

Probing BSM particles using inelastic nuclear scattering

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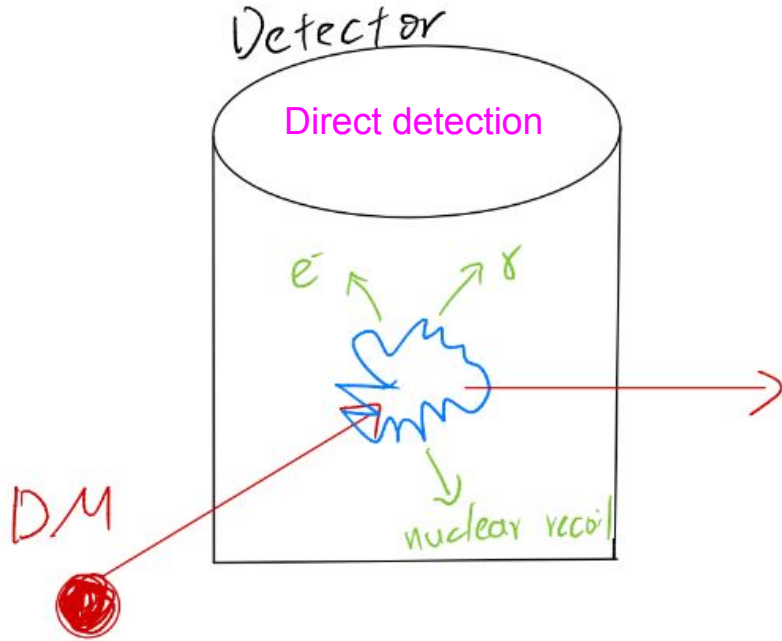
In collaboration with Bhaskar Dutta, Jayden Newstead, and Vishvas Pandey

Outline

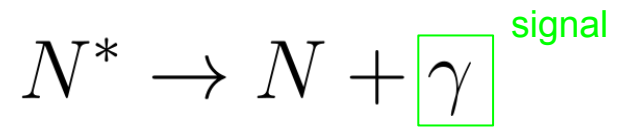
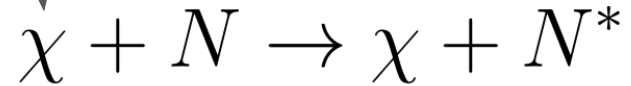
1. Overview
2. Lab produced dark matter
3. Ambient dark matter (WIMP, boosted DM)
4. Axion
5. Conclusion

B. Dutta, [W. Huang](#), J. L. Newstead 2302.10250; to appear
B. Dutta, [W. Huang](#), J. L. Newstead and V. Pandey, PRD 106 (2022) 113006;
B. Dutta, [W. Huang](#), D. Kim, J. Newstead, J. Pouk, to appear

Overview



Any particle to search



For each of DM, WIMP, and axion

- Theories and models
- Benchmark experiments
- Results

Light dark matter in stopped pion experiment

Pros

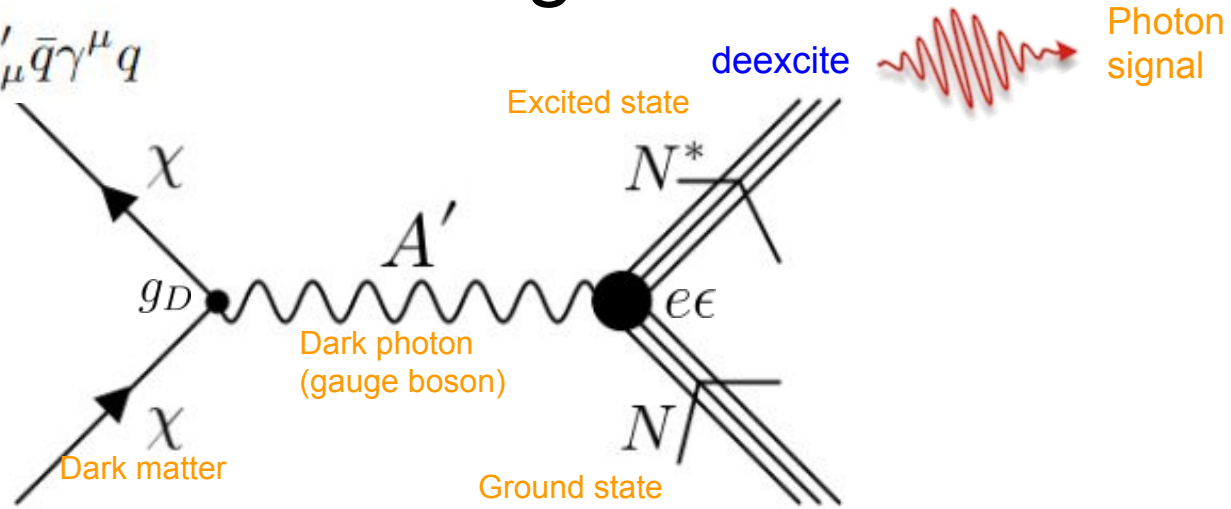
- Lower background
- Little to none threshold effect

Cons

- Lower signal
- More calculations

Inelastic DM-Nucleus Scattering

$$\mathcal{L} \supset g_D A'_\mu \bar{\chi} \gamma^\mu \chi + e\epsilon Q_q A'_\mu \bar{q} \gamma^\mu q$$



Q: Why not elastic?

A: Elastic has high signal rates but also high background (keV nuclear recoil)


$$\frac{d\sigma_{\text{el}}^{DM}}{dE_r} = \frac{e^2 \epsilon^2 g_D^2 Z^2}{4\pi (E_\chi^2 - m_\chi^2) (2m_N E_r + m_{A'}^2)^2} F^2(E_r) \times \left[2E_\chi^2 m_N \left(1 - \frac{E_r}{E_\chi} - \frac{m_N E_r}{2E_\chi^2} \right) + E_r^2 m_N \right]$$

Multipole expansion

J. D. Walecka, 2004;
W. Haxton, et al., 0706.2210

In Standard Model... $\langle f | \hat{H}_W | i \rangle = \frac{G_F}{\sqrt{2}} \int d^3x \langle f | j_\mu^{lep} \hat{J}^\mu(\vec{x}) | i \rangle$

$$= \frac{G_F}{\sqrt{2}} \int d^3x e^{-i\vec{q}\cdot\vec{x}} \left(l_0 \mathcal{J}^0(\vec{x}) - \vec{l} \cdot \mathcal{J}(\vec{x}) \right)$$


 Spherical decomposition

Multipole operator $\hat{M}_{JM}(q) \equiv \int d^3x [j_J(qx) Y_{JM}(\Omega_x)] \hat{J}^0(\vec{x})$

Nuclear response function $M_{JM}(q\vec{x}_i) \equiv j_J(qx_i) Y_{JM}(\Omega_{x_i})$

where Y_{JM} is Bessel spherical harmonics

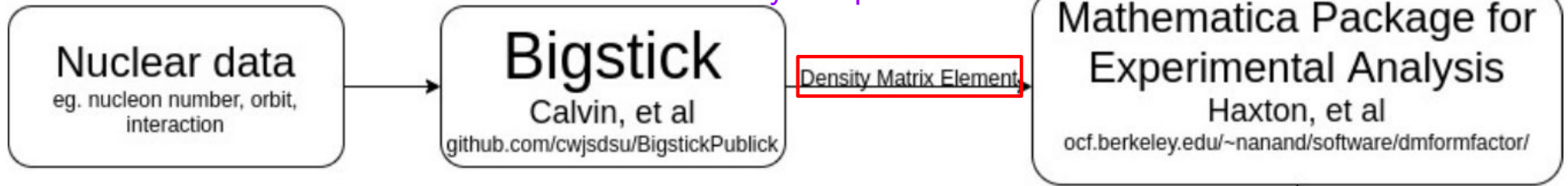
$$\frac{d\sigma_{\text{inel}}^{DM}}{dE_r} = \frac{2e^2 \epsilon^2 g_D^2 E_\chi'^2}{p_\chi p'_\chi (2m_N E_r + m_{A'}^2)^2} \frac{m_N}{2\pi} \frac{4\pi}{2J+1} \left\{ \sum_{J \geq 1, \text{spin}} \left[\frac{1}{2} (\vec{l} \cdot \vec{l}^* - l_3 l_3^*) \left(|\langle J_f | \hat{T}_J^{\text{mag}} | J_i \rangle|^2 + |\langle J_f | \hat{T}_J^{\text{el}} | J_i \rangle|^2 \right) \right] \right.$$

$l_\mu = \bar{\chi} \gamma^\mu \chi.$

$$\left. + \sum_{J \geq 0, \text{spin}} \left[l_0 l_0^* |\langle J_f | \hat{M}_J | J_i \rangle|^2 + l_3 l_3^* |\langle J_f | \hat{L}_J | J_i \rangle|^2 - 2 l_3 l_0^* \text{Re} \left(\langle J_f | \hat{L}_J | J_i \rangle \langle J_f | \hat{M}_J | J_i \rangle^* \right) \right] \right\}$$

Nuclear Shell Model Code: BIGSTICK

Heavy computation!



Nuclear Response Function

$$W_M^{\tau\tau'}(y) = \sum_{J=0,2,\dots}^{\infty} \langle j_N || M_{J;\tau}(q) || j_N \rangle \langle j_N || M_{J;\tau'}(q) || j_N \rangle$$
$$W_{\Phi''}^{\tau\tau'}(y) = \sum_{J=0,2,\dots}^{\infty} \langle j_N || \Phi''_{J;\tau}(q) || j_N \rangle \langle j_N || \Phi''_{J;\tau'}(q) || j_N \rangle$$
$$W_{\Phi''M}^{\tau\tau'}(y) = \sum_{J=0,2,\dots}^{\infty} \langle j_N || \Phi''_{J;\tau}(q) || j_N \rangle \langle j_N || M_{J;\tau'}(q) || j_N \rangle$$

Multipole Operator

$$\mathcal{M}_{LM} = F_1^N M_L^M + \frac{q^2}{4m_N^2} (F_1^N + 2F_2^N) \left(\Phi_L''^M - \frac{1}{2} M_{LM} \right)$$

It takes a long time and needs tons of RAM and CPU

Is there a shortcut?

Long Wavelength Limit

$$\hat{\mathcal{M}}_{JM}(q) = \frac{q^J}{(2J+1)!!} \int d^3x x^J Y_{JM} \hat{\mathcal{J}}_0(\mathbf{x})$$

$$\hat{\mathcal{L}}_{JM}(q) = \frac{-iq^{J-1}}{(2J+1)!!} \int d^3x x^J Y_{JM} \nabla \cdot \hat{\mathcal{J}}(\mathbf{x})$$

$$\hat{\mathcal{T}}_{JM}^{el}(q) = -i \frac{q^{J-1}}{(2J+1)!!} \left(\frac{J+1}{J} \right)^{1/2} \int d^3x x^J Y_{JM} \nabla \cdot \hat{\mathcal{J}}(\mathbf{x})$$

$$\hat{\mathcal{T}}_{JM}^{mag}(q) = i \frac{q^J}{(2J+1)!!} \left(\frac{J+1}{J} \right)^{1/2} \int d^3x \left[\frac{1}{J+1} \mathbf{r} \times \hat{\mathcal{J}}(\mathbf{x}) \right] \cdot \nabla x^J Y_{JM}$$

The surviving multipoles are $\hat{\mathcal{M}}_{00}, \hat{\mathcal{L}}_{1M}, \hat{\mathcal{T}}_{1M}^{el}$.

$$\hat{\mathcal{M}}_{00} = \frac{1}{\sqrt{4\pi}} F_1 \sum_{i=1}^A \hat{F}T (\sim \hat{\tau})$$

Fermi doesn't contribute to inelastic scattering because it's even-even operator. It only exists in elastic scattering

$$\hat{\mathcal{T}}_{1M}^{el} = \sqrt{2} \hat{\mathcal{L}}_{1M} = \frac{i}{\sqrt{6\pi}} G_A \sum_{i=1}^A \hat{G}T (\sim \hat{\sigma} \hat{\tau})$$

Is GT (Gamow Teller) the shortcut?

Strength and Multipole in BIGSTICK

BIGSTICK can calculate the strength of a given operator

$$\left| (\Psi_f : J_f || \hat{O}_J || \Psi_i J_i) \right|^2$$

	Strength	Multipole
Time	Short	Long
RAM & CPU	Light	Heavy
Output	Less detailed The strength and energy	Comprehensive Density matrix -> strength

Multipole has energy, spin, isospin, and density matrix

Strength only has energy, strength

GT Strength Lines

eg. Ar40

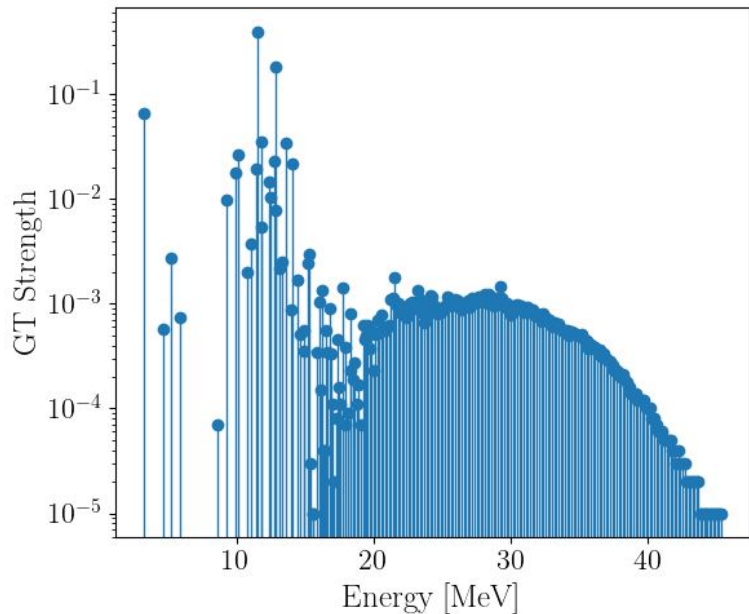
$$\Delta J = |J_f - J_i| = 1$$

$$\text{GT strength} = \left| \langle J_f \left\| \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 \right\| J_i \rangle \right|^2$$

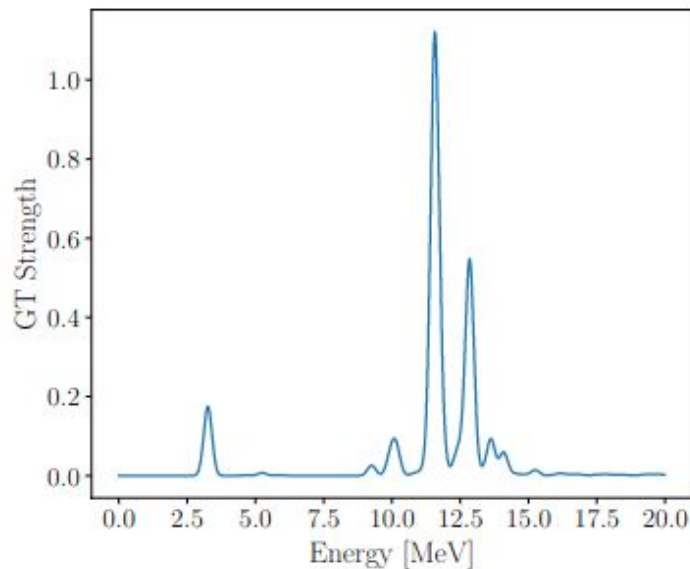
excited state

ground state

one line = one transition = one excited state



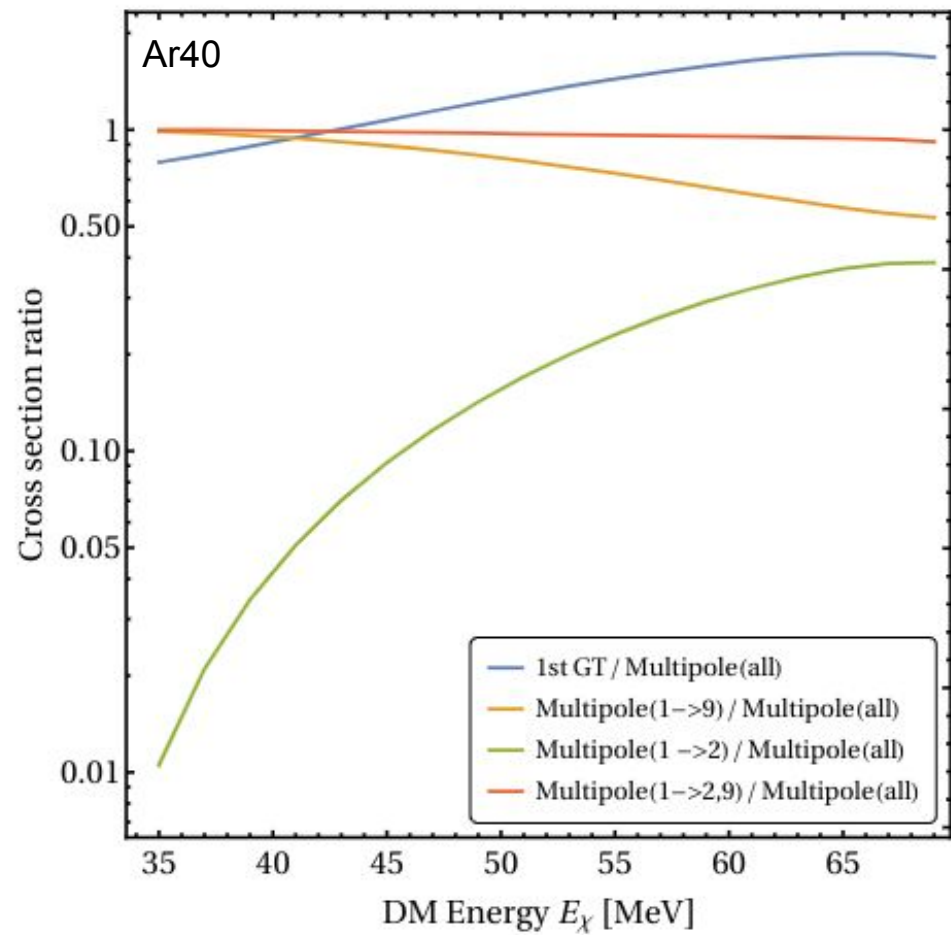
convoluted with 150 keV width Gaussian



Experiment data

E^*	J^π
[keV]	
0.0 ^a	0 ⁺
1460.851(6) ^a	2 ⁺
2120.8(3) ^b	0 ⁺
2524.1(2) ^b	2 ⁺
2892.6(1) ^a	4 ⁺
3208.0(6)	2 ⁺
3464.5(1) ^a	6 ⁺
3511.3(5)	2 ⁺
3515(1) ^b	4 ⁺
3680.8(2)	3 ⁻
3918.8(2)	2 ⁺
3941.7(3)*	
4041(1)	N
4082.5(2)	3 ⁻
4178.9(3)*	
4226(1)	4 ⁻
4229(1)	→ 1 ⁺ , 2 ⁻ , 3 ⁺
4300.8(3)	→ (1,3) ⁻
4324.5(3)	→ 2 ⁺
4358.0(3)	→ N
4420(1)	→ (0 ⁺ -4 ⁺)
4427(1)	→ (4 ⁺)
4473(1)	→ 1
4481.0(3)	→ 1 ⁻
4494(1) ^c	→ 5 ⁻
4562.3(2)	→ (1,3) ⁻
4578(1)	→ (2 ⁺ , 3 ⁻)
4602(1)	→ (0 ⁺ -4 ⁺)
4674(1)	→ 1 ⁺ , 2 ⁻ , 3 ⁺
4737.8(4)*	
4769.0(3)	→ 1 ⁻

