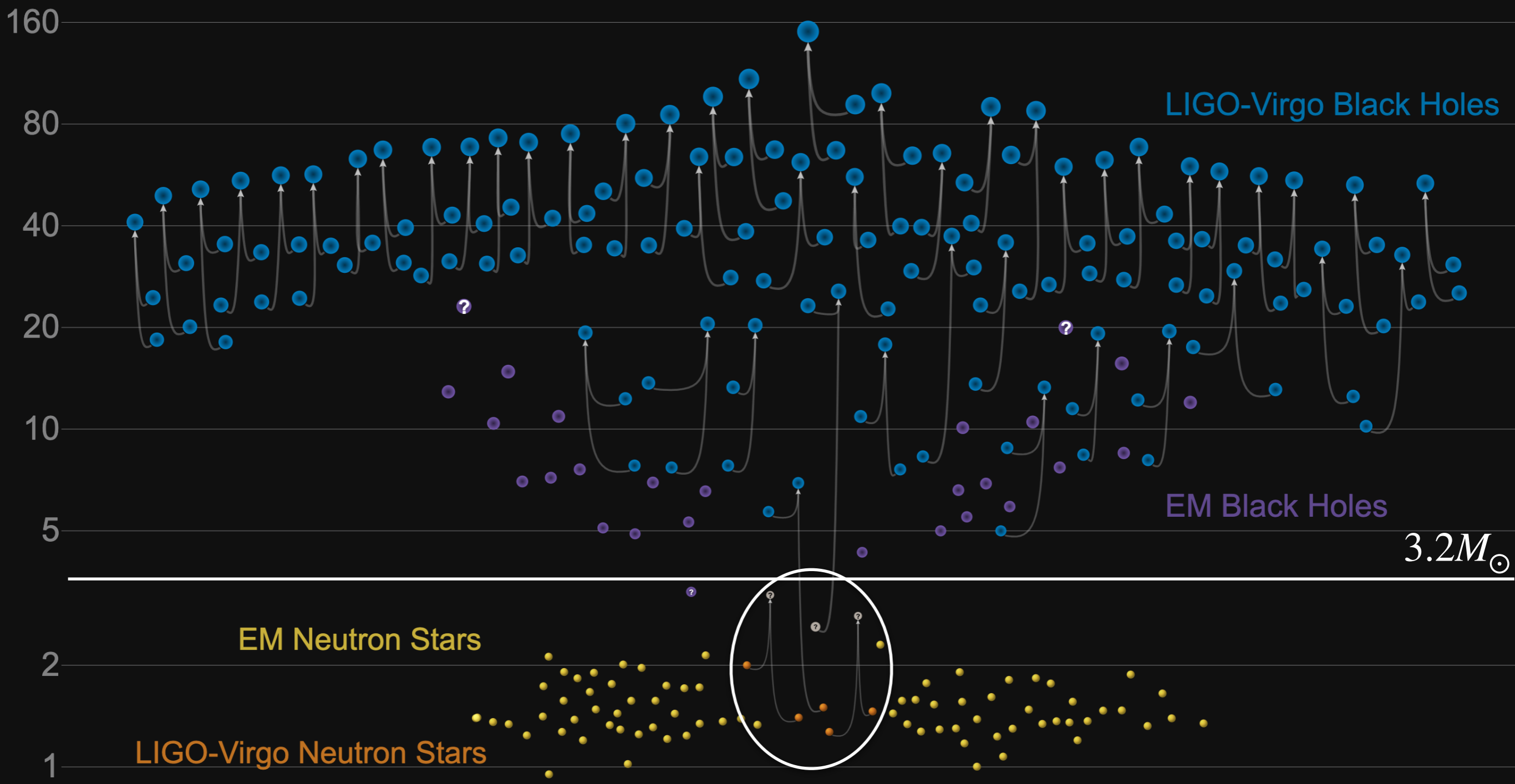
