

Electroweak structure functions as input to precision radiative corrections at low energies

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Parity-Violation and other Electroweak Physics at JLab 12 GeV
and Beyond, INT, June 2022

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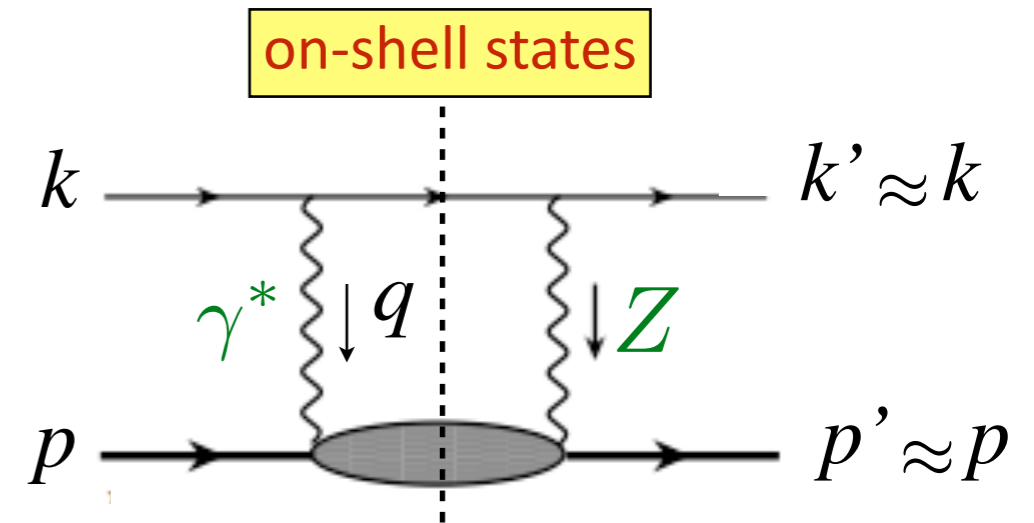
Outline

- Update AJM model predictions for inelastic PVES in SOLID kinematics
- Discuss what improvements could be made
- Discuss improvements to models of $F_3(W^2, Q^2)$ structure functions and their role in precision electroweak physics

“Modern” approach to box diagrams

Gorchtein, Horowitz, PRL 102 (2009) 091806

- GH formulated box corrections in terms of forward angle dispersion relations using γZ structure functions
- Data-driven theoretical analysis



Forward scattering limit: $|f\rangle \approx |i\rangle$

$$\text{Im}\langle i|\mathcal{M}^{\gamma Z}|i\rangle \sim \int d^3 k_1 \frac{L_{\mu\nu}W^{\mu\nu}}{q^2(q^2 - M_Z^2)}$$

hadronic tensor:

$$MW_{\gamma Z}^{\mu\nu} = -g^{\mu\nu} \underbrace{F_1^{\gamma Z}}_{\text{vector } h} + \frac{p^\mu p^\nu}{p \cdot q} \underbrace{F_2^{\gamma Z}}_{\text{vector } h} - i\varepsilon^{\mu\nu\lambda\rho} \frac{p_\lambda q_\rho}{2p \cdot q} \underbrace{F_3^{\gamma Z}}_{\text{axial } h}$$

Vector h correction

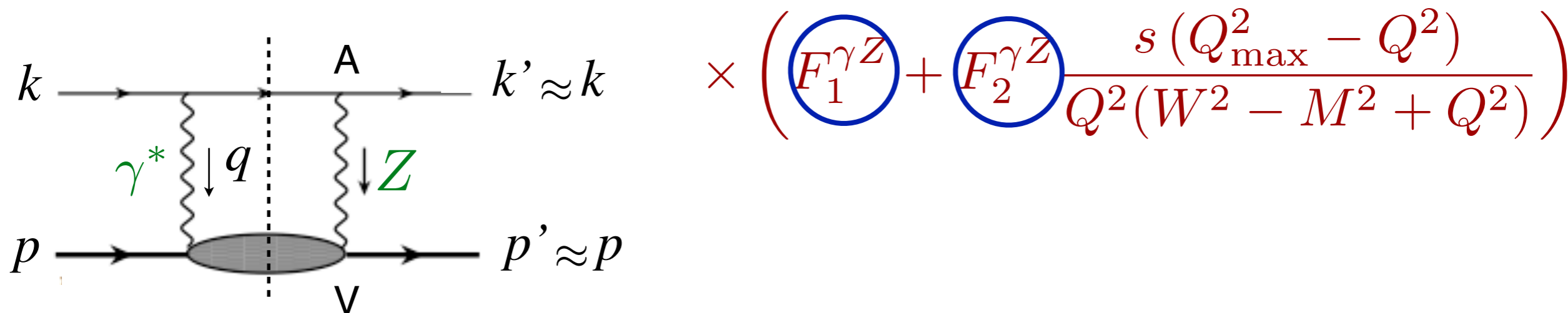
- vector h correction $\square_{\gamma Z}^V$ vanishes at $E = 0$ (atomic PV limit), but PVES is at finite energy (e.g. Qweak $E \sim 1$ GeV)

→ Use forward limit dispersion relation

$$\Re \square_{\gamma Z}^V(E) = \frac{2E}{\pi} \int_0^\infty dE' \frac{1}{E'^2 - E^2} \Im m \square_{\gamma Z}^V(E')$$

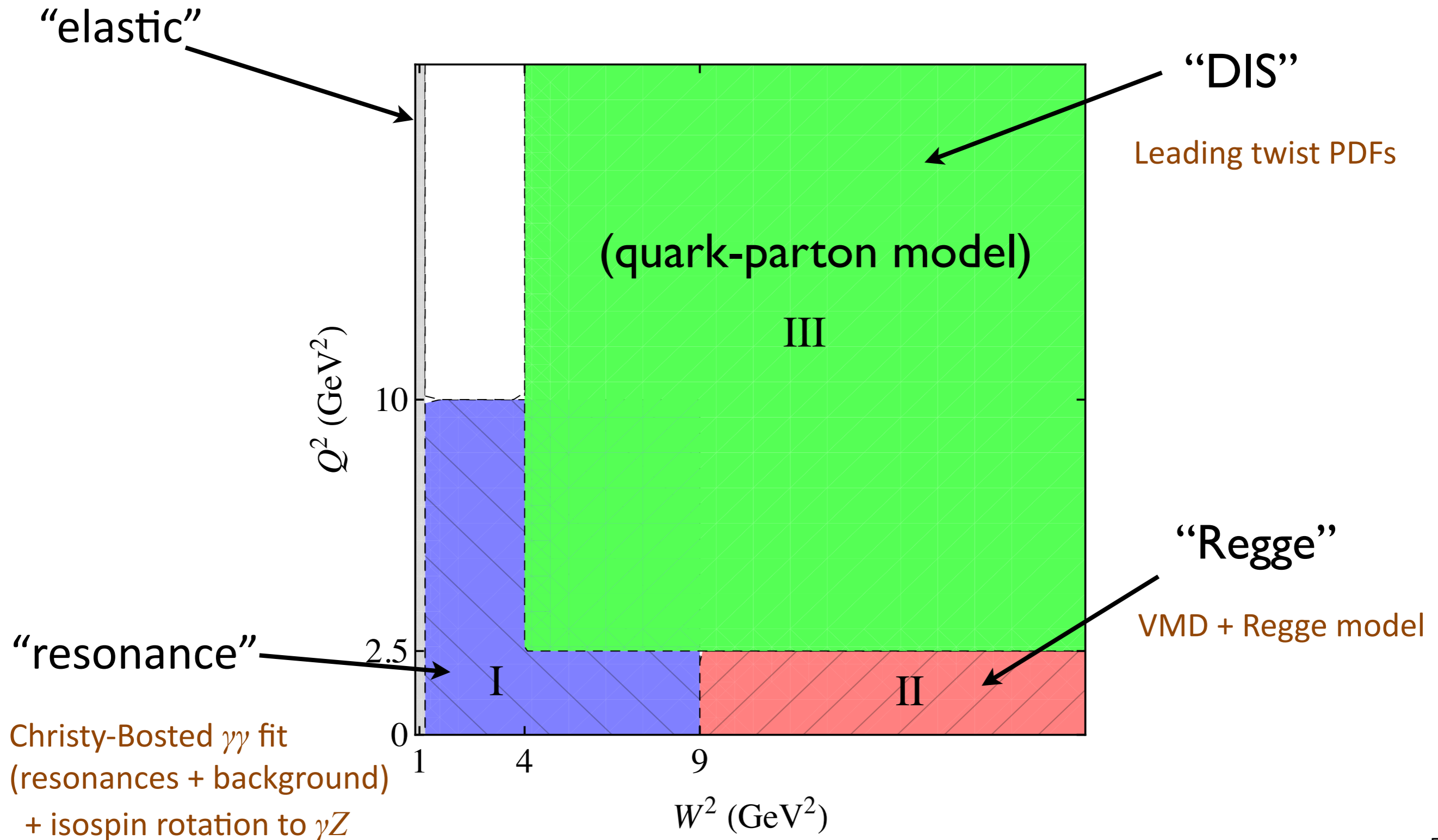
→ imaginary part given by

$$\Im m \square_{\gamma Z}^V(E) = \frac{\alpha}{(s - M^2)^2} \int_{W_\pi^2}^s dW^2 \int_0^{Q_{\max}^2} \frac{dQ^2}{1 + Q^2/M_Z^2}$$



Structure function map: 3 distinct regions

Need structure functions for $A_e \times V_h$ ($F_{1,2}^{\gamma Z}$) and $V_e \times A_h$ ($F_3^{\gamma Z}$)



Main issues

- Relatively little known about $F_i^{\gamma Z}$ interference structure functions below HERA measurements, with $Q^2 \geq 1500 \text{ GeV}^2$
- Fit $F_i^{\gamma\gamma}$ over all kinematics in Q^2 and W , then “rotate” to $F_i^{\gamma Z}$ using available theoretical/phenomenological constraints