Cross Section in Long Wavelength Limit



$E_r \rightarrow 0$

$$\frac{d\sigma_{inel}^{DM}}{d\cos\theta} = \frac{2e^2\epsilon^2 g_D^2 E_\chi'^2}{(2m_N E_r + m_{A'}^2)^2} \frac{1}{2\pi} \frac{4\pi}{2J+1}$$

$$\times \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} |\langle J_f | \left[\sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_{i0} \right] | J_i \rangle|^2$$

Gamow-Teller (GT) operator

GT strength is the shortcut

Er in keV, DM in MeV or above
 => GT is a good approximation

DM-Nucleus Scattering: Elastic vs Inelastic

$$\frac{d\sigma_{\text{el}}^{DM}}{dE_r} = \frac{e^2 \epsilon^2 g_D^2 Z^2}{4\pi(E_\chi^2 - m_\chi^2)(2m_N E_r + m_{A'}^2)^2} F^2(E_r)$$

$$\times \left[2E_\chi^2 m_N \left(1 - \frac{E_r}{E_\chi} - \frac{m_N E_r}{2E_\chi^2} \right) + E_r^2 m_N \right]$$

$$\frac{d\sigma_{\text{inel}}^{DM}}{d\cos\theta} = \frac{2e^2 \epsilon^2 g_D^2 E_\chi'^2}{(2m_N E_r + m_{A'}^2)^2} \frac{1}{2\pi} \frac{4\pi}{2J+1}$$

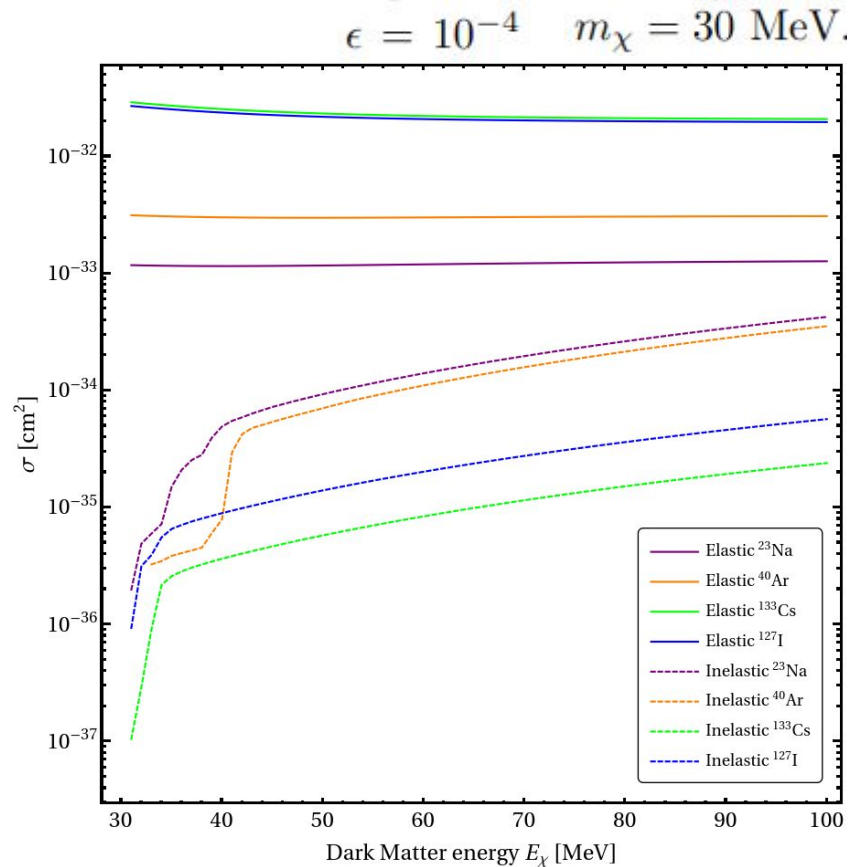
$$\times \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2$$

For Ar40

$$\left(\frac{\text{Inelastic}}{\text{Elastic}} \right)_{\text{signal}} = 10^{-2} - 10^{-1}$$

$$\left(\frac{\text{Inelastic}}{\text{Elastic}} \right)_{\text{bkg}} = 10^{-4} - 10^{-3}$$

Inelastic search can be better



B. Dutta, [W. Huang](#), J. L. Newstead and V. Pandey, PRD 106 (2022) 113006

Inelastic DM-Nucleus Scattering

Cross section has the same form for fermion and scalar DM, only difference is the **current** $l_\mu = \bar{\chi}\gamma^\mu\chi$.

$$\frac{d\sigma_{inel}^{DM}}{d\cos\theta} = \frac{2e^2\epsilon^2 g_D^2 E_\chi'^2}{(2m_N E_r + m_{A'}^2)^2} \frac{1}{2\pi} \frac{4\pi}{2J+1} \times \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2$$

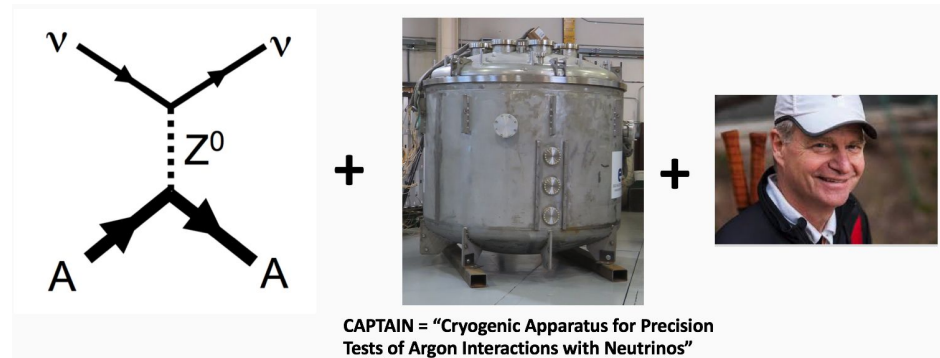
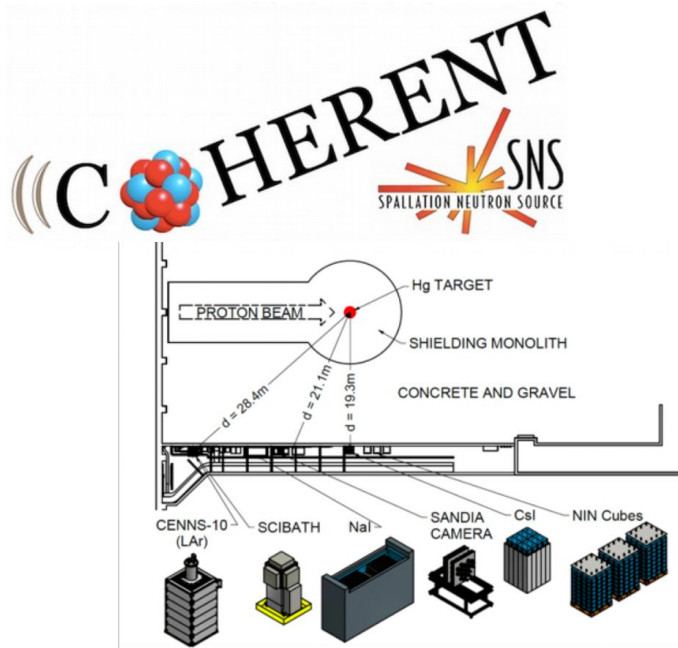
Fermion $\mathcal{L}_f \supset g_D A'_\mu \bar{\chi} \gamma^\mu \chi + e\epsilon Q_q A'_\mu \bar{q} \gamma^\mu q$

$$\sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* = 3 - \frac{1}{E_\chi E_\chi'} \left[\frac{1}{2} \left(p_\chi^2 + p_\chi'^2 - 2m_N E_r \right) + \frac{3m_\chi^2}{4} \right]$$

Scalar $\mathcal{L}_s \supset |D_\mu \phi|^2 + e\epsilon Q_q A'_\mu \bar{q} \gamma^\mu q$

$$\vec{l} \cdot \vec{l}^* = \frac{1}{2E_\chi E_\chi'} \left(p_\chi^2 + p_\chi'^2 - 2m_N E_r \right)$$

Experiments and Detectors

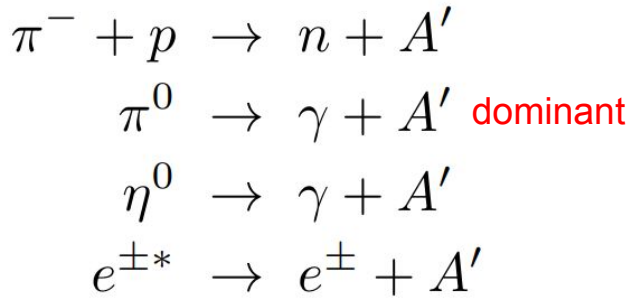


Why NaI and CCM?
They have large detector mass

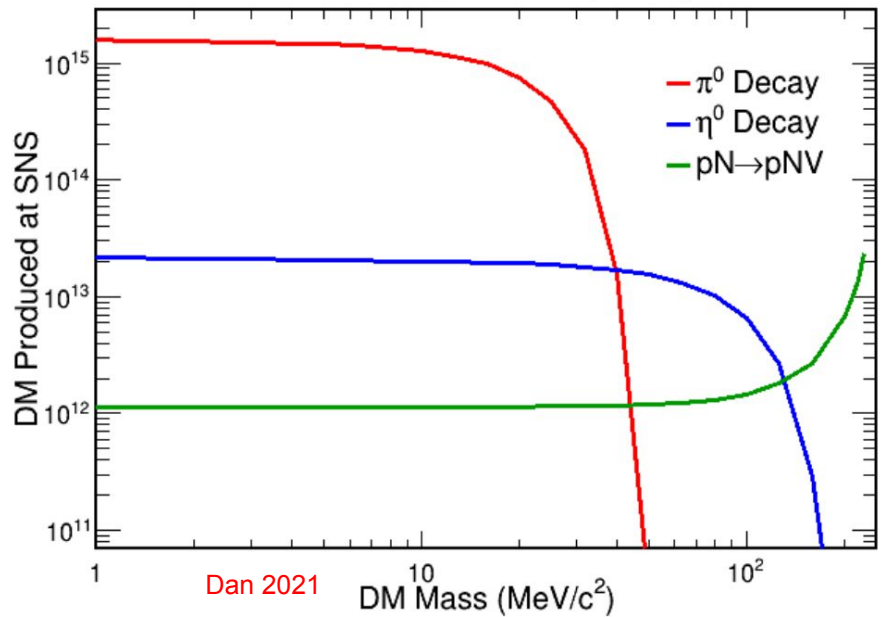
Experiment	E_{beam} [GeV]	POT [yr ⁻¹]	Target	Detector:				
				target	mass	distance	angle	E_r^{th}
COHERENT	1	1.5×10^{23}	Hg	CsI[Na]	14.6 kg	19.3 m	90°	6.5 keV
				NaI[Tl]	185 kg	22 m	120°	900 keV
				NaI[Tl]	3500 kg	22 m	120°	~few keV
CCM	0.8	1.0×10^{22}	W	Ar	7 t	20 m	90°	25 keV

DM Flux in CCM

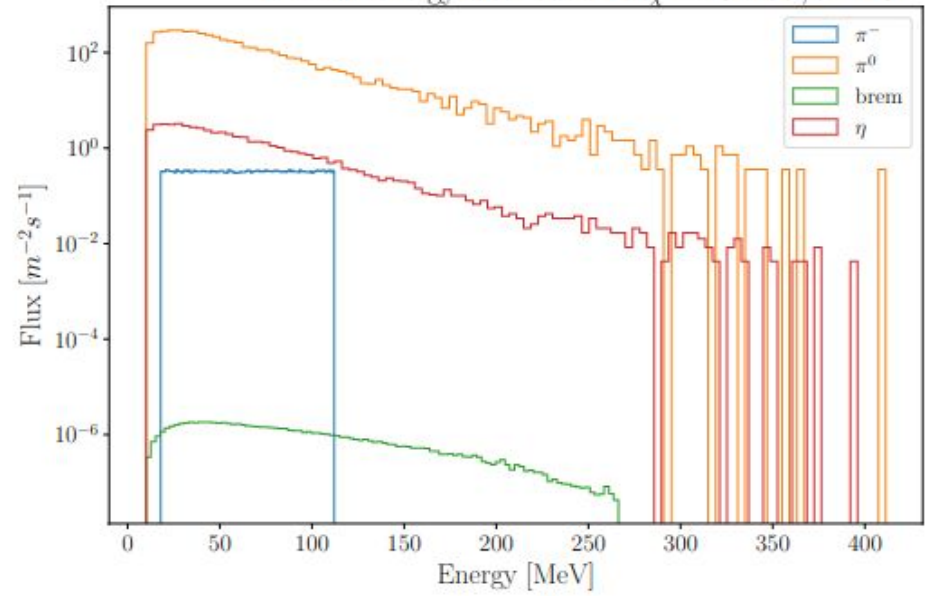
$A' \rightarrow \chi\bar{\chi}$ Decay in $< O(-10)$ ns



$\epsilon = 10^{-4} / \alpha_D = 0.5 / m_V = 3m_\chi$



DM flux energy distribution $m_\chi = 10\text{MeV}, \epsilon = 10^{-4}$



Prompt Window

Signal = inelastic DM from pion and eta decay

Bkg = inelastic prompt neutrino

Remove ν_e

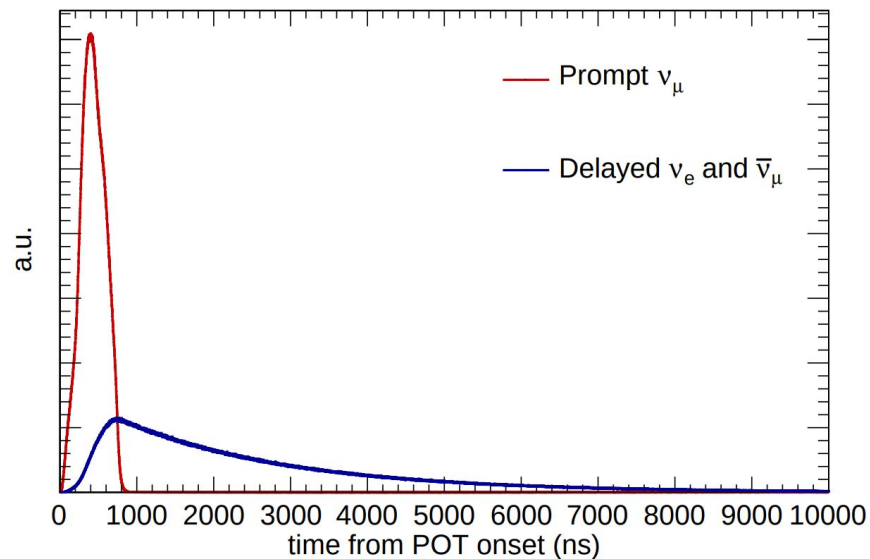
No bkg due to charge current

$$\tau_\eta \sim 10^{-10} \text{ ns}$$

$$\tau_{\pi^0} \sim 10^{-7} \text{ ns}$$

Inelastic nu background

Detector	CCM (LAr)	COHERENT (NaI)
Delayed + prompt	327	462
Prompt only ($t < 1\text{ms}$)	64.8	106
220ns cut (to 5%)	3.24	5.3



220ns = 150ns (beam arrival) + 70ns (relativistic travel for 20 meter)

Background

Detector	Bkg estimation*
COHERENT NaI	$\sim O(100)$
CCM	$\sim O(100)$

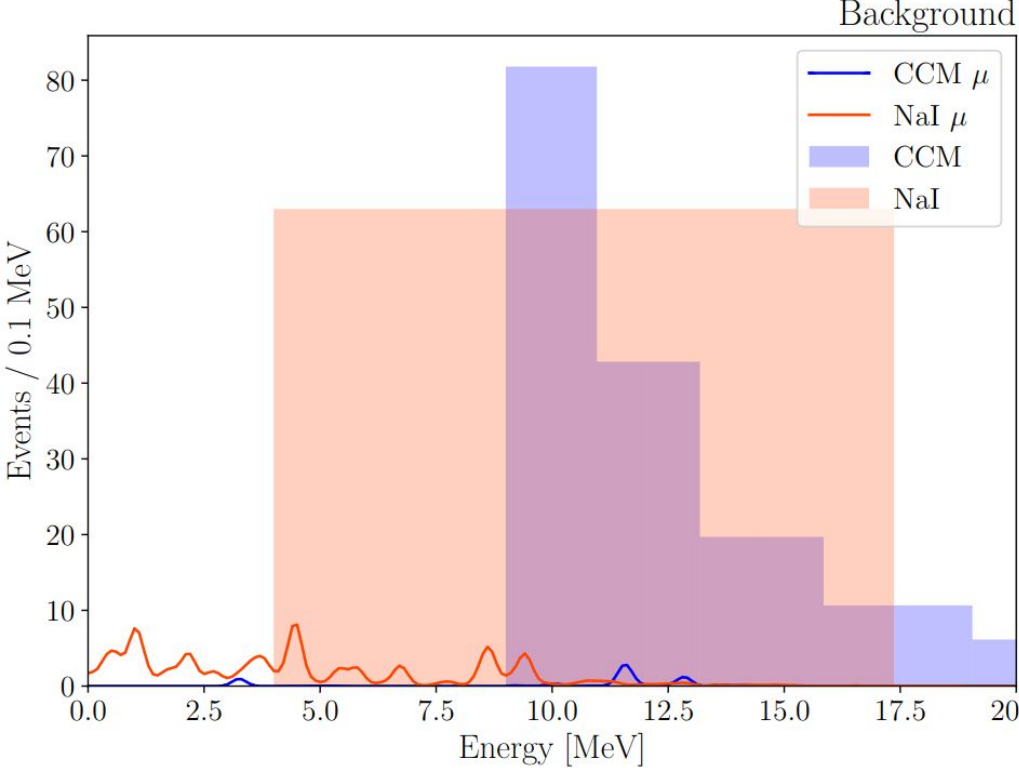
*1/100 Background reduction applied

NaI (COHERENT) is ongoing, we assumed similar background as CCM

CCM Collaboration, A. A. Aguilar-Arevalo et al., 2022

Rescale the GT strength (0.162) to be consistent with the experiment

W. Tornow et al., 2210.14316



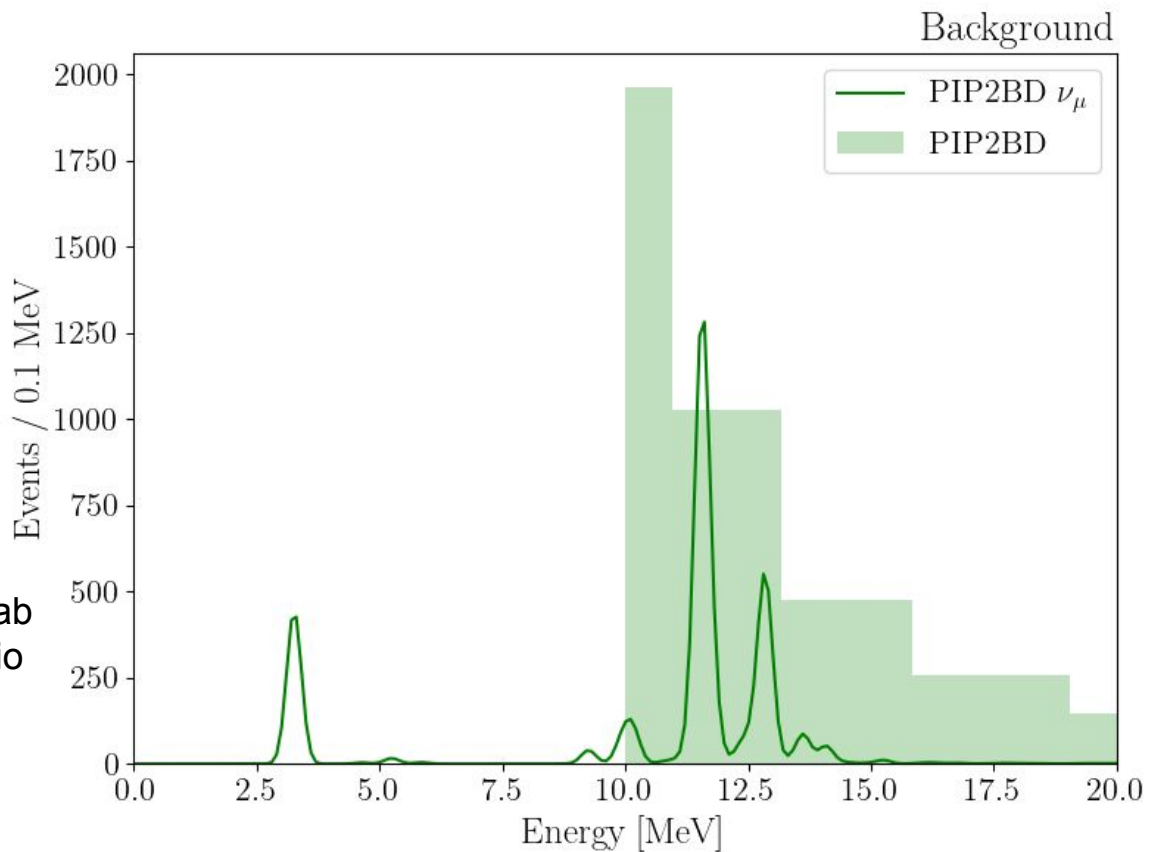
Baishan's calculation on many (~30) multipole states shows that GT indeed works