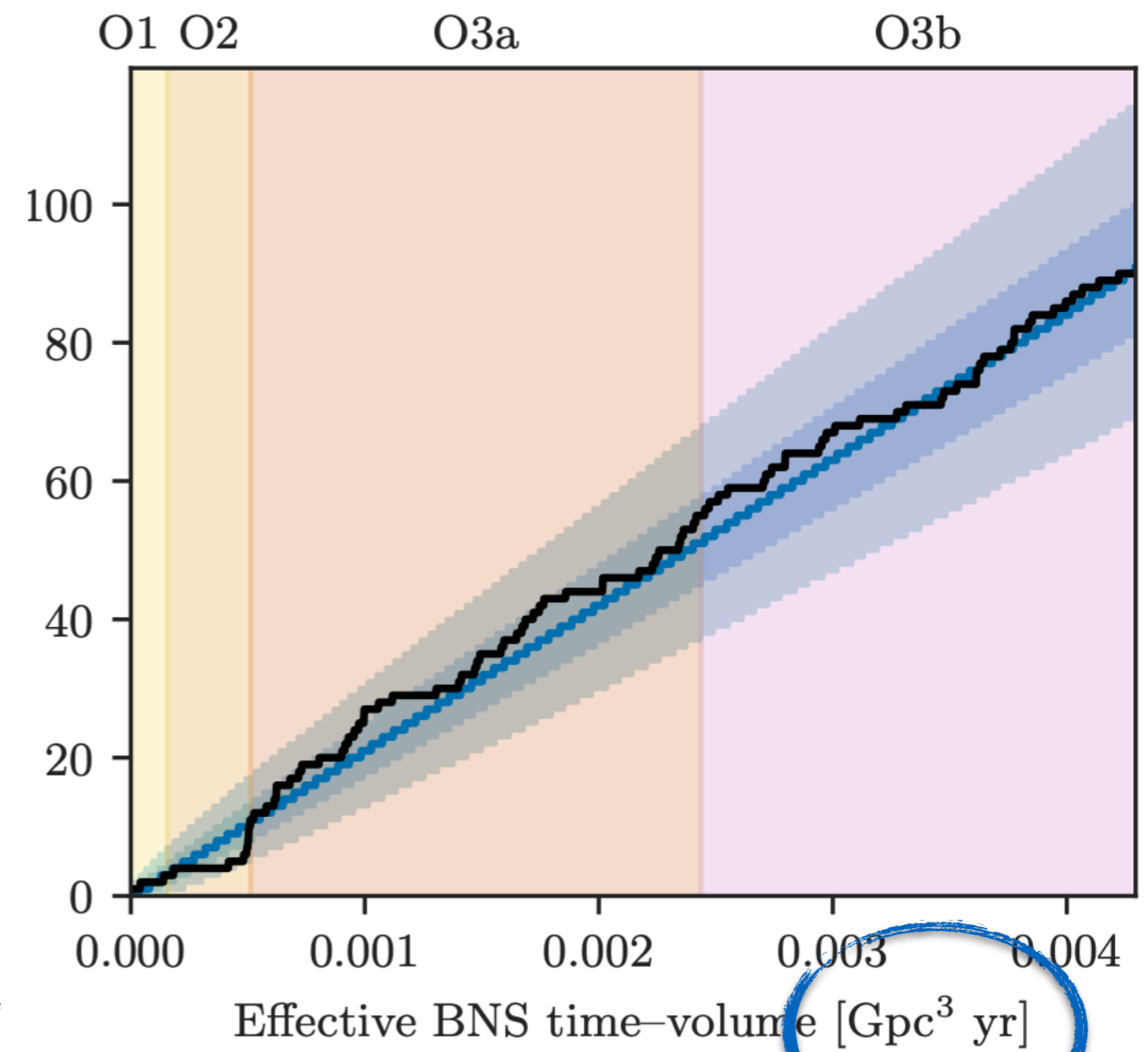
