Bose gases out of equilibrium: from Turbulence to Coarsening

Martin Gazo (mg816@cam.ac.uk)

Compressible Turbulence: From Cold Atoms to Neutron Star Mergers 24 Jun 2025



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People involved in the work







Zoran Hadzibabic



Andrey Karailiev



Gevorg Martirosyan



Jiří Etrych



Simon Fisher



Seb Morris



Maciej Gałka



Tanish Satoor



Christoph Eigen



Panos Christodoulou



Julian Schmitt



Chris Ho

Experimental tools: ${}^{39}K$ Bose–Einstein Condensate in a box

2D Box trap (BEC 2)



3D Box trap (BEC3)



Tunable interactions



≈ 30–70k atoms

 $\omega_z/(2\pi) \approx 1 - 3 \text{ kHz}$ Trap depth: up to $k_B \times 800 \text{ nK}$ up to \approx 300k atoms Trap depth: up to k_B × 250 nK

Feshbach resonance (with zero crossing)

Magnetic Field, B (G)

Turbulence in Bose gases

1. Driven steady-state turbulence (2D gas)

Galka et.al., PRL **129**, 190402 (2022) Karailiev et. al., PRL **133**, 243402 (2024)



2. Free turbulence & universal coarsening (2D & 3D gas)

Gazo et.al., arXiv:2312.09248 (2023) Martirosyan et. al., arXiv: 2411.19948 (2024)



(Wave) turbulence



Wavenumber k (μm^{-1})

Big whorls have little whorls Which feed on their velocity, And little whorls have lesser whorls And so on to viscosity. Lewis Richardson

Here:

whirls viscosity → (dominantly) waves
 → trapdepth

Key signature: Power law $n_k \sim k^{-\gamma}$

(Vortex) turbulence *elsewhere*



Weak wave turbulence

S Dipole pair



Steady-state turbulence in 2D gas



Usually $\gamma = d$, here finite-size effects \rightarrow log-corrections

Emergence of the cascade



Key concepts:

Emergent isotropy
 Dynamic scaling of momentum distributions

Cascade front: $k_{cf} \sim t^{-\beta}$

More generally:

In its wake: $n_k \sim k^{-\gamma}$ $n_k \qquad \approx k_{cf}$ $n_k(k,t) = \left(\frac{t}{t_0}\right)^{\alpha} n_k \left[\left(\frac{t}{t_0}\right)^{\beta} k, t_0 \right]$

Here : $\alpha = \gamma \beta$ $\beta = 1/(\gamma - d - 2)$... also known in: KPZ, NTFP, phase-ordering

1) Isotropy of the cascade

Low *k*: PCA (in-situ)





-0.04 0.04



anisotropic acoustic cascade?

1) Isotropy of the cascade



2) Dynamic scaling – large k

Plot $k^{\gamma}n_k$ ($\gamma = 2.9$)

4

3

2

1.5

2

 $n_k k^\gamma$ (arb. units)



 $(t/t_0)^{\beta}k \ (\mu m^{-1})$

 $k \,(\mu m^{-1})$

Inverse wave cascade in 2D gas



Isotropic excitation $k_{\rm F} \approx 2 \ \mu {\rm m}^{-1}$ > $k_{\xi} = 0.8 \ \mu {\rm m}^{-1}$





Time

Inverse wave cascade in 2D gas



Look at: $k_{\xi} < k < k_F$



Dual cascade: Inverse & Direct Dissipation k_D finite \rightarrow direct cascade with Π_N Injection k_F finite \rightarrow inverse cascade with Π_E Experiment: $\gamma = 1.55(15)$ Theory: $\gamma = 4/3$ (ignoring all problems)

(Weakly) nonlocal cascade, reflected fluxes



Get still ~ $k^{-1.5}$ along the drive

Nonlocality of the cascade





Qualitatively similar

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How does a Bose–Einstein condensate form?

Typical situation: a rapid <u>quench</u> through a phase transition into the <u>ordered phase</u>



Preparing far-from-equilibrium states



Navon et. al, Nature (2016)

Chaotic driving @ zero interactions, but disorder important

Destroys condensate faster



Experiment: Martirosyan et.al., PRL, 132, 113401 2024 Theory: Zhang et.al., 2024

Bidirectional relaxation



Dynamic self-similar scaling (IR and UV):

$$n_k(k,t) = \left(\frac{t}{t_0}\right)^{\alpha} n_k \left[\left(\frac{t}{t_0}\right)^{\beta} k, t_0 \right]$$

Berges et.al., 101, 041603 (2008). Schmied et.al., 2019



(+ more recent ones...)