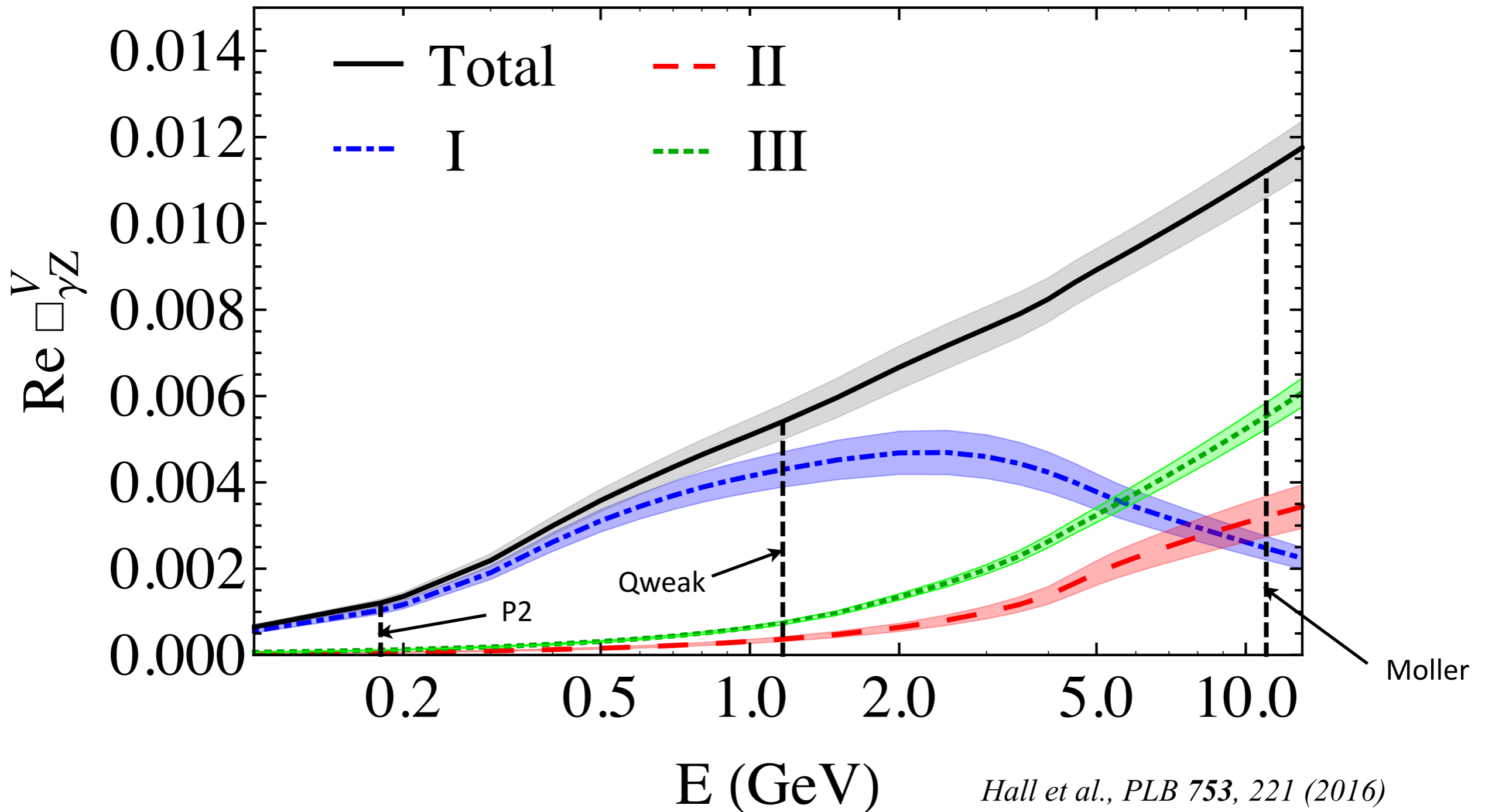
→ e.g. isospin symmetry

$$\langle N^* | J_Z^\mu | p \rangle = (1 - 4 \sin^2 \theta_W) \langle N^* | J_\gamma^\mu | p \rangle - \langle N^* | J_\gamma^\mu | n \rangle$$

- AJM model (approach): require structure functions to match across different kinematic regions (constrain free parameters)

Contribution from different regions to $\square_{\gamma Z}^V$

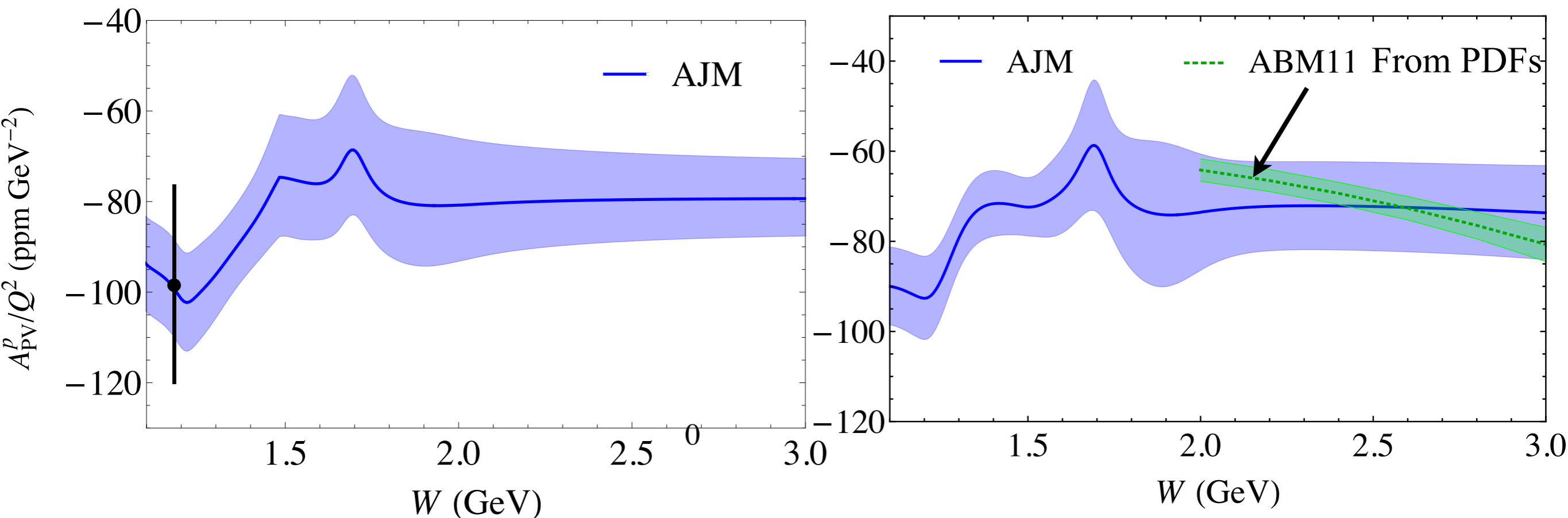
(relative to weak charge of 0.0713)



γZ model direct test

Parity-violating Deep Inelastic Scattering (PVDIS) asymmetry allows a direct measurement of the γZ structure functions

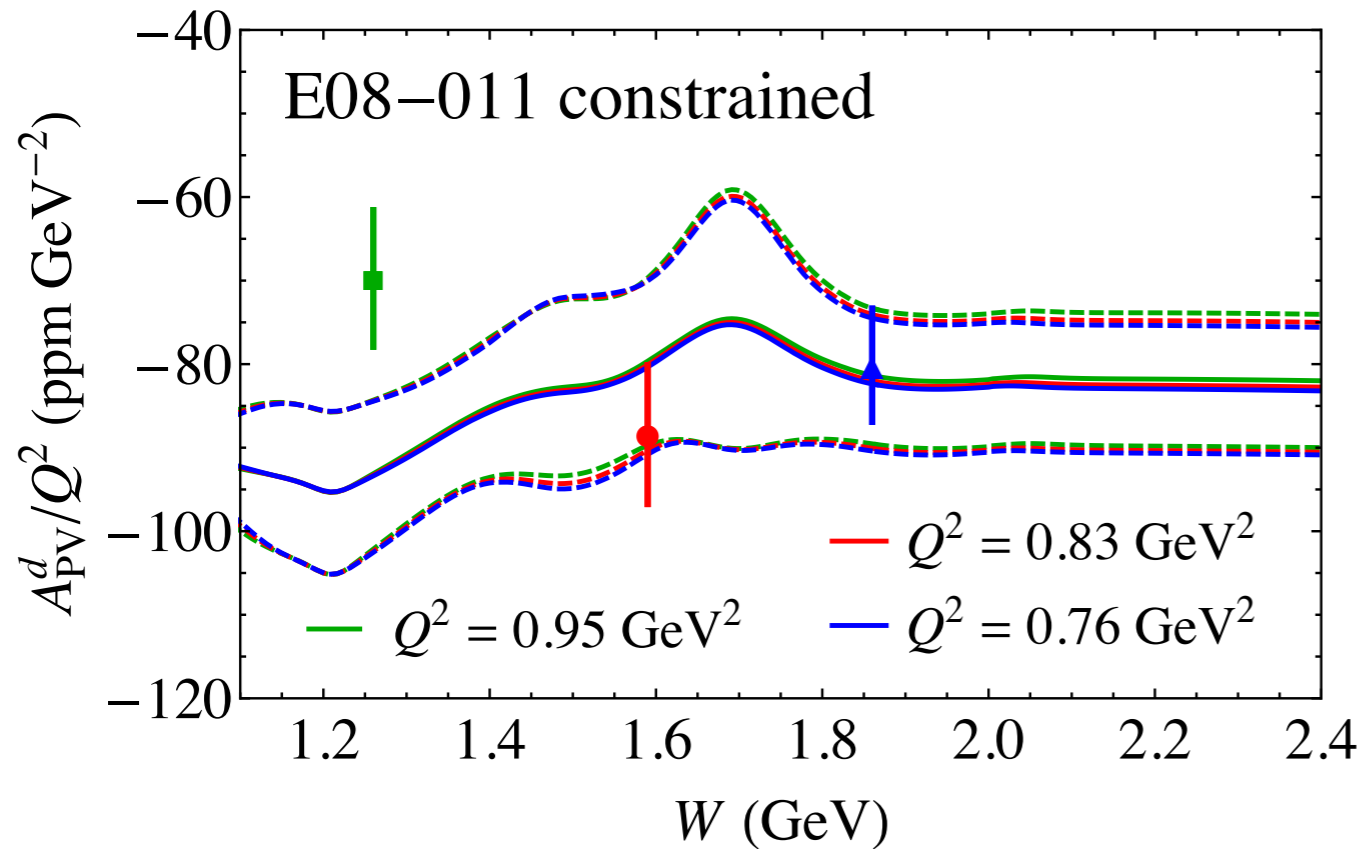
$$A_{PV} = g_A^e \left(\frac{G_F Q^2}{2\sqrt{2}\pi\alpha} \right) \frac{xy^2 F_1^{\gamma Z} + (1-y)F_2^{\gamma Z} + \frac{g_V^e}{g_A^e} (y - y^2/2)x F_3^{\gamma Z}}{xy^2 F_1^{\gamma\gamma} + (1-y)F_2^{\gamma\gamma}}$$