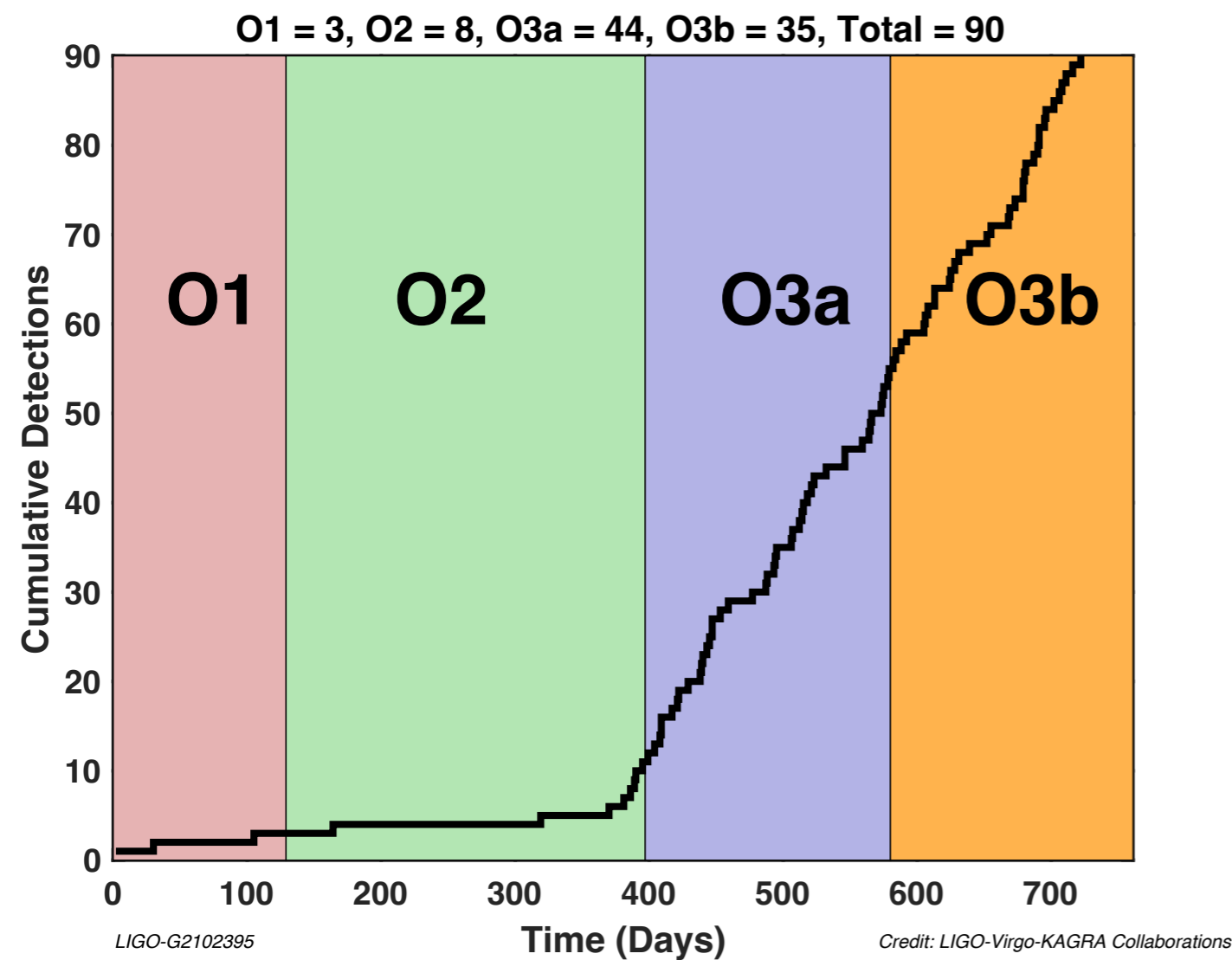