Baishan Hu, to be appear

Background

Detector	Bkg estimation
PIP2BD	~700

~GeV beam
100t LAr
4.95e23 POT total
Proposed beam dump experiment at Fermilab
Bkg is scaled from CCM by relative POT ratio



Sensitivities Plot

Dashed is our calculation

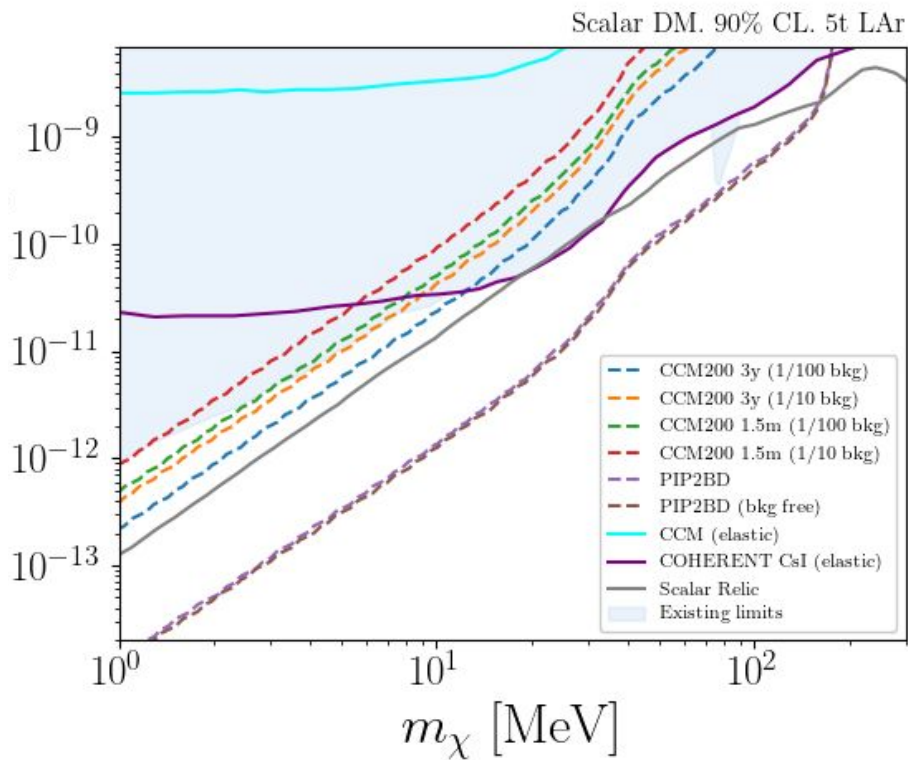
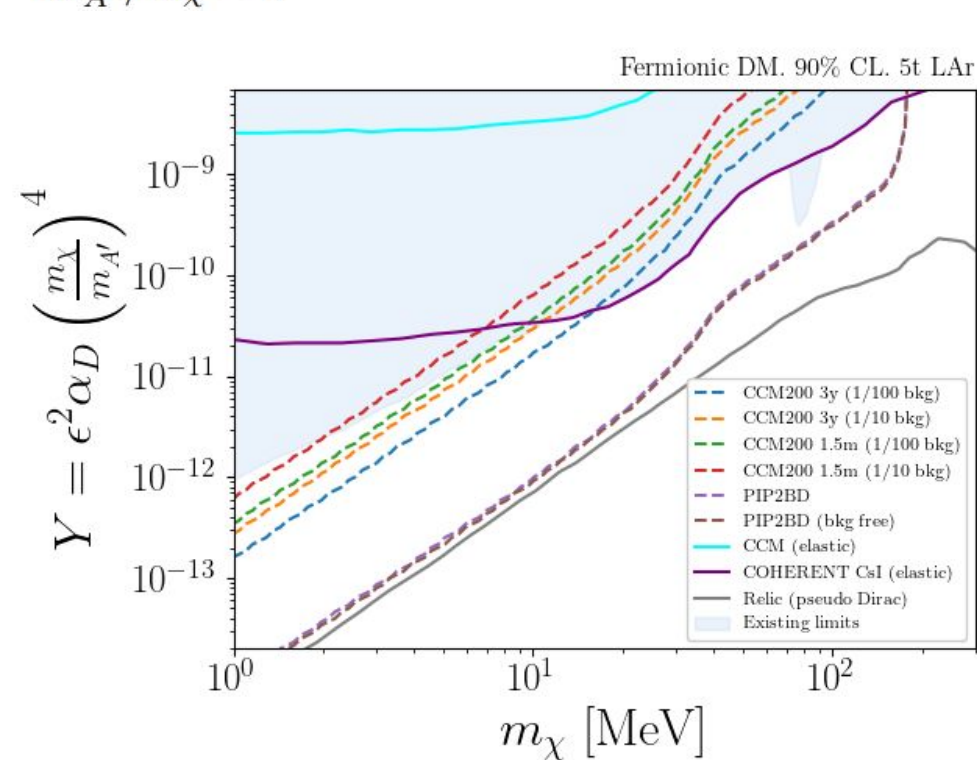
$$g_D = \sqrt{2\pi}$$

$$m_{A'}/m_\chi = 3$$

$$\sigma_{\text{elastic}} \sim 1/(2m_N E_r + m_{A'}^2)^2$$

$$\approx 1/(2m_N E_r)^2 \text{ if } m_{A'} < 2m_N E_r$$

$E_r \sim \mathcal{O}(10)$ keV causes the flatness in the sensitivity curve for elastic nuclear scattering

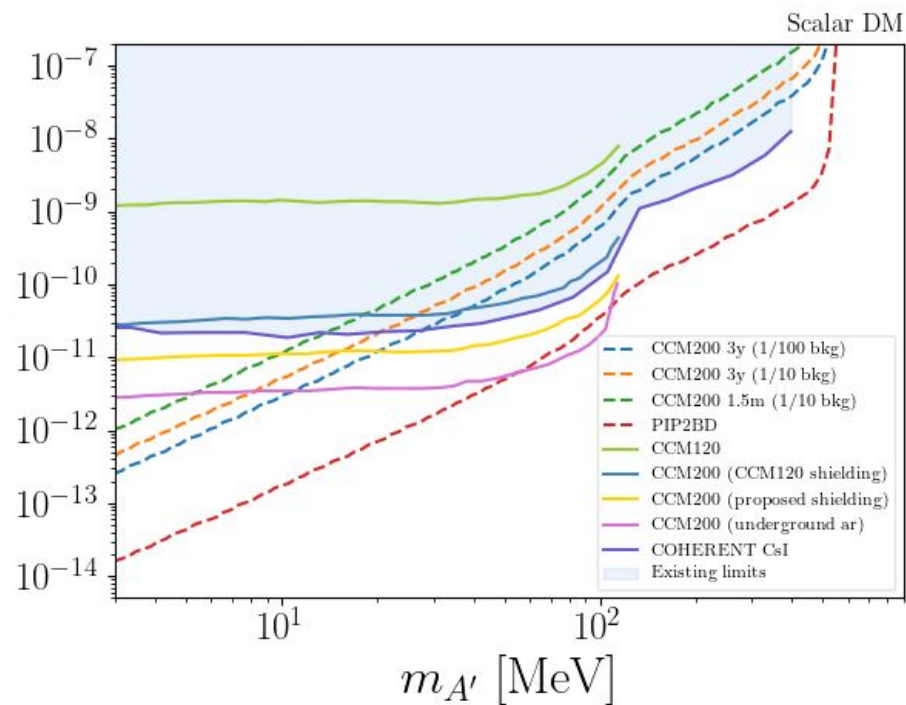
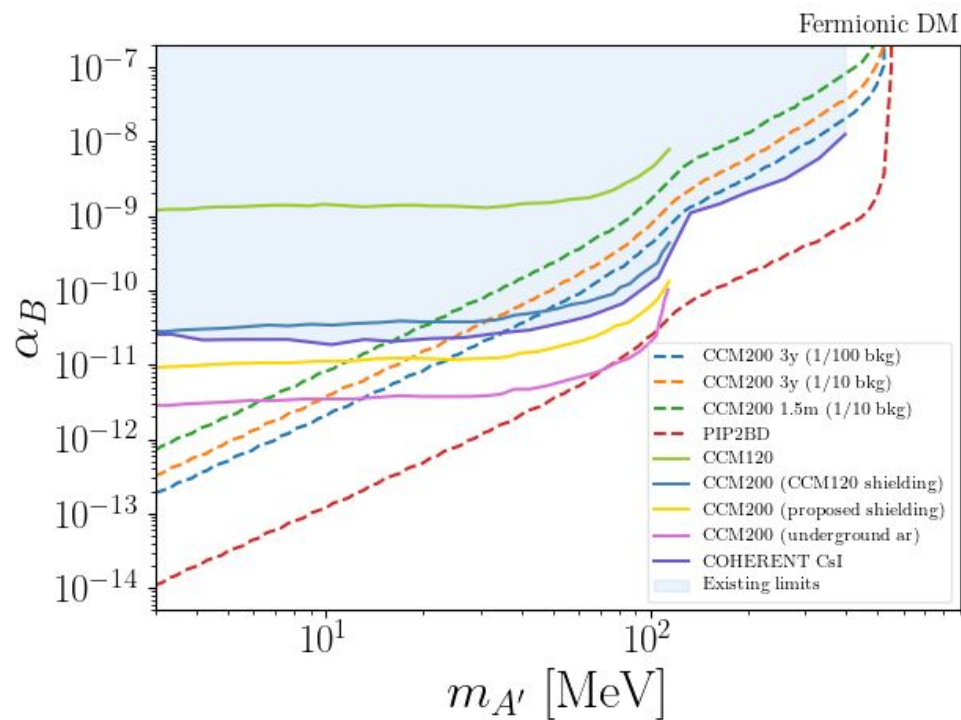


Sensitivities Plot

COHERENT Collaboration, 2205.12414;
A. A. Aguilar-Arevalo et al, 2109.14146

Dashed is our calculation

$$m_{A'} / m_\chi \approx 2$$



WIMP & Ambient DM

Interaction and cross section

$$\mathcal{L} \supset g_D A'_\mu \bar{\chi} \gamma^\mu \chi + e \epsilon Q_q A'_\mu \bar{q} \gamma^\mu q$$

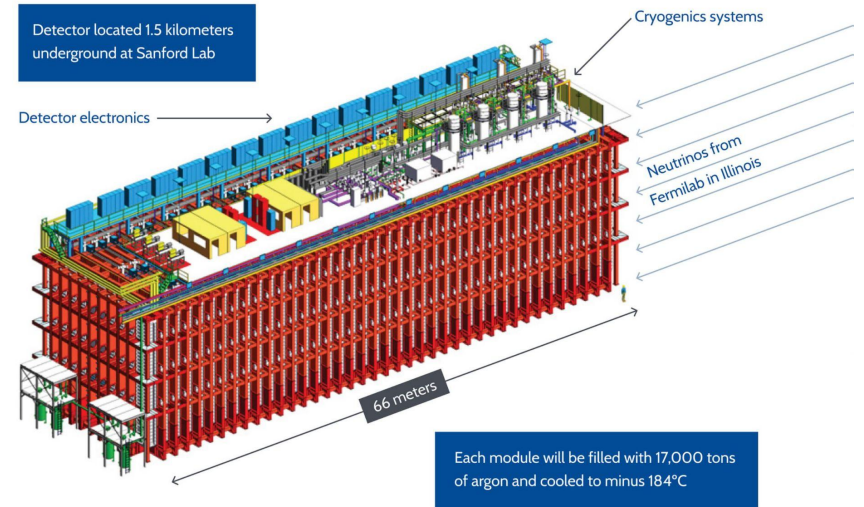
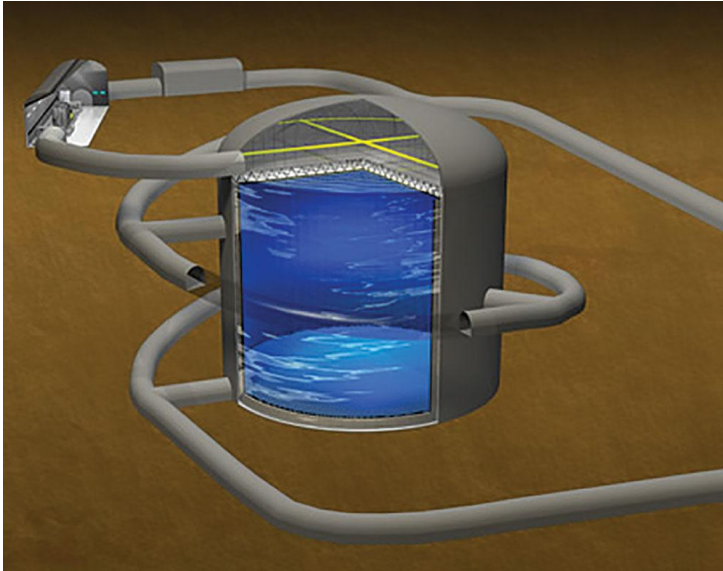
Same as light DM

$$\frac{d\sigma_{inel}^{DM}}{d\cos\theta} = \frac{2e^2\epsilon^2 g_D^2 E'_\chi p'_\chi}{(2m_N E_r + m_{A'}^2)^2} \frac{1}{2\pi} \frac{4\pi}{2J+1}$$
$$\times \boxed{r_\chi} \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2$$

A correction factor
due to large mass
of DM

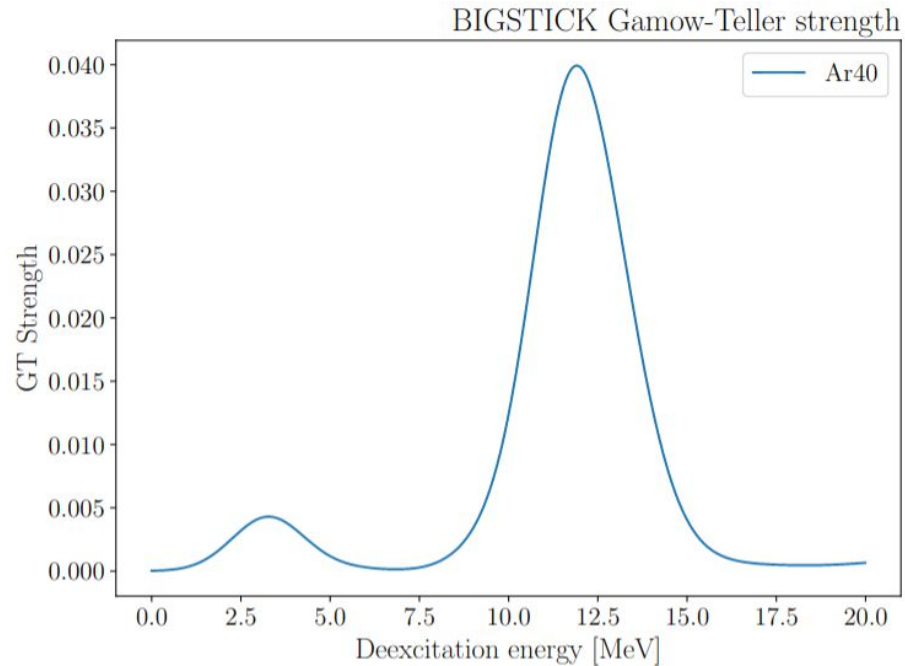
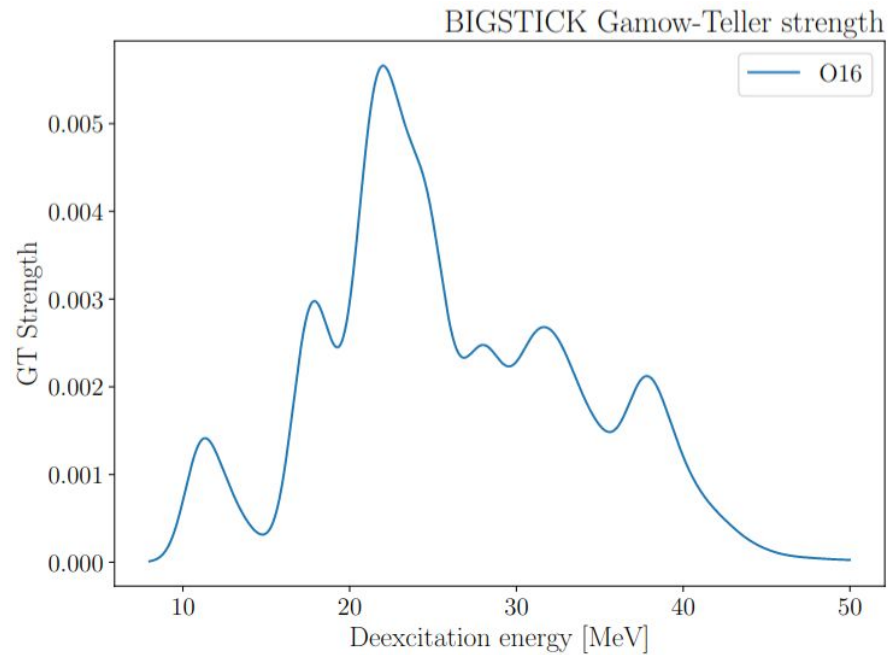
$$r_\chi = \left[1 + \frac{E_\chi}{m_N} (1 - \cos\theta) \right]^{-1}$$
$$r_\chi \rightarrow 1 \text{ if } m_N \gg m_\chi$$