$Q^2 = 0.34 \text{ GeV}^2, E = 0.69 \text{ GeV}$

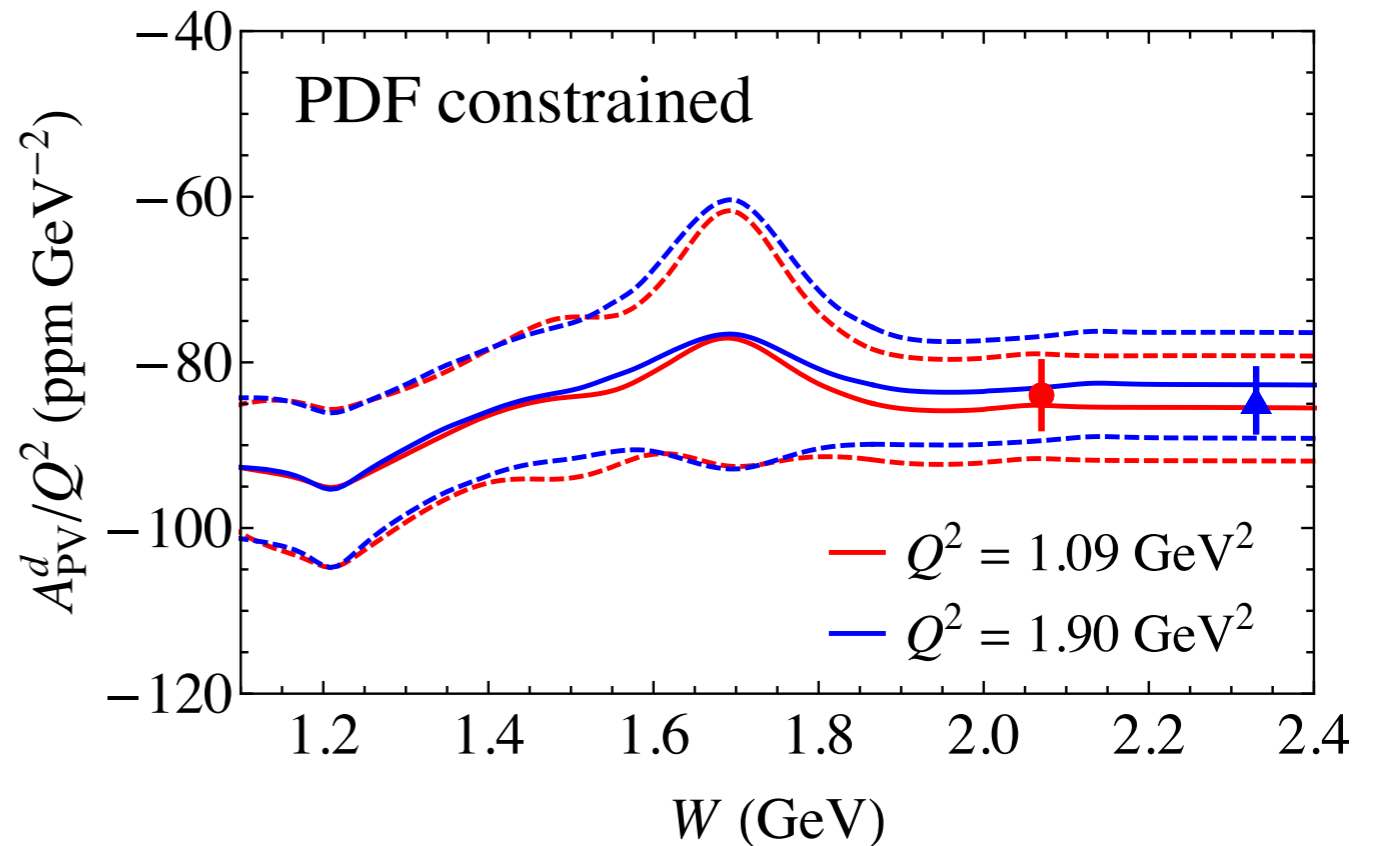
$Q^2 = 2.5 \text{ GeV}^2, E = 6 \text{ GeV}$

Constraints from PV inelastic asymmetries



Wang et al. PRL 111, 082501 (2013)

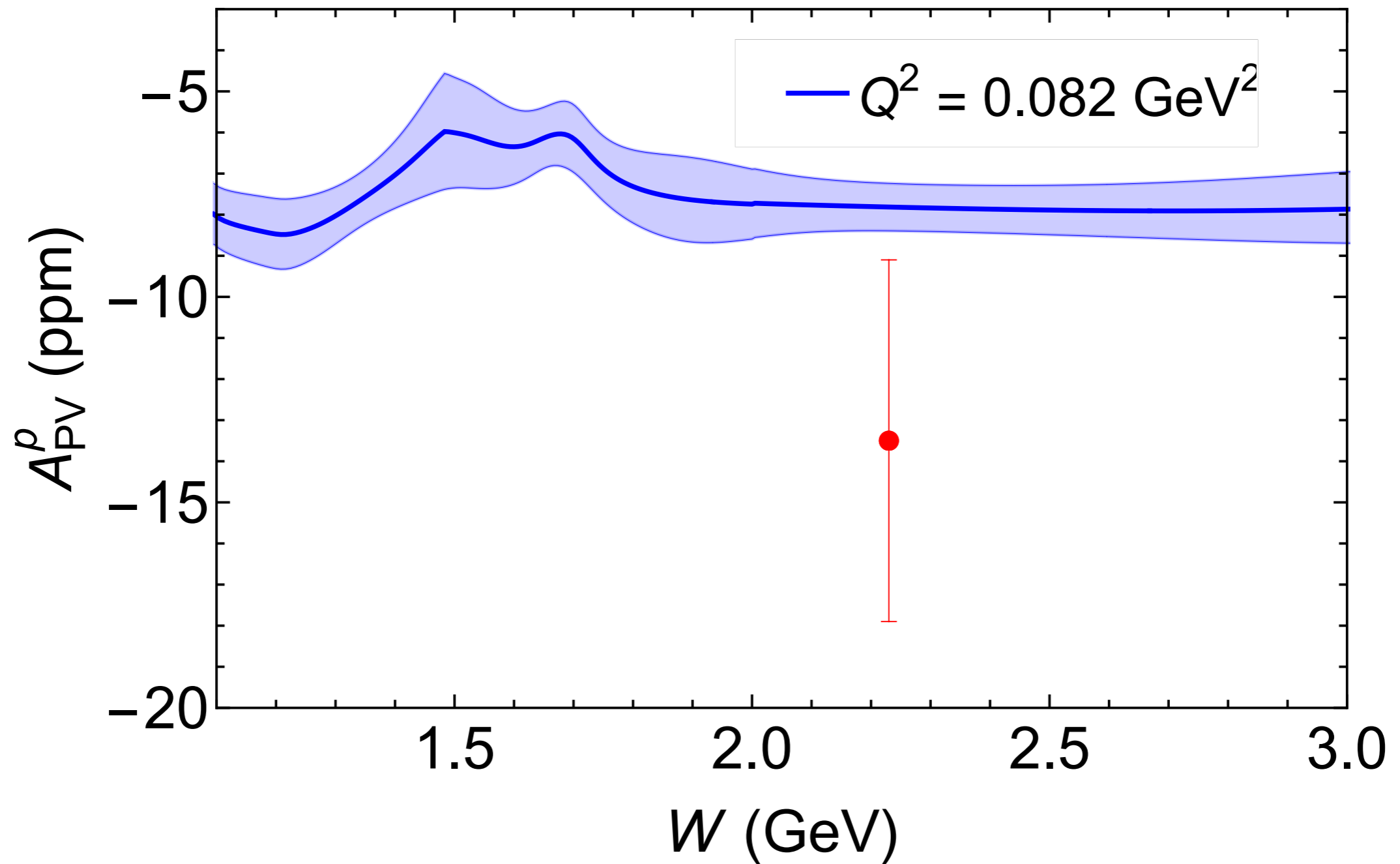
AJM model asymmetries and uncertainties for PV deuteron asymmetry constrained by fit to E08-011 data *Hall et al. (2013)*



E08-011: *Wang et al. Nature 506, 67 (2014)*

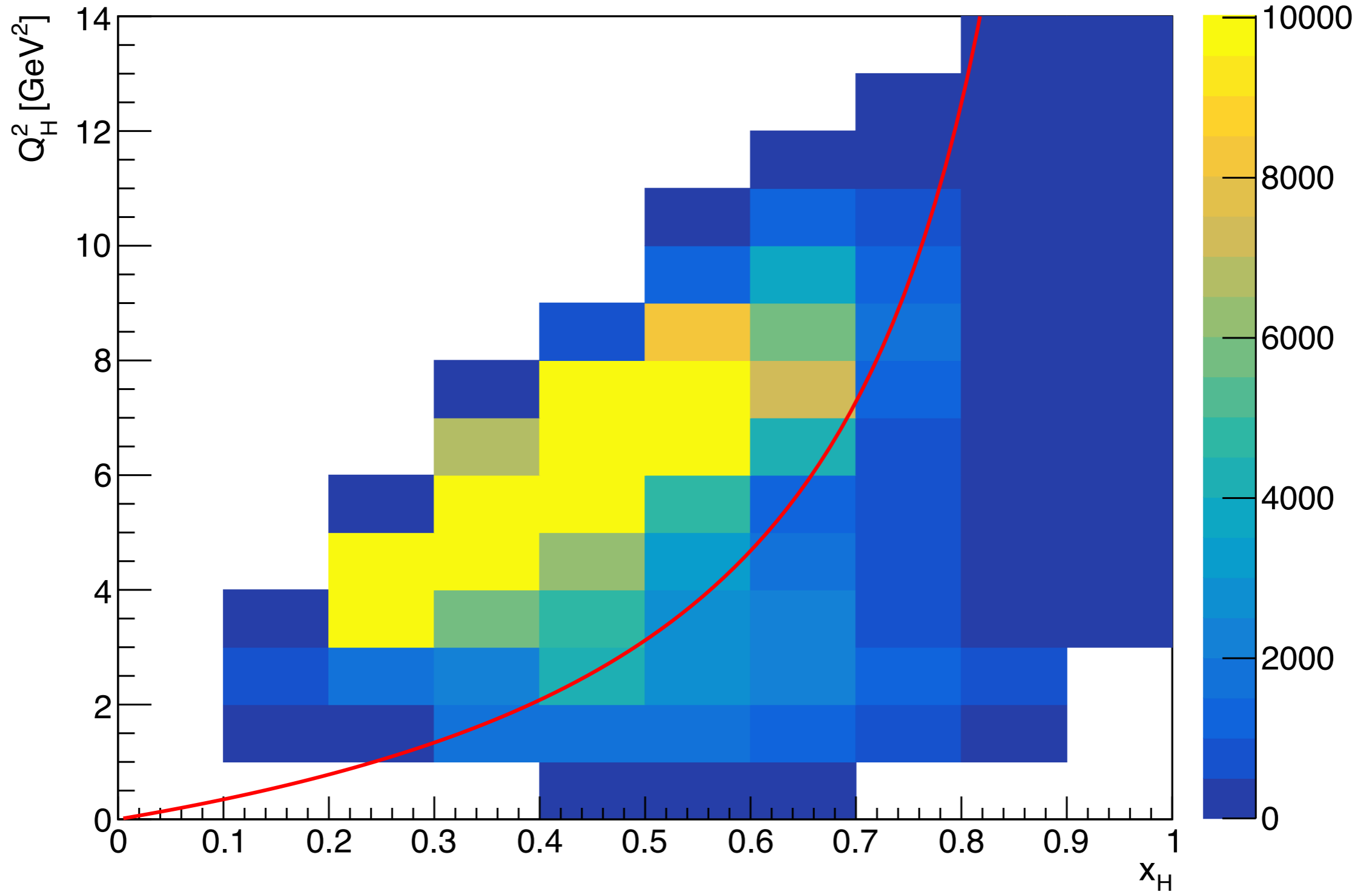
Inelastic asymmetry data from Qweak

Androić et al., PRC 101, 055503 (2020)

