

Non-Gaussian Fluctuations in Fluid

Xin An

Based on work with Basar, Stephanov and Yee

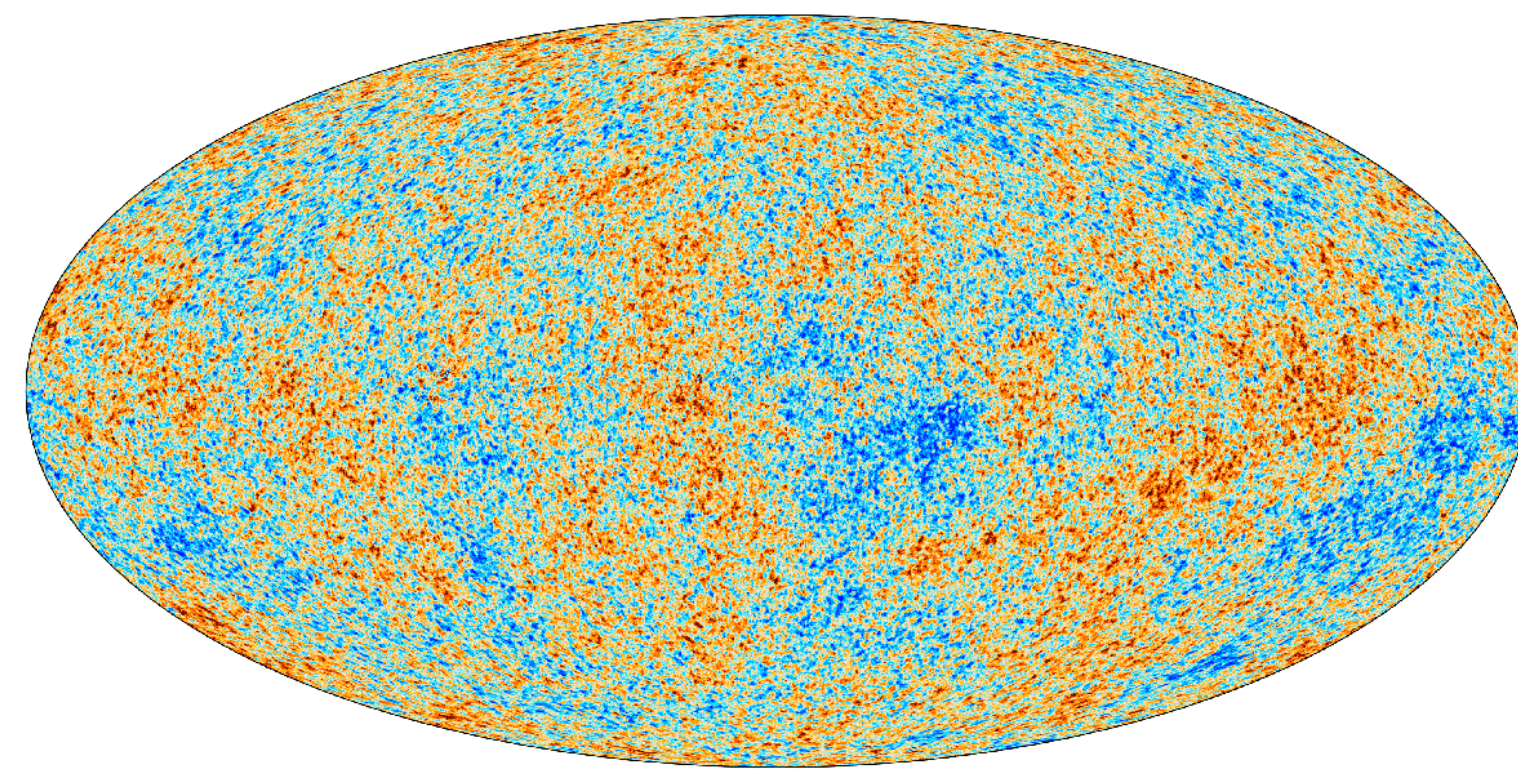
INT workshop on criticality and chirality