Masses in the Stellar Graveyard

in Solar Masses



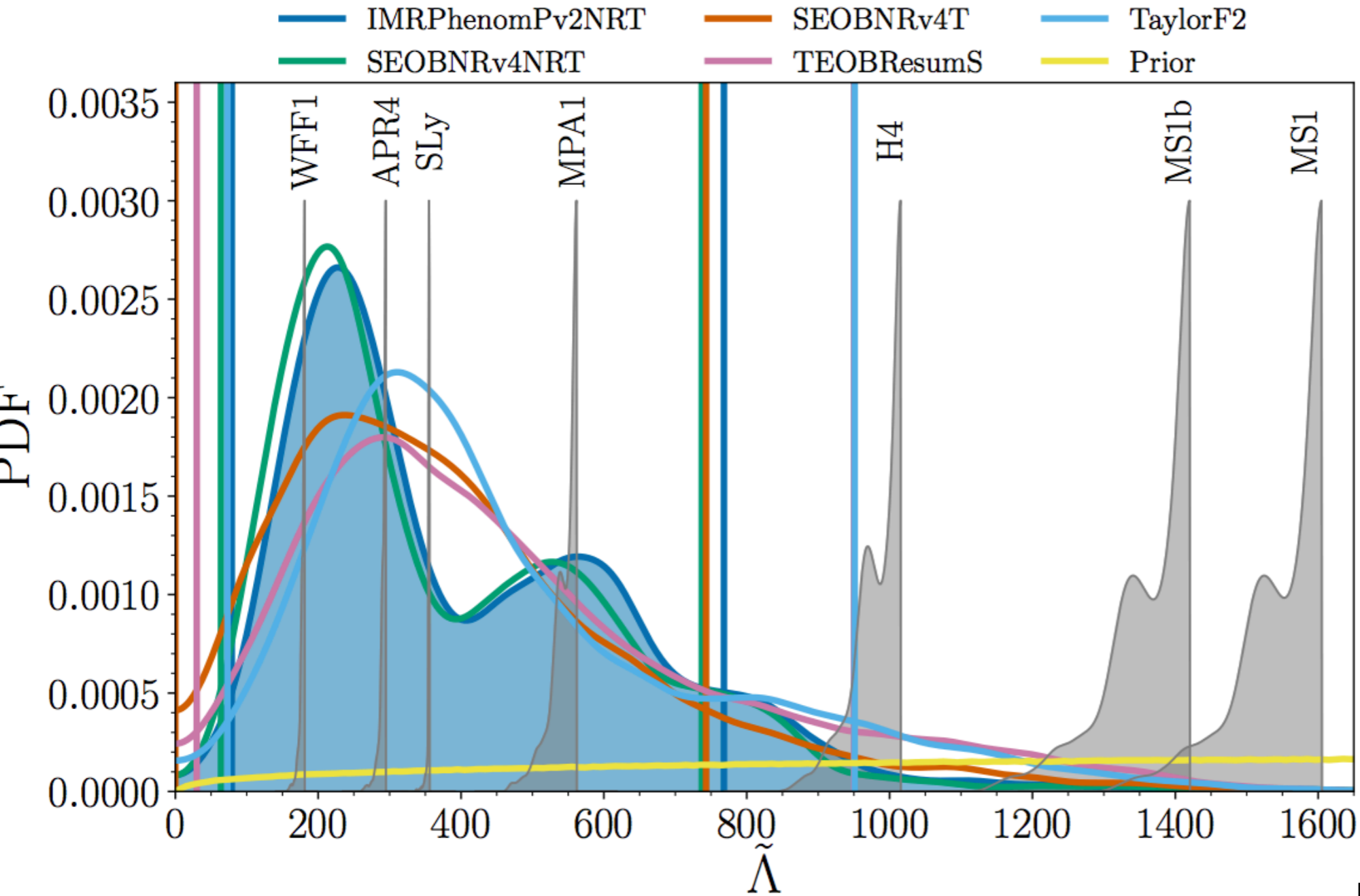
Accelerating growth



$$\text{Sensitivity} \sim \frac{1}{r}$$

$$\# \text{ of detections} \sim r^3$$

Tides in neutron stars: low mass GW170817



LVC (arxiv:1811.12907)

PE: Veitch+ (arxiv:1409.7215)

