# Microscopic and empirical constraints of the nuclear equation of state

**Christian Drischler** (drischler@ohio.edu) INT-23-1a: Intersection of nuclear structure & high-energy nuclear collisions February 6, 2023







## Ab initio workflow (idealized)





CD & Bogner, Few Body Syst. 62, 109

#### Here: nuclear equation of state (EOS) energy per particle (and derived quantities)

$$\frac{E}{A}(n,\delta,T)$$

baryon density nneutron excess  $\delta$ temperature T



#### theory of strong interactions

QCD is nonperturbative at the low energies relevant for nuclear physics (cf. pQCD & LQCD)

CD, Haxton, McElvain, Mereghetti et al., PPNP 121, 103888

## Ab initio workflow (idealized)

uncertainty quantification



nuclear observables (structure, reactions, astrophysics, ...) many-body theory exact QMC, NCSM, ... approximate CC, IMSRG, MBPT, SCGF, ... phenomenological SM, DFT, ... renormalization group

(SRG, Okubo-Lee-Suzuki, ...)

chiral forces & currents (Weinberg, van Kolck, Kaiser, LENPIC, Idaho, ...)

> quantum chromodynamics (CalLat, HALQCD, NPLQCD, ...)

> > CD & Bogner, Few Body Syst. 62, 109

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#### computational framework

solves the (many-body) Schrödinger equation requires a nuclear potential as input

#### chiral effective field theory

provides microscopic interactions consistent with the symmetries of *low-energy* QCD

#### theory of strong interactions

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CD, Holt & Wellenhofer, Ann. Rev. NPS **71**, 403

#### <sup>15</sup> **Tremendous progress** <sup>10</sup> in CEFT, many-body theory, UQ & HPC

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CEFT enables *ab initio* calculations of **finite nuclei** & **nuclear matter** at  $T \ge 0$ and arbitrary proton fractions ( $n \le 2n_{sat}$ )

Computational & algorithmic advances in *all* many-body frameworks

#### **Many-Body Perturbation Theory**

- fully automated | GPU-accelerated
- full N<sup>3</sup>LO calculations

Quantum Monte Carlo with local CEFT interactions: Lonardoni, Tews, Gandolfi, Carlson, PRR 2, 022033(R)

Bayesian statistics allows for **rigorous UQ & propagation** in EFT-based calculations (use emulators!)

## At what density does CEFT ?

## **Rigorous UQ for nuclear matter**



CD, Furnstahl, Melendez, Phillips, PRL **125**, 202702

**Chiral Effective Field Theory** (nucleons & pions)

dominant approach for deriving *microscopic* interactions consistent with the symmetries of *low-energy* QCD

three- and four-*neutron* forces predicted through  $N^{3}LO$ 

Correlated EFT truncation errors from **order-by-order calculations** & physics-informed GP truncation error model

model checking is important ("trust but verify")

Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Kaiser, Krebs, Machleidt, Meißner, ...



Open-source software & tutorials (Jupyter): https://buqeye.github.io



Bayesian Uncertainty Quantification: Errors for our EFT

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An example: symmetric matter

$$y = \frac{E}{A}$$
,  $k = 4$  (N<sup>3</sup>LO)

Uncertainty bands depict 68% credibility regions

 $y = y_k + \delta y_k$ 









### More details? Recent review article

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# Chiral Effective Field Theory and the High-Density Nuclear Equation of State

#### **Annual Review of Nuclear and Particle Science**

Vol. 71:403-432 (Volume publication date September 2021) First published as a Review in Advance on July 6, 2021 https://doi.org/10.1146/annurev-nucl-102419-041903



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#### Keywords:

Chiral EFT | neutron stars | MBPT nuclear matter at zero and finite temperature Bayesian uncertainty quantification recent neutron star observations

see also in the same journal: James Lattimer, Annu. Rev. Nucl. Part. Sci. **71**, 433

see also Sorensen et al., arXiv:2301.13253

**Open Access** 

## **Neutron matter & symr**

## ric matter



saturation point: **fine-tuned cancellation** between the kinetic and interaction contributions (ideal testbed for chiral EFT)

empirical constraints provide important benchmarks of chiral interactions, esp. 3NF Annotations:  $(\lambda / \Lambda_{3N})$  in fm<sup>-1</sup> or  $(\Lambda)$  in MeV

CD, Hebeler et al., PRL 122, 042501; Hoppe, CD et al., PRC 100, 024318; Simonis, Stroberg et al., PRC 96, 014303; Ekström et al., PRC 97, 024332; Atkinson et al., PRC 102, 044333; and many more