Aug 22 2023, Seattle

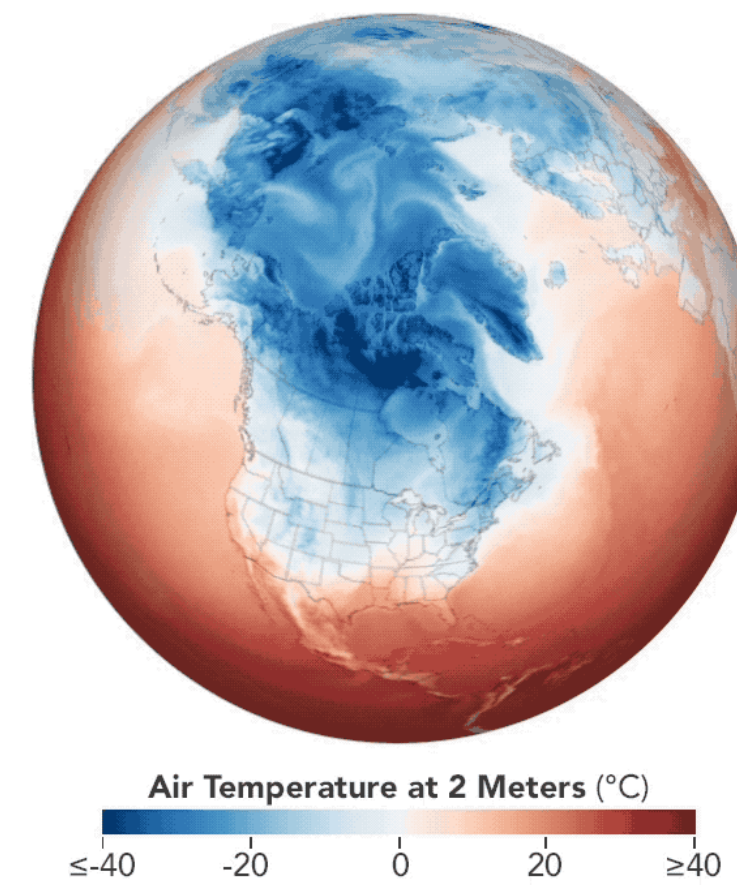


Fluctuations on all length scales

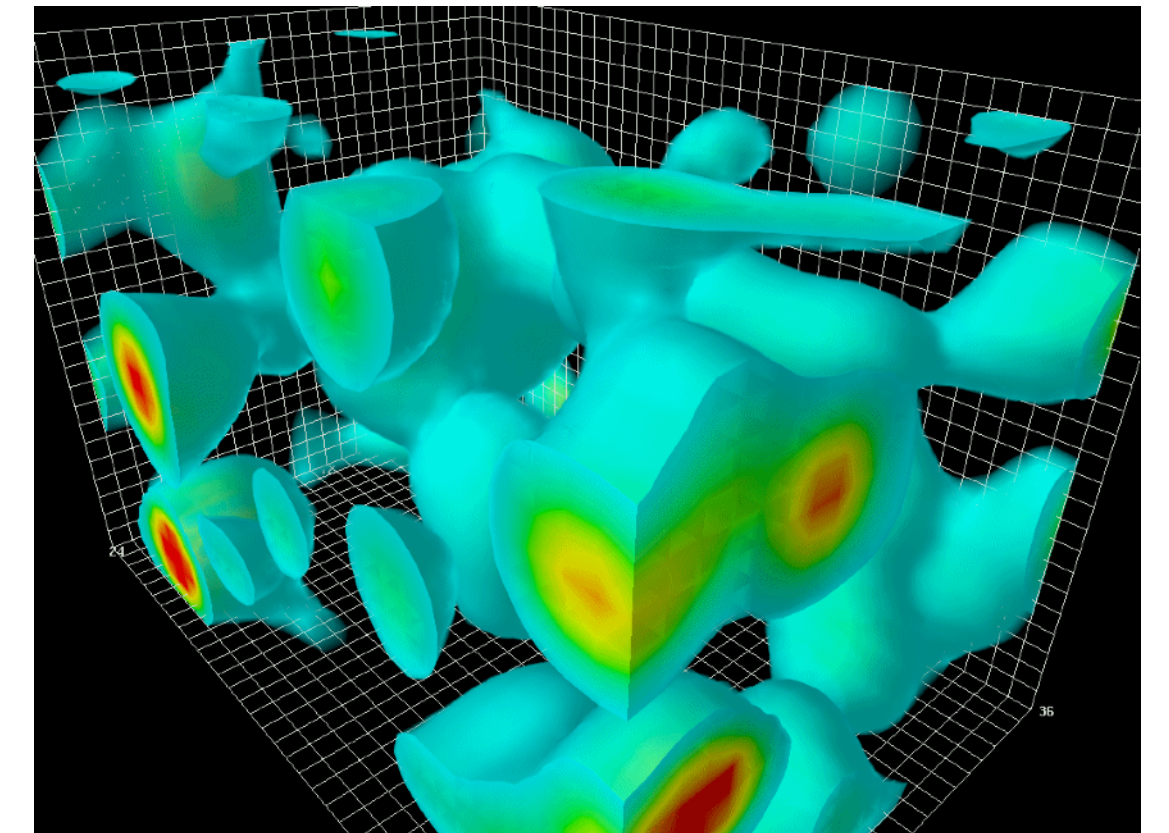
- Fluctuations are ubiquitous phenomena emerging on all length scales.