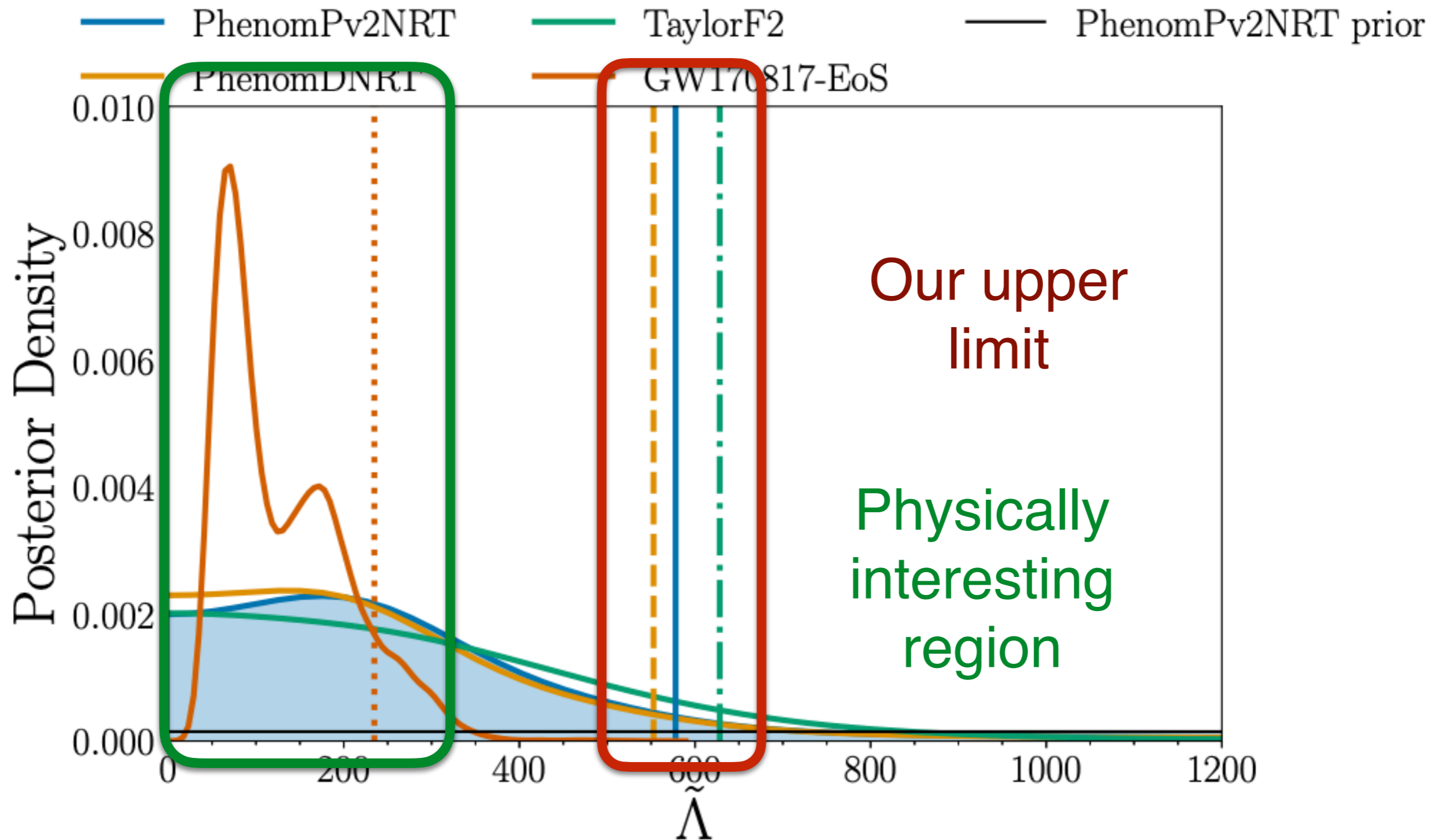
Waveforms:

Dietrich+ (arxiv:1804.02235),

Nagar+ (arxiv:1806.01772),

Hinderer+ (arxiv:1602.00599)