#### Huth et al., PRC 103, 025803

neutron matter below saturation density is **well-constrained** by NN scattering phase shifts

## **Neutron matter & symr**

### ric matter





Hebeler et al., ApJ (2013) Tews *et al.*, PRL (2013) Lynn *et al.*, PRL (2016) 20Drischler *et al.*, PRL (2019) Drischler et al., GP-B (2020) Gezerlis, Carlson, PRC (2010) Unitary gas ( $\xi = 0.376$ )  $\begin{bmatrix} 15 \\ MeV \end{bmatrix}$ neutron matter **3N forces important** 0.050.150.20.1 $n \, [{\rm fm}^{-3}]$ 

CD, Holt, and Wellenhofer, Annu. Rev. Nucl. Part. Sci. 71, 403

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## **Nuclear symmetry energy**





## Nuclear symmetry energy (at saturation density)

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## Select empirical constraints from DFT (overview)

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#### **Empirical saturation box (2016):**

- based on 14 (out of 240) functionals that reproduce well selected nuclear properties
- often used to benchmark chiral interactions
- limited statistical meaning at best

#### **Significant progress in UQ for DFT:**

Schunck, O'Neal, Grosskopf, Lawrence, Wild, JPG: NP **47**, 074001 McDonnell, Schunck, Higdon, Sarich, Wild, Nazarewicz, PRL **114**, 122501 Neufcourt, Cao, Nazarewicz, Olsen, Viens, PRL **122**, 062502 Chen & Piekarewicz, PRC **90**, 044305; **and more** 

#### Recently: UQ driven by projection-based emulators (promising!)

Bonilla, Giuliani, Godbey, Lee, PRC **106**, 054322 Giuliani, Godbey, Bonilla, Viens, Piekarewicz, arXiv:2209.13039

#### **Empirical constraints are** *precise* **but** *not very*

*accurate* (systematic uncertainties are *difficult* to estimate)

# Unclear how to use these scattered DFT constraints to benchmark saturation properties of chiral interactions rigorously



Dutra *et al.*, PRC **85**, 035201 Kortelainen *et al.*, PRC **89**, 054314 Brown & Schwenk, PRC C **89**, 011307

## **Bayesian inference: empirical saturation point**





**Model assumption:** DFT samples are random draws from a bivariate normal distribution with *unknown* mean vector  $\mu$  and covariance matrix  $\Sigma$ 

$$\mathbf{y}^* = [n_0, E(n_0)/A] \sim \mathcal{N}(\boldsymbol{\mu}, \boldsymbol{\Sigma})$$



(marginalization)

(evaluates to a bivariate t-distribution)

## **Analysis: Saturation Box (2016)**

(preliminary)





## All DFT constraint: joint MC analysis

(preliminary)





## **Exploring the limits of chiral EFT**





CD, Holt *et al.*, ARNPS **71**, 403 exp. constraints compiled by B. Tsang

**experimental & observational constraints are important** for testing and improving chiral EFT interactions

**needed:** improved predictions with **novel NN+3N interactions** and robust uncertainty quantification See also Betty Tsang's talk (tomorrow) Nuclear structure constraints on EOS

See also Bill Lynch's talk (Thursday) The role of Nuclear structure in reaction experiments

## **Exploring the limits of chiral EFT**



CD, Holt *et al.*, ARNPS **71**, 403 exp. constraints compiled by B. Tsang

experimental & observational constraints are important for testing and improving chiral EFT interactions

**needed:** improved predictions with **novel NN+3N interactions** and robust uncertainty quantification derived **bounds on the neutron star radius** (and sound speed) assuming chiral EFT is valid up to a given critical density (here:  $2n_0$ ) could already be challenged by NICER  $R_{2.0} = (11.4 - 16.1) \text{ km}$  Riley *et al.*, AJL **918**, L27 Miller *et al.*, AJL **918**, L28

CD, Han, Lattimer et al., PRC 103, 045808

CD, Han, and Reddy, PRC 105, 035808

Han & Prakash, APJ **899**, 2 Alford *et al.*, JPG: NPP **46**, 114001

extend EFT EOS at  $n_m$  to linear EoS with finite discontinuity (softening)

#### continuous match sets upper bound

use **lower limit on**  $M_{max}$  from observation to adjust  $\Delta \varepsilon$  and constrain  $R_{min}$ 

## **Emulators: mining scattering & reaction data**

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Kohn VP: spurious *singularities* can be mitigated **Proof of principle:** fast & accurate emulation of scattering observables for parameter estimation **Goal:** improving next-generation optical models &

**Goal:** improving next-generation optical models & chiral interactions in the FRIB era König, Ek

Construct **reduced-order models** by removing superfluous information in high-fidelity models

Emulators enable applications *thought* to be prohibitively slow

For bound-state emulators, see Frame *et al.*, PRL **121**, 032501; König, Ekström *et al.*, PLB **810**, 135814; Ekström & Hagen, PRL **123**, 252501, and more

## Model (Order) Reduction for nuclear emulators

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**Emulators** (surrogate models)

data-driven (non-intrusive) Gaussian Processes, Artificial Neural Networks, etc.