Dynamic scaling observed in experiments, but with exponents not fully understood

Experiment (2D)



Spectral particle density $\mathbf{N}_k \propto k^{d-1} \ n_k$

Experiment (2D)



Bidirectional relaxation

(particles to IR, energy to UV)

First, focus on the IR:

$$n_k(k,t) = \left(\frac{t}{t_0}\right)^{\alpha} n_k \left[\left(\frac{t}{t_0}\right)^{\beta} k, t_0 \right]$$
$$\alpha = d\beta = d/z = d/2 \quad \Rightarrow \alpha = 1$$

"decreasing level of confidence"

For condensate population n_0 :

$$n_0 = n_k (k = 0) \propto t^d$$

Condensate growth $n_0(t)$





the universal highway has its one clock

$$t_1 \neq t_2$$
$$t^* = t_1 - t_2$$
$$t_{\text{uni}} = t - t^*$$

 $n_{0} \propto (t - t^{*})^{\alpha}$ $t^{*} < 0 \quad t^{*} > 0$ $lgnoring t^{*}:$ $\frac{d \ln n_{0}}{d \ln t} = \frac{\alpha}{1 - t^{*}/t}$ $\log t$

Theorist:

prescaling (... and just waits)

Mazeliauskas et. al., **122**, 122301 (2019) Schmied et. al., **122**, 170404 (2019) Heller et. al., **132**, 071602 (2024)

Experimentalist:

gets wrong exponents (... and cannot wait)

A solution



Moreover, this reveals values of t^*

Collapse full IR n_k curves with theoretical exponents:



Exponent κ in the shape:

Vortices: $\kappa = 4$ Porod's law, phase-ordering kineticse.g. Bray, Adv. Phys. 51, 481 (2002)Waves: $\kappa = 3$ Resummed kinetic theoryChantesana et.al., PRA 99, 043620 (2019)
Rosenhaus, Falkovich, Arxiv: (2024)

Experiment: $\kappa = 2.9(2) \rightarrow$ wave-dominated dynamics!

UV dynamics: characteristic lengthscale



$$n_k(k,t) = \left(\frac{t}{t_0}\right)^{\alpha} n_k \left[\left(\frac{t}{t_0}\right)^{\beta} k, t_0 \right]$$

weak-wave free turbulence

 $\beta = -1/6$ $\alpha = (d+2)\beta$ But no special k givin a simple power law...

Peak of $\epsilon \sim k^3 n_k$



UV dynamics: dynamic scaling



Different *E* cannot collapse



Can we understand different *E*?

UV dynamics: relation to fluxes

 $\left| \Pi_arepsilon = -rac{eta}{t_{ ext{uni}}} k \, arepsilon_k \, .
ight.$

$$n_k(k,t) = \left(\frac{t}{t_0}\right)^{\alpha} n_k \left[\left(\frac{t}{t_0}\right)^{\beta} k, t_0 \right]$$

Energy spectral density $\epsilon_k \sim k^3 n_k$

Energy flux Π_{ε} : $\nabla \Pi_{\varepsilon} = \dot{\epsilon}_k$ (k-dependent energy flux)

 n_k reveals fluxes

 $\Pi_{\epsilon} \sim \epsilon_k \sim n_k$? EoS: $n_0 \sim \Pi_{\varepsilon}^{\kappa}$ with $\kappa < 1$ t_{uni} has to be different!



Estimate EoS exponent: $\kappa \approx 0.6$



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3D Bose gas

Key universal features also **work** in 3D:

1) Coarsening:

$$\alpha = d\beta = d/z = 3/2$$

$$\Rightarrow n_0^{2/3} \text{ is linear in time}$$

2) Universal shape:





Martirosyan et. al., arXiv: 2411.19948



Interaction strength dependence

In WWT: timescales $\propto g^2$

Coarsening: ?

 $g_{\rm eff} \sim p^2$ independent on bare g



Chantesana et.al., PRA 99, 043620 (2019) Rosenhaus, Falkovich, arXiv:2501.12451 (2025)



- 1. Non-scaling part (open symbols): ...*depends on g*
- 2. Universal NTFP part (solid symbols): ...*independent on bare g!*

Universal slope:



Effective interaction strength

Martirosyan et. al., arXiv: 2411.19948

Transition from the weakly interacting (perturbative) regime

So far: IR dynamics $k \leq 1/\xi$ (strongly-interacting regime)

Opposite regime also possible $k > 1/\xi$: coarsening \rightarrow weak-wave turbulence



Predictions for inverse cascade: ≈ 2.5 (not 7/3)



Thank you for you attention!

Martin Gazo mg816@cam.ac.uk











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