Tides in neutron stars: high mass GW190425

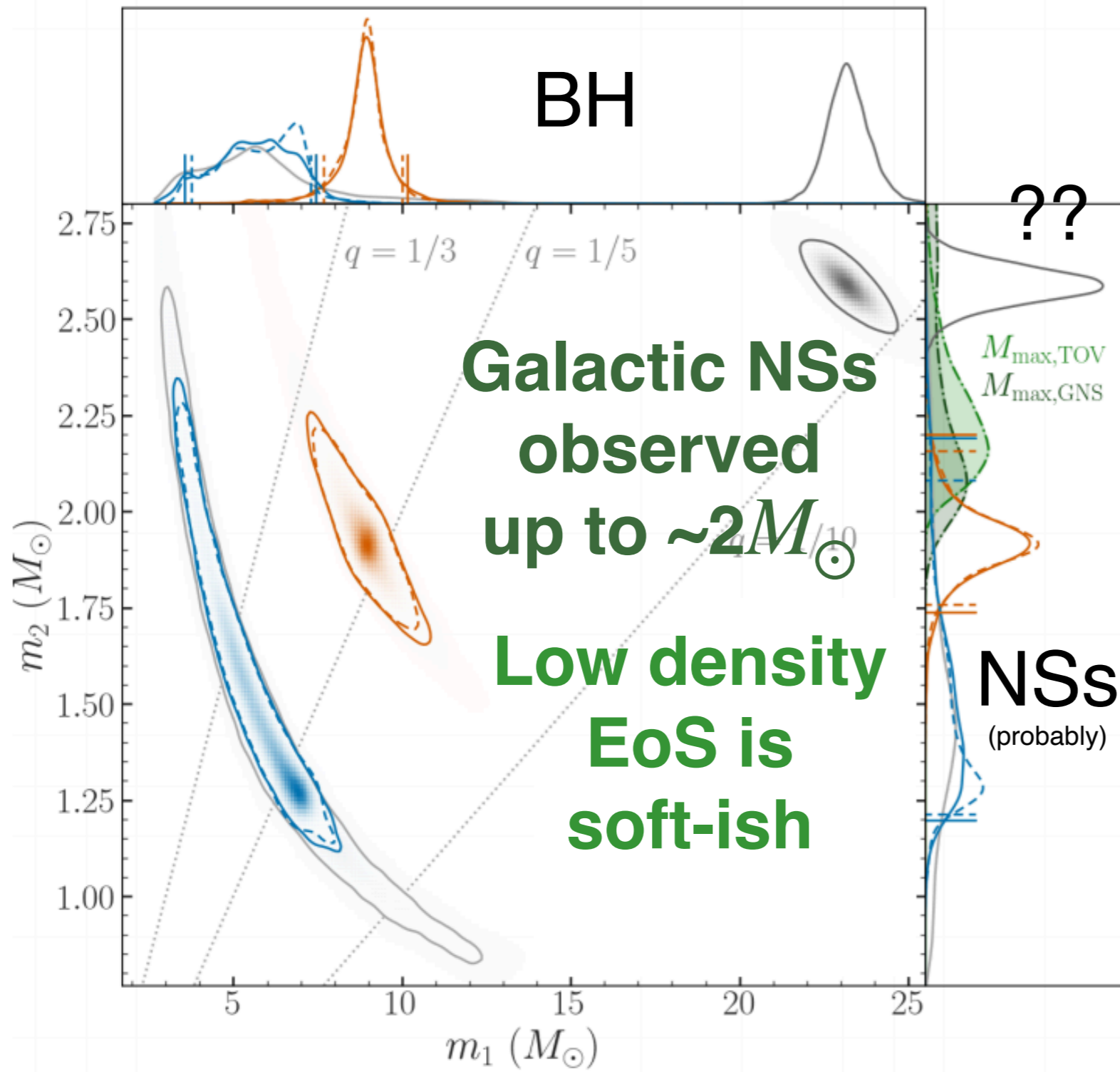


LVC (arxiv:2001.01761)

PE: Veitch+ (arxiv:1409.7215)

Waveform: Dietrich+ (arxiv:1804.02235)

Observations



No **tidal** signature
in any event,
need **external** input?

A $2.6 M_{\odot}$ object:

- BH
- Spinning NS
- Phase transitions
- Statistical outlier
- ...