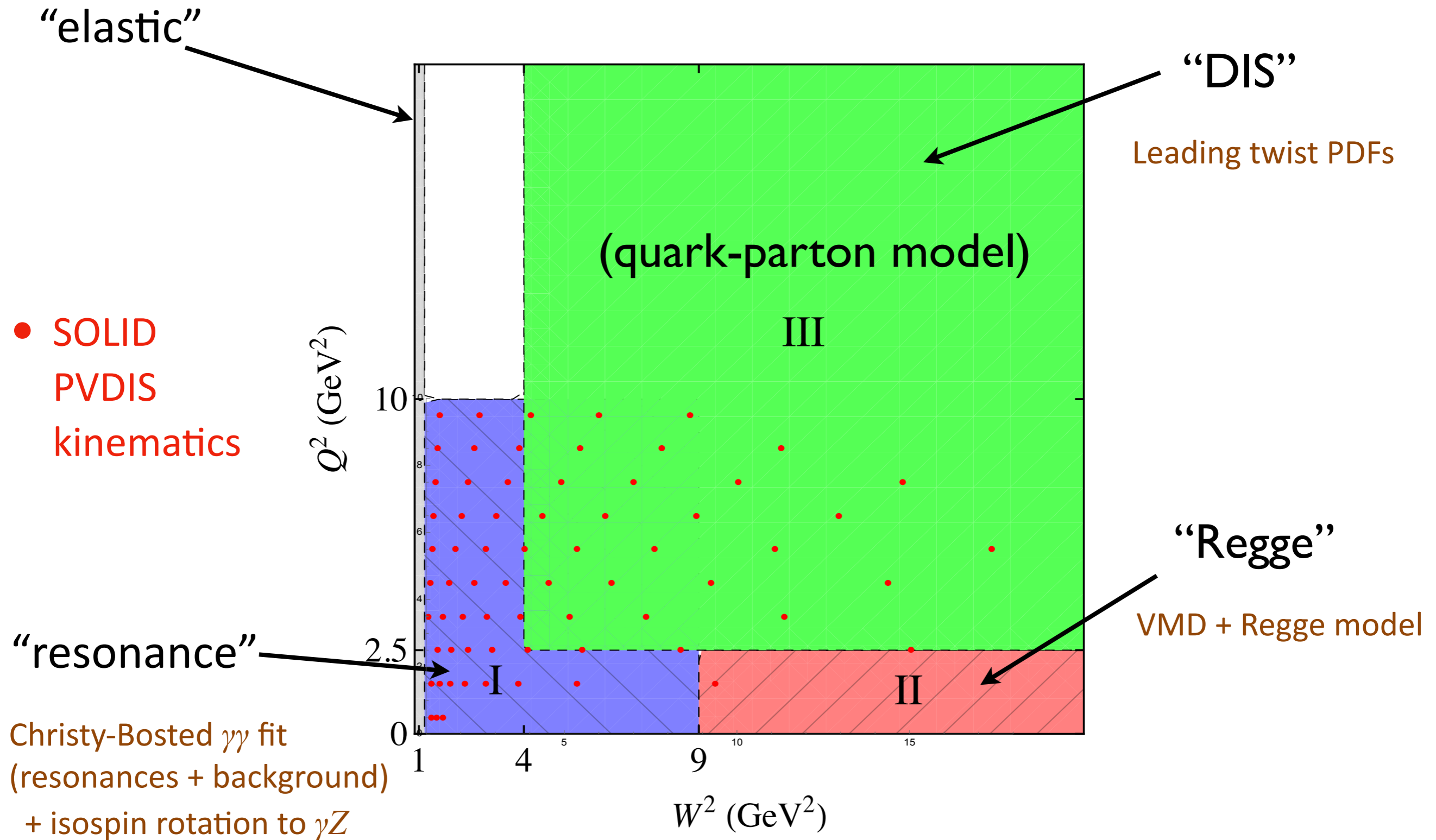


SOLID $A_{PV}(d)$ kinematics

rate_Q2x_h

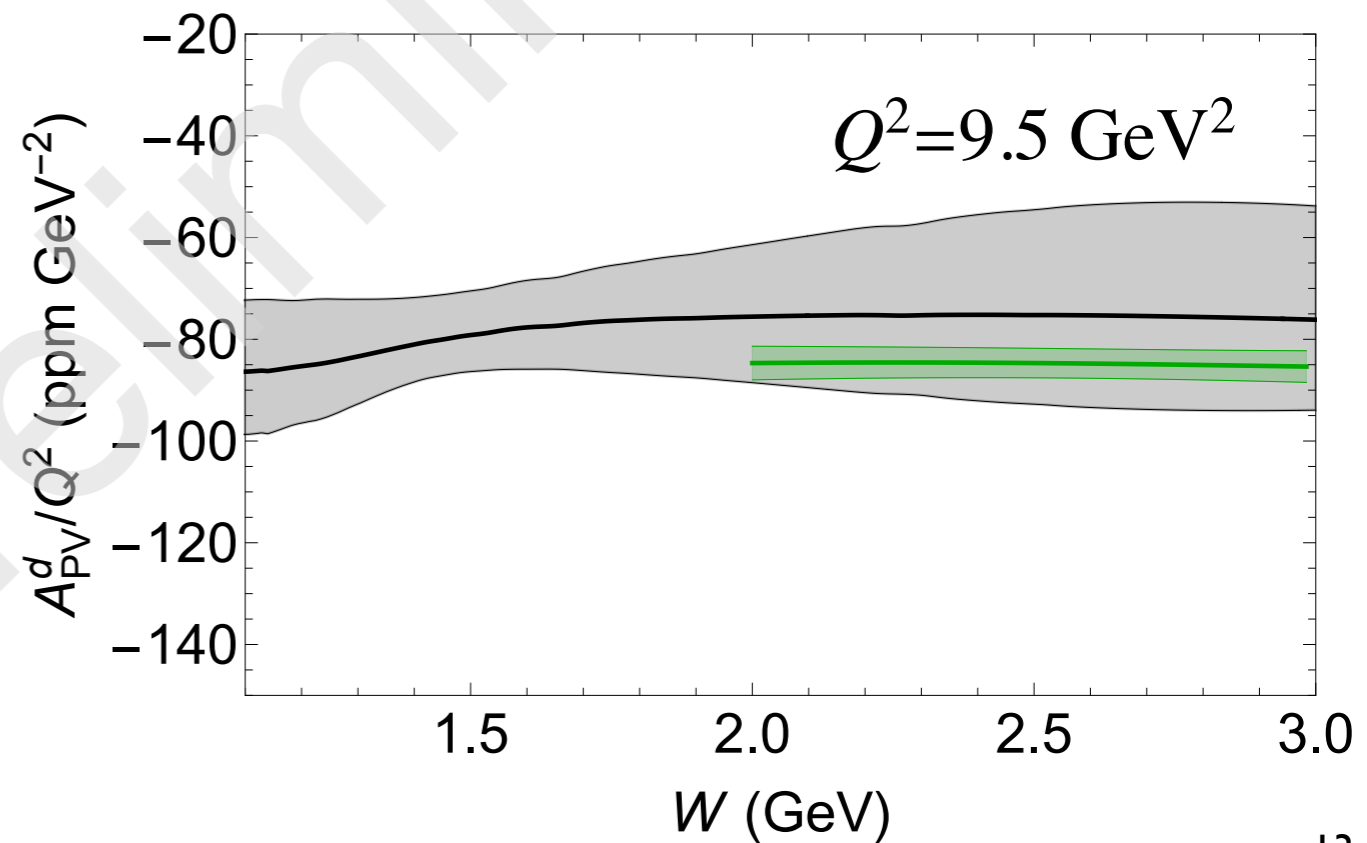
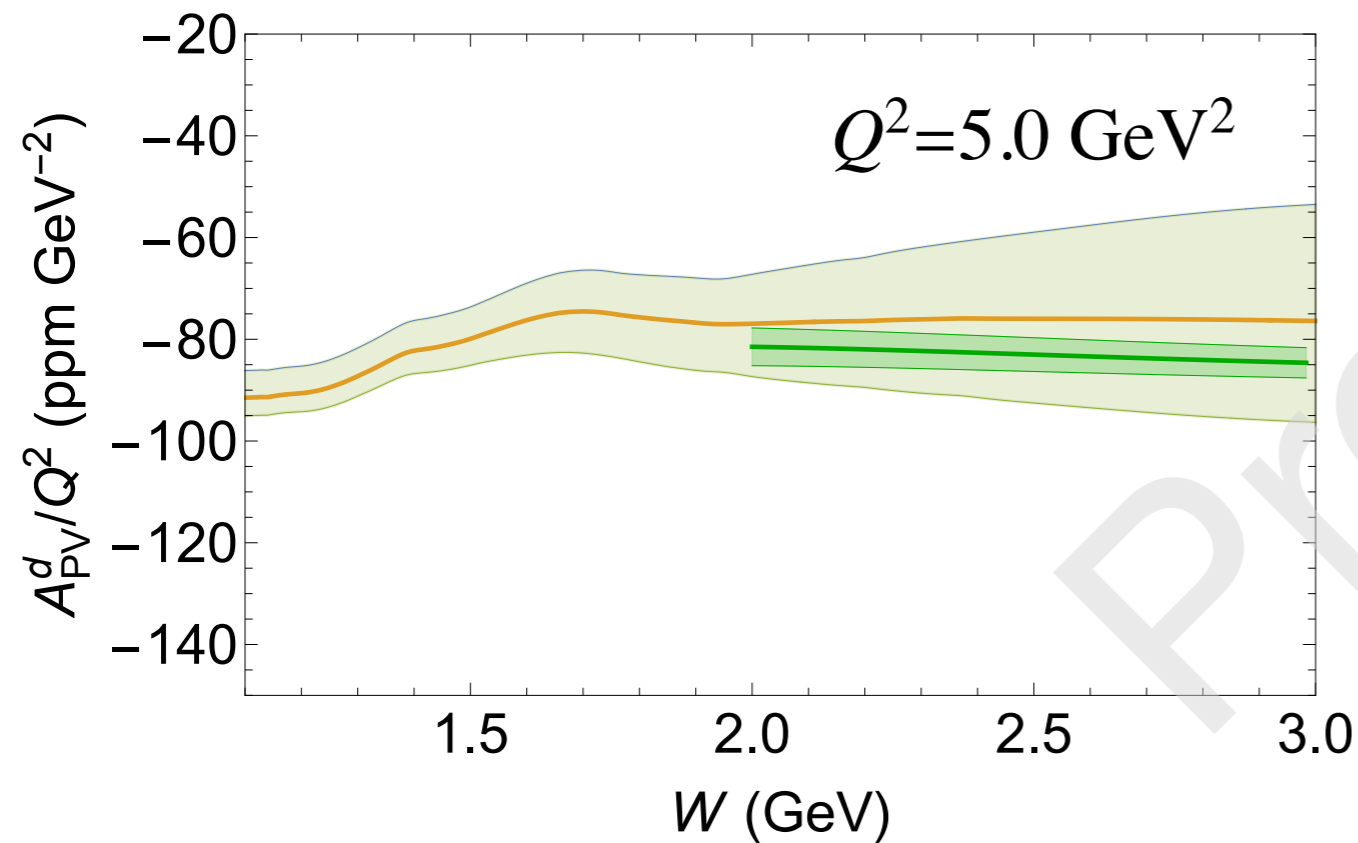
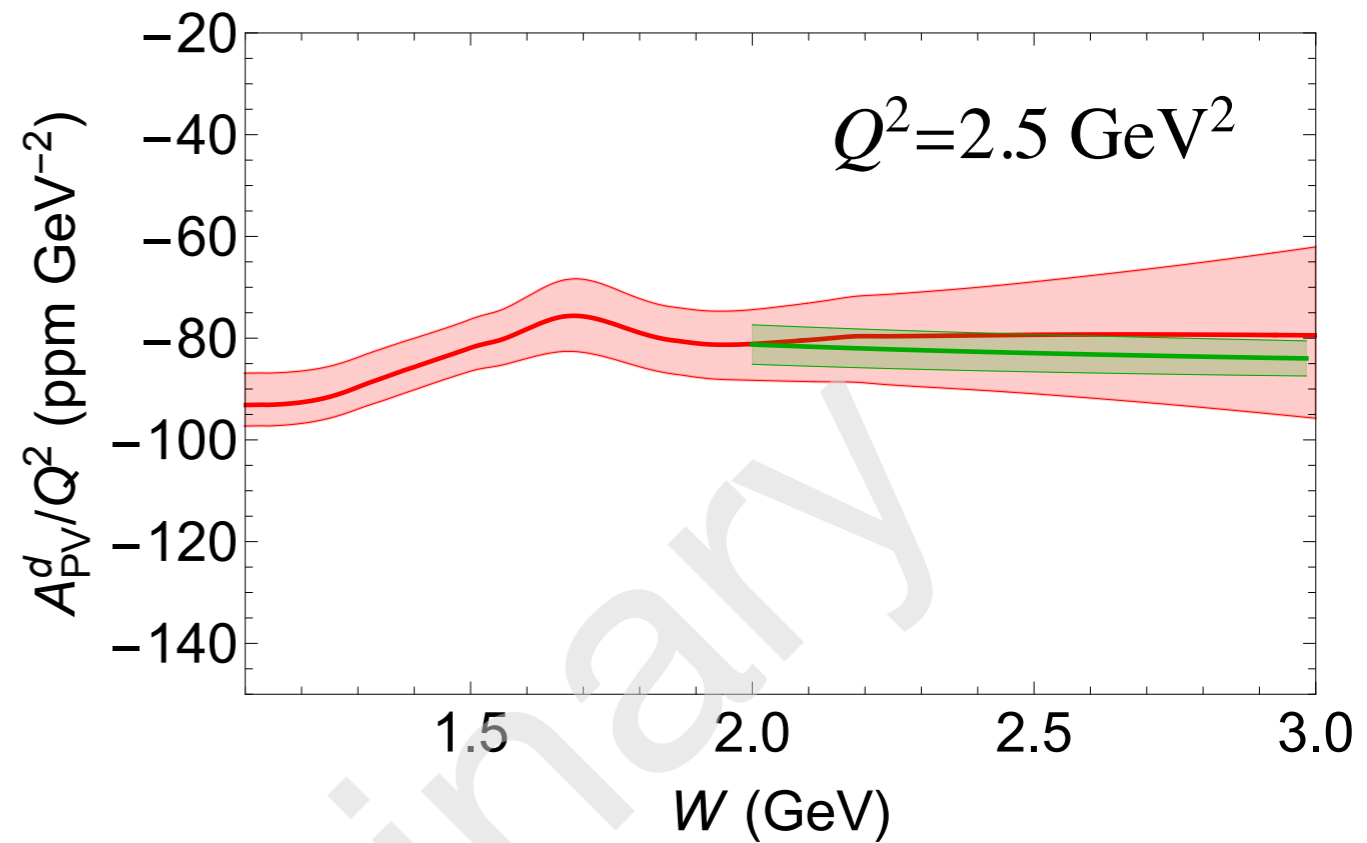
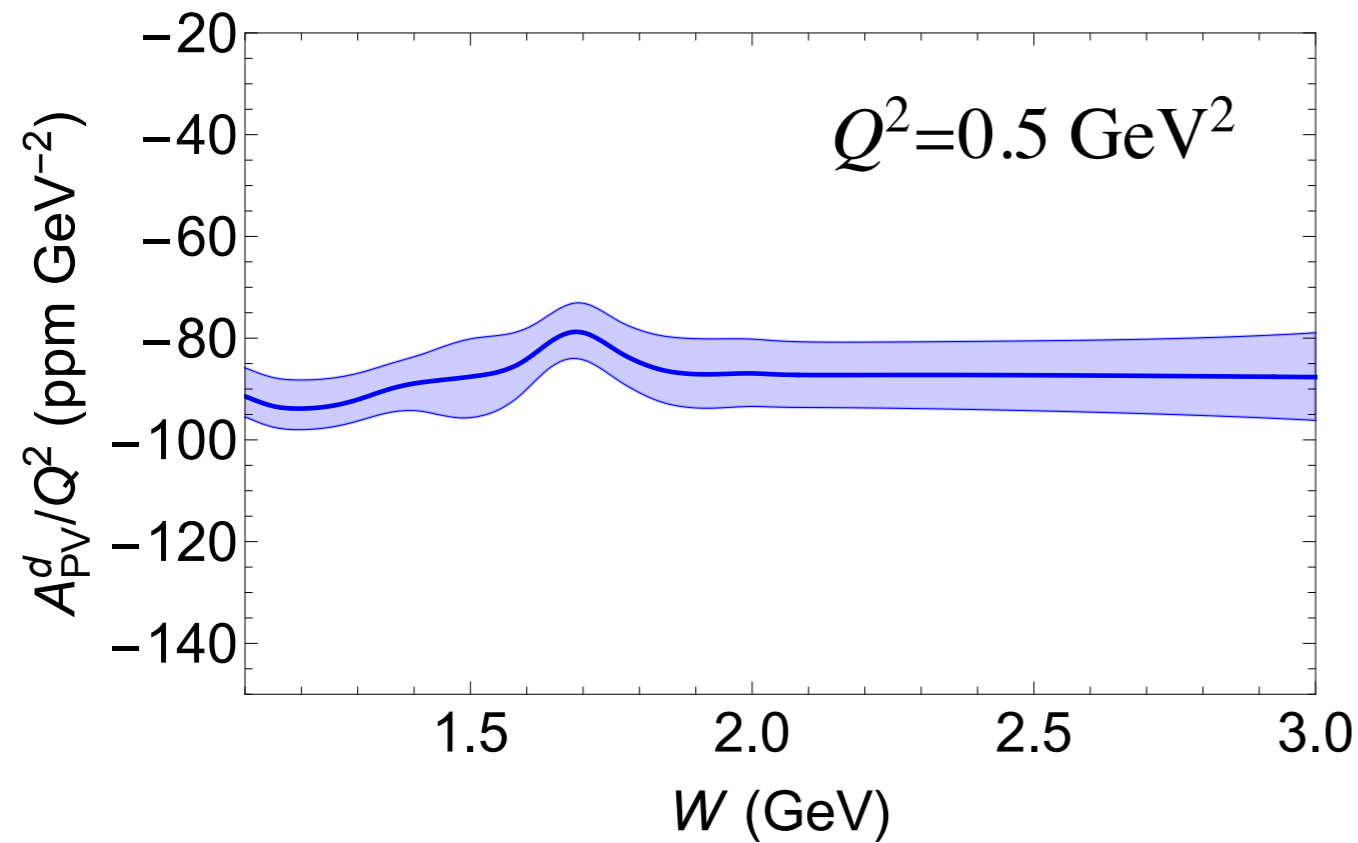


Structure function map with SOLID kinematics



AJM model $A_{PV}(d)$ for selected SOLID kinematics

($E=11$ GeV)



“To do” list

- Use other PDF datasets (*e.g.* JAM)
- All analyses (Gorchtein *et al.*, Rislow-Carlson, AJM collaboration) use Christy-Bosted parametrization of resonance region (pre 2010).
 - This is a parametrization of N^* states to fit inclusive cross section data
 - It is **not** based on transition form factors and branching ratios of produced resonant states, for which exclusive meson production is suited.
 - Suggests that other more recent parametrizations should be explored (*e.g.* CLAS)
- Incorporate recent improvements to $F_3(W^2, Q^2)$ structure function into the predictions

Axial h correction

$$\square_A^{\gamma Z}(E=0) \sim g_V^e \int_0^\infty dQ^2 \frac{\alpha(Q^2)}{Q^2(1+Q^2/M_Z^2)} \int_0^{x_{\max}} dx F_{3p}^{\gamma Z}(x, Q^2) \frac{1+2r}{(1+r)^2},$$
$$x = \frac{Q^2}{W^2 - M^2 + Q^2}, \quad r = \sqrt{1 + 4M^2 x^2 / Q^2}.$$

Blunden *et al.* PRL **107**, 181861 (2011)

→ in DIS region ($Q^2 > Q_0^2$) integrand can be expressed in terms of moments of structure functions by letting $x_{\max} \rightarrow 1$

$$M_3^{\gamma Z(n)} = \int_0^1 dx x^{n-1} F_3^{\gamma Z}(x, Q^2)$$

Allows for systematic small corrections to a model-independent leading term

$$\underline{n=1} \quad M_3^{\gamma Z(1)}(Q^2) = \frac{5}{3} \left(1 - \frac{\alpha_s(Q^2)}{\pi} \right)$$

