

Statistical evaluation of $0\nu\beta\beta$ Nuclear Matrix Elements uncertainties

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Support from DOE grant DE-SC0022538 is acknowledged

Plan of talk

- Neutrinoless double beta decay: Brief introduction
- Nuclear matrix elements
 - Shell model approach
 - Statistical analysis for ^{136}Xe
 - Statistical analysis for ^{76}Ge
- Conclusions and outlook

Two-Neutrino Double Beta Decay

SEPTEMBER 15, 1935

PHYSICAL REVIEW

VOLUME 48

Double Beta-Disintegration

M. GOEPPERT-MAYER, *The Johns Hopkins University*

(Received May 20, 1935)

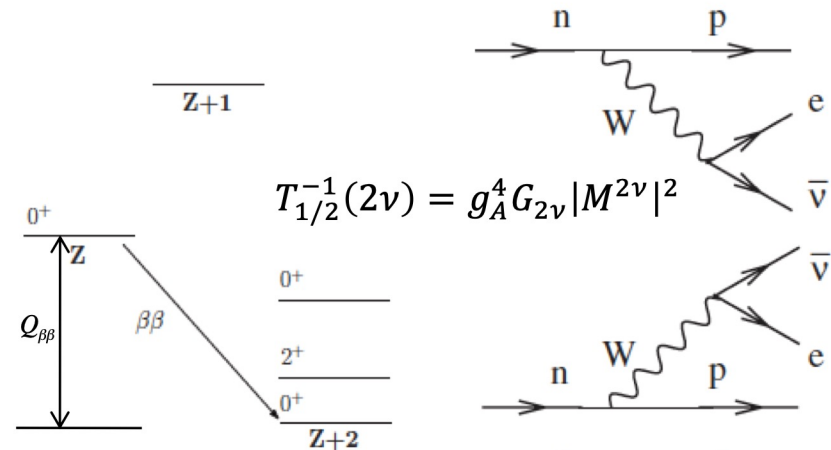
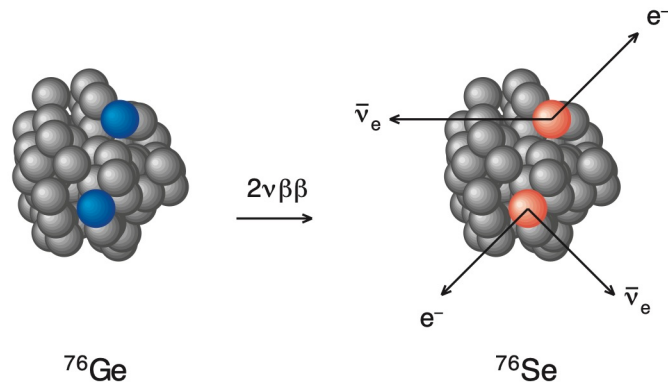
From the Fermi theory of β -disintegration the probability of simultaneous emission of two electrons (and two neutrinos) has been calculated. The result is that this process occurs sufficiently rarely to allow a half-life of over 10^{17} years for a nucleus, even if its isobar of atomic number different by 2 were more stable by 20 times the electron mass.

$$P = 1.15 \times 10^{-35} F(\epsilon - 2) \text{ sec.}^{-1}$$

$$= 3.6 \times 10^{-28} F(\epsilon - 2) \text{ year}^{-1}. \quad (12)$$

$$F(\epsilon + 2) = \begin{matrix} \epsilon = & 4 & 6 & 8 & 10 \\ 0.37 \times 10^2 & 9.2 \times 10^4 & 3.4 \times 10^6 & 4.2 \times 10^7 \end{matrix}$$

$$P(\epsilon - 2) = \begin{matrix} \epsilon = & 12 & 20 \\ 3.3 \times 10^8 & 1 \times 10^{11} \end{matrix}$$



Neutrinoless Double Beta Decay

DECEMBER 15, 1939

PHYSICAL REVIEW

VOLUME 56

On Transition Probabilities in Double Beta-Disintegration

W. H. FURRY

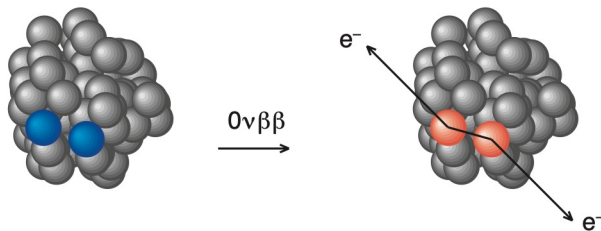
Physics Research Laboratory, Harvard University, Cambridge, Massachusetts

(Received October 16, 1939)

THE probability of double β -disintegration was calculated some years ago by Goeppert-Mayer¹ on the basis of the Fermi theory.^{2, 3} The result obtained was extremely small, corresponding to a lifetime of the order of 10^{25} years in the case of two isobars whose masses differ by 0.002

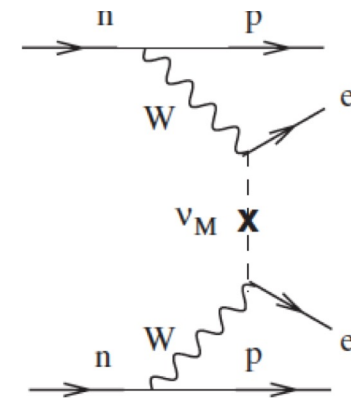
IV. CONCLUSION

We have seen that the phenomenon of double beta-disintegration is one for which there is a decided difference between the results of the Majorana theory and those of the older theory of the neutrino. According to the older theory it seemed certain that double beta-disintegration could never be capable of observation because of its extremely minute probability, but the Majorana theory indicates that this is by no means necessarily the case. Indeed, if the inter-



$$\langle m_{\beta\beta} \rangle = \left| \sum_k m_k U_{ek}^2 \right|$$

$$T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta})^2 [M^{0\nu}(0^+)]^2 \left(\frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$



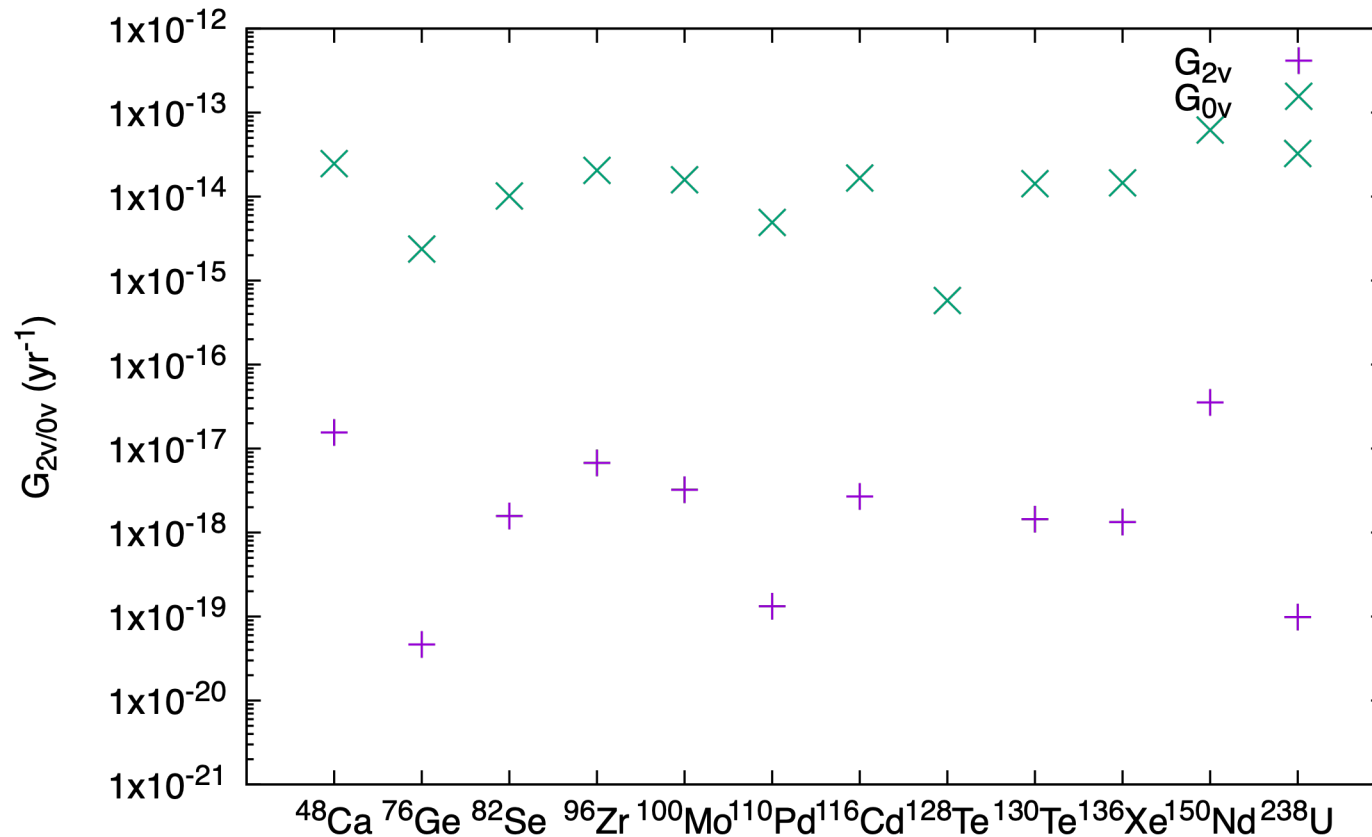
$$\nu_e(x) = \sum U_{ei} \nu_i(x)$$

$$|\nu_e\rangle = \sum_i U_{ei}^* |\nu_i\rangle$$

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DBD Phase Space Factors



$$T_{1/2}^{-1}(2\nu) = g_A^4 G_{2\nu} |M^{2\nu}|^2$$

$$T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 \left(\frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$

First Searches of $0\nu\beta\beta$

Phys. Rev 75, 323 (1949) Letter to the Editor

A Measurement of the Half-Life of Double Beta-Decay from $_{50}\text{Sn}^{124}$ *

E. L. FIREMAN

Department of Physics, Princeton University, Princeton, New Jersey
November 29, 1948

TABLE II. $L(A)$ gives the coincidences from specimen A between counters L , and $R(B)$ gives the coincidences from B between counters R . Holder 0° and 180° are the two holder positions. Positions 1 and 2 are positions for the specimens in the holder.

Pos. 1	Holder 0°	$L(A)$	$R(B)$ ^u
	Coin. counts/hr.	16.4 ± 0.3	14.3 ± 0.3
Pos. 2	Holder 180°	$L(B)$	$R(A)$
	Coin. counts/hr.	14.4 ± 0.3	15.9 ± 0.3
Pos. 2	Holder 0°	$L(B)$	$R(A)$
	Coin. counts/hr.	14.6 ± 0.3	16.4 ± 0.3
Pos. 2	Holder 180°	$L(A)$	$R(B)$
	Coin. counts/hr.	16.4 ± 0.3	13.9 ± 0.3

A detailed report of this work is being prepared for publication in the Physical Review.

The author is grateful to Professor R. Sherr for accepting the supervision of this research and to Professor E. P. Wigner for many profitable discussions.

* This work is assisted by the Office of Naval Research.

** These isotopes were obtained from Oak Ridge.

¹Maria Goeppert-Mayer, Phys. Rev. **48**, 512 (1935).

²W. H. Furry, Phys. Rev. **56**, 1184 (1939).

PHYSICAL REVIEW

VOLUME 86, NUMBER 4

MAY 15, 1952

A Re-Investigation of the Double Beta-Decay from Sn^{124}

E. L. FIREMAN AND D. SCHWARZER

Brookhaven National Laboratory, Upton, New York

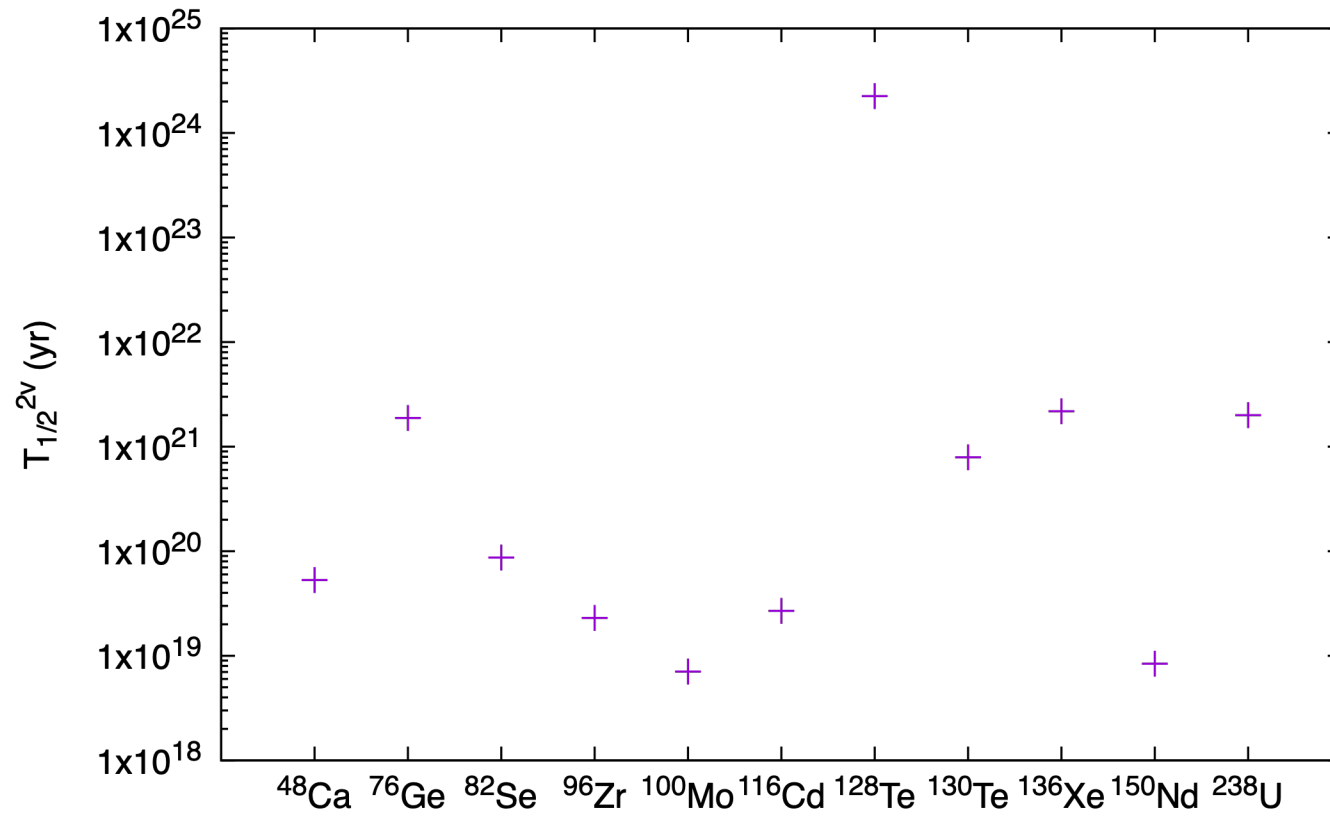
(Received February 5, 1952)

Radiations from natural tin and tin enriched with the 124 isotope are examined in a magnetic field with a helium filled cloud chamber that is triggered by internal counters. Only three pictures out of more than four thousand photographs are pictures of two electrons coming out of the same point in the tin and entering the counters, and even these may be pictures of multiply scattered electrons passing through the tin. However, one may set a lower limit to the half-life of double beta-decay from Sn^{124} as 10^{17} years. This is a decay rate less than one-tenth of a value previously reported by one of the authors.