LVC (arxiv:2106.15163)

LVC (arxiv:2006.12611)

PE: Veitch+ (arxiv:1409.7215), Ashton+ (1811.02042)

Waveforms: Khan+ (arxiv:1911.06050), Ossokine+ (arxiv:2004.09442), Pratten+ (2004.06503)

Tan+ (arxiv:2006.16296)

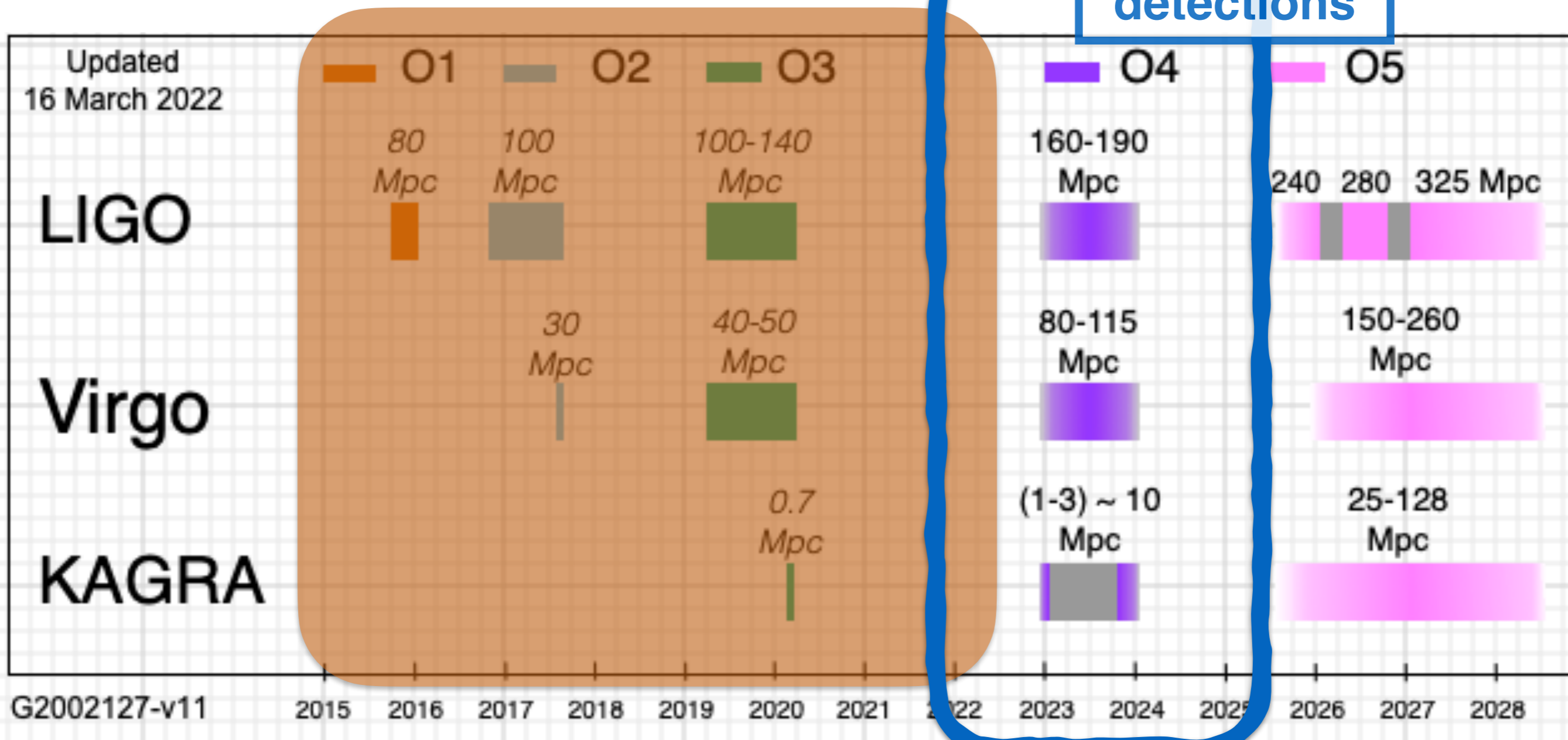
Essick+ (arxiv:2007.01372)

Dexheimer+ (arxiv:2007.08493)

Tews+ (arxiv:2007.06057)