Experiments



Experiment	Material	Geometry
HyperK	188kton water	78m tall, 74m diameter
DUNE	40kton Ar40	L: 65.8m, W: 18.9m, H: 17.8m

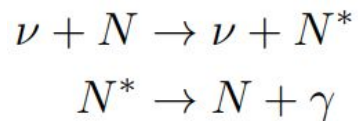
BIGSTICK GT strength



Backgrounds

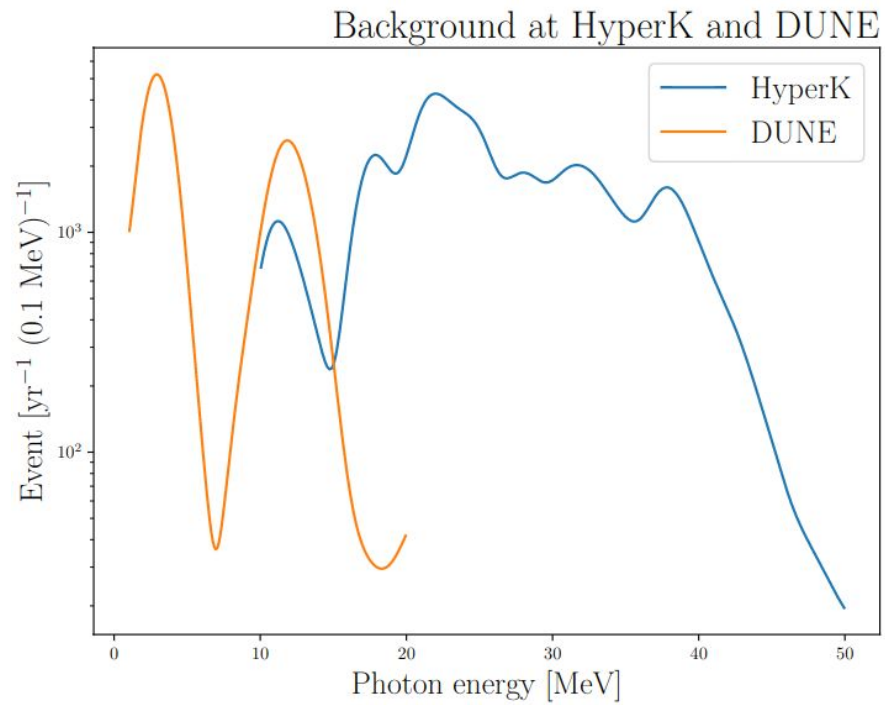
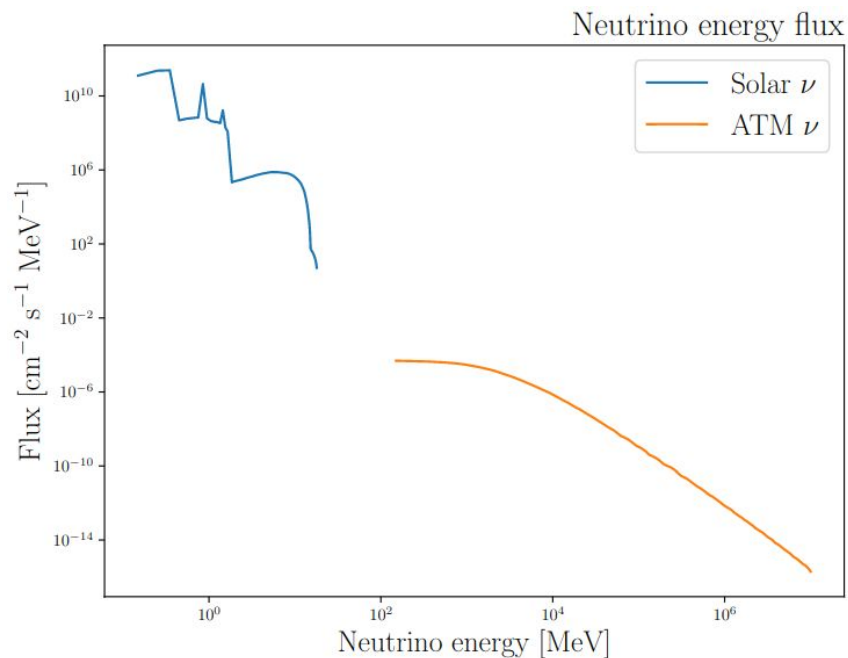
M. Honda et al., PRD 92, 023004, 2015;

Yoichiro Suzuki, 2000



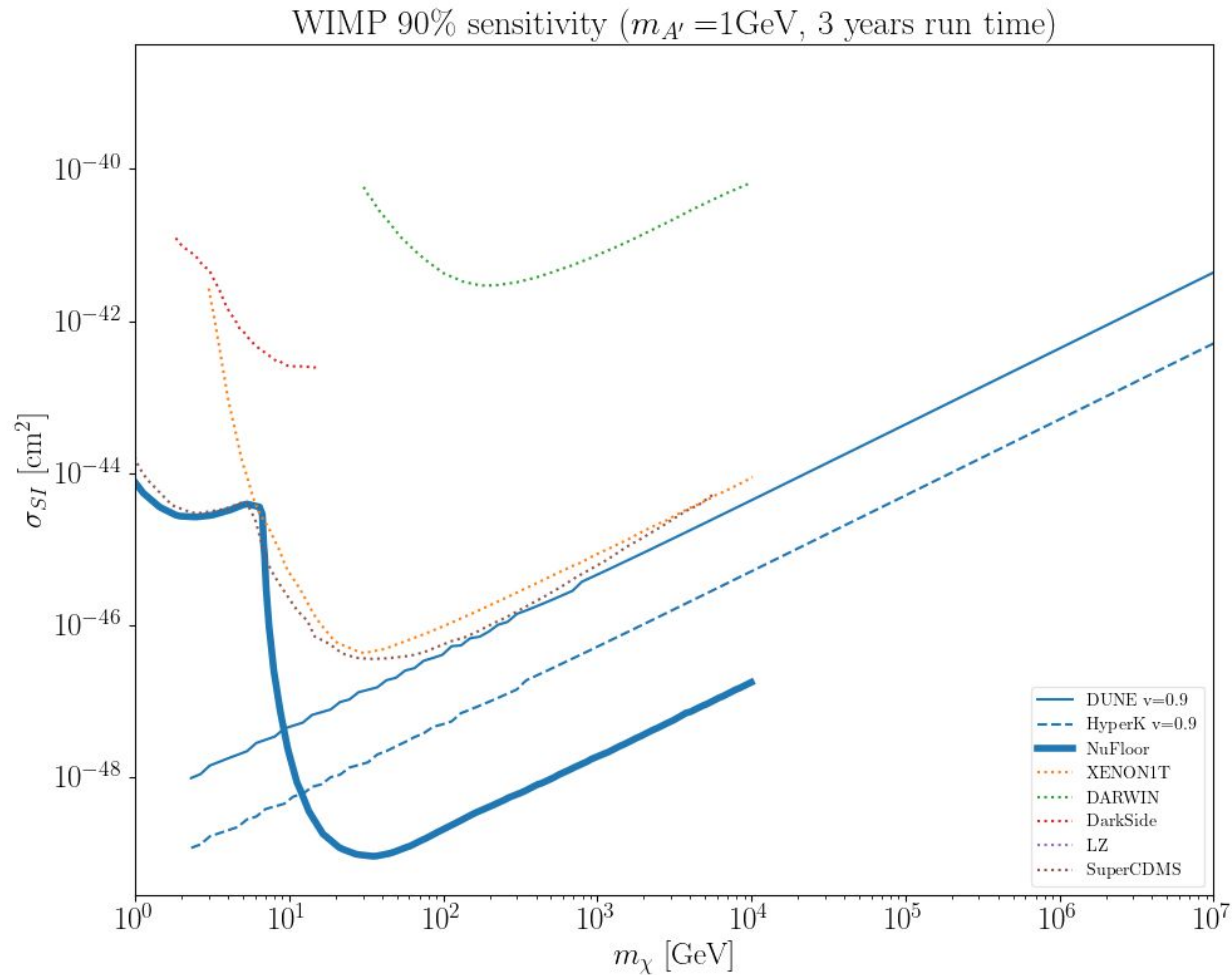
Inelastic neutrino scattering

$$\sigma_{\nu}^{\text{GT}} \approx \frac{G_f^2 g_A^2}{\pi(2J+1)} (E_{\nu} - \Delta E)^2 \left| \left\langle J_f \left\| \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 \right\| J_i \right\rangle \right|^2$$



Sensitivities Plot

Ambient WIMP
Boosted DM



Axion

Interaction

$$\mathcal{L}_{\text{int}} \supset \frac{\partial_\mu a}{f_a} \bar{N} \gamma^\mu \gamma^5 N - \frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

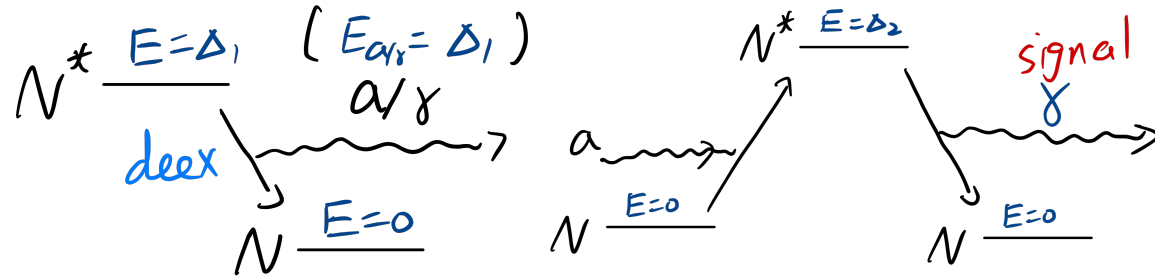
$$g_{ann} = \frac{1}{f_a}$$

a is axion (pseudoscalar)

g_{ann} and $g_{a\gamma\gamma}$ have dimension E^{-1}

$\Delta_1 =$ deexcitation energy $\ll m_N$ (source/production)

$\Delta_2 =$ excitation energy $\ll m_N$ (absorption/detection)



Production

Nuclear deexcitation $N^* \rightarrow N + \gamma/a$

Primakoff $\gamma + e^- / N \rightarrow a + e^- / N$

Detection

Nuclear absorption $a + N \rightarrow N^* \rightarrow N + \gamma$

Inverse Primakoff $a + N \rightarrow \gamma + N$

Decay $a \rightarrow 2\gamma$

Coupling

g_{ann}

$g_{a\gamma\gamma}$

g_{ann}

$g_{a\gamma\gamma}$

$g_{a\gamma\gamma}$

Production cross sections

Primakoff $\frac{d\sigma_P^p}{d\cos\theta} = \frac{1}{4}g_{a\gamma\gamma}^2\alpha Z^2 F^2(t) \frac{|\vec{p}_a|^4 \sin^2\theta}{t^2}$ $t = m_a^2 + E_\gamma(E_a - p_a \cos\theta)$

Channel	E_γ (keV)	Transitions	Φ_γ (fission ⁻¹)	(GCi)
p(n, γ)d	2230	Isovector M1	0.25	0.61
¹⁰ B(n, α) ⁷ Li*	478	M1 ($\frac{1}{2}^- \rightarrow \frac{3}{2}^-$)	0.28	0.68
⁹¹ Y*	555	M4 ($\frac{9}{2}^+ \rightarrow \frac{1}{2}^-$)	0.024	0.058
⁹⁷ Nb*	743	M4 ($\frac{1}{2}^- \rightarrow \frac{9}{2}^+$)	0.055	0.13
¹³⁵ Xe*	526	M4 ($\frac{11}{2}^- \rightarrow \frac{3}{2}^+$)	0.0097	0.023
¹³⁷ Ba*	662	M4 ($\frac{11}{2}^- \rightarrow \frac{3}{2}^+$)	0.0042	0.010

Nuclear lines
at target

Detection cross sections

$$\Gamma(a \rightarrow 2\gamma) = \frac{g_{a\gamma\gamma}^2 m_a^3}{64\pi}$$

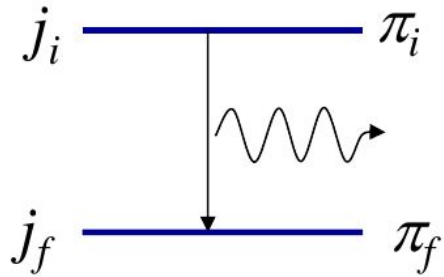
$$\frac{d\sigma_{\text{invPri}}}{d\cos\theta} = \frac{1}{2}g_{a\gamma}^2\alpha Z^2 F^2(t) \frac{p_a^4 \sin^2\theta}{t^2}$$

$$\sigma_{\text{abs}} = g_{ann}^2 2\pi \delta(E_a - E_r - \Delta_2) p_a \cdot F$$

Gamow Teller cross section

$$\sigma_{\text{abs}}(E_a, \Delta_2) = \frac{g_A^2 \pi}{6(2J+1)} g_{ann}^2 \delta(E_a - E_r - \Delta_2) p_a |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2$$

M and E transitions



$$|J_i - J_f| \leq J \leq J_i + J_f$$

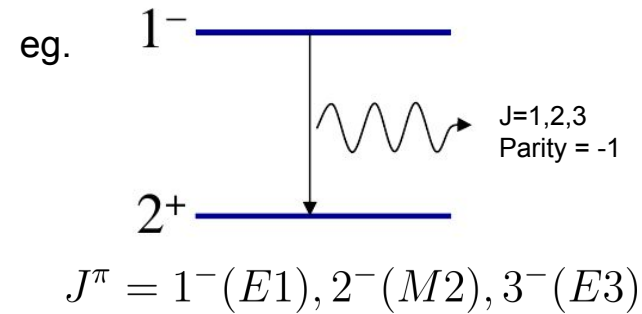
$$\pi = \pi_i \pi_f$$

Axion is pseudo scalar => -1 parity => M trans

$\pi = (-1)^I$ for E transitions

$\pi = (-1)^{I+1}$ for M transitions

multipolarity = 2^I



I	π_γ	Radiation type	Label
1	-1	Electric Dipole radiation	E1
1	+1	Magnetic Dipole radiation	M1
2	-1	Magnetic Quadrupole radiation	M2
2	+1	Electric Quadrupole radiation	E2
3	-1	Electric Octapole radiation	E3
3	+1	Magnetic Octapole radiation	M3

Nuclear deexcitation lines

$$N^* \rightarrow N + \gamma \quad \text{Photon energy } E_\gamma, \text{ Flux } \Phi_\gamma$$



Nucleon coupling model g_{ann}

$$N^* \rightarrow N + a \quad \text{Axion energy } E_a, \text{ Flux } \Phi_a$$

This table uses 2.9GW power reactor, but we use 1MW, so scaling is required

1Ci = 3.7×10^{10} decay per second

Channel	E_γ (keV)	Transitions	Φ_γ (fission ⁻¹)	$1\text{Gi} = 10^9\text{Ci}$ (GiCi)
p(n, γ)d	2230	Isvector M1	0.25	0.61
¹⁰ B(n, α) ⁷ Li*	478	M1 ($\frac{1}{2}^- \rightarrow \frac{3}{2}^-$)	0.28	0.68
⁹¹Y*	555	M4 ($\frac{9}{2}^+ \rightarrow \frac{1}{2}^-$)	0.024	0.058
⁹⁷Nb*	743	M4 ($\frac{1}{2}^- \rightarrow \frac{9}{2}^+$)	0.055	0.13
¹³⁵ Xe*	526	M4 ($\frac{11}{2}^- \rightarrow \frac{3}{2}^+$)	0.0097	0.023
¹³⁷ Ba*	662	M4 ($\frac{11}{2}^- \rightarrow \frac{3}{2}^+$)	0.0042	0.010

BIGSTICK can calculate the parameters of the branching ratio. However....

p/n orbits are different

too many nucleons

Branching ratio

F. T. Avignone III, et al., PRD.37.618, 1988

Nucleus	transition	ΔE [keV]	J	gs = 0	gs = 1	beta	eta
				n (exp)	n (BIGSTICK)		
Li7	M1	478	1/2- => 3/2-	1 => 0	2 => 1	1	-3.4 * 10^-3
Xe135	M4	526	11/2- => 3/2+	2 => 0			
Ba137	M4	662	11/2- => 3/2+	2 => 0			

$$\left(\frac{\Gamma_a}{\Gamma_\gamma}\right)_{\text{MJ}} = \frac{1}{\pi\alpha} \frac{1}{1 + \delta^2} \frac{J}{J + 1} \left(\frac{|\vec{p}_a|}{|\vec{p}_\gamma|}\right)^{2J+1} g_{ann}(\text{at left}) = g_{ann}(\text{in 1st slide}) \times GeV$$

$g_{ann}(\text{at left})$ is dimensionless

$$\times \left(\frac{g_{ann}(1 + \beta)}{(\mu_0 - 1/2)\beta + \mu_1 - \eta}\right)^2, \delta \approx 0$$

$$\beta = \frac{\langle J_f \parallel \sum_{i=1}^A \sigma^{(i)} \parallel J_i \rangle}{\langle J_f \parallel \sum_{i=1}^A \sigma^{(i)\tau_3(i)} \parallel J_i \rangle}$$

$$\eta = -\frac{\langle J_f \parallel \sum_{i=1}^A I^{(i)\tau_3(i)} \parallel J_i \rangle}{\langle J_f \parallel \sum_{i=1}^A \sigma^{(i)\tau_3(i)} \parallel J_i \rangle}$$