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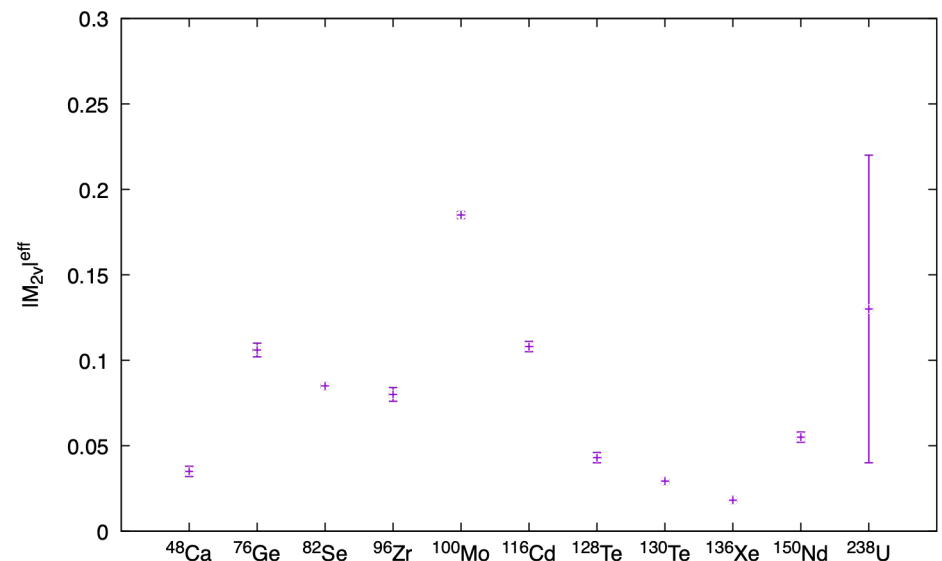
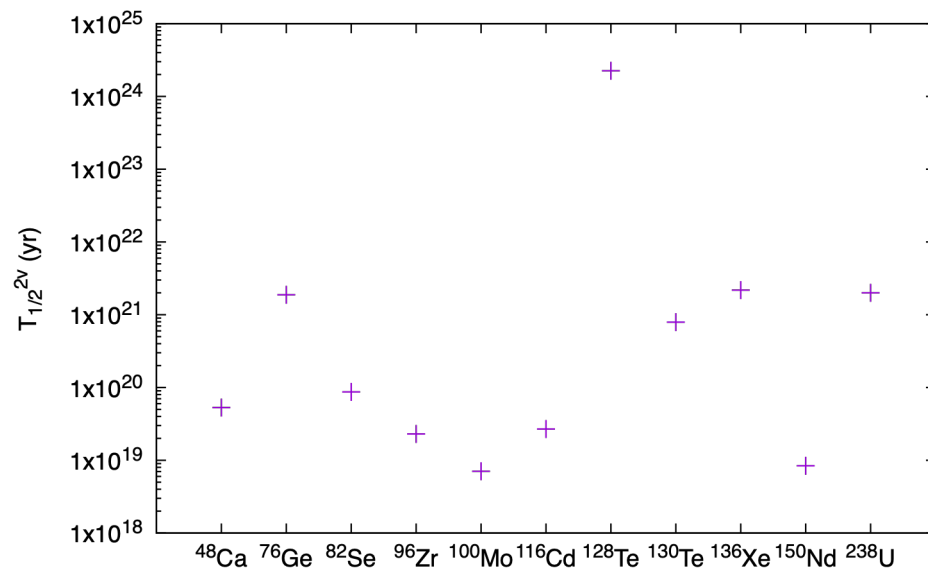
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$2\nu\beta\beta$ Half-Lives



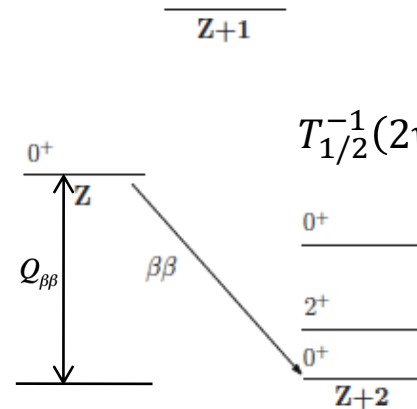
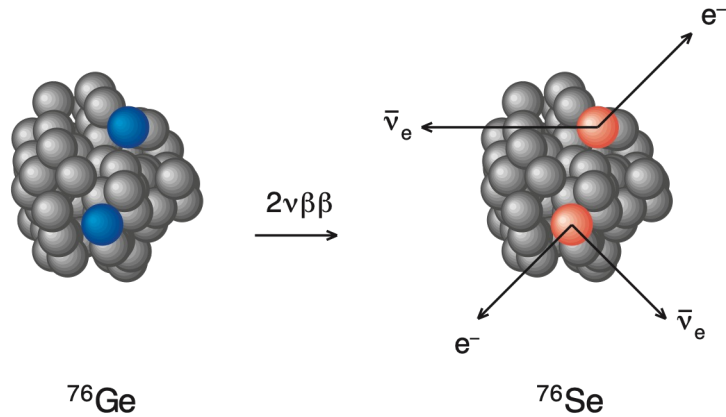
$2\nu\beta\beta$ Half-Lives and NME

$$\left[T_{1/2}^{2\nu}\right]^{-1} = G_{2\nu} \cdot \left[g_A^2 (m_e c^2 \cdot M_{2\nu})\right]^2 \equiv G_{2\nu} \left(M_{2\nu}^{eff}\right)^2$$

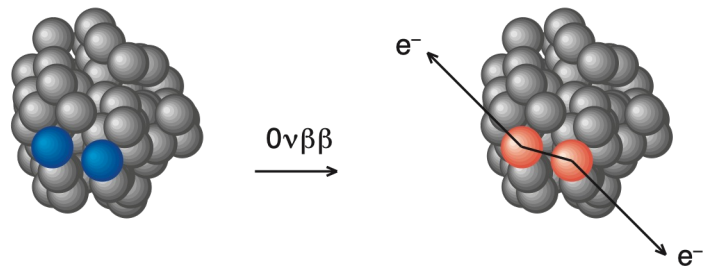
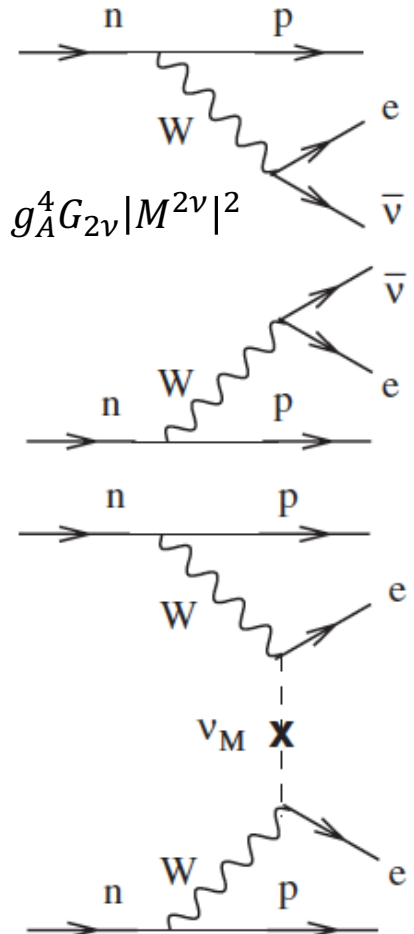


A. Barabash, Universe 6, 159 (2020)

Classical Double Beta Decay Problem



$$T_{1/2}^{-1}(2\nu) = g_A^4 G_{2\nu} |M^{2\nu}|^2$$



$$\langle m_{\beta\beta} \rangle = \left| \sum_k m_k U_{ek}^2 \right|$$

$$T_{1/2}^{-1}(0\nu) = G^{0\nu} (Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 \left(\frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$

$$\nu_e(x) = \sum_i U_{ei} \nu_i(x)$$

$$|\nu_e\rangle = \sum_i U_{ei}^* |\nu_i\rangle$$

$$|\nu_\alpha\rangle = \sum U_{\alpha i}^* |\nu_i\rangle$$

$$|\nu_i\rangle = \sum_\alpha U_{\alpha i} |\nu_\alpha\rangle$$

Neutrino Masses

PMNS – matrix

$$U = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{bmatrix} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$c_{12} \equiv \cos\theta_{12}, s_{12} = \sin\theta_{12}, \text{ etc}$$

- Tritium decay:



$$m_{\nu_e} = \sqrt{\sum_i |U_{ei}|^2 m_i^2} < 2.2 \text{ eV (Mainz exp.)}$$

KATRIN (to take data): goal $m_{\nu_e} < 0.3 \text{ eV}$

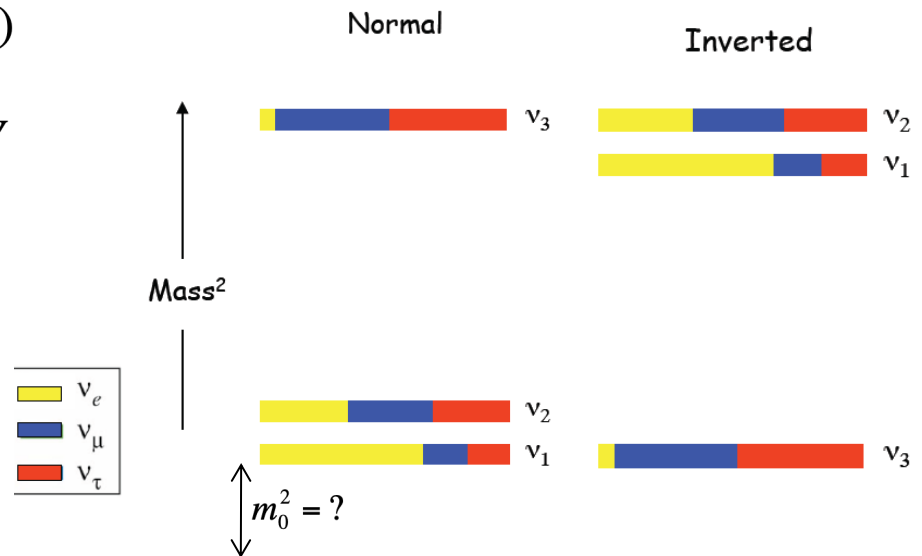
- Cosmology: CMB power spectrum, BAO, etc,

$$\sum_{i=1}^3 m_i < 0.23 \text{ eV}$$

Goal: 0.01 eV (5 – 10 y)

$$\Delta m_{21}^2 \approx 7.5 \times 10^{-5} \text{ eV}^2 (\text{solar})$$

$$|\Delta m_{32}^2| \approx 2.4 \times 10^{-3} \text{ eV}^2 (\text{atmospheric})$$

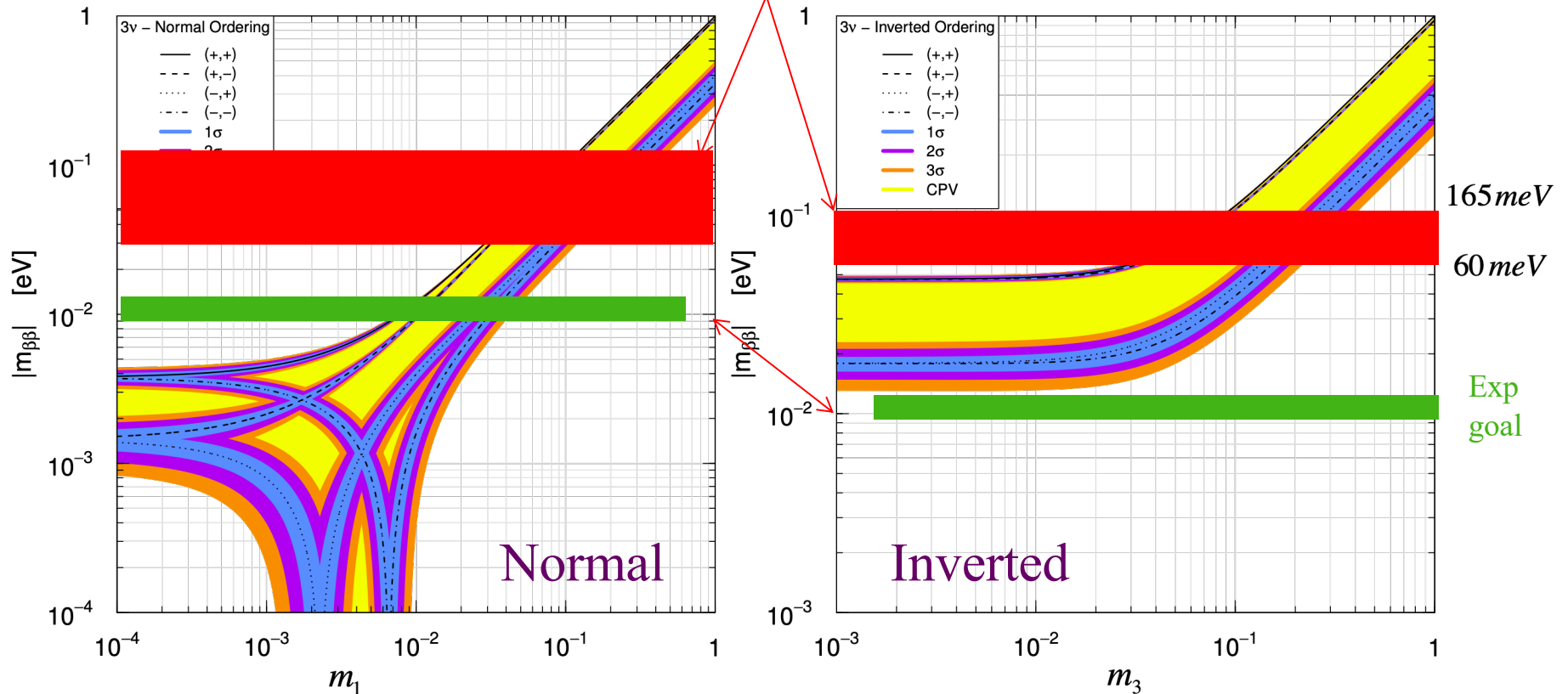


Two neutrino mass hierarchies

Neutrino $\beta\beta$ effective mass

arxiv:1507.08204

KamLAND – Zen, PRL 117, 082503 (2016): ^{136}Xe



$$|m_{\beta\beta}| = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

$$\phi_2 = \alpha_2 - \alpha_1 \quad \phi_3 = -\alpha_1 - 2\delta$$

$$\Leftrightarrow T_{1/2}^{-1}(0\nu) = G^{0\nu}(Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 (\eta_{0\nu})^2$$

$$\eta_{0\nu} = \frac{|m_{\beta\beta}|}{m_e}$$

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NSAC 2023 Long Range Plan

A NEW ERA OF DISCOVERY THE 2023 LONG RANGE PLAN FOR NUCLEAR SCIENCE



EXECUTIVE SUMMARY

ial pursuits. The workforce that enables discovery in nuclear science also makes breakthroughs in technologies with tremendous impact on the nation's

RECOMMENDATION 1

The highest priority of the nuclear science community is to capitalize on the extraordinary opportunities for scientific discovery made possible by the substantial and sustained investments of the United States. We must draw on the talents of all in the nation to achieve this goal.