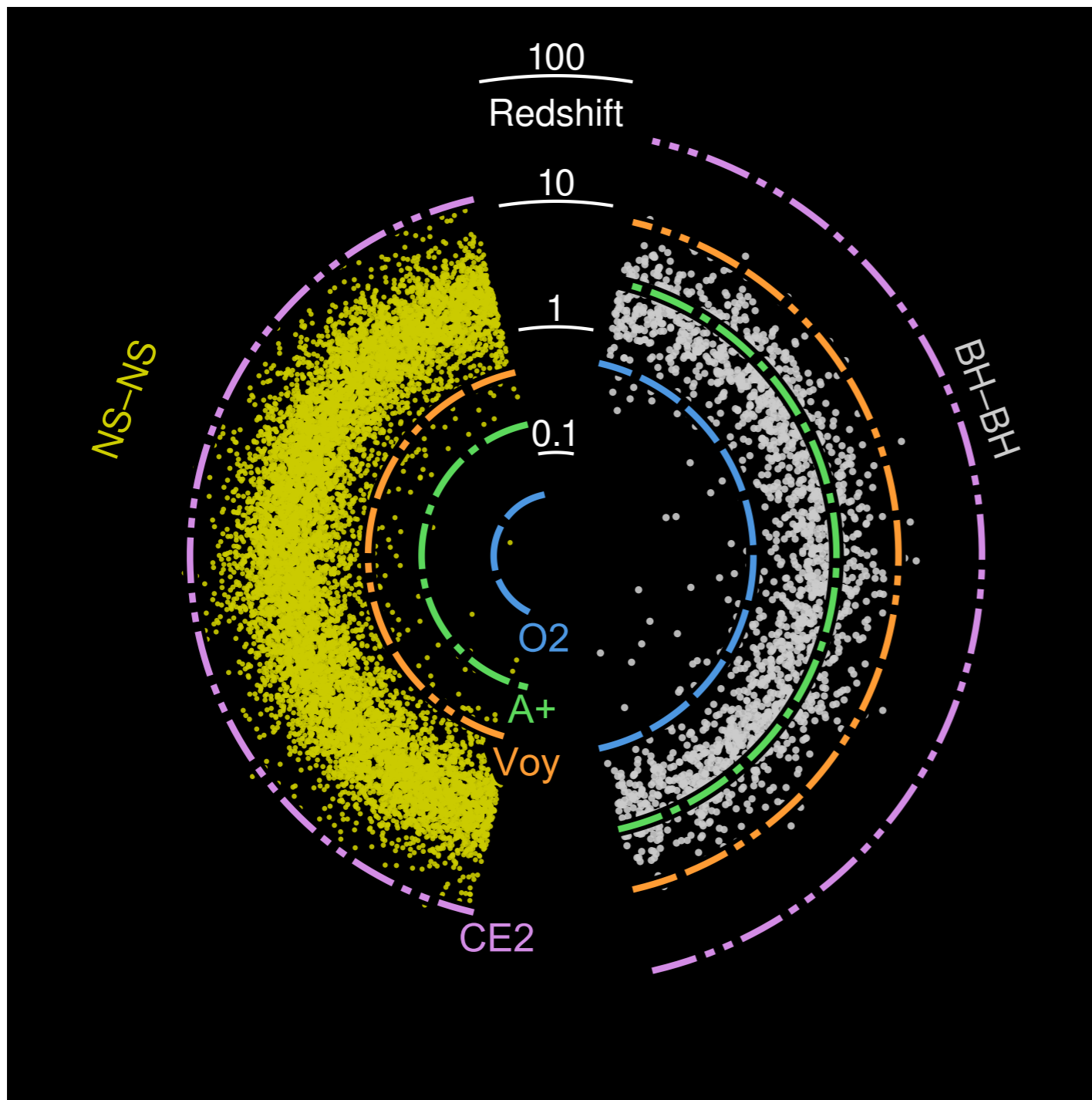
Fattoyev+ (arxiv:2007.03799)

The next steps



**Pulsar timing arrays
(nHz; 10⁹ M_⊙)**

Even further ahead



**O2: 10 binary black holes,
1 binary neutron star**

O5/A+: 2xLIGO

**Voyager: ceiling for
current sites**

**CE2: 3rd gen detectors,
science case**

**LISA (mHz; $10^6 M_{\odot}$,
WD, EMRIs,...)**