CMB



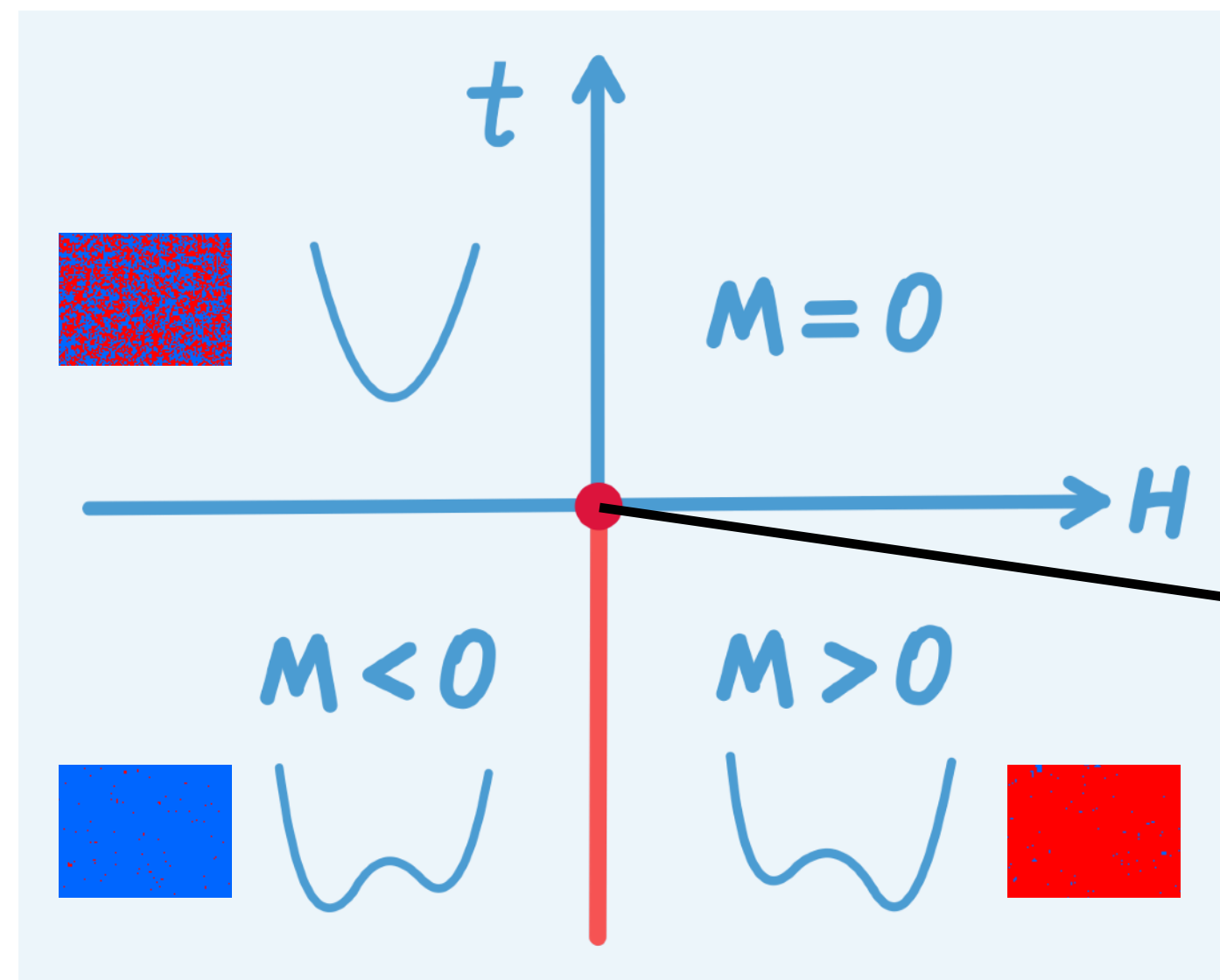
Atmosphere



Quantum fluctuations

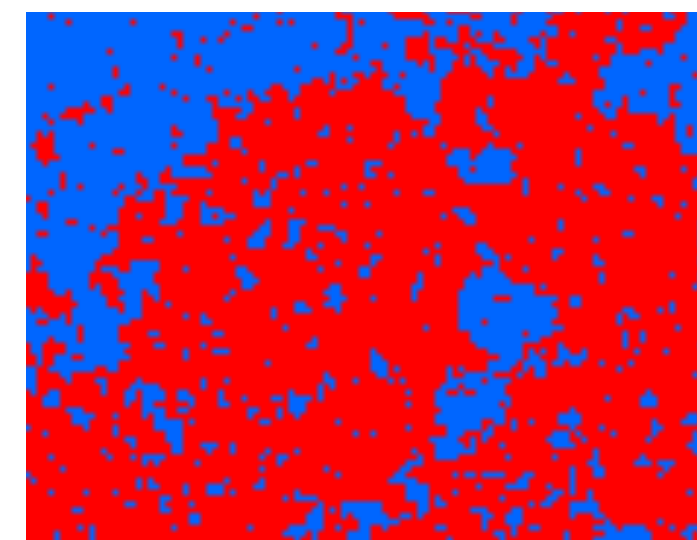
Thermal fluctuations and critical point

- Thermal fluctuations can be described by EOS dealing with *large* number of DOFs *in equilibrium*.
- Fluctuation correlation length ξ diverges at the critical point.