This recommendation requires

- Increasing the research budget that advances the science program through support of theoretical and experimental research across the country, thereby expanding discovery potential, technological innovation, and workforce development to the benefit of society.
- Continuing effective operation of the national user facilities ATLAS, CEBAF, and FRIB, and completing the RHIC science program, pushing the frontiers of human knowledge.
- Raising the compensation of graduate researchers to levels commensurate with their cost of living—without contraction of the workforce—lowering barriers and expanding opportunities in STEM for all, and so boosting national competitiveness.
- Expanding policy and resources to ensure a safe and respectful environment for everyone, realizing the full potential of the US nuclear workforce.

RECOMMENDATION 2

As the highest priority for new experiment construction, we recommend that the United States lead an international consortium that will undertake a neutrinoless double beta decay campaign, featuring the expeditious construction of ton-scale experiments, using different isotopes and complementary techniques.

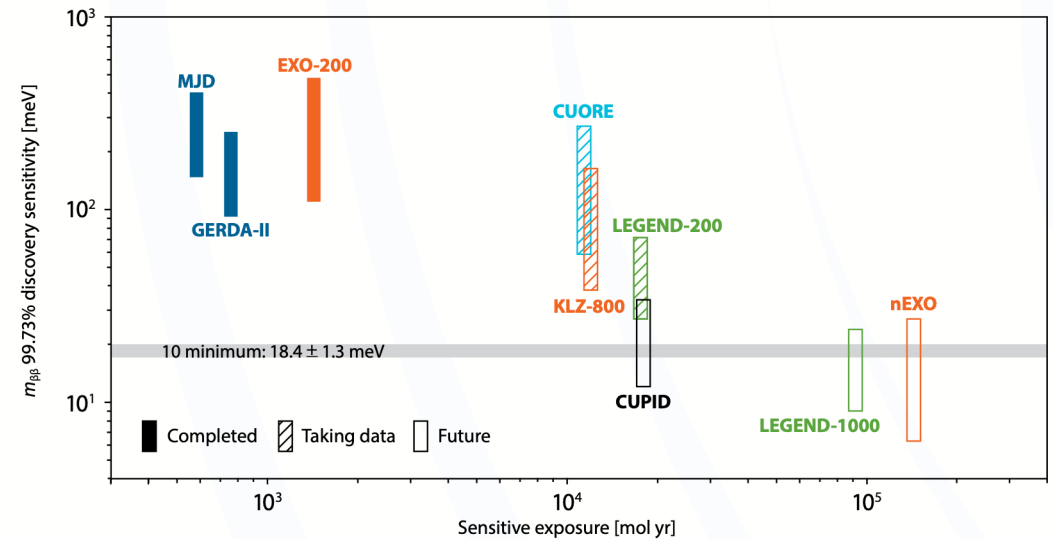
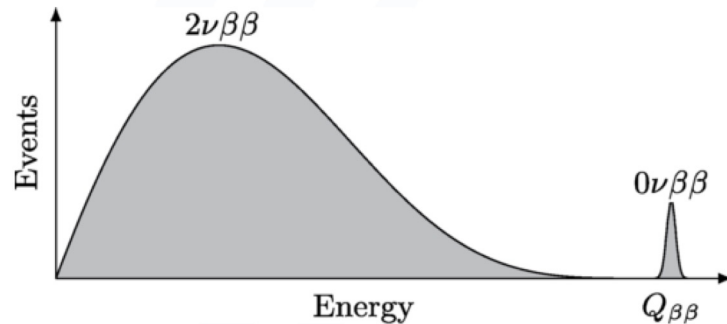
One of the most compelling mysteries in all of science is how matter came to dominate over antimatter in the universe. Neutrinoless double beta decay, a process that spontaneously creates matter, may hold the key to solving this puzzle. Observation of this rare nuclear process would unambiguously demonstrate that neutrinos are their own antiparticles and would reveal the origin and scale of neutrino mass. The nucleus provides the only laboratory through which this fundamental physics can be addressed.

The importance of the physics being addressed by neutrinoless double beta decay has resulted in worldwide excitement and has catalyzed the international cooperation essential to carrying out a successful campaign. An extraordinary discovery of this magnitude requires multiple experiments using different techniques for a select set of isotopes. Such measurements demand unprecedented sensitivity and present unique challenges. Since the 2015 Long Range Plan, the US-led CUPID, LEGEND, and nEXO international collaborations have made remarkable progress with three distinct technologies. An independent portfolio review committee has deemed these experiments ready to proceed now.

Neutrinoless double beta decay is sensitive to new physics spanning very different scales and physical mechanisms. The identification of the underlying physics will pose a grand challenge and opportunity for theoretical research. An enhanced theoretical effort is an integral component of the campaign and is essential for understanding the underlying physics of any signal.

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NSAC 2023 Long Range Plan



$$T_{1/2}^{-1}(0\nu) = G^{0\nu}(Q_{\beta\beta}) \left[M^{0\nu}(0^+) \right]^2 \left(\frac{\langle m_{\beta\beta} \rangle}{m_e} \right)^2$$

$$|m_{\beta\beta}| = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

Experiment	Isotope	Half-life limit (10^{26} years)	$m\beta\beta$ limit (meV)
MAJORANA	Germanium-76	0.83	113–269
GERDA	Germanium-76	1.8	79–180
EXO-200	Xenon-136	0.35	93–286
KamLAND-Zen	Xenon-136	2.3	36–156
CUORE	Tellurium-130	0.22	90–305

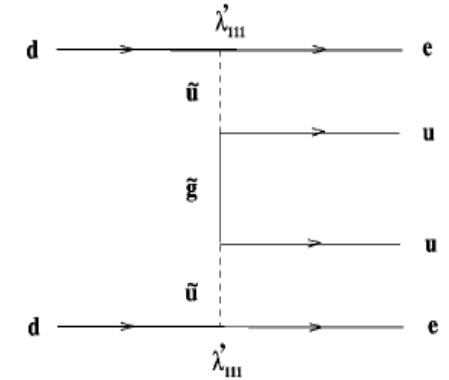
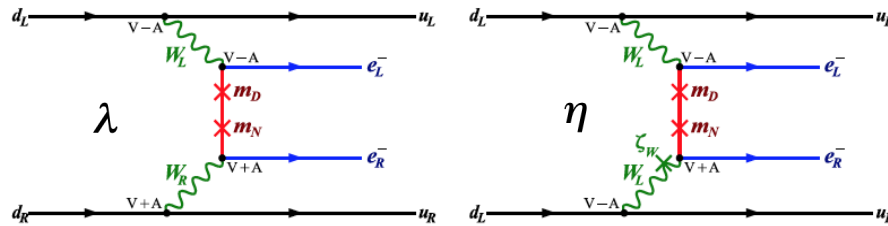
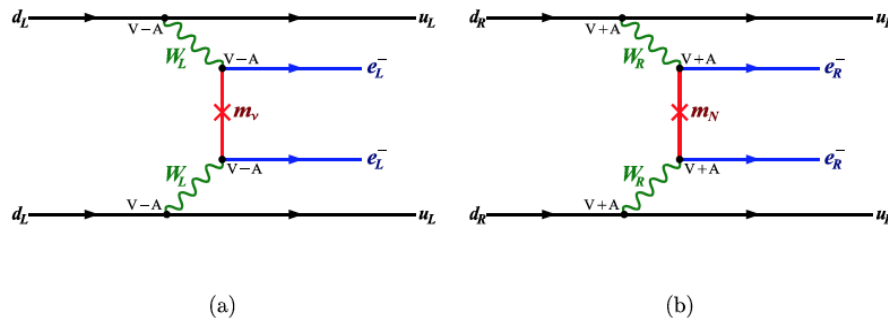
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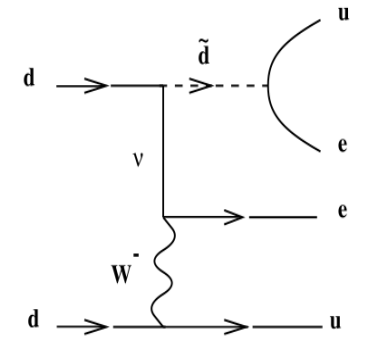
Other models: Left-Right Symmetric Model (LRSM) and SUSY R-parity violation

DAS *et al.*

PHYSICAL REVIEW D 86, 055006 (2012)



Gluino exchange



Squark
exchange

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{01} g_A^4 \left| \eta_{0\nu} M_{0\nu} + \left(\eta_{N_R}^L + \eta_{N_R}^R \right) M_{0N} + \eta_{\tilde{q}} M_{\tilde{q}} + \eta_{\lambda'} M_{\lambda'} + \eta_{\lambda} X_{\lambda} + \eta_{\eta} X_{\eta} \right|^2.$$

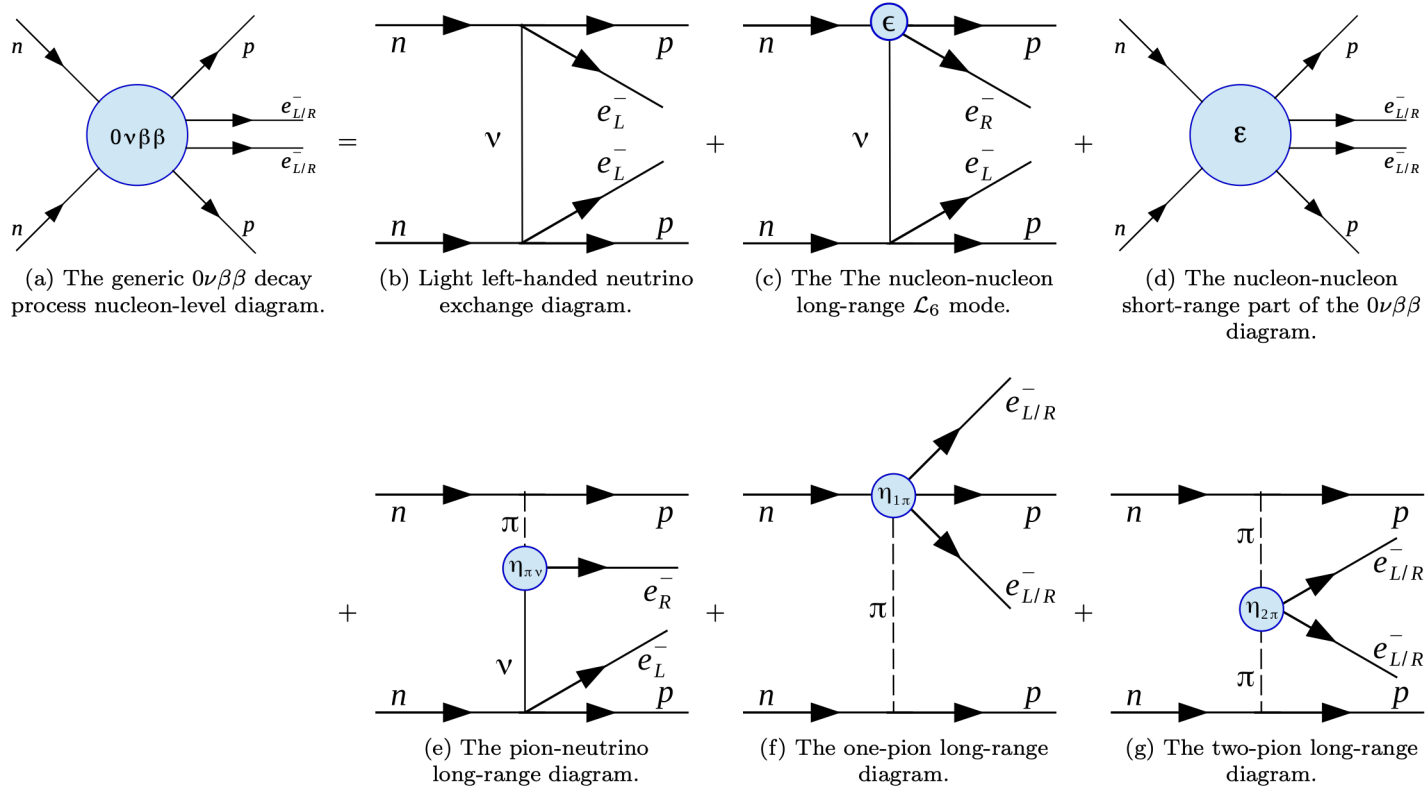
(e)

M. Horoi, A. Neacsu, PRD 93, 113014 (2016)

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Effective field theory after hadronization



$$\left[T_{1/2}^{0\nu}\right]^{-1} = g_A^4 \left[\sum_i |\mathcal{E}_i|^2 \mathcal{M}_i^2 + \text{Re} \left[\sum_{i \neq j} \mathcal{E}_i \mathcal{E}_j \mathcal{M}_{ij} \right] \right]$$

M. Horoi, A. Neacsu, PRC 98, 035502 (2018)

INT-26-2a, 7/13/2026

$$\mathcal{E}_{2-7} = \{ \epsilon_{V-A}^{V+A}, \epsilon_{V+A}^{V+A}, \epsilon_{S \pm P}^{S+P}, \epsilon_{TR}^{TR}, \eta_{\pi\nu} \}$$

$$\mathcal{E}_{8-15} = \{ \epsilon_1, \epsilon_2, \epsilon_3^{LLz(RRz)}, \epsilon_3^{LRz(RLz)}, \epsilon_4, \epsilon_6, \eta_{1\pi}, \eta_{2\pi} \}$$

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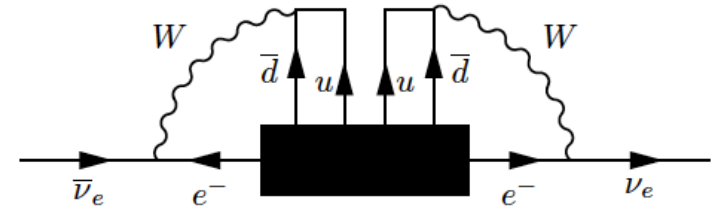
The Black Box Theorems

Black box I (electron neutrino)

J. Schechter and J.W.F Valle, PRD 25, 2951 (1982)

E. Takasugi, PLB 149, 372 (1984)

J.F. Nieves, PLB 145, 375 (1984)



$0\nu\beta\beta$ observed $\Leftrightarrow$ (i) Lepton number conservation is violated by 2 units.
 at some level (ii) Electron neutrinos are Majorana fermions (with $m > 0$).

