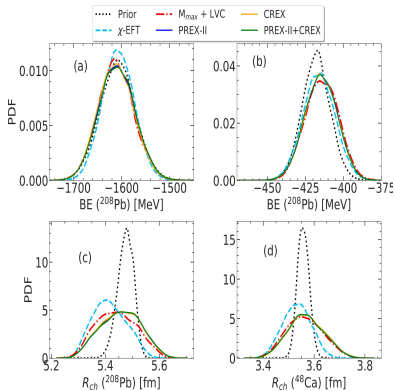
Nucleonic metamodel



Nuclear Matter Properties

# The Null hypothesis

Nucleonic metamodel

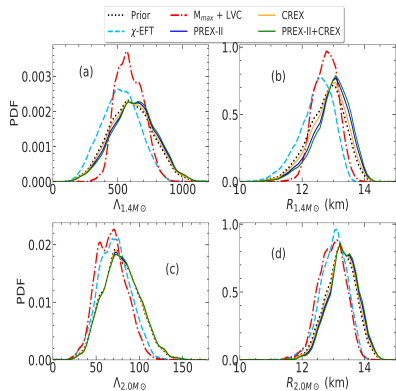
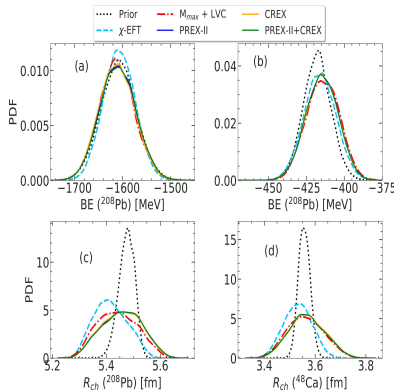


Finite Nuclei



# The Null hypothesis

Nucleonic metamodel

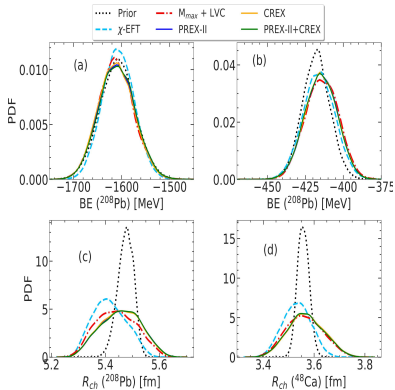


Properties of Neutron Star

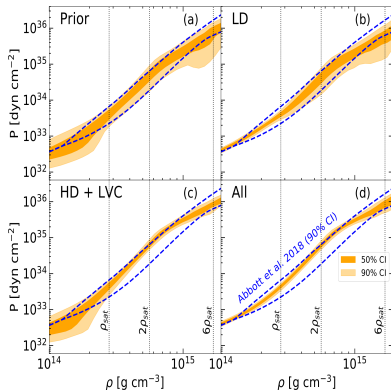
Finite Nuclei

# The Null hypothesis

Nucleonic metamodel



Finite Nuclei



EoS

H. Dinh Thi 2021.

# Looking for phase transition signal

Post-o5 LIGO-Virgo

$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}} \tilde{\Lambda}^{1/5}.$$

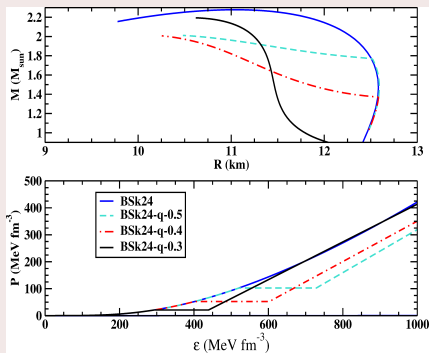
# Looking for phase transition signal

Post-o5 LIGO-Virgo

$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}} \tilde{\Lambda}^{1/5}.$$

## M-R & EoS



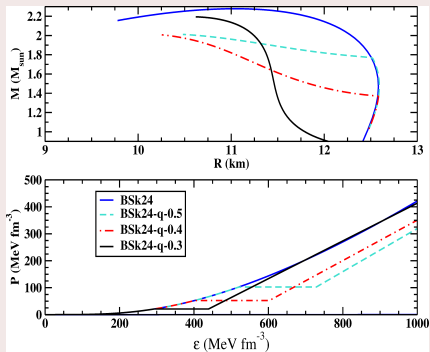
# Looking for phase transition signal

Post-o5 LIGO-Virgo

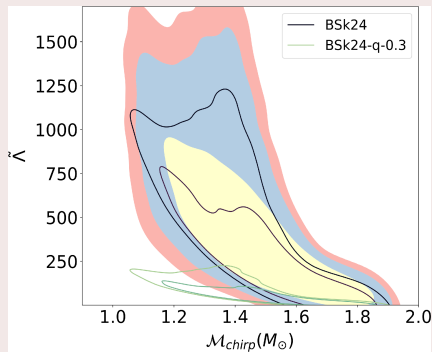
$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}} \tilde{\Lambda}^{1/5}.$$