**model-driven (intrusive)** reduced-order equations from high-fidelity models

J. A. Melendez,<sup>1,\*</sup> C. Drischler,<sup>2,†</sup> R. J. Furnstahl,<sup>1,‡</sup> A. J. Garcia,<sup>1,§</sup> and Xilin Zhang<sup>2,¶</sup> <sup>1</sup>Department of Physics, The Ohio State University, Columbus, OH 43210, USA <sup>2</sup>Facility for Rare Isotope Beams, Michigan State University, MI 48824, USA

#### Many pointers to the model reduction literature ("Guide"): J. Phys. G 49, 102001 (2022)

The field of model order reduction (MOR) is growing in importance due to its ability to extract the key insights from complex simulations while discarding computationally burdensome and superfluous information. We provide an overview of MOR methods for the creation of fast & accurate emulators of memory- and compute-intensive nuclear systems. As an example, we describe how "eigenvector continuation" is a special case of a much more general and well-studied MOR formalism for parameterized systems. We continue with an introduction to the Ritz and Galerkin projection methods that underpin many such emulators, while pointing to the relevant MOR theory and its successful applications along the way. We believe that this will open the door to broader applications in nuclear physics and facilitate communication with practitioners in other fields.

see also: CD & Zhang, Chap. 8, pp. 29–36, in Few Body Syst. **63**, 67 (collective pieces of the INT program 21–1b); for RBMs applied to DFT, see Melendez's GitHub; Bonilla, Giuliani *et al.*, PRC **106**, 054322; Anderson *et al.*, PRC **106**, L031302









## Model (Order) Reduction for nuclear emulators

#### frontiers (accepted, review-like article in press) high-fidelity space reduced space $|\tilde{\psi}\rangle = \sum$ **BUQEYE** Guide to Projection-Based Emulators in $\psi_2$ $|\psi_1|$ **Nuclear Physics** arXiv:2212.04912 $|\psi(\boldsymbol{\theta})$ C. Drischler,<sup>1,2,\*</sup> J. A. Melendez,<sup>3</sup> R. J. Furnstahl,<sup>3</sup> A. J. Garcia,<sup>3</sup> and Xilin Zhang<sup>2</sup> Parametric eigenvalue problem $H(\boldsymbol{\theta}_i) |\psi_i\rangle = E(\boldsymbol{\theta}_i) |\psi_i\rangle$

The BUQEYE collaboration (Bayesian Uncertainty Quantification: Errors in Your EFT) presents a pedagogical introduction to projection-based, reduced-order emulators for applications in low-

energy nuclear physics. The term emulator refers here to a fast surrogate model capable of reliably

approximating high-fidelity models. well-known in the nuclear physics cor emphasize the benefits of offline-on emulators for bound and scattering different model parameter sets. We a for nuclear physics, guided by the n here and more are available as intera adapt projection-based emulators for Keywords: emulators, reduced-order models, mod variational principles, Galerkin projection

ABSTRACT











Pedagogical Jupyter notebooks for learning Model (Order) Reduction: https://github.com/bugeye/frontiers-emulator-review



- Chiral EFT enables *ab initio* calculations of finite nuclei & nuclear matter at  $T \ge 0$  & arbitrary proton fractions ( $n \le 2n_{sat}$ ). Where does it break down and why?
- 2 B
  - Bayesian statistics allows for rigorous UQ & propagation in EFT-based calculations (in many cases facilitated by new emulators!)



Need for improved constraints of the nuclear matter EOS in the density regime  $1 \leq n/n_{sat} \leq 2$ . Many advances can be expected across multiple disciplines.



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Our preliminary analysis suggests for the empirical saturation point:  $n_0 \approx 0.157 \pm 0.009 \text{ fm}^{-3}$ , with  $E_0/A \approx -16.02 \pm 0.33 \text{ MeV}$  (95%, correlated!)

Many thanks to: R. Furnstahl A. Garcia P. Giuliani S. Han J. W. Holt J. Lattimer A. Lovell K. McElvain J. Melendez F. Nunes D. Phillips M. Prakash S. Reddy C. Wellenhofer X. Zhang T. Zhao