However:

M. Duerr et al, JHEP 06 (2011) 91

$$(\delta m_{\nu_e})_{BB} \sim 10^{-24} \text{ eV} \ll \sqrt{|\Delta m_{32}^2|} \approx 0.05 \text{ eV}$$

Black box II (all flavors + oscillations)

M. Hirsch, S. Kovalenko, I. Schmidt, PLB 646, 106 (2006)

$0\nu\beta\beta$ observed $\Leftrightarrow$ (i) Lepton number conservation is violated by 2 units.
 at some level (ii) Neutrinos are Majorana fermions.

Regardless of the dominant $0\nu\beta\beta$ mechanism!

$$(iii) \quad \langle m_{\beta\beta} \rangle = \left| \sum_{k=1}^3 m_k U_{ek}^2 \right| = \left| c_{12}^2 c_{13}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right| > 0$$

NME: mass mechanism

Rev. Mod. Phys. 95, 025002 (2023)

Rep. Prog. Phys. **80** (2017) 046301

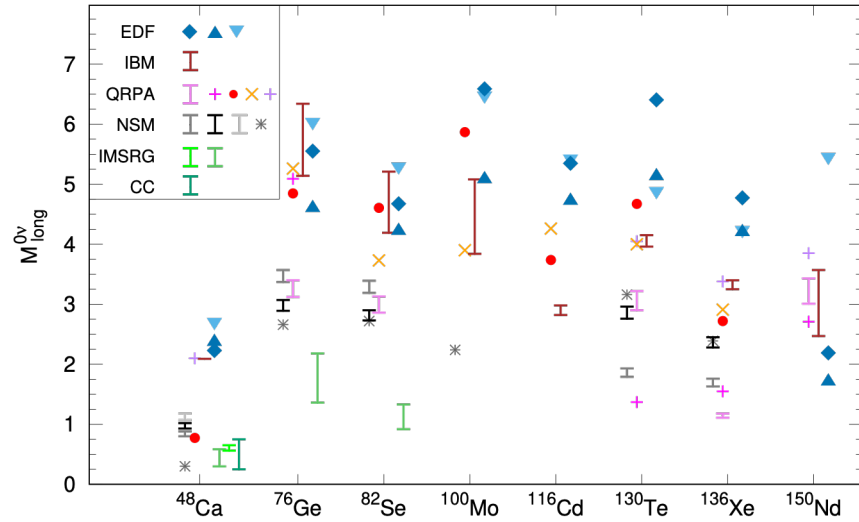


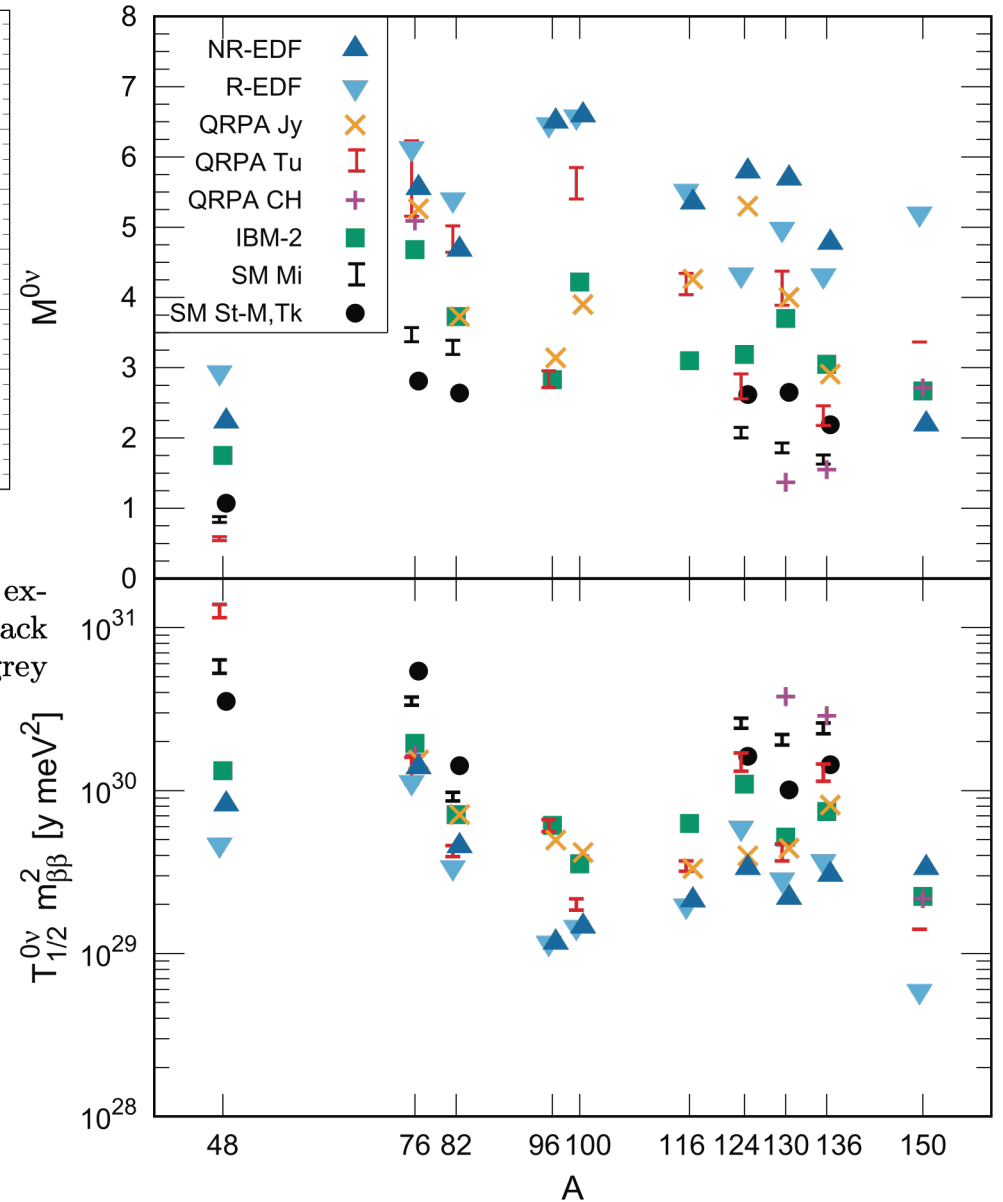
FIG. 9 Nuclear matrix elements $M^{0\nu}$ for light-neutrino exchange from different many-body methods. NSM: black (Menéndez, 2018), grey (Horoi and Neacsu, 2016b), light-grey

$$M_{0\nu} = M_{GT} + M_F + M_T + \dots$$

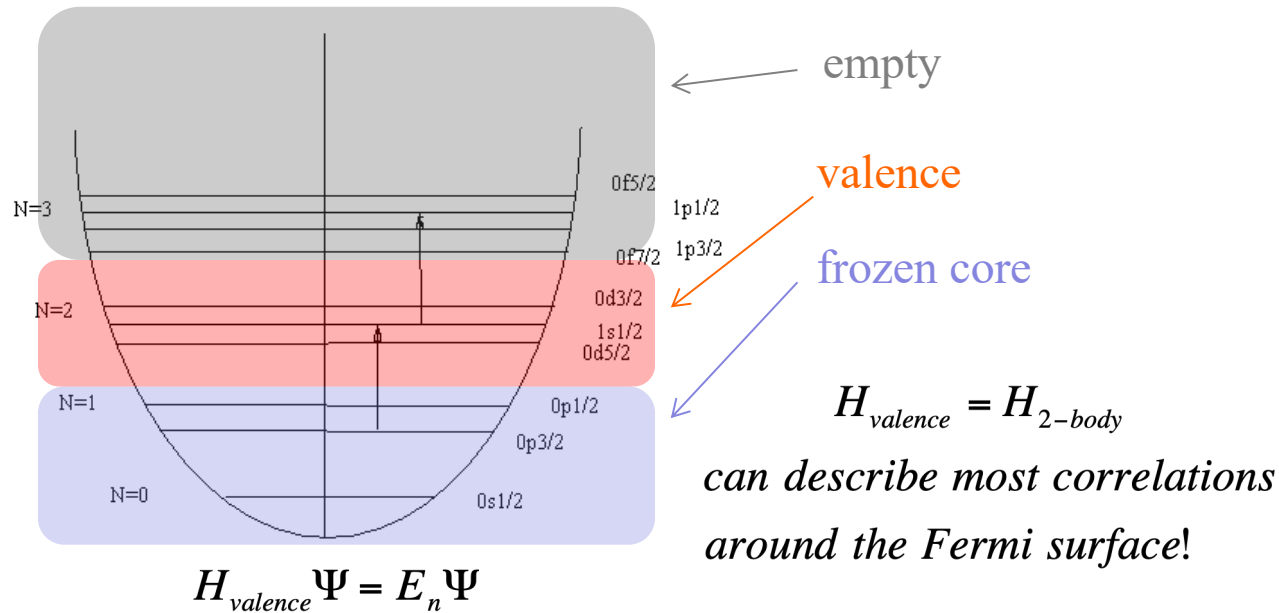
$$(T_{1/2}^{0\nu})^{-1} = G_{01} g_A^4 |M_{0\nu}|^2 |m_{\beta\beta}|^2$$

INT-26-2a, 7/13/2026

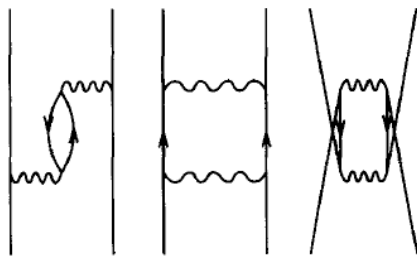
M



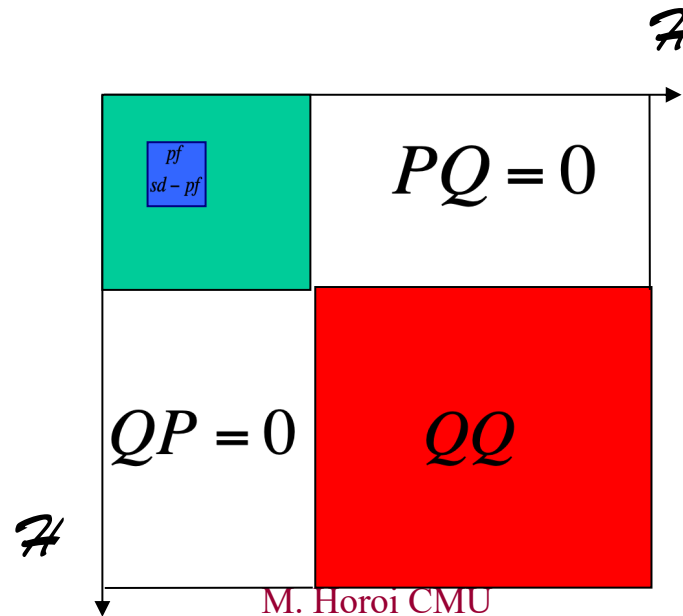
Shell Model Effective Hamiltonians



core polarization:
Phys.Rep. **261**, 125
(1995)

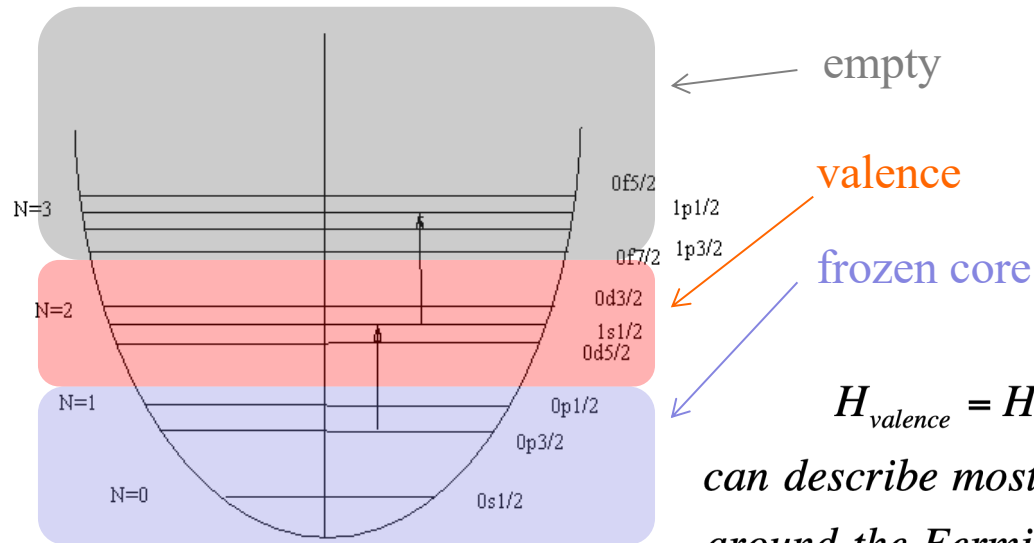


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Shell Model Effective Hamiltonians



$$H_{valence} = H_{2-body}$$

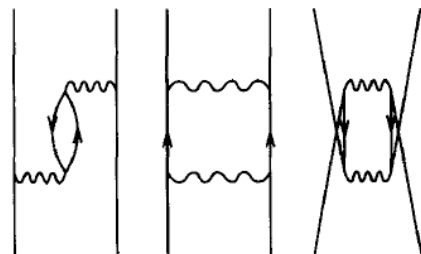
*can describe most correlations
around the Fermi surface!*

$$H_{valence} \Psi = E_n \Psi$$

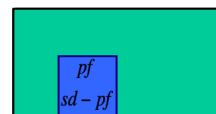
7



core polarization:
Phys.Rep. **261**, 125
(1995)