→ γZ analog of Gross-Llewellyn Smith sum rule for F_3^W

→ leads to old result of Marciano & Sirlin

F_3 structure function landscape

F_3 structure functions	PDF (DIS region)	Observable/Experiment
$F_{3p}^{\gamma Z}$	$\frac{2}{3}u_v + \frac{1}{3}d_v + \frac{2}{3}c_v + \frac{1}{3}s_v$	$A_{\text{PV}}(p)$ (Qweak): $\square_A^{\gamma Z}(E)$
$F_{3n}^{\gamma Z}$	$\frac{1}{3}u_v + \frac{2}{3}d_v + \frac{2}{3}c_v + \frac{1}{3}s_v$	$A_{\text{PV}}(\text{Cs})$ (atomic PV)
$F_{3p}^{\gamma Z} + F_{3n}^{\gamma Z}$	$u_v + d_v + \frac{4}{3}c_v + \frac{2}{3}s_v$	$A_{\text{PV}}(d)$ (SOLID)
$F_3^W = \frac{1}{2} \left(F_{3p}^{W^+} + F_{3p}^{W^-} \right)$	$u_v + d_v + c_v + s_v$	$\nu p + \bar{\nu} p$: GLS sum rule
$F_{3(0)}^{\gamma W} = F_{3p}^{\gamma Z} - F_{3n}^{\gamma Z}$	$\frac{1}{3}u_v - \frac{1}{3}d_v$	superallowed β -decay: $\square_A^{\gamma W}(E=0)$
$F_{3(1)}^{\gamma W} = F_{1p}^\gamma - F_{1n}^\gamma$	$(u + \bar{u}) - (d + \bar{d})$	Gottfried sum rule: $\square_A^{\gamma W}(E \neq 0)$

$\square_A^{\gamma Z} (\times 10^{-3})$

P2

Qweak

E (GeV)

0.155

1.165

Marciano and Sirlin (1984)

5.2(5)

5.2(5)

Blunden *et al.* (2011)

4.4(4)

3.7(4)

No background term in resonance region

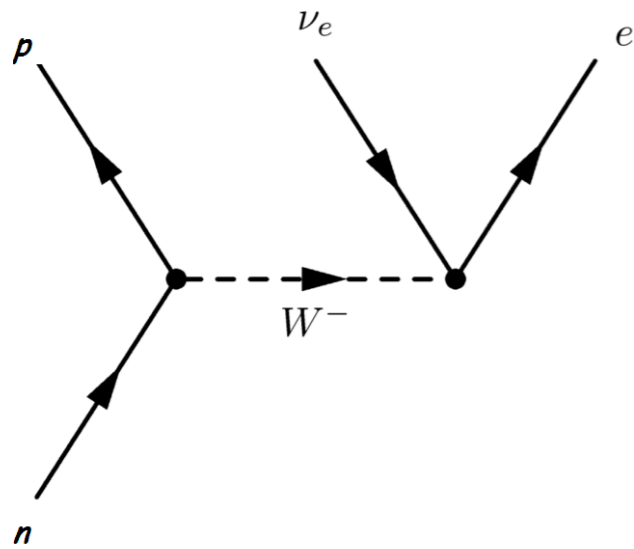
Erler *et al.* (2019)

4.46(21)

3.97(22)

Rewrite in terms of isoscalar + isovector currents and use experimental constraints from F_3^W

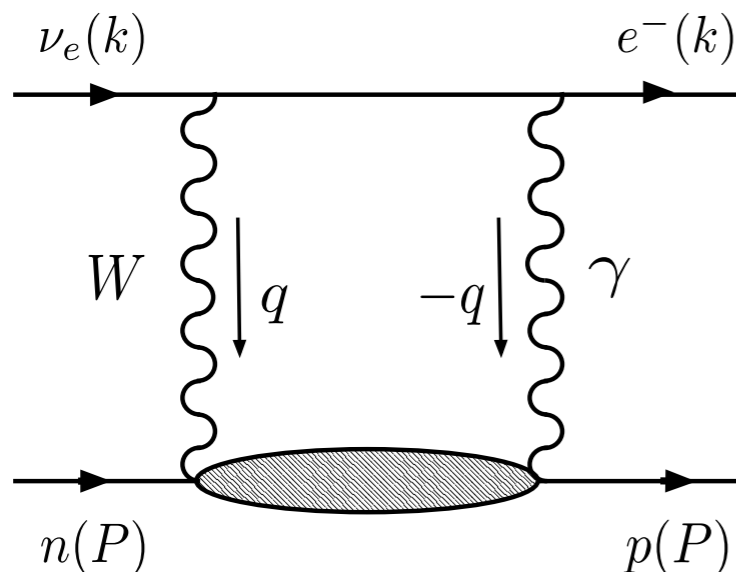
Superallowed 0^+ to 0^+ β decays



$$|V_{ud}|^2 = \frac{0.97148(20)}{1 + \Delta_R^V},$$

Inner radiative correction term (nucleus-independent)

$$\Delta_R^V = \frac{\alpha}{2\pi} \left[3 \ln \frac{M_Z}{M} - \ln \frac{M_W}{M_Z} \right] + 2 \square_A^{\gamma W},$$



γW box is largest source of hadronic uncertainty

Seng *et al.* (2018, 2019) found the correct dispersion representation of $\square_A^{\gamma W}$

Analogous to the contribution to the weak axial charge. At zero energy

$$\square_A^{\gamma W} = \frac{\alpha}{2\pi} \int_0^\infty dQ^2 \frac{1}{Q^2(1 + Q^2/M_W^2)} \int_0^{x_{\max}} dx F_3^{(0)}(x, Q^2) \frac{1 + 2r}{(1 + r)^2}.$$

Isospin symmetry implies $F_3^{(0)} = F_{3p}^{\gamma Z} - F_{3n}^{\gamma Z}$

In DIS region: $F_3^{(0)} = \frac{1}{3}u_v - \frac{1}{3}d_v$ (not $F_3^{(0)} = \frac{1}{6}(u - \bar{d})$ as originally written)

$$F_3^W \approx u_v + d_v$$

Used sum rules plus data on F_3^W from $\nu p + \bar{\nu} p$

→ Found a shift in $\square_A^{\gamma W}$ from $3.52(11) \times 10^{-3}$ to $3.79(10) \times 10^{-3}$.

Shiells *et al.* (2021): calculation based on constructing $F_3^{(0)}(W^2, Q^2)$ and imposing continuity across kinematic boundaries (AJM approach)

Elastic contribution

$$F_{3(\text{el})}^{(0)}(Q^2) = -[G_M^p(Q^2) + G_M^n(Q^2)] G_A(Q^2) Q^2 \delta(W^2 - M^2)$$

- Isoscalar magnetic form factor from recent fits (including TPE)
- Axial form factors from νN scattering (major source of uncertainty)

$$G_A(Q^2) = -\frac{g_A}{(1 + Q^2/m_A^2)^2} \quad m_A = 1.05(5) \text{ GeV}$$

Resonance contributions (isoscalar, so very small)

- Electromagnetic $e+p$ and $e+n$ helicity amplitudes from MAID2009 (can also use CLAS for $e+p$ with similar results)
- Axial transition form factors from Leitner (νN scattering)

DIS contribution

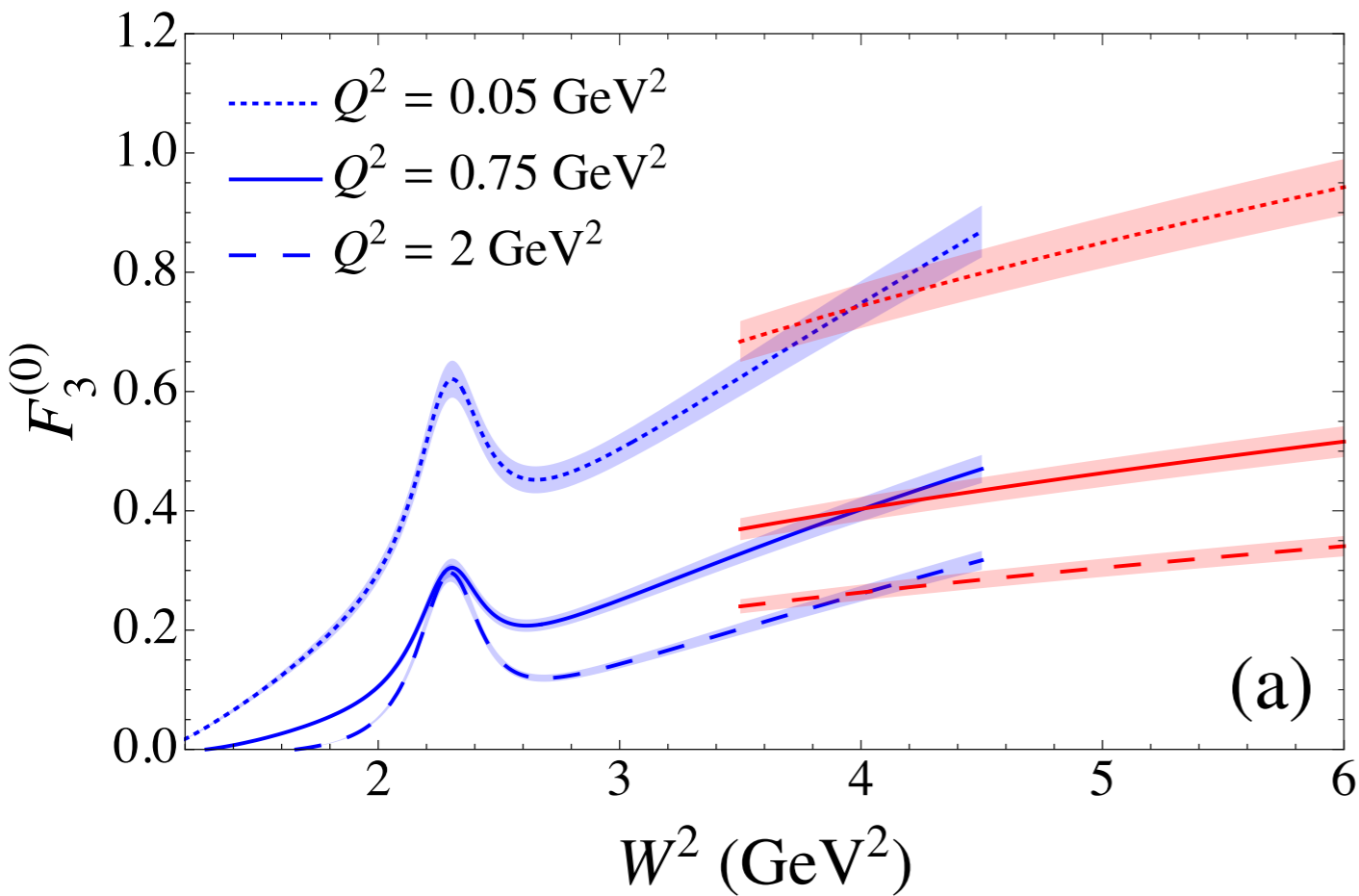
- Convolution at NLO

$$F_{3,\text{DIS}}^{(0)}(x, Q^2) = \int_x^1 \frac{dz}{z} C_3^{(1)}(z) \times \frac{1}{3} (u_v(x/z, Q^2) - d_v(x/z, Q^2))$$

	$\square_A^{\gamma W} (\times 10^{-3})$	
	Q_0^2 (GeV ²)	
	2.0	1.0
elastic	1.05(4)	1.05(4)
resonance	0.04(1)	0.04(1)
DIS + ($Q^2 > Q_0^2$) bgd	2.29(3)	2.43(5)
Regge + ($Q^2 < Q_0^2$) bgd		
Total	3.90(9)	3.91(9)