Ising phase diagram

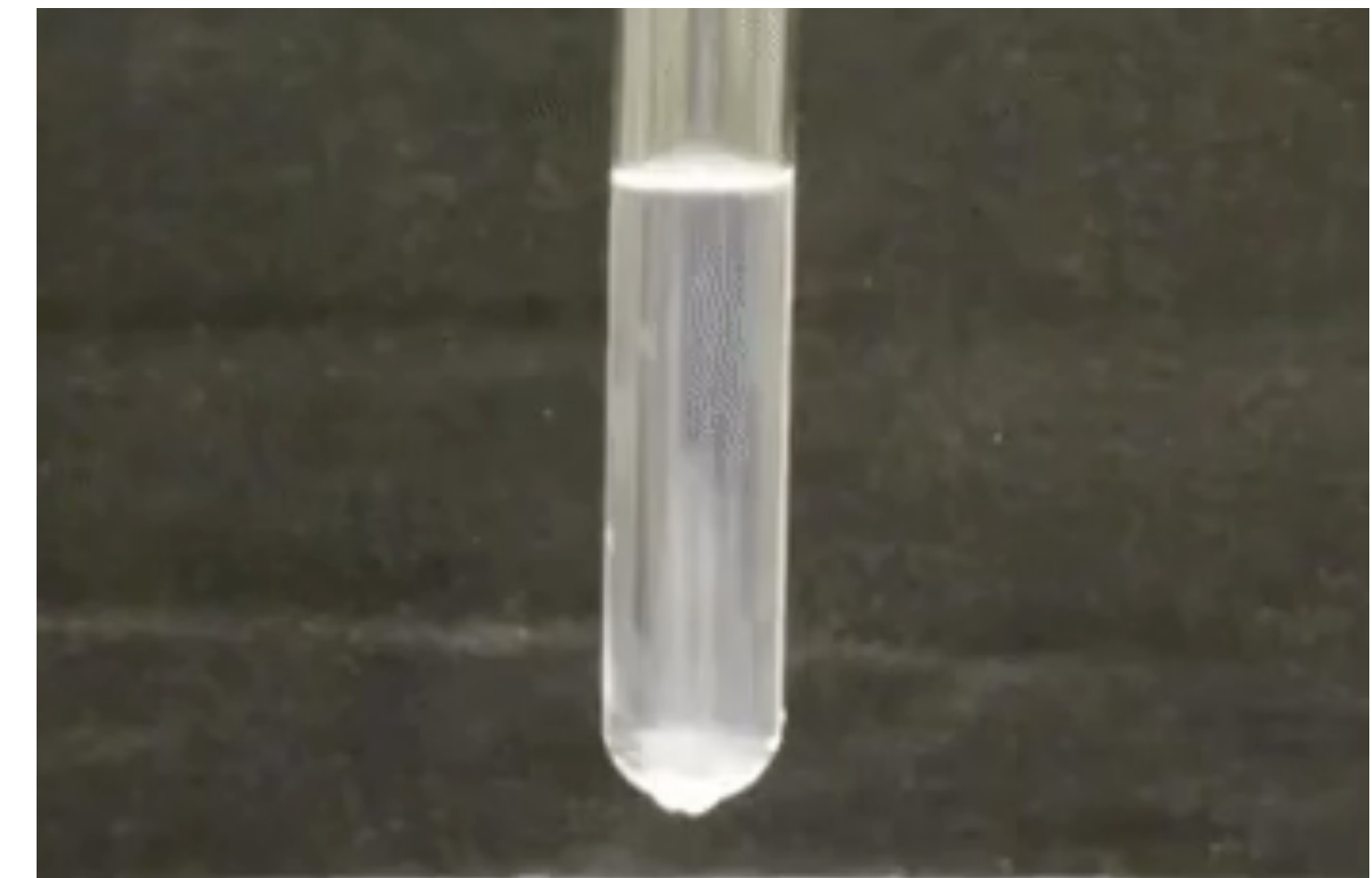
$$\langle \phi(x_1)\phi(x_2) \rangle \sim e^{-\frac{|x_1-x_2|}{\xi}}$$