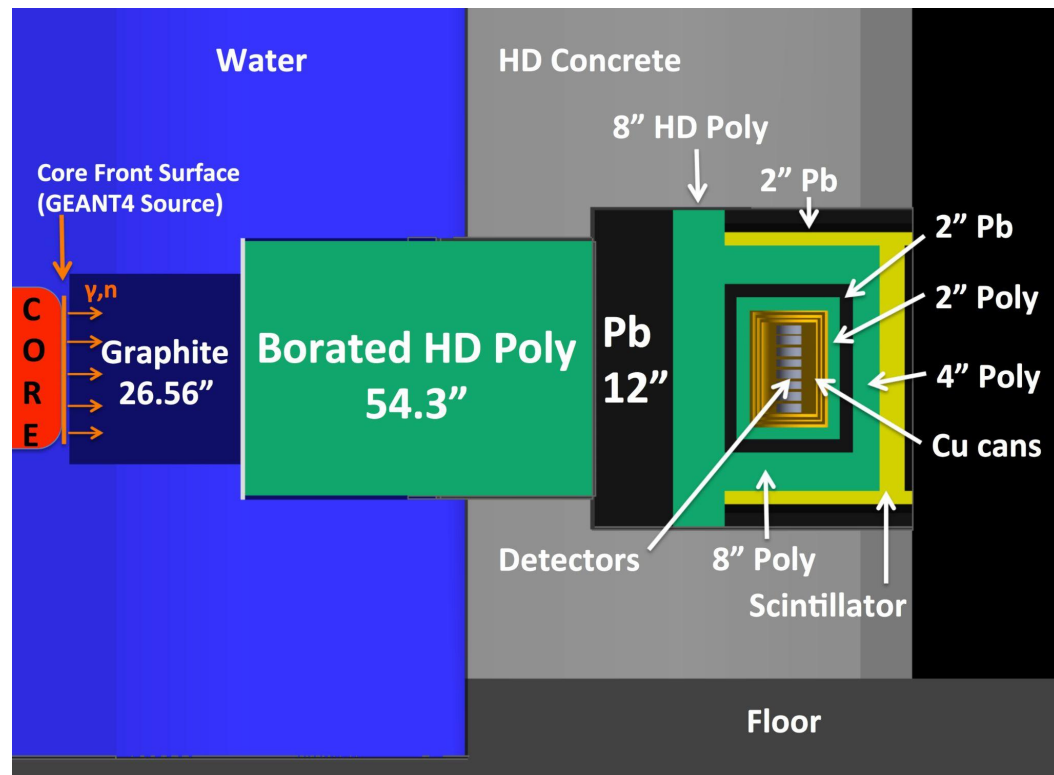
$$p + n \rightarrow D + \gamma/a$$

$$\frac{\Gamma_a}{\Gamma_\gamma}(pn \rightarrow d\gamma) = \frac{1}{2\pi\alpha} \left(\frac{p_a}{p_\gamma}\right)^3 \left(\frac{g_{ann}^1}{\mu_1}\right)^2 \quad \text{equivalently } \delta = \beta = \eta = 0, J = 1$$

Experiments

- Beam dump: CCM, IsoDAR, PIP2BD, ...
- Reactor: MINER at Texas A&M

MINER Mitchell Institute Neutrino Experiment at Reactor



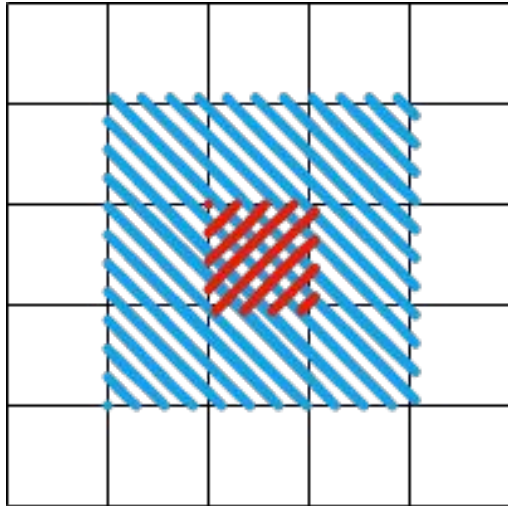
MINER CsI detector

Background reduction: 10%

Exposure: 1 year

Detector threshold: 50 keV

Name	Mass [kg]	Area [m ²]	Length [m]	Distance [m]
1x1	3.547	0.0155	0.0508	4
3x3	31.923	0.0464	0.1524	4



Detectors array

Blue: 3x3

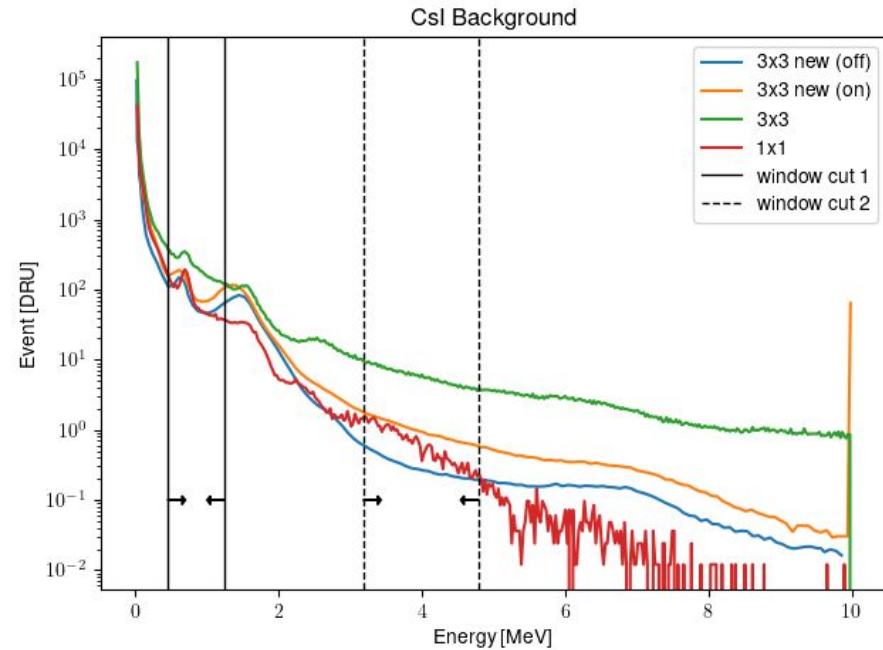
Red: 1x1

MINER Background and photon flux

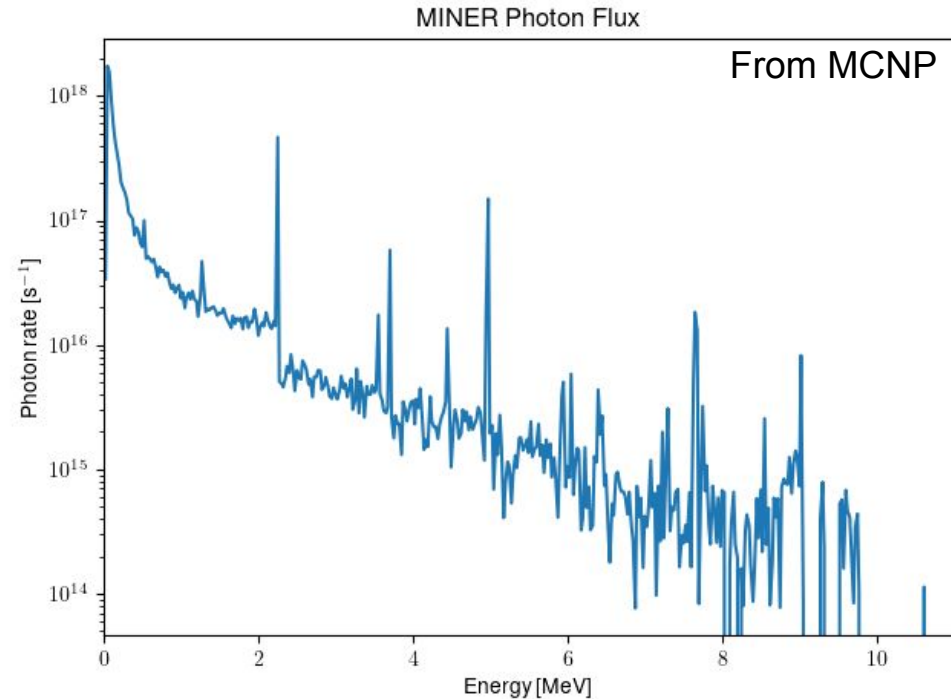
Window cuts on background

cut 1: 0.45-1.25 MeV

cut 2: 3.18-4.79 MeV



DRU = counts/kg/keV/day



BIGSTICK GT strength and cross section

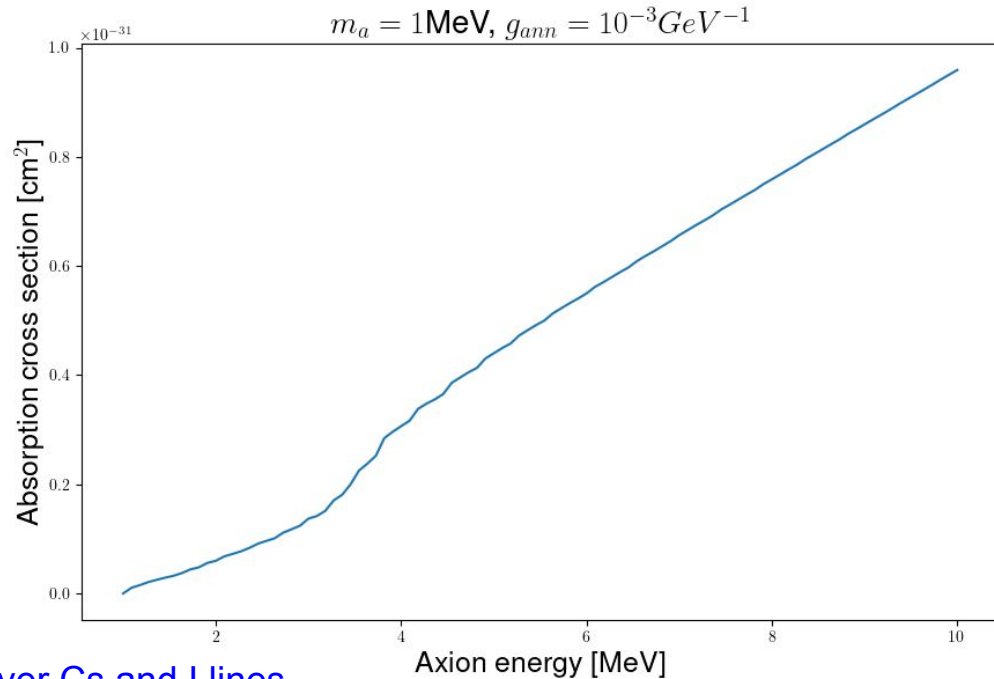
$$\sigma_{\text{abs}}(E_a, \Delta_2) = \frac{g_A^2 \pi}{6(2J+1)} g_{\text{ann}}^2 \delta(E_a - E_r - \Delta_2) p_a |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2$$

$$\sigma_{\text{abs}}(E_a) = \sum_{\Delta_2 \leq E_a} \sigma_{\text{abs}}(E_a, \Delta_2)$$

Δ_2 = excitation energy

BIGSTICK output (Cs133)

Energy	Strength	
-300.05300	0.00020	ground state
-300.01663	0.00471	
-299.81764	0.00114	
-299.69193	0.00000	
-299.57699	0.00000	
-299.50052	0.00000	
-299.45278	0.00314	
-299.41415	0.00000	
-299.39827	0.00013	
-299.33202	0.00419	
-299.27586	0.00008	
-299.21456	0.00368	
-299.18414	0.00000	

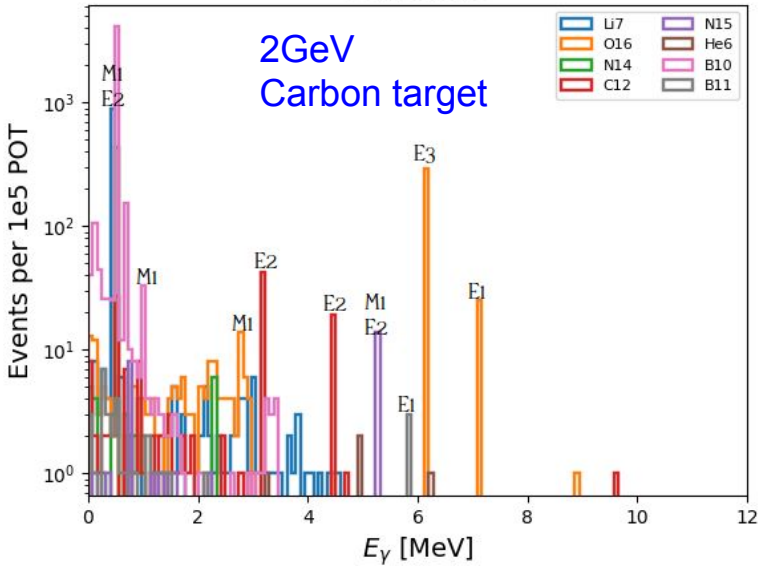


Nuclear lines in PIP2BD

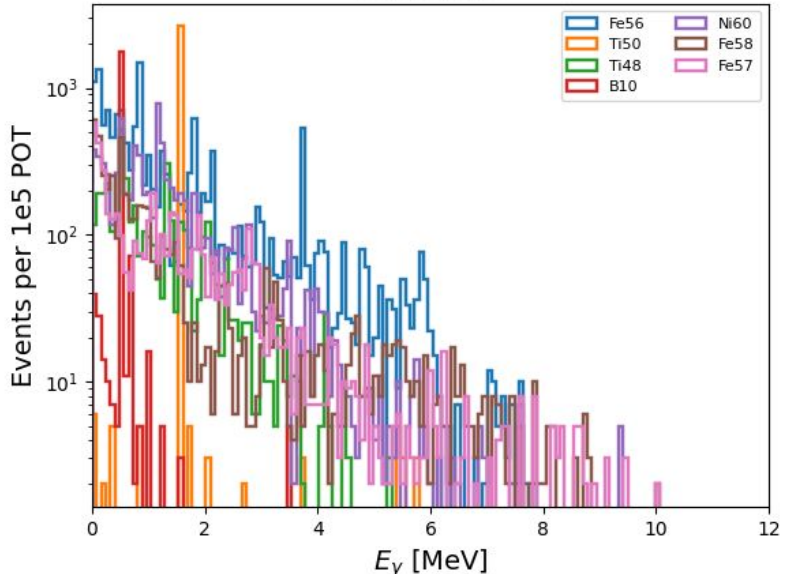
Nucleus	Energy [MeV]	per 10 ⁵ POT
Li7	0.435	887.041
N15	5.266	13.836
O16	2.774	13.836

There are too many nuclei, but most of them are not matched with experiment data. Ni60 seems to be the best match

Gammas



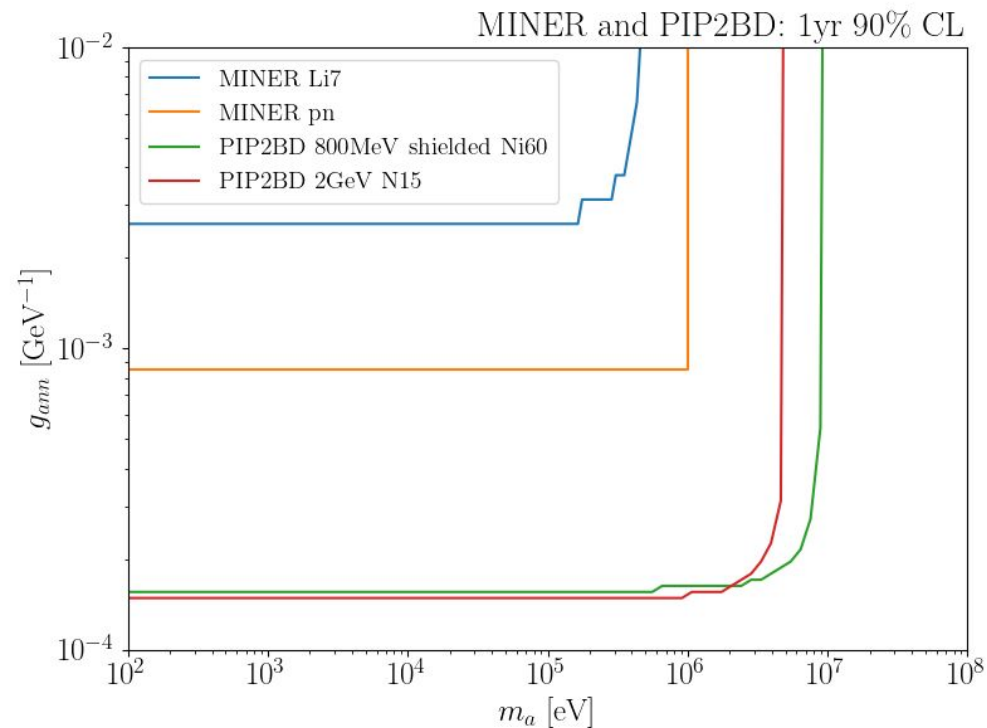
Gammas



Source: nuclear deexcitation

Detection: nuclear absorption

*axions don't decay



PIP2BD

100t LAr, 1yr exposure
1e23 total POT
100 bkg reduction

Lines:

- Ni60 9.38MeV (800MeV beam, shielded target)
- N15 5.266MeV (2GeV beam, C target)

MINER

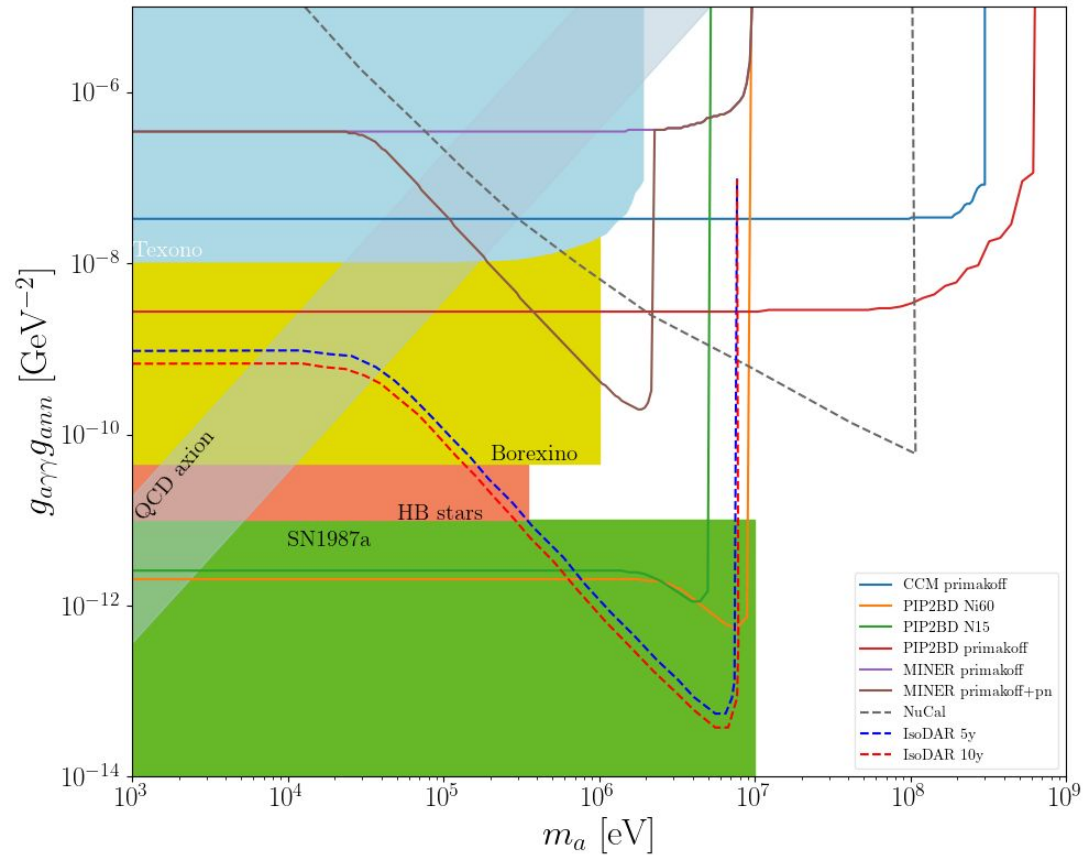
32kg CsI, 1yr exposure
No bkg reduction

Lines:

- Li7 478keV
- pn 2230keV

Source: primakoff AND/OR nuclear

Detection: inv. primakoff, decay, absorption



IsoDAR

8e24 POT per 5 yr
 2.26kton liquid scintillator
 17 meter away from target

Nucleus	Energy in MeV	Type	β	η
Fe57	7.606	M1	0.7071	-0.3111
Li8	1.009	M1	1	-0.0260
Li8	2.053	M1	1	-0.1034
O15	5.281	M2	1	0.5

L. Waites et al., 2207.13659

Conclusion

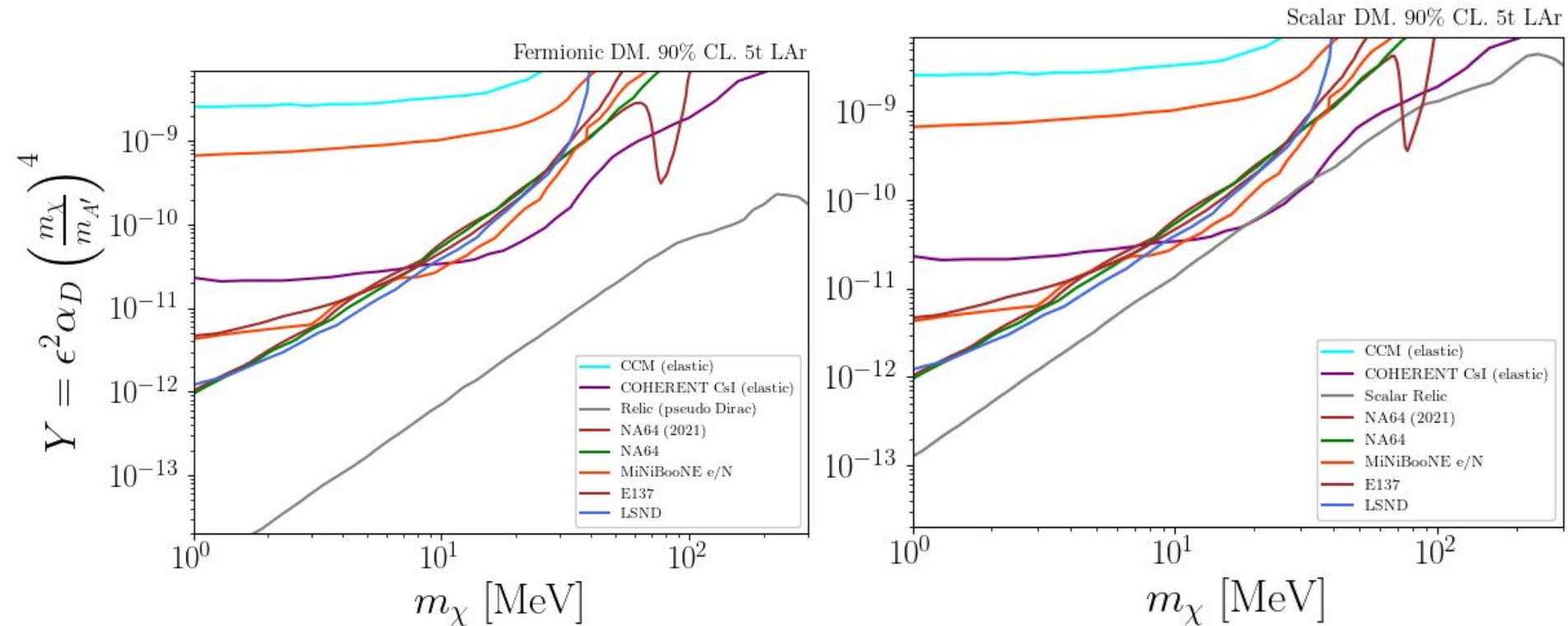
- We calculate the inelastic cross-section and event rates for DM/WIMP/axion nucleus scattering.
- Gamow-Teller transitions in long wavelength limit dominate the cross section.
- With inelastic scattering, we explore new region on the parameter space.
- We can remove most of the neutrino background efficiently with prompt timing cut in DM search in beam dump experiments

Next steps

- Search in MicroBoone and SBND
- Apply efficiency to have better estimation of the search

Backup slides

The shaded region in DM sensitivities plots



Multipole expansion (Standard Model)

Current-current interaction in Hamiltonian

$$\begin{aligned}
 \langle f | \hat{H}_W | i \rangle &= -\frac{G}{\sqrt{2}} l_\mu \int d^3x e^{-i\mathbf{q}\cdot\mathbf{x}} \langle f | \hat{\mathcal{J}}_\mu(\mathbf{x}) | i \rangle \\
 &= -\frac{G}{\sqrt{2}} \int d^3x e^{-i\mathbf{q}\cdot\mathbf{x}} [\mathbf{l} \cdot \mathcal{J}(\mathbf{x})_{fi} - l_0 \mathcal{J}_0(\mathbf{x})_{fi}]
 \end{aligned}$$

↙ Lepton current
↘ Nucleus current

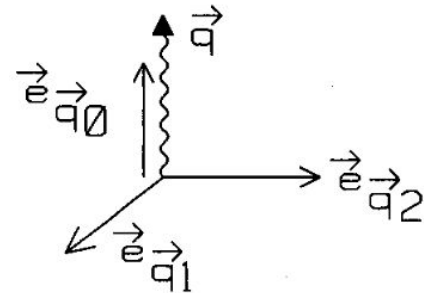
$$e_{\mathbf{q}\lambda}^\dagger e^{-i\mathbf{q}\cdot\mathbf{x}} = -\sum_{J \geq 1}^{\infty} \sqrt{2\pi(2J+1)} (-i)^J \left\{ \lambda j_J(\kappa x) \mathcal{Y}_{JJ1}^{-\lambda} + \frac{1}{\kappa} \nabla \times [j_J(\kappa x) \mathcal{Y}_{JJ1}^{-\lambda}] \right\}$$

; for $\lambda = \pm 1$

$$e_{\mathbf{q}0}^\dagger e^{-i\mathbf{q}\cdot\mathbf{x}} = \frac{i}{\kappa} \sum_{J \geq 0}^{\infty} \sqrt{4\pi(2J+1)} (-i)^J \nabla [j_J(\kappa x) Y_{J0}]$$

Spatial decomposition

$$\begin{aligned}
 \langle f | j_\mu^{\text{lept}}(\mathbf{x}) | i \rangle &= l_\mu e^{-i\mathbf{q}\cdot\mathbf{x}} \\
 \mathbf{l} &= \sum_{\lambda=0,\pm 1} l_\lambda \mathbf{e}_\lambda^\dagger
 \end{aligned}$$



Multipole expansion (Standard Model)

$$\langle f | \hat{H}_W | i \rangle = \frac{-G}{\sqrt{2}} \langle f | \left\{ - \sum_{J \geq 1} \sqrt{2\pi(2J+1)} (-i)^J \sum_{\lambda=\pm 1} l_\lambda \left[\lambda \hat{T}_{J-\lambda}^{\text{mag}}(\kappa) + \hat{T}_{J-\lambda}^{\text{el}}(\kappa) \right] \right. \\ \left. + \sum_{J \geq 0} \sqrt{4\pi(2J+1)} (-i)^J \left[l_3 \hat{\mathcal{L}}_{J0}(\kappa) - l_0 \hat{\mathcal{M}}_{J0}(\kappa) \right] \right\} | i \rangle$$

Multipole Operators

$$\hat{\mathcal{M}}_{JM}(\kappa) \equiv \hat{M}_{JM} + \hat{M}_{JM}^5 = \int d^3x [j_J(\kappa x) Y_{JM}(\Omega_x)] \hat{\mathcal{J}}_0(\mathbf{x})$$

$$\hat{\mathcal{L}}_{JM}(\kappa) \equiv \hat{L}_{JM} + \hat{L}_{JM}^5 = \frac{i}{\kappa} \int d^3x \{ \nabla [j_J(\kappa x) Y_{JM}(\Omega_x)] \} \cdot \hat{\mathcal{J}}(\mathbf{x})$$

$$\hat{T}_{JM}^{\text{el}}(\kappa) \equiv \hat{T}_{JM}^{\text{el}} + \hat{T}_{JM}^{\text{el}5} = \frac{1}{\kappa} \int d^3x [\nabla \times j_J(\kappa x) \mathcal{Y}_{JJ_1}^M(\Omega_x)] \cdot \hat{\mathcal{J}}(\mathbf{x})$$

$$\hat{T}_{JM}^{\text{mag}}(\kappa) \equiv \hat{T}_{JM}^{\text{mag}} + \hat{T}_{JM}^{\text{mag}5} = \int d^3x [j_J(\kappa x) \mathcal{Y}_{JJ_1}^M(\Omega_x)] \cdot \hat{\mathcal{J}}(\mathbf{x})$$

Multipole operator and response function

Multipole
Operators

Response
functions

$$\mathcal{M} = \mathcal{M}_{LM} + \mathcal{M}_{LM}^5 = \left\{ F_1^N M_L^M + \frac{\mathbf{q}^2}{4m_N^2} (F_1^N + 2F_2^N) (\Phi_L''^M - \frac{1}{2} M_{LM}) \right\} + \left\{ -i \frac{|\mathbf{q}|}{m_N} G_A^N \left[\Omega_L^M + \frac{1}{2} \Sigma_L''^M \right] \right\}$$

$$\mathcal{L} = \mathcal{L}_{LM} + \mathcal{L}_{LM}^5 = \left\{ \frac{q^0}{|\mathbf{q}|} \mathcal{M} \right\} + \left\{ i \left[G_A^N \left(1 - \frac{\mathbf{q}^2}{8m_N^2} \right) - \frac{\mathbf{q}^2}{4m_N^2} G_P^N \right] \Sigma_L''^M \right\}$$

$$\mathcal{T}^{el} = \mathcal{T}_{LM}^{el} + \mathcal{L}_{LM}^{el5} = \left\{ \frac{|\mathbf{q}|}{m_N} \left[F_1^N \Delta_L'^M + \frac{F_1^N + F_2^N}{2} \Sigma_L^M \right] \right\} + \left\{ i G_A^N \left(1 - \frac{\mathbf{q}^2}{8m_N^2} \right) \Sigma_L'^M \right\}$$

$$\mathcal{T}^{mag} = \mathcal{T}_{LM}^{mag} + \mathcal{L}_{LM}^{mag5} = \left\{ -i \frac{|\mathbf{q}|}{m_N} \left[F_1^N \Delta_L^M - \frac{F_1^N + F_2^N}{2} \Sigma_L'^M \right] \right\} + \left\{ G_A^N \left(1 - \frac{\mathbf{q}^2}{8m_N^2} \right) \Sigma_L^M \right\}$$

Multipole expansion (Standard Model)

$$\frac{1}{(2J_i + 1)} \sum_{M_i} \sum_{M_f} |\langle f | \hat{H}_W | i \rangle|^2 = \frac{G^2}{2} \frac{4\pi}{(2J_i + 1)} \times$$

$$\left\{ \sum_{J \geq 1} \left[\frac{1}{2} (1 \cdot \Gamma^* - l_3 l_3^*) \left(|\langle J_f | \hat{T}_J^{\text{mag}} | J_i \rangle|^2 + |\langle J_f | \hat{T}_J^{\text{el}} | J_i \rangle|^2 \right) \right. \right. \\ \left. \left. - \frac{i}{2} (1 \times \Gamma^*)_3 \left(2 \operatorname{Re} \langle J_f | \hat{T}_J^{\text{mag}} | J_i \rangle \langle J_f | \hat{T}_J^{\text{el}} | J_i \rangle^* \right) \right] \right. \\ \left. + \sum_{J \geq 0} \left[l_3 l_3^* |\langle J_f | \hat{\mathcal{L}}_J | J_i \rangle|^2 + l_0 l_0^* |\langle J_f | \hat{\mathcal{M}}_J | J_i \rangle|^2 \right. \right. \\ \left. \left. - 2 \operatorname{Re} \left(l_3 l_0^* \langle J_f | \hat{\mathcal{L}}_J | J_i \rangle \langle J_f | \hat{\mathcal{M}}_J | J_i \rangle^* \right) \right] \right\}$$

See details at the book
Theoretical Nuclear and
Subnuclear Physics (2ed) by
Walecka

$$\left(\frac{d\sigma}{d\Omega} \right)_{\nu \bar{\nu}}^{\text{ERL}} = \frac{G^2 \varepsilon^2}{2\pi^2} \frac{4\pi}{2J_i + 1} \left\{ \cos^2 \frac{\theta}{2} \sum_{J=0}^{\infty} \left| \langle J_f | \hat{\mathcal{M}}_J - \frac{q_0}{|\mathbf{q}|} \hat{\mathcal{L}}_J | J_i \rangle \right|^2 \right.$$

$$\left. + \left[\frac{q_\mu^2}{2q^2} \cos^2 \frac{\theta}{2} + \sin^2 \frac{\theta}{2} \right] \sum_{J=1}^{\infty} \left[|\langle J_f | \hat{T}_J^{\text{mag}} | J_i \rangle|^2 + |\langle J_f | \hat{T}_J^{\text{el}} | J_i \rangle|^2 \right] \right.$$

$$\left. \mp \frac{\sin \theta / 2}{|\mathbf{q}|} \sqrt{q_\mu^2 \cos^2 \frac{\theta}{2} + q^2 \sin^2 \frac{\theta}{2}} \sum_{J=1}^{\infty} 2 \operatorname{Re} \left[\langle J_f | \hat{T}_J^{\text{mag}} | J_i \rangle \langle J_f | \hat{T}_J^{\text{el}} | J_i \rangle^* \right] \right\}$$

Cross
section

Inel DM scattering cross section

$$\frac{d\sigma_{inel}^{DM}}{dE_r} = \frac{2e^2\epsilon^2 g_D^2 E'_\chi{}^2}{p_\chi p'_\chi (2m_N E_r + m_{A'}^2)^2} \frac{m_N}{2\pi} \frac{4\pi}{2J+1}$$

$$\times \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} \left| \langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle \right|^2$$

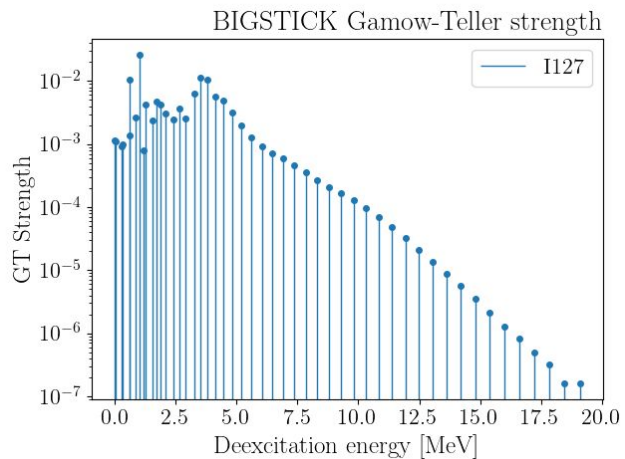
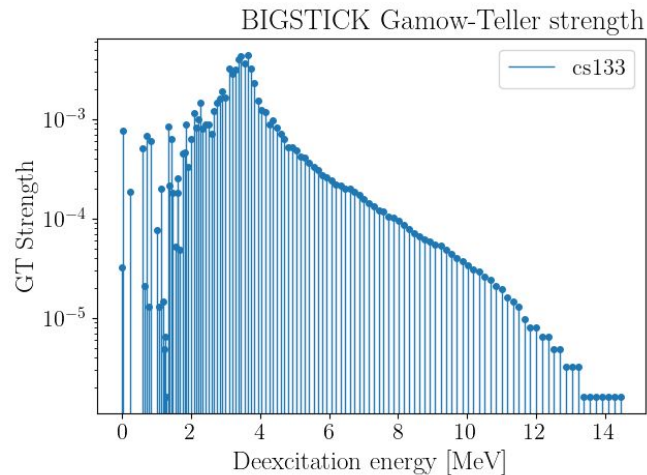
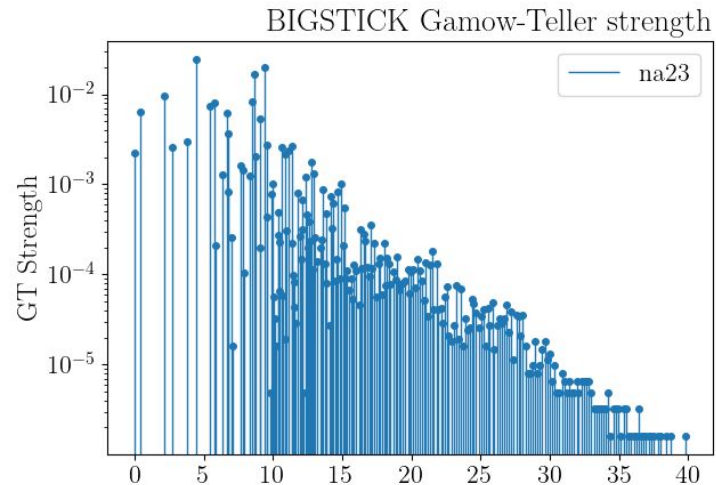
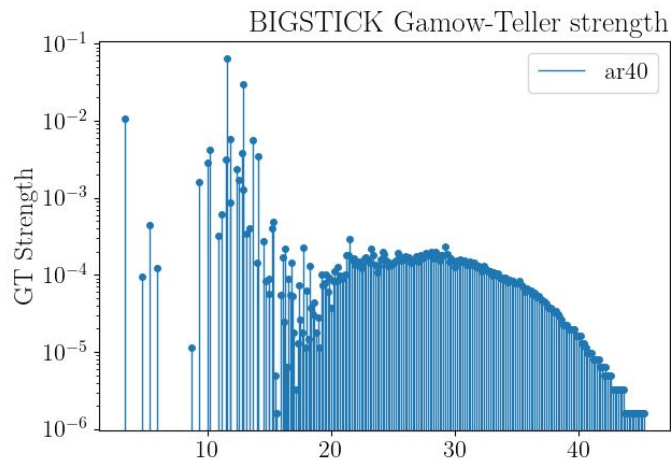
Rewrite in scattering angle with

$$\frac{d\sigma}{dE_r} = \frac{d\sigma}{d\Omega} \frac{d\Omega}{dE_r} = 2\pi \frac{d\sigma}{d\Omega} \frac{m_N}{p_\chi p'_\chi} = \frac{d\sigma}{d\cos\theta} \frac{m_N}{p_\chi p'_\chi}$$