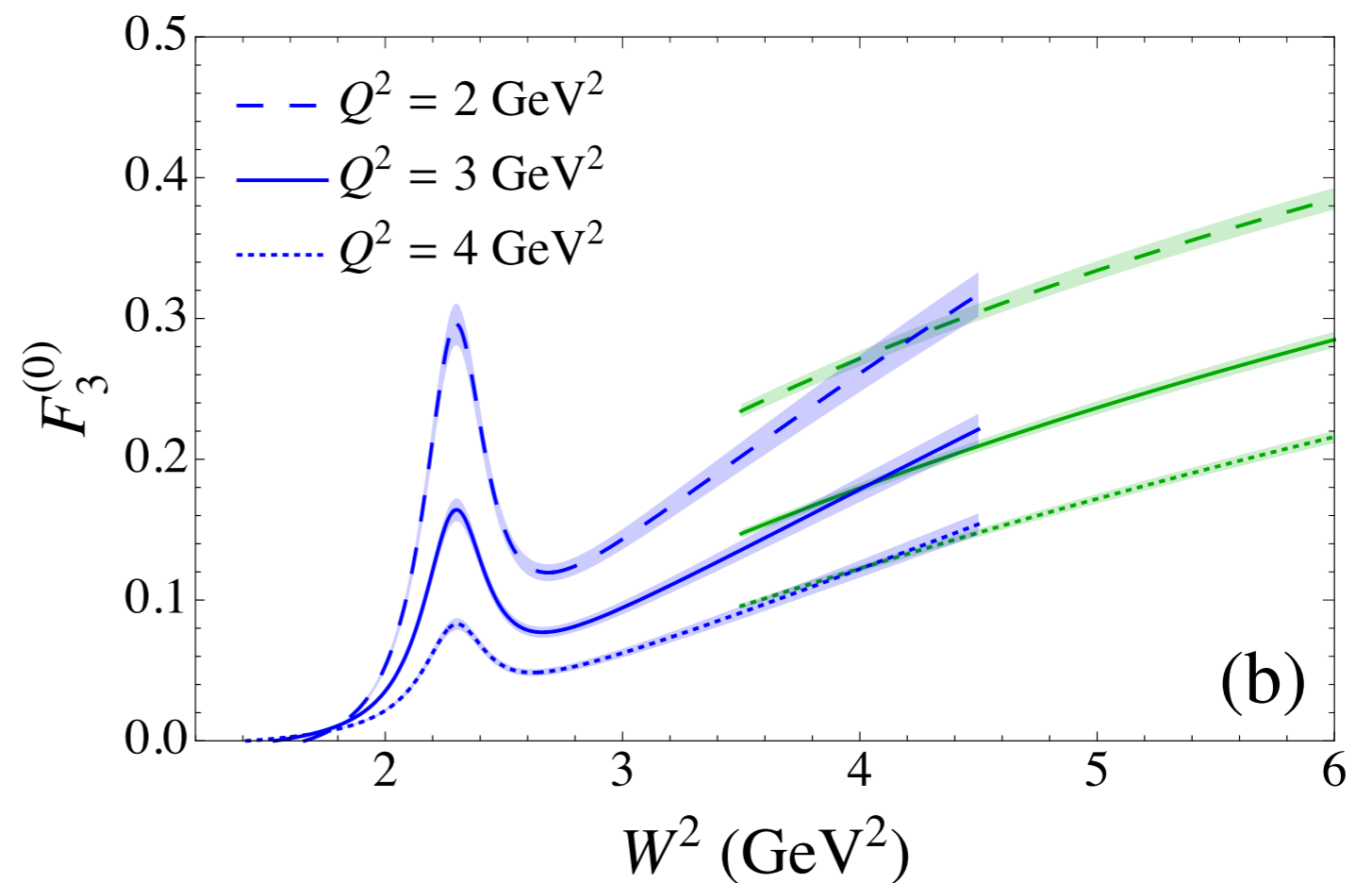
Matching CB resonant background to Regge and DIS

Shiells, PB & Melnitchouk, PRD **104**, 033101 (2021)



(a)

Match to DIS



(b)

Regge contribution

At low Q^2 and high W^2 , the strong interaction becomes nonperturbative.

Model of Capella et al. from ν - N scattering fits:

$$F_{3(\text{Reg})}^{(0)}(x, Q^2) = A^{p-n} x^{-\alpha_R} (1-x)^c \left(\frac{Q^2}{Q^2 + \Lambda_R^2} \right)^{\alpha_R}$$

The true Q^2 dependence of this structure function is not well-determined by theory.

Match this function to the well-known value in the DIS region around $Q^2 = 2 \text{ GeV}^2$ AND constrain it from available data on F_3^W from $\nu p + \bar{\nu} p$

$$A^{p-n} = A^{p+n} \frac{1}{9} \quad \text{OR} \quad A^{p-n} = A^{p+n} \left. \frac{F_{3(\text{Reg})}^{(0)}}{F_{3(\text{Reg})}^W} \right|_{Q^2=Q_0^2}$$