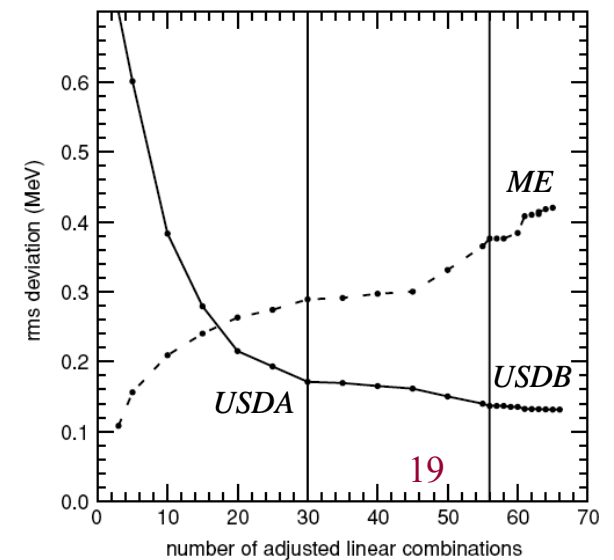
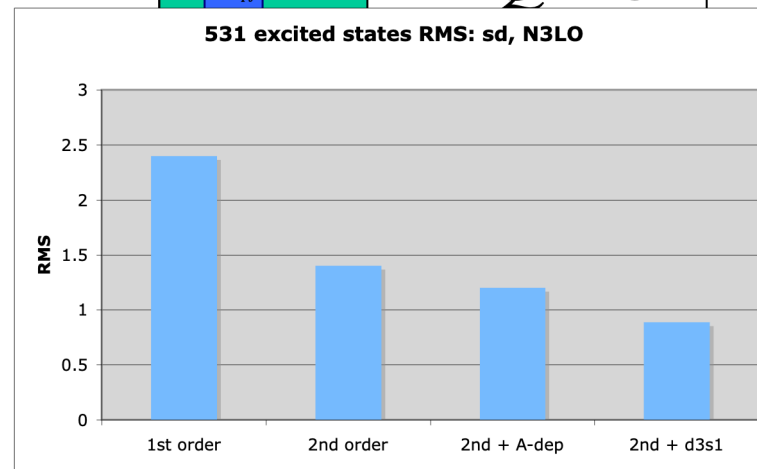


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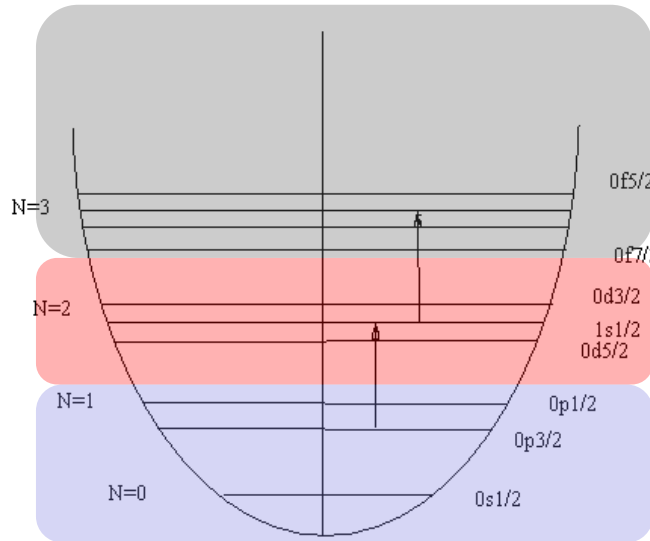


$$PQ = 0$$

► PRC 74, 34315 (2006), 78, 064302 (2008)



Shell Model Effective Hamiltonians



$$H_{valence} \Psi = E_n \Psi$$

$$g_A \sigma \tau \xrightarrow{\text{quenched}} g_A 0.77 \sigma \tau$$

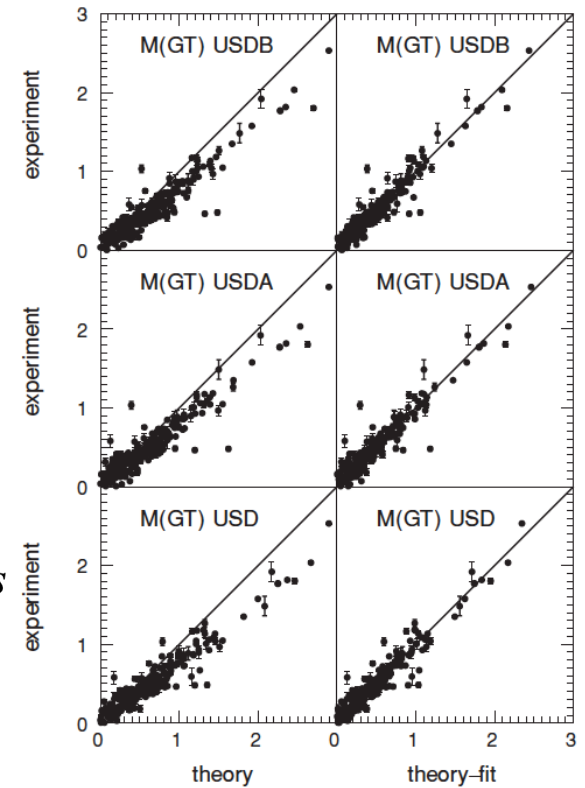
q

valence

frozen core

$$H_{valence} = H_{2-body}$$

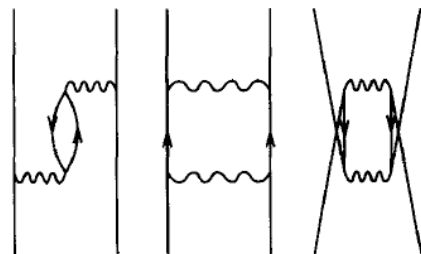
can describe most correlations
around the Fermi surface!



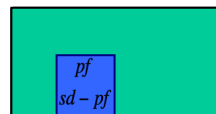
7



core polarization:
Phys.Rep. **261**, 125
(1995)

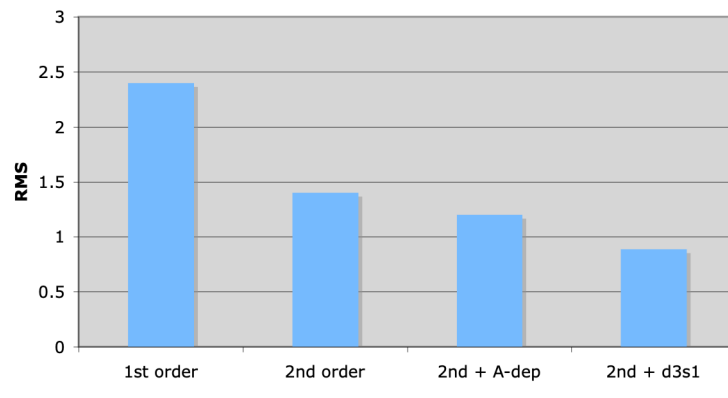


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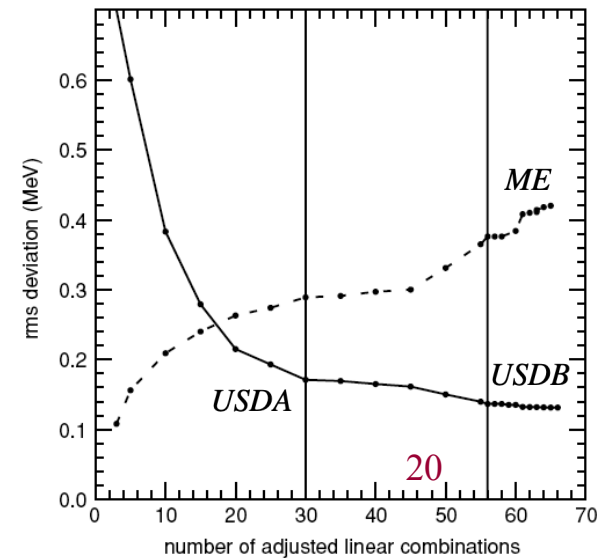


$$PQ = 0$$

531 excited states RMS: sd, N3LO

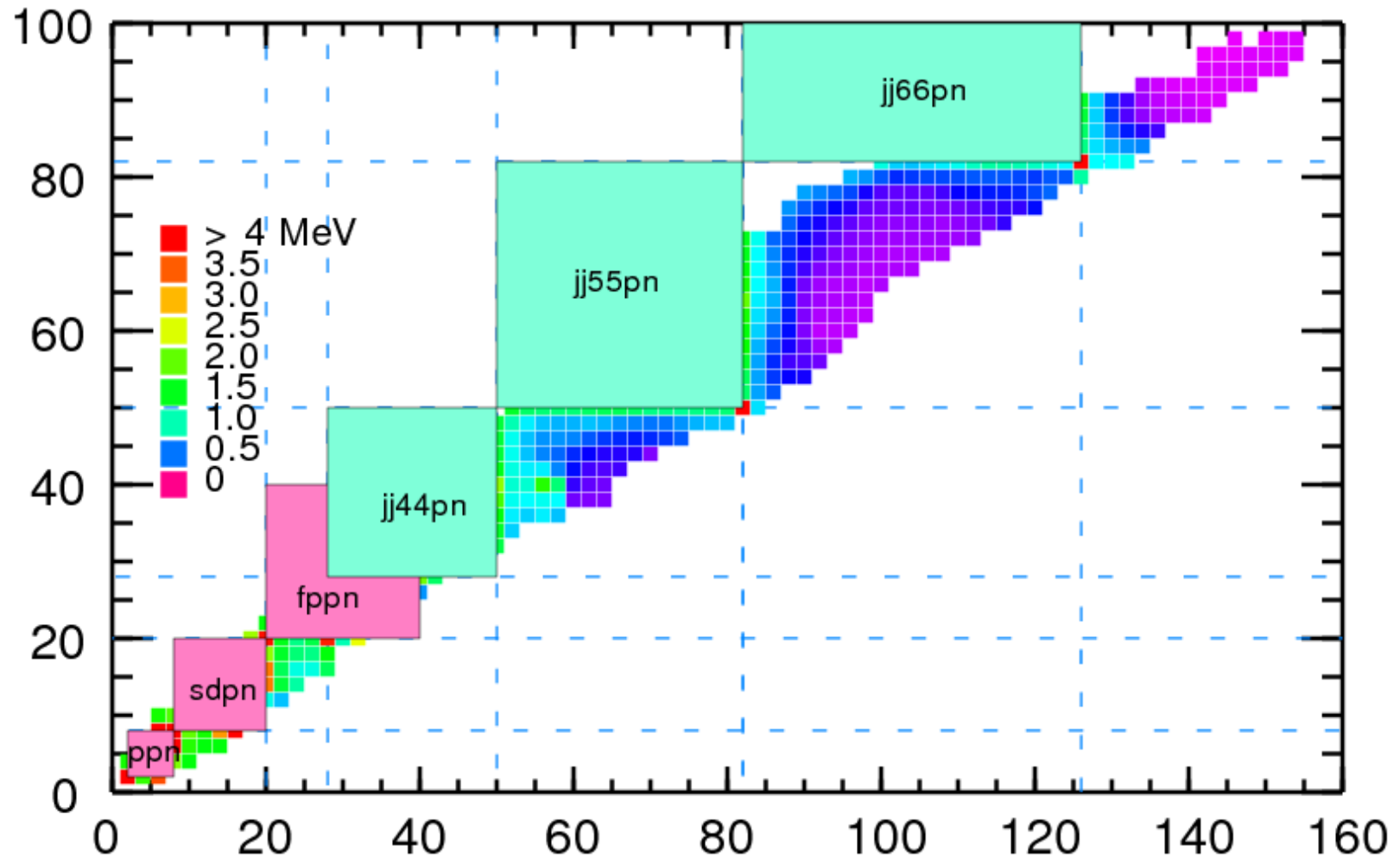


PRC 74, 34315 (2006), 78, 064302 (2008)



20

Shell Model Spaces



Statistical analysis of the NME

PHYSICAL REVIEW C **107**, 045501 (2023)

Predicting the neutrinoless double- β -decay matrix element of ^{136}Xe using a statistical approach

M. Horoi¹, A. Neacsu^{2,3} and S. Stoica²

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PHYSICAL REVIEW C **106**, 054302 (2022)

Statistical analysis for the neutrinoless double- β -decay matrix element of ^{48}Ca

M. Horoi¹, A. Neacsu^{2,3} and S. Stoica²

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IOP Publishing

Journal of Physics G: Nuclear and Particle Physics

J. Phys. G: Nucl. Part. Phys. **49** (2022) 120502 (30pp)

<https://doi.org/10.1088/1361-6471/aca03e>

Towards precise and accurate calculations of neutrinoless double-beta decay

V Cirigliano¹, Z Davoudi², J Engel^{3,*}, R J Furnstahl⁴,
G Hagen^{5,6}, U Heinz⁴, H Hergert⁷, M Horoi⁸,
C W Johnson⁹, A Lovato^{10,11,12}, E Mereghetti¹³,
W Nazarewicz^{14,*}, A Nicholson^{3,15}, T Papenbrock^{5,6},
S Pastore¹⁶, M Plumlee^{17,18}, D R Phillips^{19,*},
P E Shanahan²⁰, S R Stroberg¹⁰, F Viens²¹,
A Walker-Loud¹⁵, K A Wendt²² and S M Wild^{18,23}

INT-26-2a, 7/13/2026

Statistical Model: why ^{136}Xe ?