$\xi \rightarrow \infty$



Landau Kadanoff Wilson



Critical opalescence: $\xi \leftrightarrow \lambda_{\text{light}}$

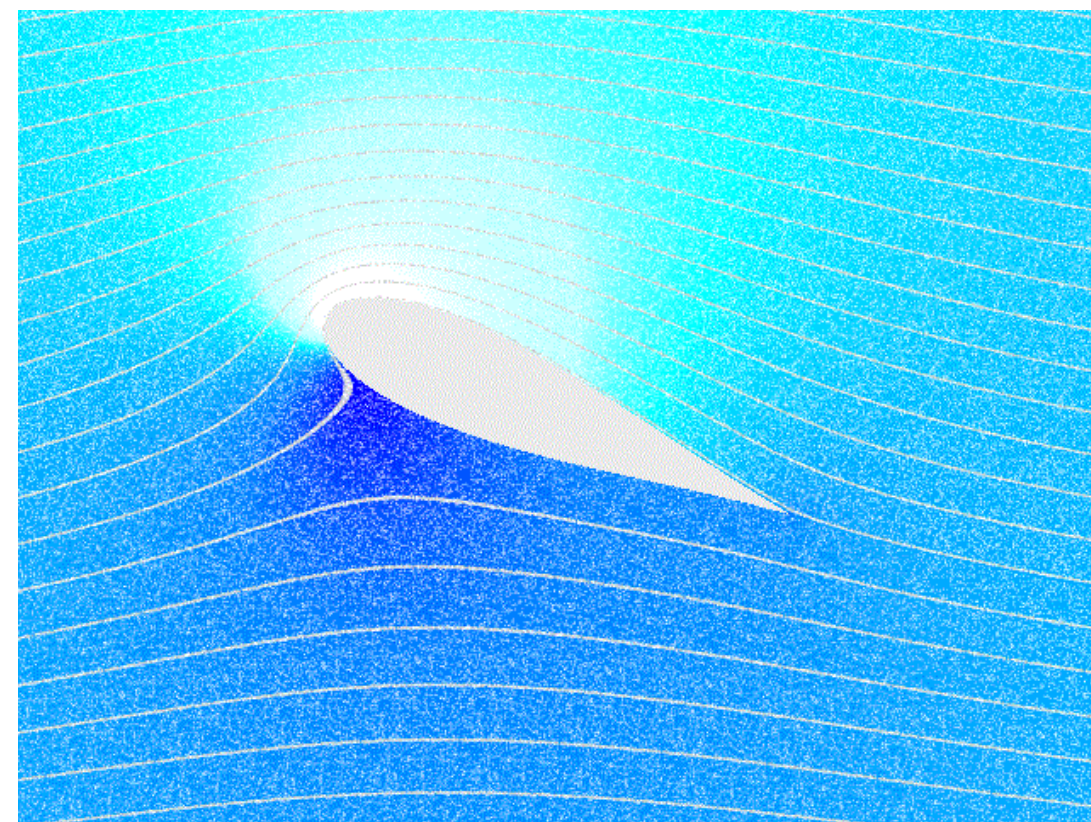
See also Skokov's talk

Fluctuations in hydrodynamics