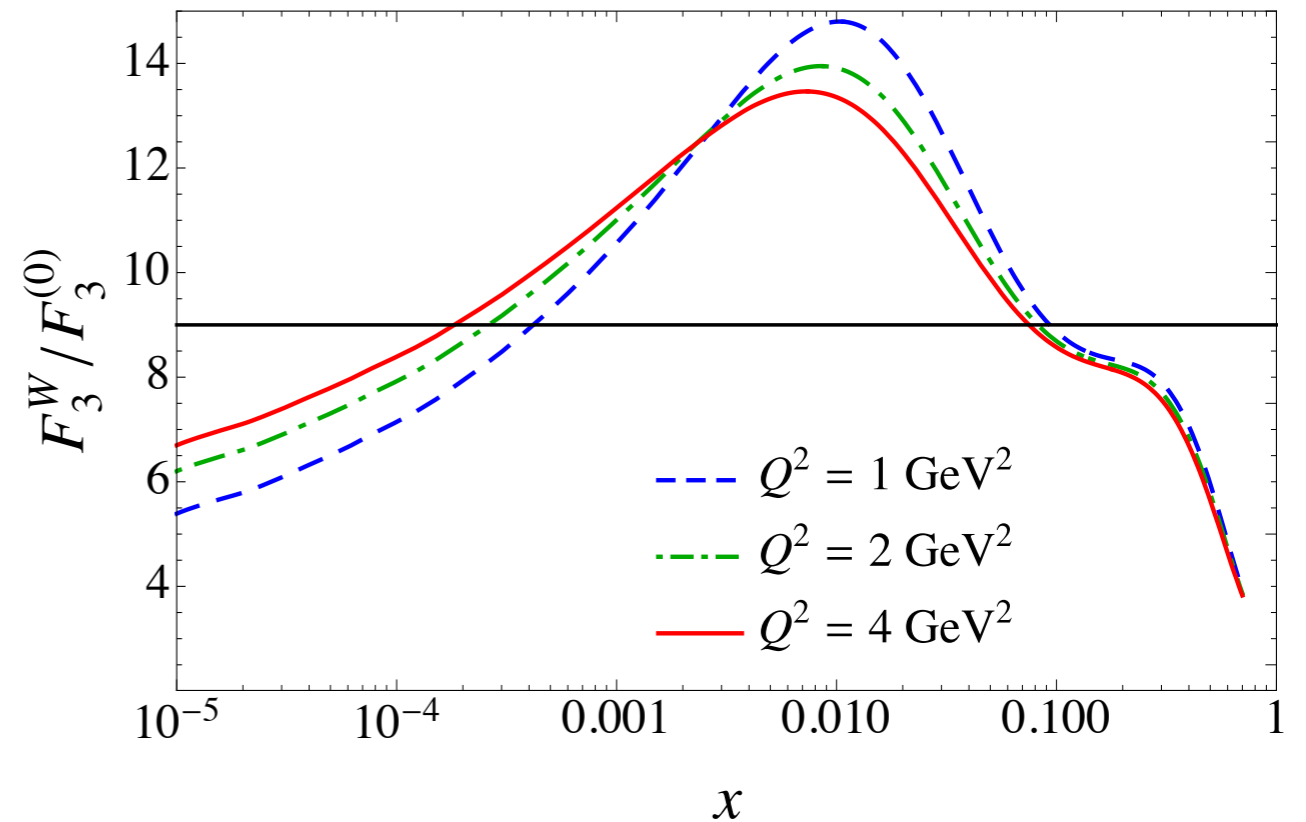
Sum-rule constrained

PDF constrained

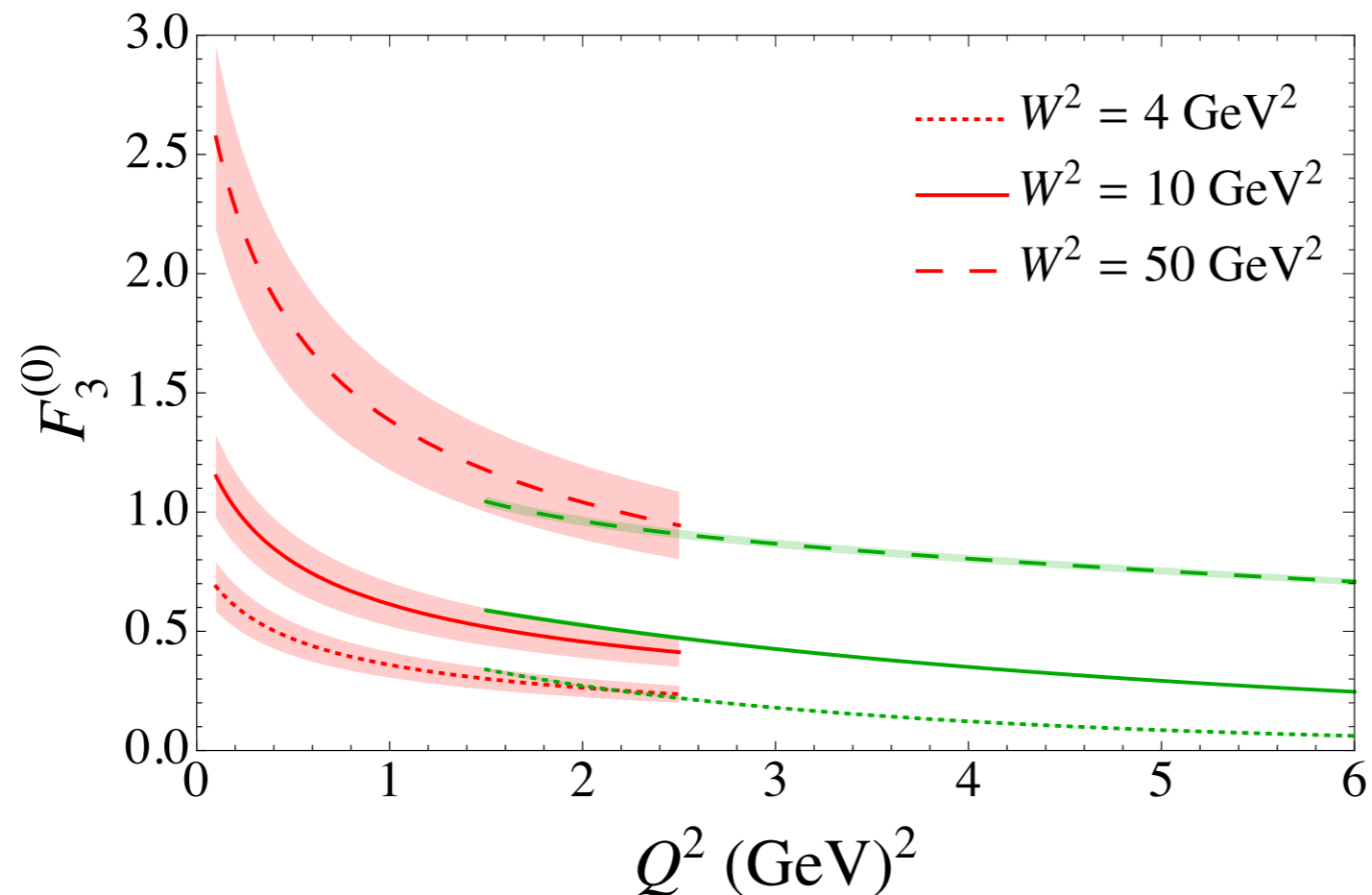
Regge contribution

Is the ratio 9 as a function of x ?

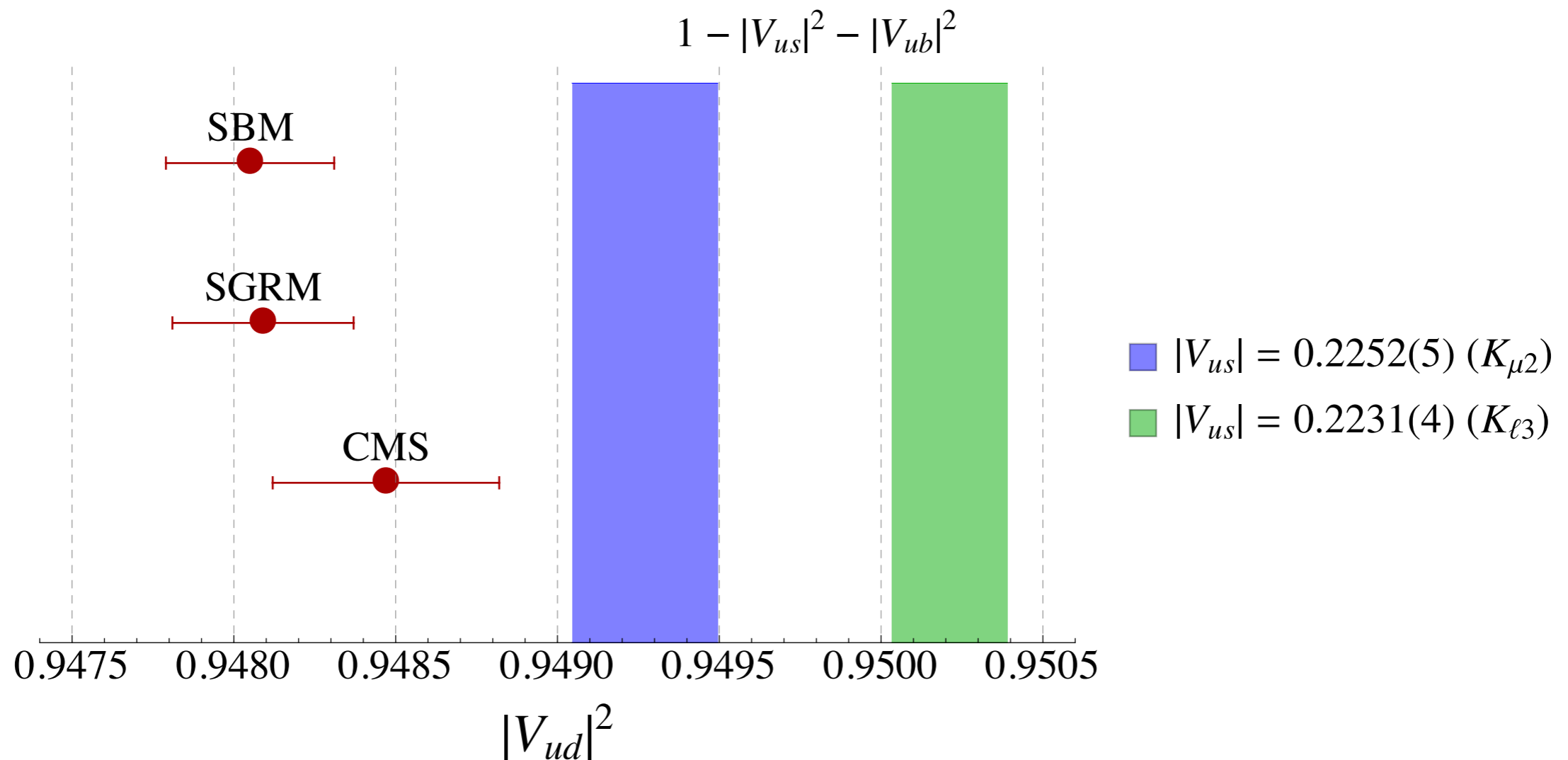
No, but the observable correction is proportional to $\int_0^1 dx F_3(x, Q^2)$ so taking a ratio of 9 is still meaningful.



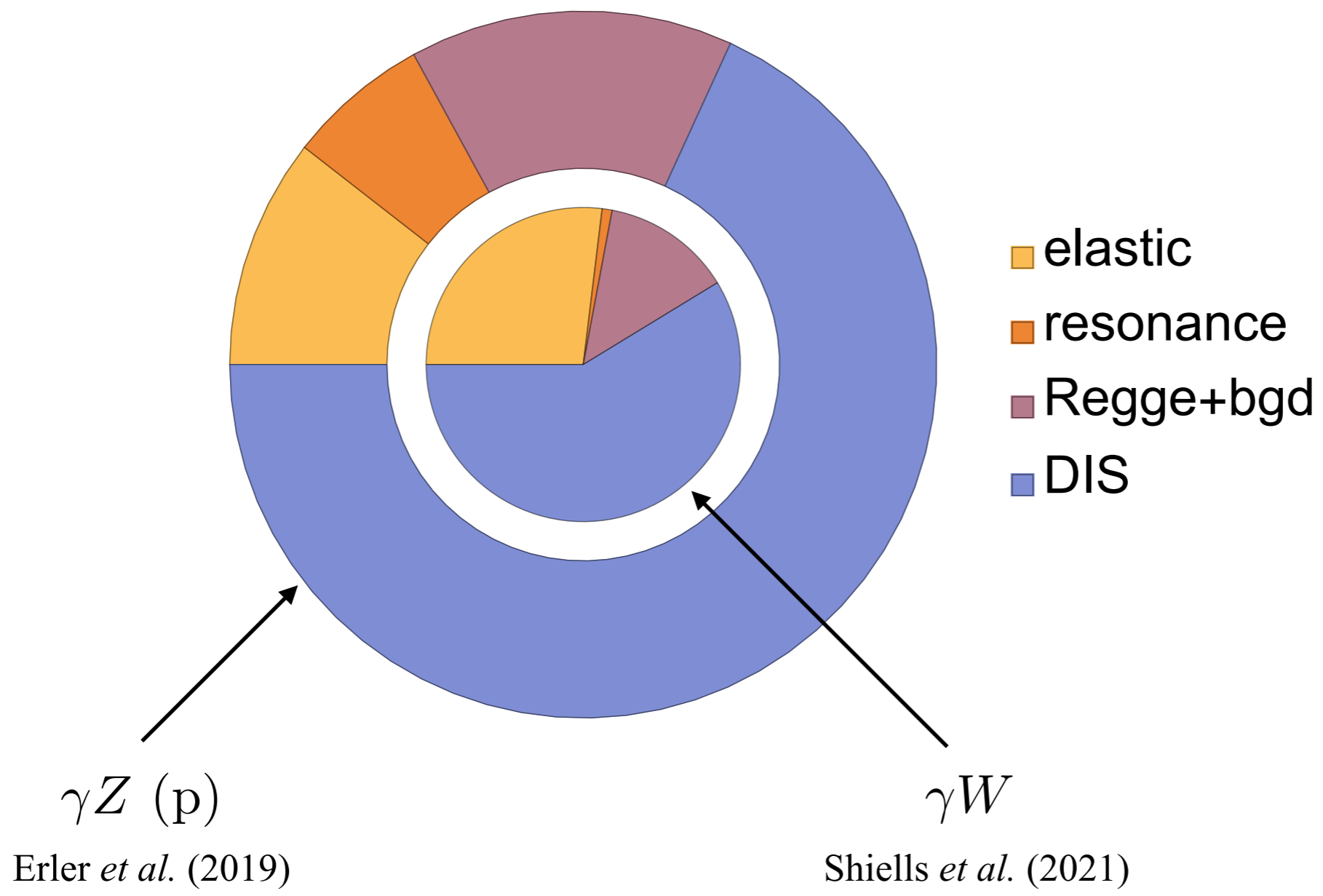
Matching Regge to DIS at low Q^2



$\square_A^{\gamma W} (\times 10^{-3})$	SBM		SGRM [7]	CMS [5]
Q_0^2 [GeV ²]	2.0	1.0	2.0	1.1
Elastic	1.05(4)	1.05(4)	1.06(6)	† 0.99(10)
Resonance	0.04(1)	0.04(1)
DIS + ($Q^2 > Q_0^2$) bgd	2.29(3)	2.43(5)	*2.17(0)	*2.29(2)
Regge + ($Q^2 < Q_0^2$) bgd	0.52(7)	0.39(5)	0.56(8)	0.25(2)
Total	3.90(9)	3.91(9)	3.79(10)	3.52(11)
Δ_R^V	0.02472(18)	0.02474(18)	0.02467(22)	0.02426(32)
$ V_{ud} ^2$	0.94805(26)	0.94803(26)	0.94809(28)	0.94847(35)



Comparison of $\square_A^{\gamma^Z}(0)$ and $\square_A^{\gamma^W}(0)$ contributions from different regions



Summary

- Dispersion approach significant improvement over old methods
- PDF region provides constraints on model-dependence of isospin rotation
- Direct comparison with PV inelastic data in resonance and DIS regions
- *e-d* PVDIS asymmetry strongly constrains the uncertainty in $\square_V^{\gamma Z}$
- checking Δ region at Mainz or JLab would be useful
- Regge and background contributions show largest uncertainties in $\square_A^{\gamma Z}$