- One of the longest half-life limit measured, $\sim 3 \times 10^{26}$ years
- One of the potential isotopes chosen for the next “tonne experiment”
- Relatively low shell model dimensions in the $jj55$ ($0g_{7/2}, 2s, 1d, 0h_{11/2}$) model space

Statistical Model: Shell Model Effective Hamiltonians

- Isospin conserving effective Hamiltonians in $jj55$ model space ($0g_{7/2}$, $2s1d$, $0h_{11/2}$)
- 327 TBME of starting Hamiltonians $\pm 10\%$ uniformly random changes
- Three starting effective Hamiltonians:
 - $\Rightarrow$ SVD, PRC 86 044323 (2012)
 - $\Rightarrow$ GCN5082, PRC 82 064304 (2010)
 - $\Rightarrow$ $jj55t$, PRL 110 222502 (2013)

Statistical Model: Observables (24)

- 0nbb NME (1): M_{0v} , short-range correlator CD-Bonn
- 2nbb NME (1): M_{2v} , $q = 0.7$
- Energies 2+, 4+, 6+ (6): PE_{2+} , PE_{4+} , PE_{6+} , DE_{2+} , DE_{4+} , DE_{6+}
- B(GT) to 1+1 state (2): PGT, DGT, $q = 0.7$
- B(E2) $2^+ \rightarrow 0^+$ (2): PBE2, DBE2 $e_{eff}^p = 1.5$ $e_{eff}^n = 0.5$
- Proton occupation (8): POP_{g7}, POP_{s1}, POP_d, POP_{h11}, DOP_{g7}, DOP_{s1}, DOP_d, DOP_{h11}
- Neutron vacancies (4): DVNg₇, DVNs₁, DVNd, DVNh₁₁
- Muon Capture rate (0)
- B(M1) (0)

Overview

Observable	Data	Error	svd_s	gcn_s	$j5t_s$	μ_{svd}	σ_{svd}	μ_{gcn}	σ_{gcn}	μ_{j5t}	σ_{j5t}
$M_{0\nu}$	N/A	N/A	1.763	2.645	2.314	1.749	0.111	2.632	0.135	2.306	0.156
$M_{2\nu}$	0.018	0.001	0.025	0.069	0.060	0.022	0.003	0.061	0.007	0.052	0.007
PGT	0.150	0.021	0.163	0.545	0.512	0.141	0.059	0.457	0.105	0.333	0.220
$PBE2$	0.286	0.081	0.154	0.121	0.096	0.153	0.009	0.122	0.013	0.099	0.012
PE_{2+}	1.313	0.150	1.498	1.363	1.513	1.494	0.089	1.352	0.089	1.507	0.098
PE_{4+}	1.694	0.150	2.073	1.747	2.012	2.070	0.089	1.740	0.107	2.011	0.107
PE_{6+}	1.892	0.150	2.178	1.892	2.254	2.192	0.088	1.884	0.125	2.212	0.096
POP_{g7}	2.930	0.100	2.705	2.716	3.143	2.702	0.187	2.705	0.209	3.082	0.267
POP_{s1}	0.057	0.006	0.089	0.025	0.020	0.090	0.018	0.025	0.006	0.021	0.006
$POPh_{11}$	0.400	0.040	0.190	0.375	0.265	0.189	0.020	0.373	0.050	0.265	0.045
$POPd$	0.520	0.030	1.016	0.884	0.572	1.019	0.180	0.896	0.197	0.632	0.250
DGT	0.012	0.005	0.001	0.009	0.004	0.001	0.000	0.008	0.003	0.003	0.003
$DBE2$	0.413	0.011	0.342	0.194	0.158	0.337	0.023	0.195	0.026	0.163	0.028
DE_{2+}	0.819	0.150	0.662	0.842	0.917	0.660	0.067	0.836	0.056	0.919	0.049
DE_{4+}	1.867	0.150	1.389	1.873	2.113	1.403	0.131	1.861	0.116	2.087	0.082
DE_{6+}	2.207	0.150	2.157	2.196	2.502	2.171	0.151	2.197	0.090	2.507	0.117
DVN_{g7}	0.000	0.150	0.102	0.174	0.130	0.100	0.010	0.172	0.014	0.132	0.023
DVN_{s1}	0.080	0.020	0.271	0.251	0.415	0.286	0.117	0.255	0.058	0.407	0.110
DVN_{h11}	1.680	0.130	1.205	0.726	0.347	1.177	0.237	0.724	0.132	0.385	0.162
$DVNd$	0.240	0.050	0.423	0.850	1.108	0.437	0.132	0.850	0.118	1.076	0.158
DOP_{g7}	3.860	0.100	3.189	3.475	4.145	3.187	0.209	3.477	0.249	4.078	0.436
DOP_{s1}	0.200	0.020	0.263	0.083	0.049	0.264	0.047	0.084	0.020	0.052	0.017
DOP_{h11}	0.620	0.060	0.264	0.658	0.625	0.269	0.049	0.658	0.093	0.613	0.121
$DOPd$	1.290	0.080	2.285	1.785	1.181	2.280	0.227	1.781	0.265	1.258	0.447

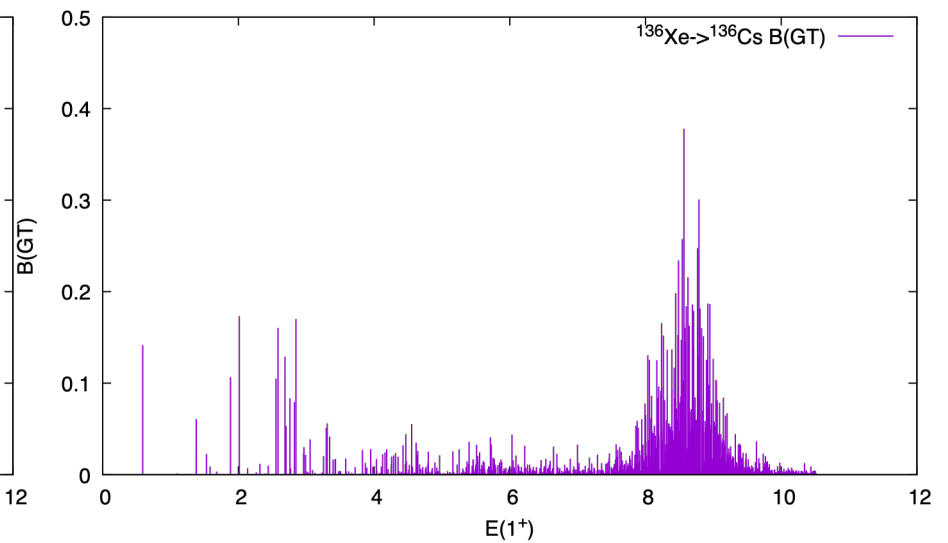
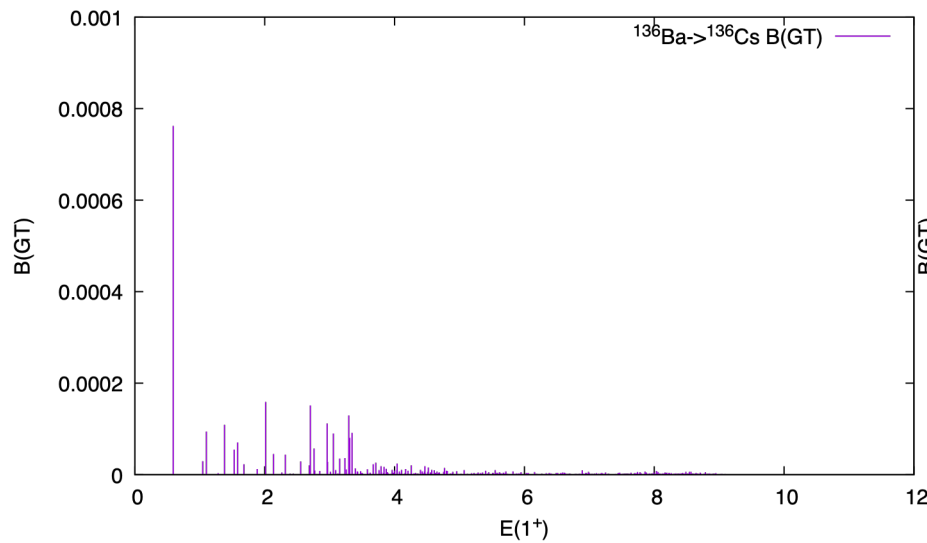
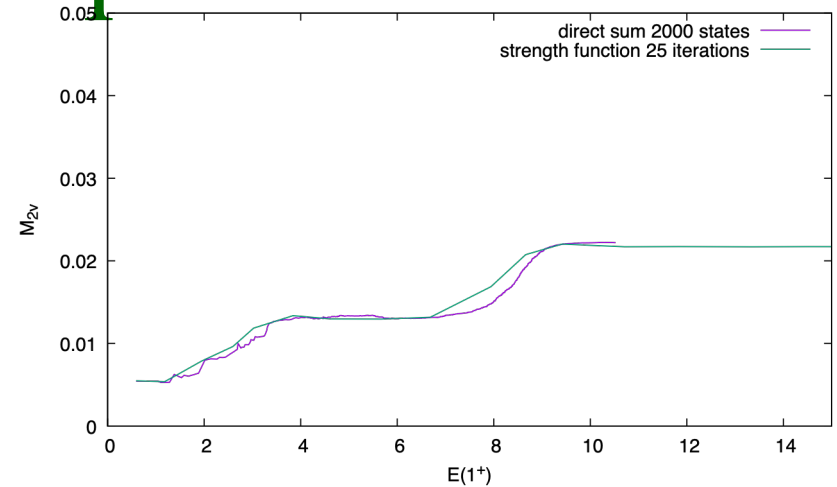
Table I. All relevant data and statistics for all selected observables.

The Late Bump in ^{136}Xe

$$M_{2v}(E_k) = \sum_{n=1}^k \frac{\langle \Psi_f | q\sigma\tau^- | 1_n^+ \rangle \langle 1_n^+ | q\sigma\tau^- | \Psi_i \rangle}{E(1_n^+) + \frac{Q_{\beta\beta}}{2} - Q_{\beta}}$$