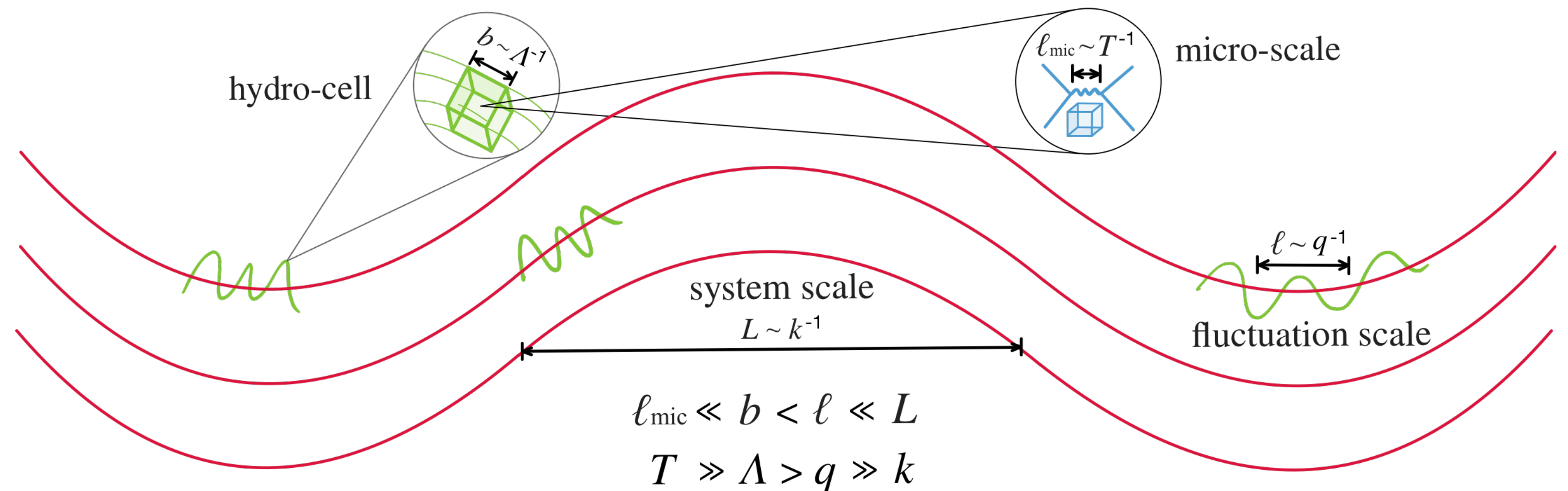
- Hydrodynamic fluctuations: *locally thermalized* thermal fluctuations comoving with fluid, may *not equilibrate* at large scales, described by a set of conservation equations