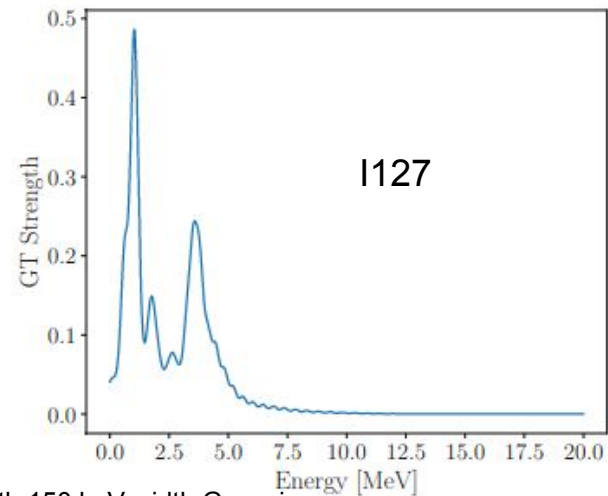
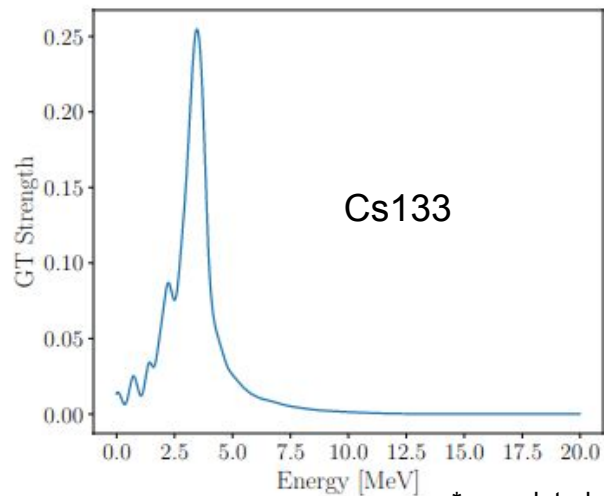
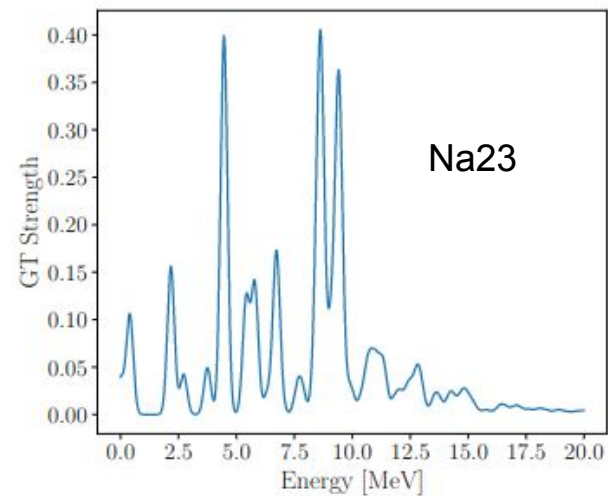
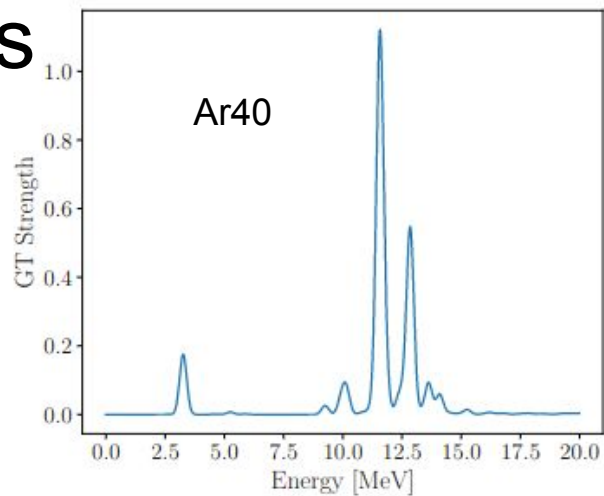
$$\frac{d\sigma_{inel}^{DM}}{d\cos\theta} = \frac{2e^2\epsilon^2 g_D^2 E'_\chi{}^2}{(2m_N E_r + m_{A'}^2)^2} \frac{1}{2\pi} \frac{4\pi}{2J+1}$$

$$\times \sum_{s_i, s_f} \vec{l} \cdot \vec{l}^* \frac{g_A^2}{12\pi} \left| \langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle \right|^2$$

Raw Discrete Strength Lines



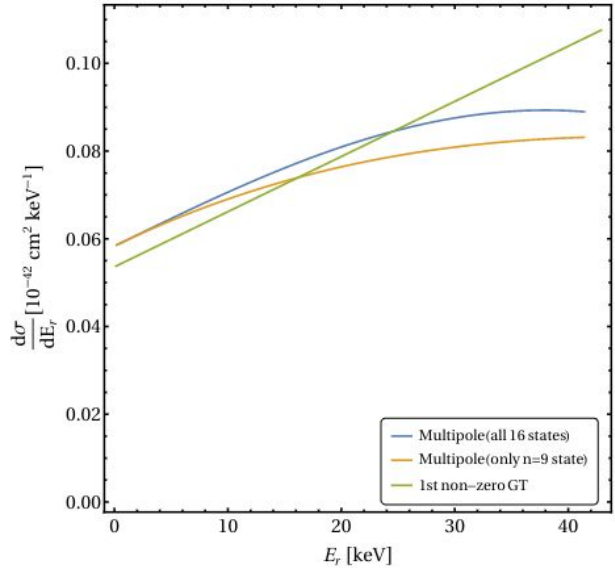
Convolution Plots



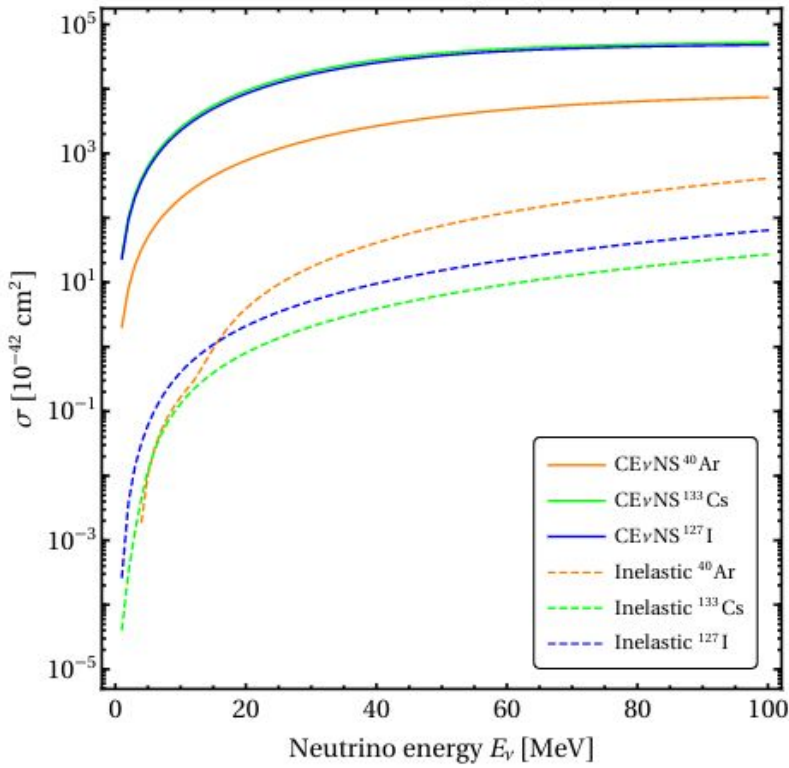
*convoluted with 150 keV width Gaussian

Inelastic neutrino-nucleus scattering

Similar to DM scattering, GT also dominates



$$\sigma_{\nu}^{GT} \approx \frac{G_f^2 g_A^2}{\pi(2J+1)} (E_{\nu} - \Delta E)^2 |\langle J_f || \sum_{i=1}^A \frac{1}{2} \hat{\sigma}_i \hat{\tau}_0 || J_i \rangle|^2.$$

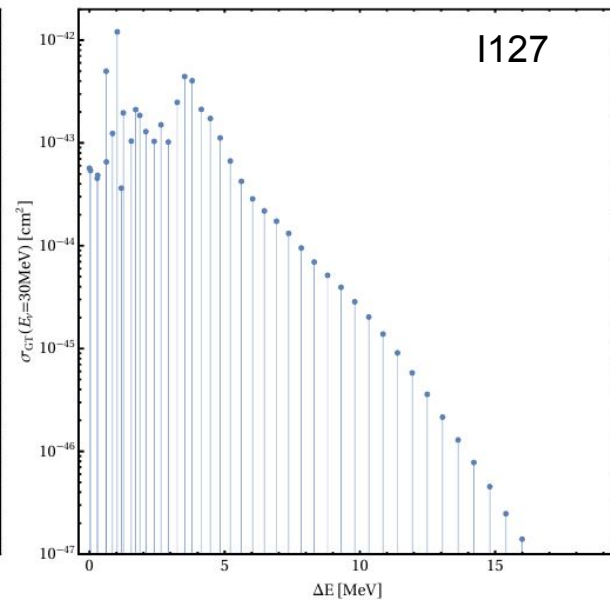
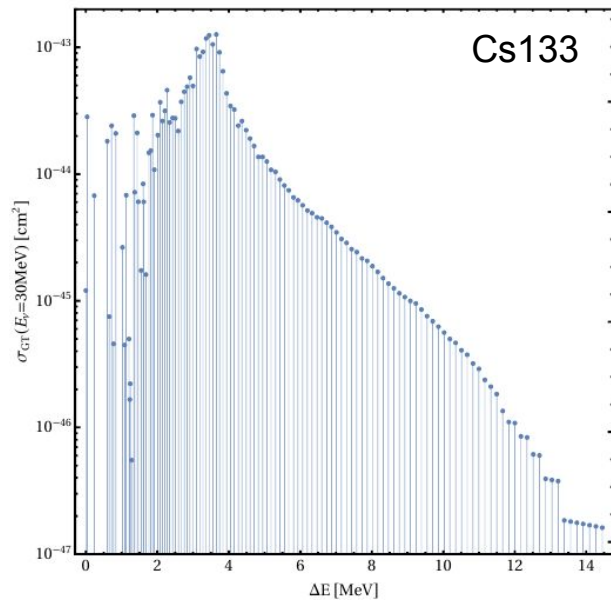
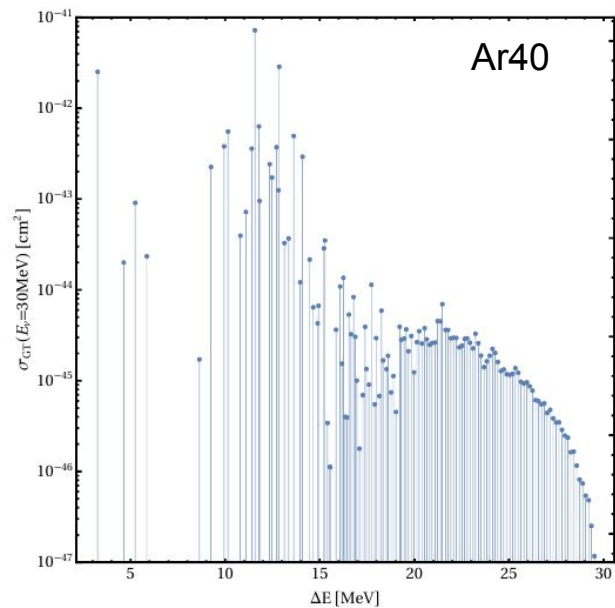


Events

Scattering	Experiment	Elastic	Inelastic	Ratio
ν - ^{40}Ar	COHERENT	2.27×10^2	3.15	7.21×10
ν - ^{40}Ar	CCM	1.91×10^4	2.65×10^2	7.21×10
ν - ^{133}Cs	COHERENT	1.16×10^3	1.52×10^{-2}	7.65×10^3
ν - ^{127}I	COHERENT	1.06×10^3	3.75×10^{-1}	2.81×10^3

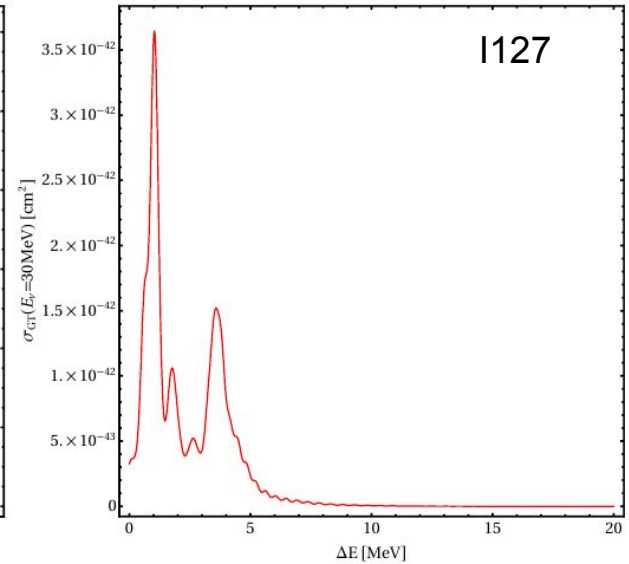
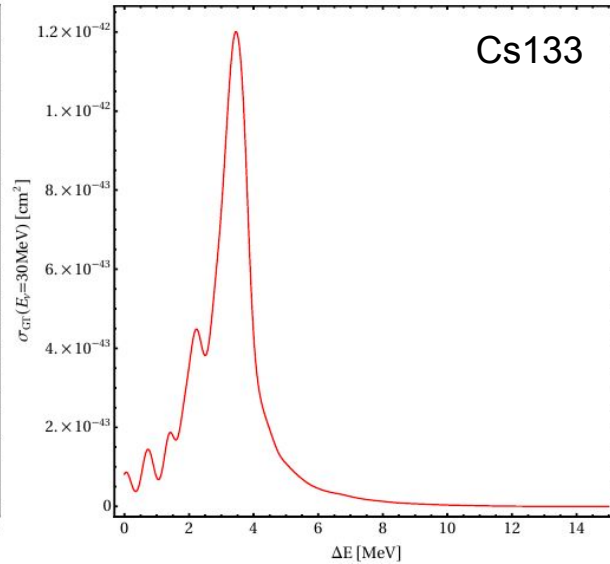
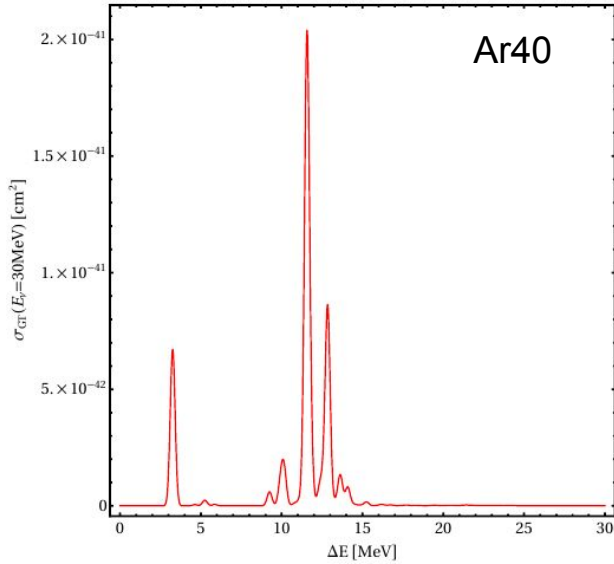
GT strength for neutrino scattering

30MeV nu energy



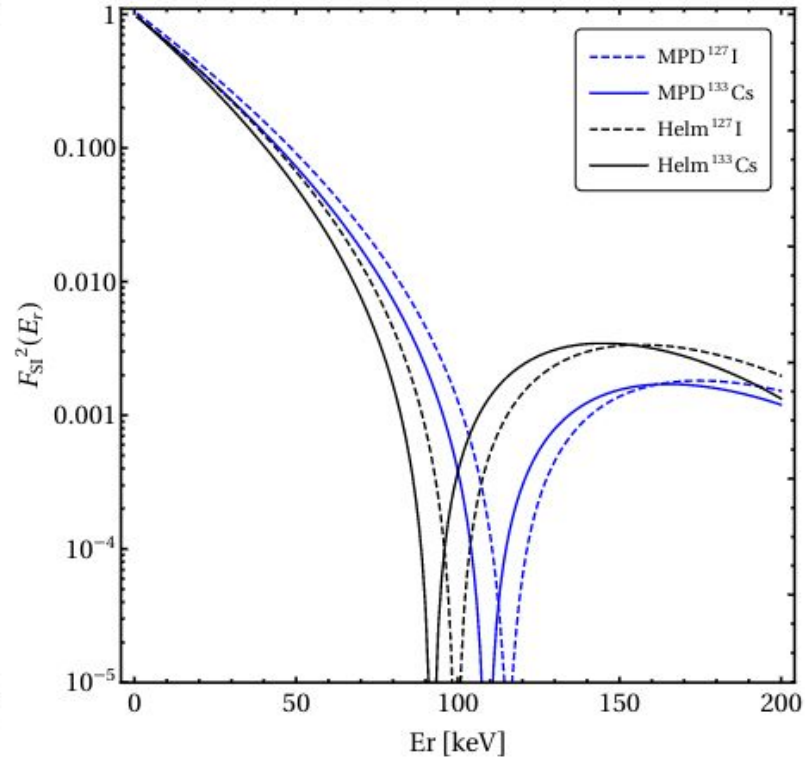
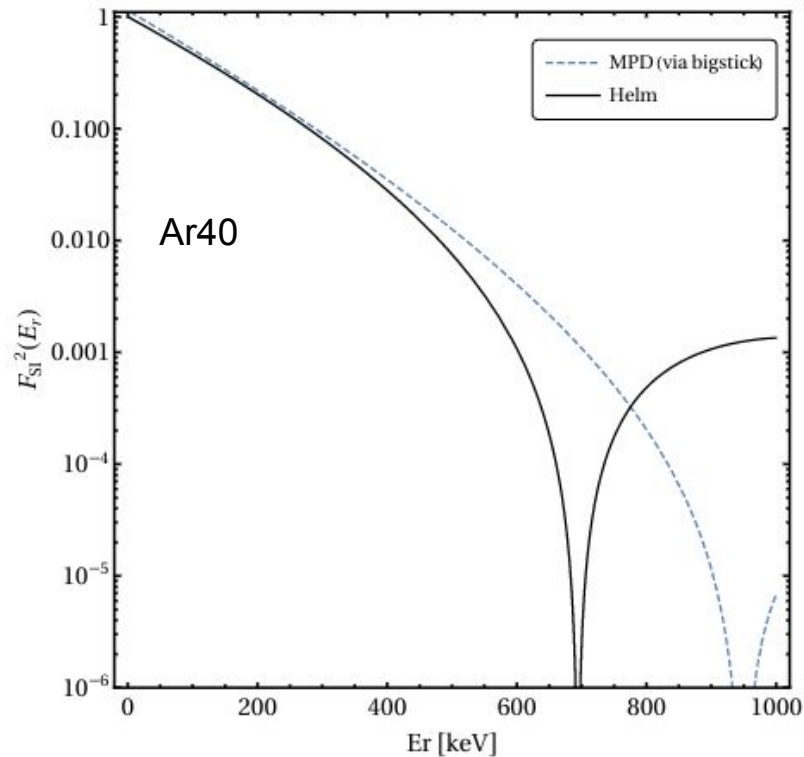
GT strength for neutrino scattering