^{136}Xe – SVD

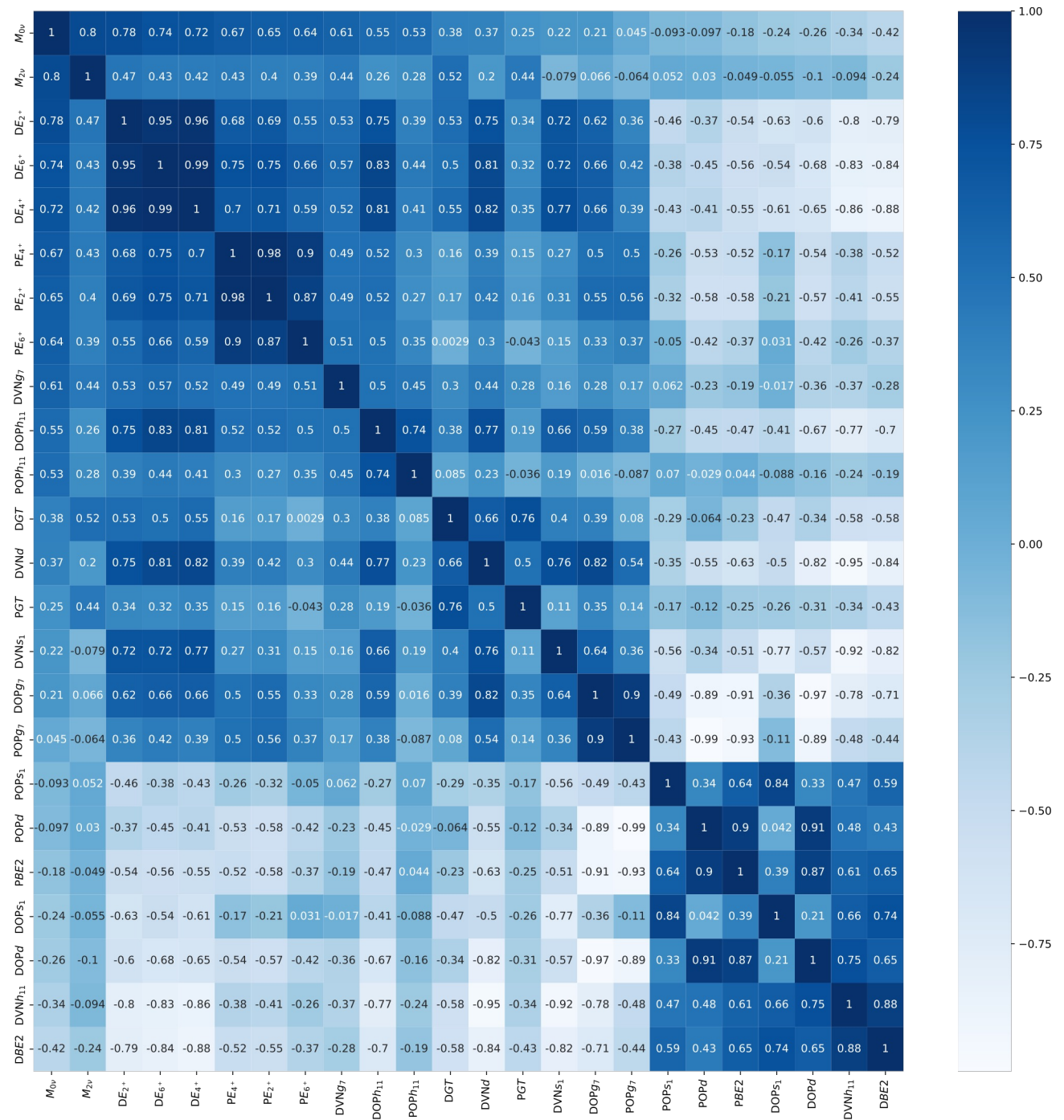
M. Horoi arXiv:2606.28224



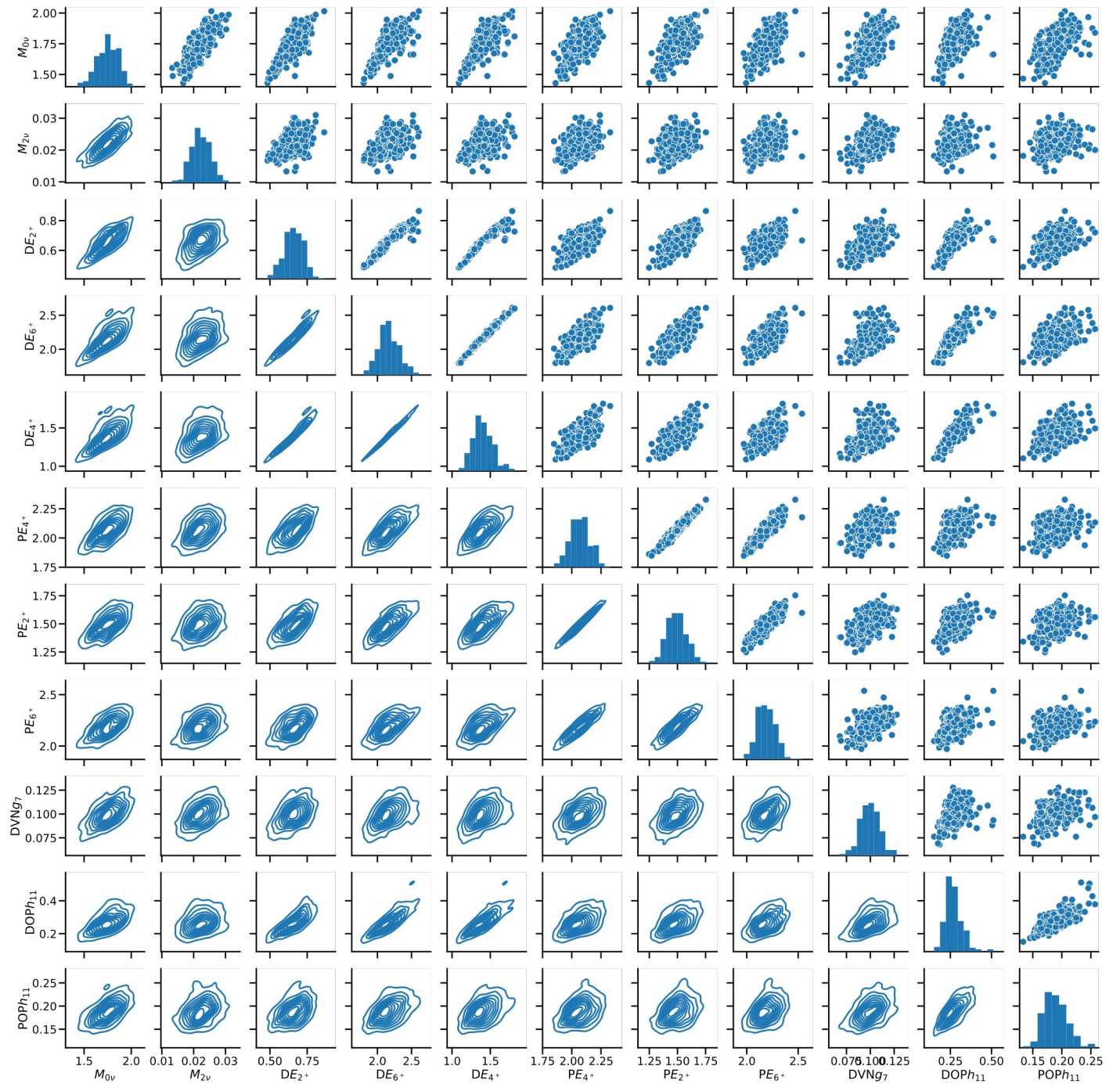
Statistical Model: Heat Map

samples=1,000

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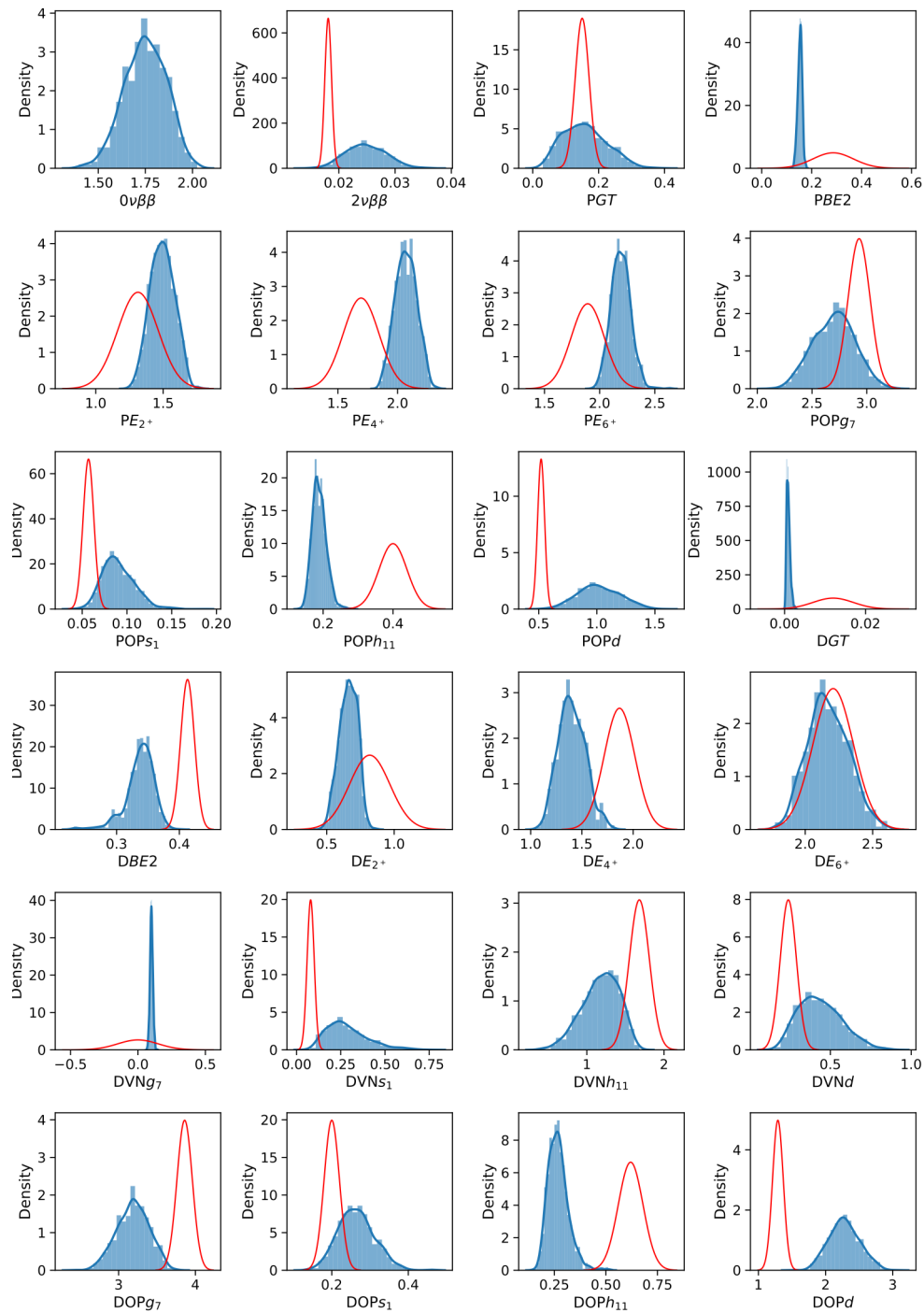


Statistical Model: Correlations Map



INT-26-2a, 7/13/202

Statistical Model: Density Distributions



INT-26-2a, 7/13/2026

Statistical Model: joint distribution

$$P(x = M_{0\nu}) = W_{svd}P_{svd}(x) + W_{gcn}P_{gcn}(x) + W_{j5t}P_{j5t}(x) ,$$

Statistical Model: Bayesian Model Averaging

$$p(x = M_{0\nu}|y_e, \sigma_e) = \sum_{k=svd}^{j5t} p(x = M_{0\nu}|y_e, \sigma_e, \mathcal{M}_k)p(\mathcal{M}_k|y_e, \sigma_e),$$

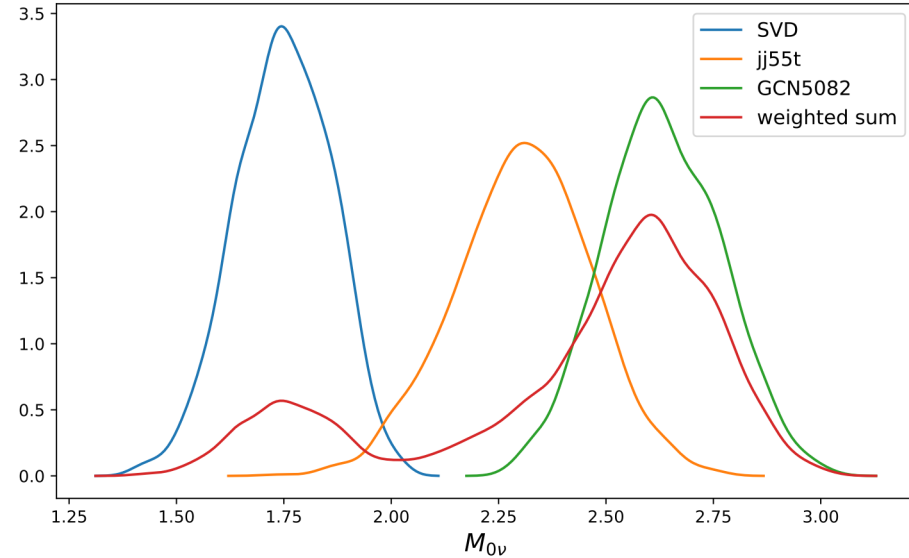
$$p(\mathcal{M}_k|y_e, \sigma_e) = \frac{p(y_e, \sigma_e|\mathcal{M}_k)\pi(\mathcal{M}_k)}{\sum_{k=svd}^{j5t} p(y_e, \sigma_e|\mathcal{M}_k)\pi(\mathcal{M}_k)}$$

$$p(y_e, \sigma_e|\mathcal{M}_k) = \int \prod_i^{N_{obs}} dy_i p_{y_e, \sigma_e}(y_i) \left[\int dq \prod_j^{N_{TBME}} d\theta_j p(y_i|\theta_j, q, \mathcal{M}_k)\pi(\theta_j, q|\mathcal{M}_k) \right],$$

Evidence integrals

$$p_{y_e, \sigma_e}(y_i) \propto \prod_i^{N_{obs}} \exp[-(y_i - y_{e_i})^2/(2\sigma_{e_i}^2)]$$

M. Horoi, Universe 10, 252 (2024)



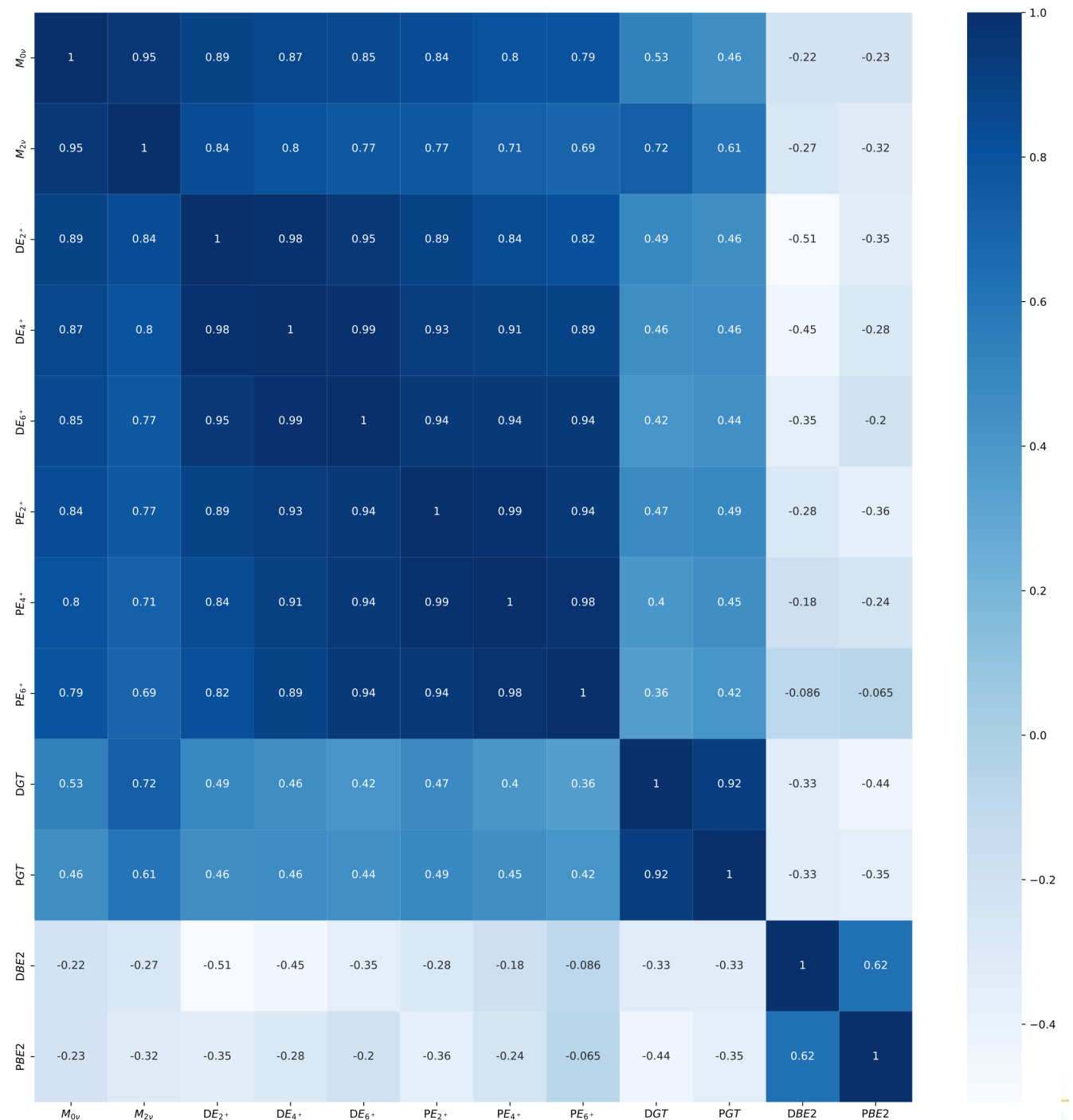
$$M_{0\nu} = 2.2 \pm 0.48$$

← W_H

^{76}Ge : Heat Map

samples=200

INT-26-2a, 7/13/2026



2νββ NME: Direct Sum vs Strength Function Approach

Direct sum:

$$M_{2\nu} = \sum_k \frac{\langle 0_f^+ | (\sigma\tau^-)^{eff} | 1_k^+ \rangle \langle 1_k^+ | (\sigma\tau^-)^{eff} | 0_i^+ \rangle}{E_k - E_0} \quad (\sigma\tau^-)^{eff} = q\tau^- \sigma$$