$$\partial_t \psi = \nabla \cdot (J[\psi])$$

\swarrow \searrow
 $\psi = (n, \epsilon, \pi_i)$ conserved current

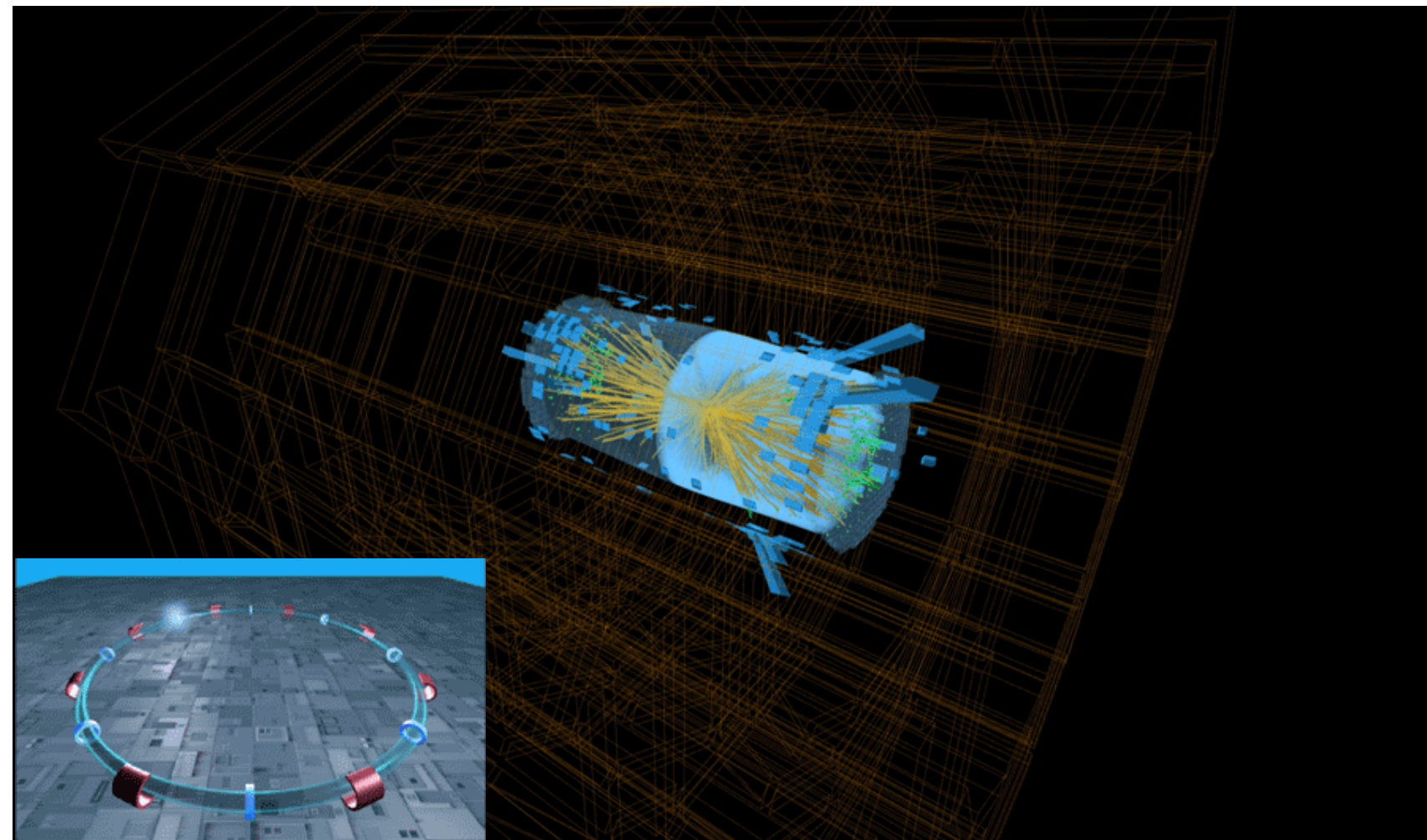


Flow around a wing
(Wikipedia)