30MeV nu energy
150 keV width Gaussian



BIGSTICK ground state to ground state comparing to Helm form factor

MPD = multipole decomposition



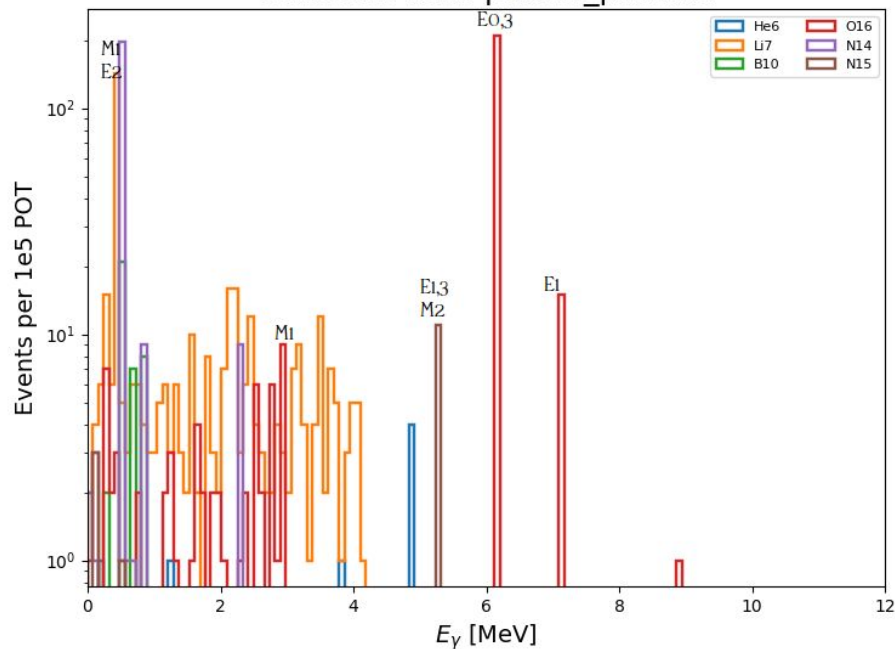
M lines from GEANT

PIP2BD

800MeV Be

Nucleus	Energy [MeV]	per 10 ⁵ POT
Li7	0.435	141.2

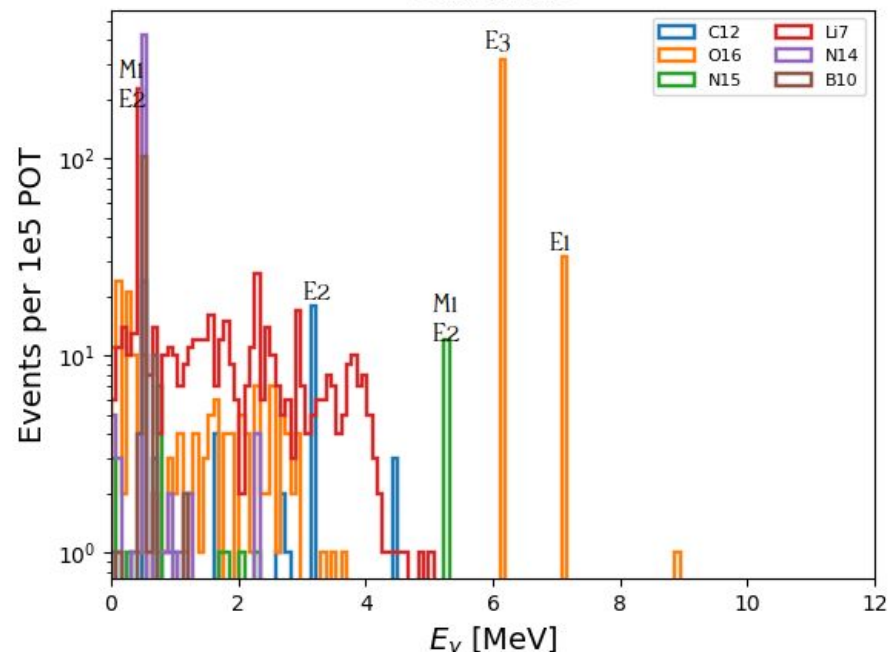
Gammas from parent_particle



2GeV Be

Nucleus	Energy [MeV]	per 10 ⁵ POT
Li7	0.435	222.54
N15	5.266	11.89

Gammas



Nucleus	Energy [MeV]	per 10 ⁵ POT
Li7	0.435	465.8
N15	2.29	3.974
O16	2.774	13.836
B10	1	21.675

2GeV C

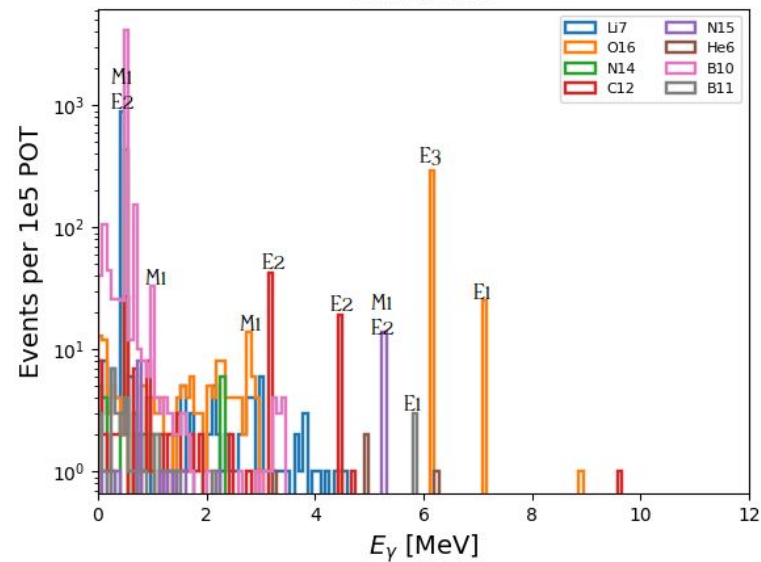
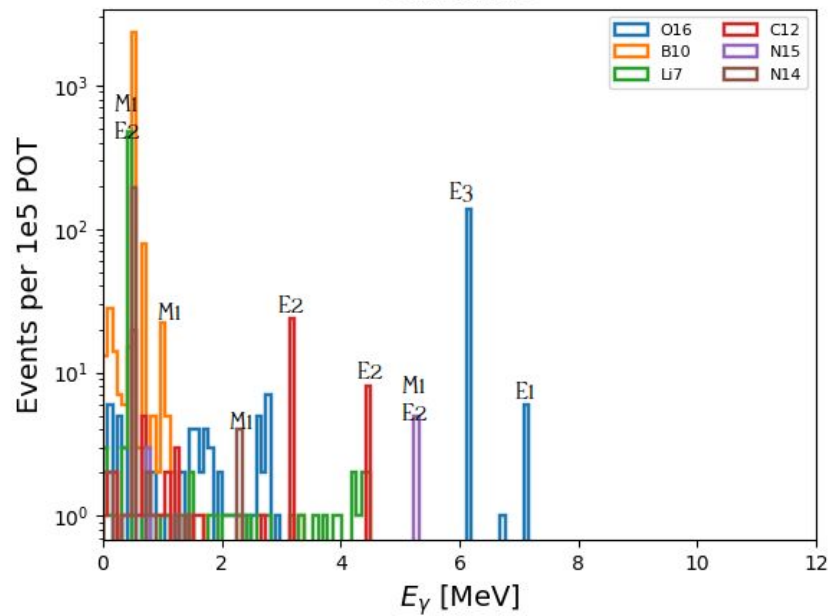
PIP2BD

Nucleus	Energy [MeV]	per 10 ⁵ POT
Li7	0.435	887.041
N15	5.266	13.836
O16	2.774	13.836
B10	1	32.79

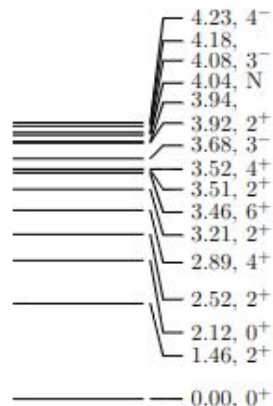
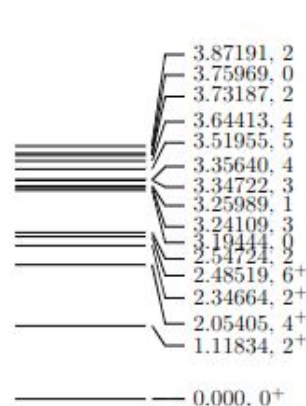
800MeV C

Gammas

Gammas

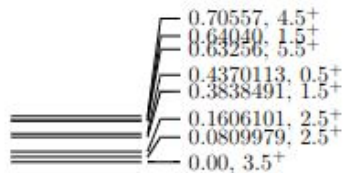
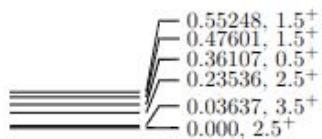


BIGSTICK energy level and spin



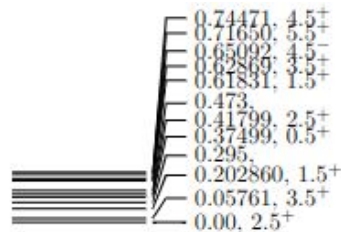
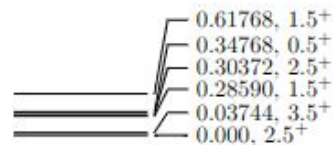
^{40}Ar (BIGSTICK)

^{40}Ar (Exp)



^{133}Cs (BIGSTICK)

^{133}Cs (Exp)



^{127}I (BIGSTICK)

^{127}I (Exp)

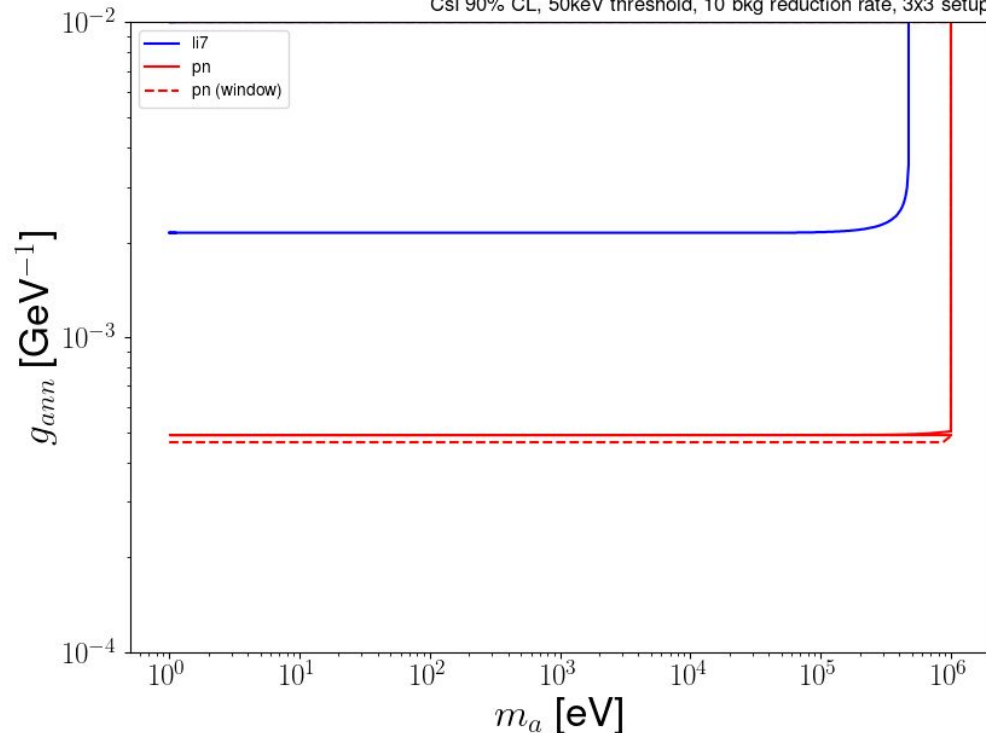
BIGSTICK nuclear magnetic moments

in keV						
Nucleus	level n	J^π	μ	Expt.	E_x	Expt.
^{127}I	1	$5/2^+$	3.851	2.813	0	0
	2	$7/2^+$	3.007	2.54	37.44	57.61
	3	$3/2^+$	0.9155	0.97	285.9	202.86
^{133}Cs	1	$7/2^+$	3.007	2.582	0	0
	2	$5/2^+$	3.851	3.45	36.37	80.9979
	3	$5/2^+$	2.5849	2.0	235.36	160.6101
^{40}Ar	1	0^+	0	N/A	0	0
	2	2^+	0	-0.04	1118.33	1460.85

Source: Li7 (M1 478keV) / pn (M1 2230keV)

Detection: nuclear absorption

g_{ann} only. Nuclear deexcitation+excitation
CsI 90% CL, 50keV threshold, 10 bkg reduction rate, 3x3 setup

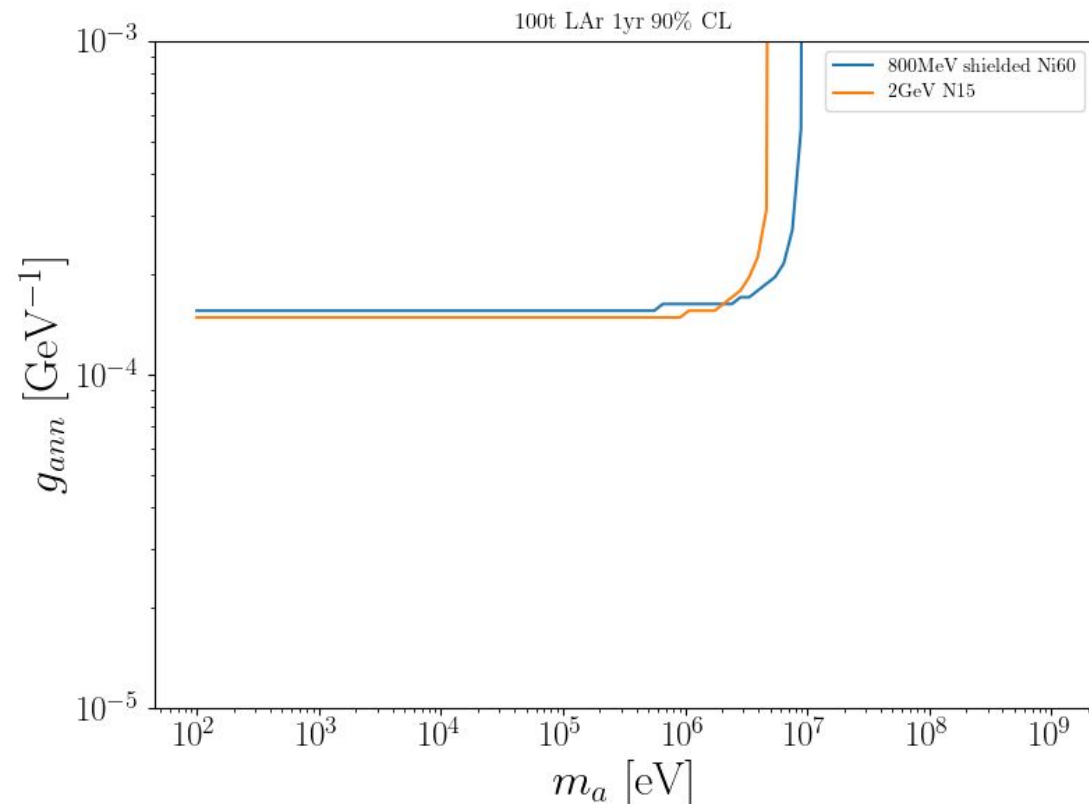


*axions don't decay

Energy window: 0.45-1.25MeV
Li7 is skipped because it doesn't have
enough of bins

Source: Ni60 9.38MeV / N15 5.266MeV

Detection: nuclear absorption



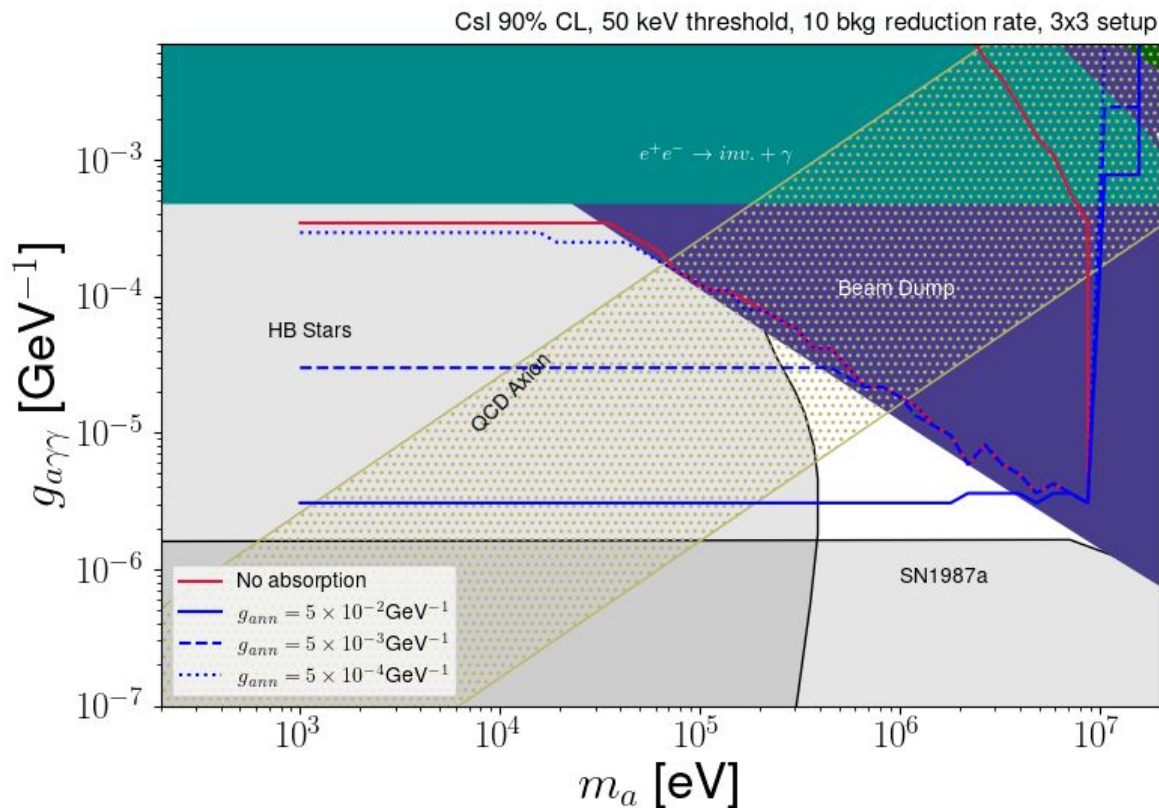
*axions don't decay

100t LAr, 1yr exposure
1e23 total POT

800MeV beam, shielded target: Ni60 9.38MeV
2GeV beam, C target: N15 5.266MeV

Source: primakoff

Detection: inv. primakoff, decay, absorption



WIMP existing constraints