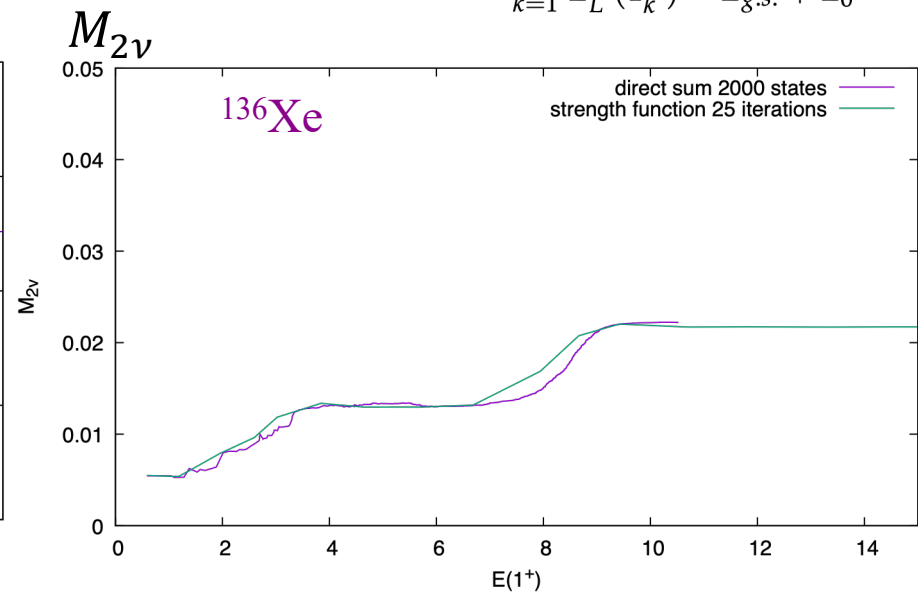
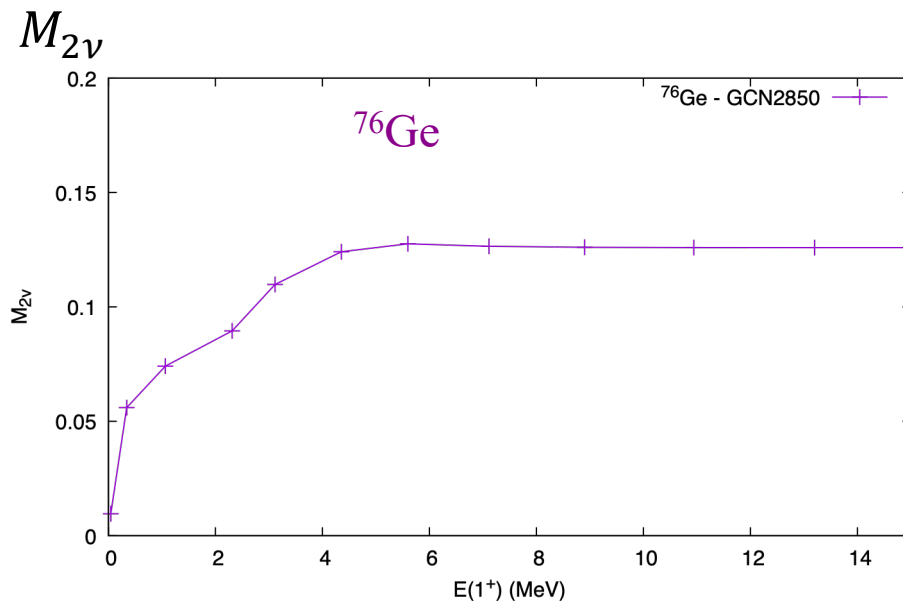
Strength function (Physics 2022, 4, 1135) :

$$|\sigma\tau^- 0_i^+ \rangle = c_- |dw_- \rangle \equiv c_- |L_1^- \rangle$$

$$M_{2\nu}(0^+) \approx 3c_+ c_- \sum_m g_m^- \langle dw_+ | L_m^- \rangle$$

$$|\sigma\tau^+ 0_f^+ \rangle = c_+ |dw_+ \rangle \equiv c_+ |L_1^+ \rangle$$

$$g_m^\pm = \sum_{k=1}^N \frac{V_{1k}^\pm V_{mk}^\pm}{E_L^N(1_k^+) - E_{g.s.} + E_0}$$

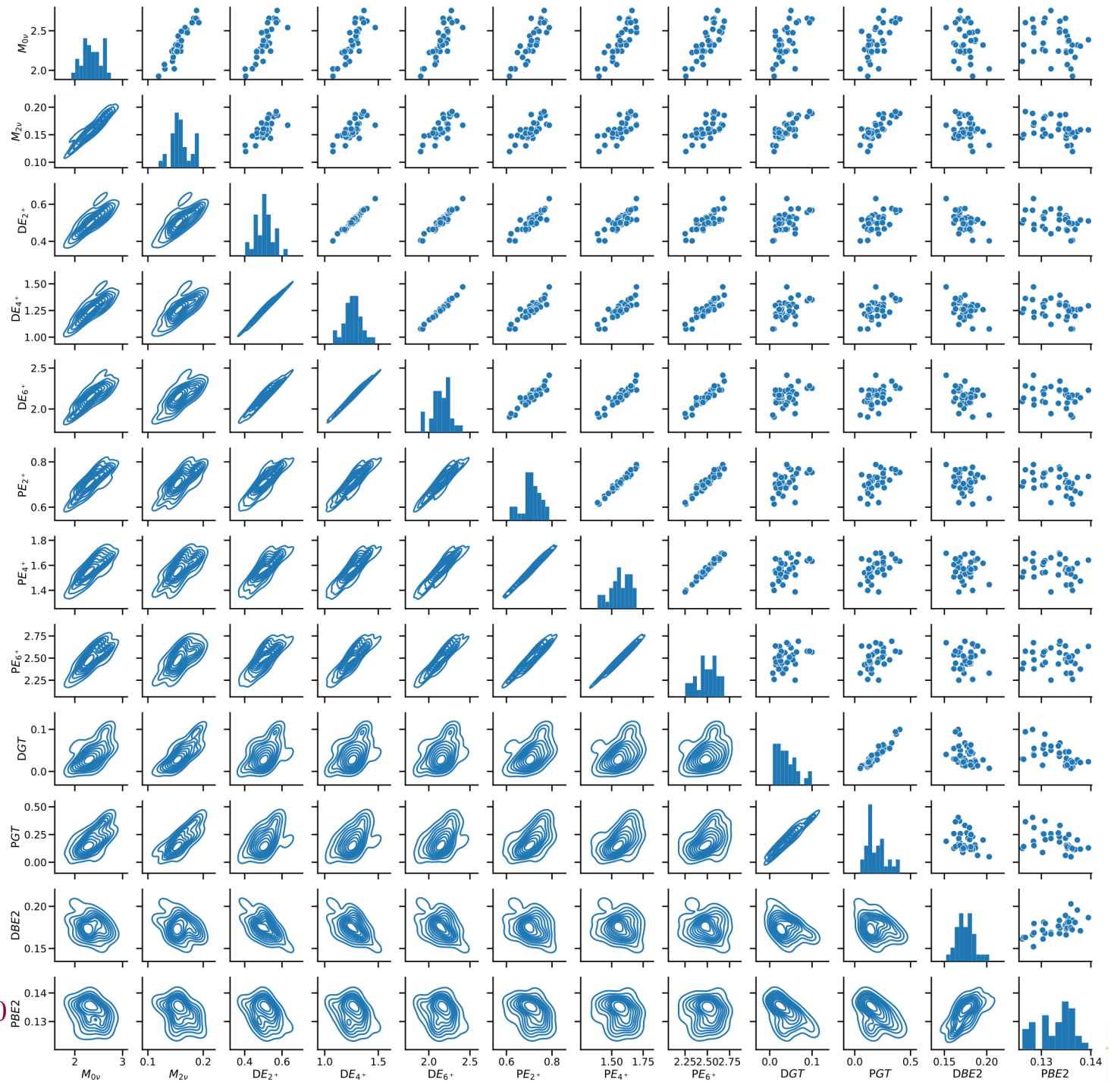


M. Horoi arXiv:2606.28224

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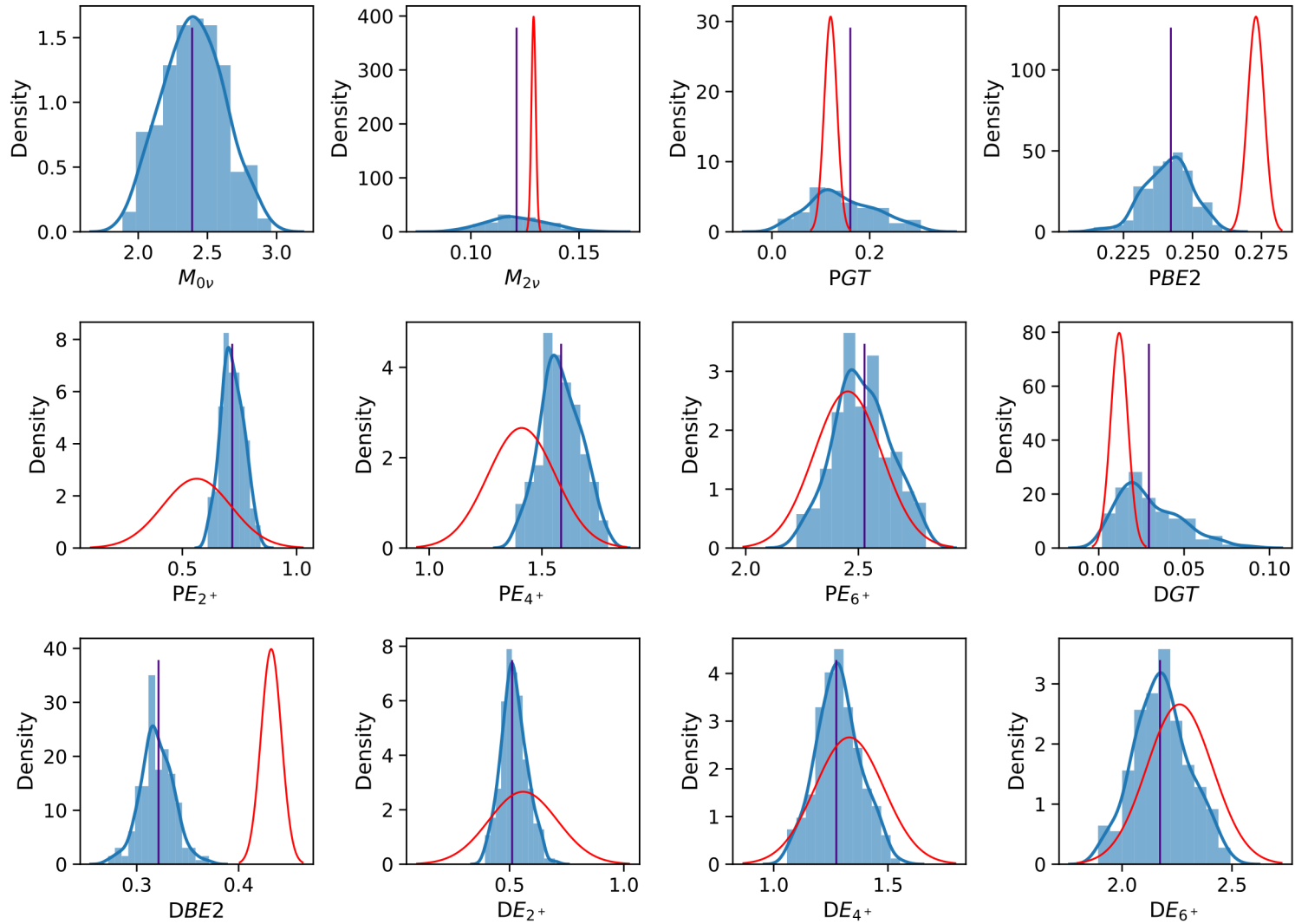
M. Horoi CMU

^{76}Ge : Correlations Map



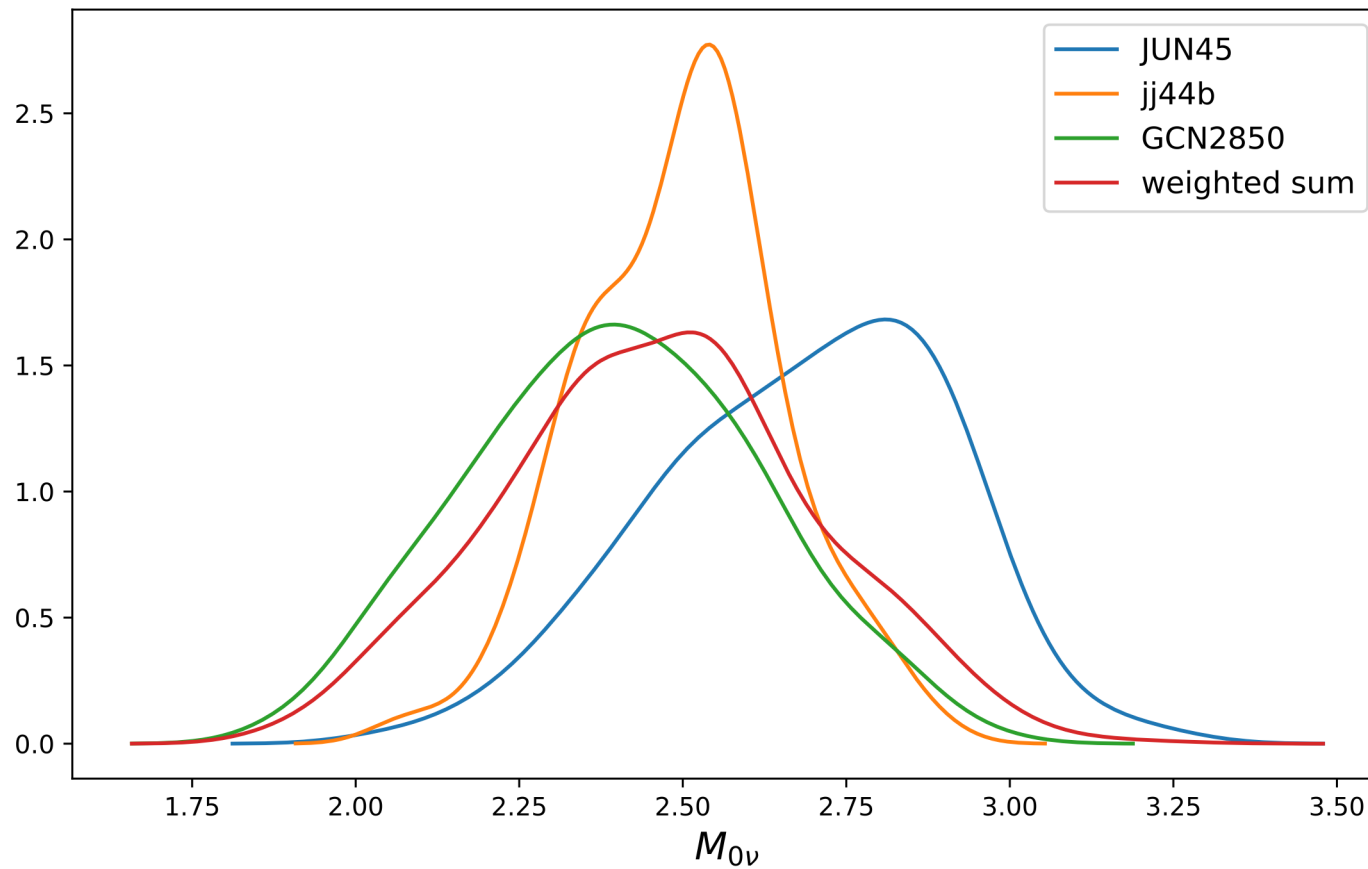
INT-26-2a, 7/13/20

^{76}Ge : Density Distributions



^{76}Ge : Joint Distributions

M. Horoi and A. Neacsu, arXiv:2605.21657



$$M_{0\nu} = 2.46 \pm 0.26$$

$$1.93-3.02 @90\%$$

Statistical Model: Results

- Shell model observables show robustness relative to small random changes of the effective Hamiltonian
- M_{0v} shows significant correlations with about half of the observables considered
- The outcome of our **Bayesian Averaging model** is a joint distribution of the ^{76}Ge M_{0v} with a range of 1.93 – 3.02 at 90% confidence level
- We propose a M_{0v} NME of 2.46 ± 0.26 for ^{76}Ge

Future work:

- Statistical investigation of the robustness of ^{76}Ge results
- Statistical investigation of the maximal interference effects