## M-R & EoS



## Distribution of tidal signal from GW



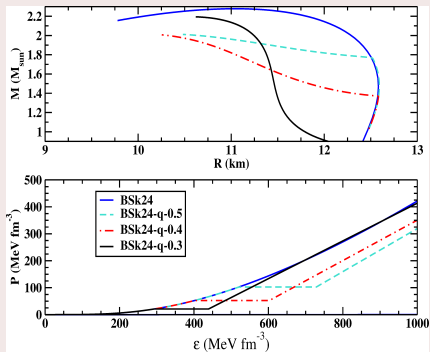
# Looking for phase transition signal

Post-o5 LIGO-Virgo

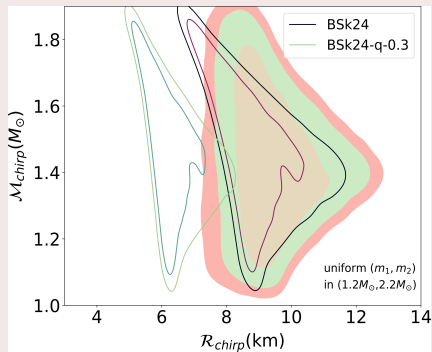
$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}} \tilde{\Lambda}^{1/5}.$$

## M-R & EoS



## Distribution of tidal signal from GW



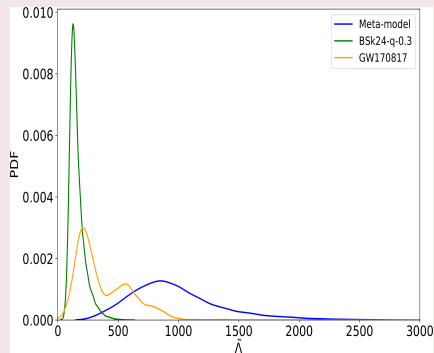
# Looking for phase transition signal

Post-o5 LIGO-Virgo

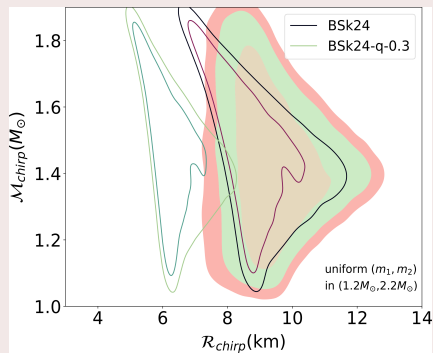
$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}}\tilde{\Lambda}^{1/5}.$$

## M-R & EoS



## Distribution of tidal signal from GW



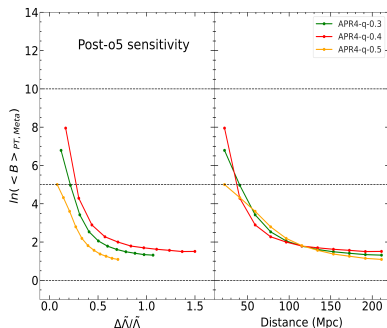
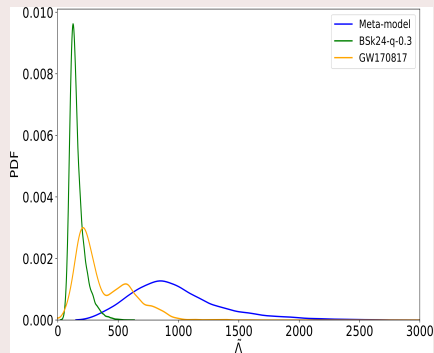
# Looking for phase transition signal

Post-o5 LIGO-Virgo

$$\tilde{\Lambda} = \frac{16}{13} \frac{(12q + 1)\Lambda_1 + (12 + q)q^4\Lambda_2}{(1 + q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5} m_1 / (1 + q)^{1/5}; \quad \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}} \tilde{\Lambda}^{1/5}.$$

## M-R & EoS





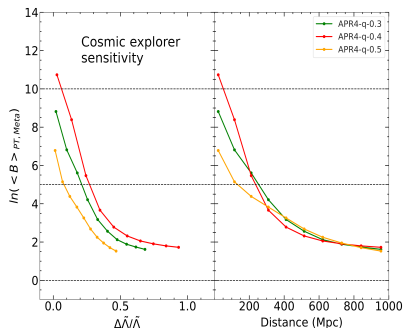
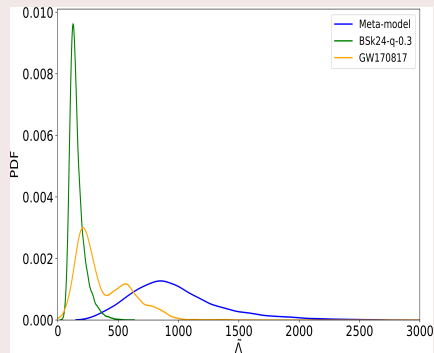
# Looking for phase transition signal

Post-o5 LIGO-Virgo

$$\tilde{\Lambda} = \frac{16(12q+1)\Lambda_1 + (12+q)q^4\Lambda_2}{(1+q)^5}.$$

$$\mathcal{M}_{\text{chirp}} = q^{3/5}m_1/(1+q)^{1/5}; \mathcal{R}_{\text{chirp}} = 2\mathcal{M}_{\text{chirp}}\tilde{\Lambda}^{1/5}.$$

## M-R & EoS



## Caen-Meudon Virgo group

- Hoa Dinh-Thi
- Gaël Servignat
- Lami Suleiman
- Marco Antonelli
- Anthea Fantina
- Philip Davis
- Francesca Gulminelli
- Micaela Oertel
- Jérôme Novak