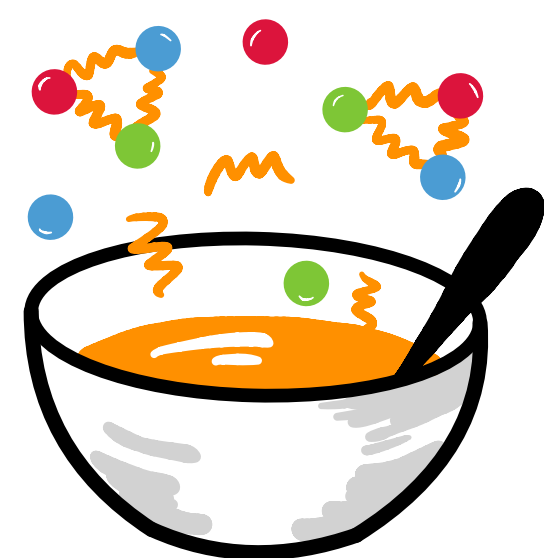


Hydrodynamic description of QGP

- Quark-Gluon Plasma in heavy-ion collisions: *small* enough for fluctuations to be important; and *large* enough for hydrodynamics to be applicable.

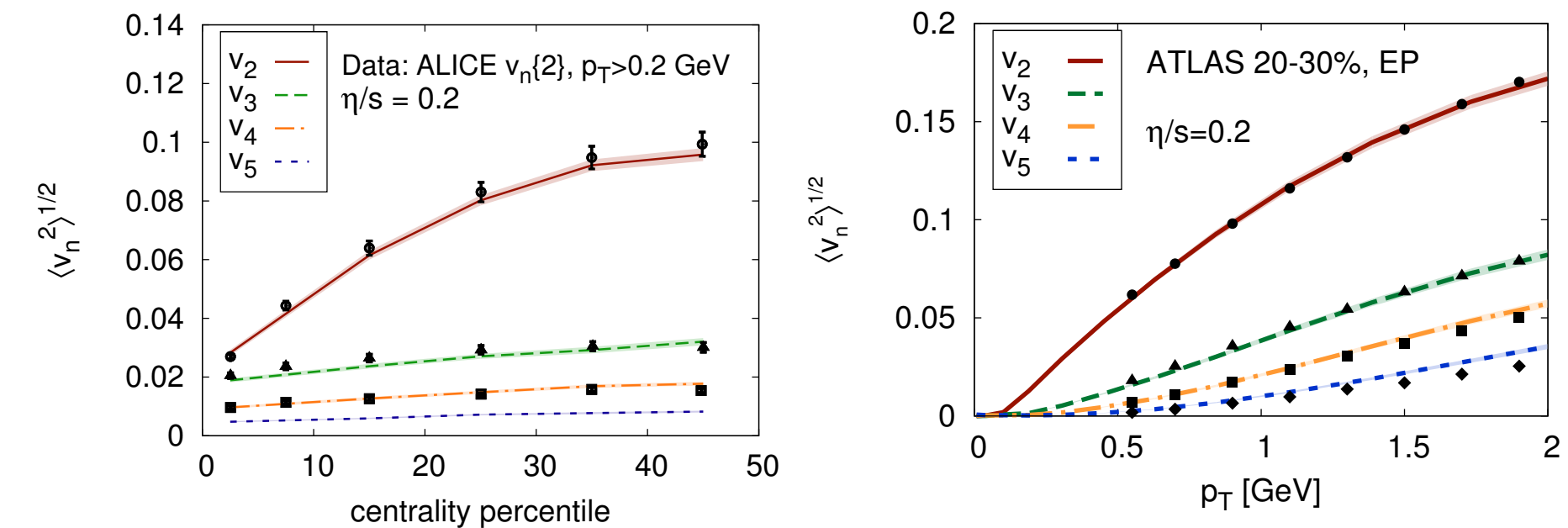


Collision event simulation at LHC (CERN)

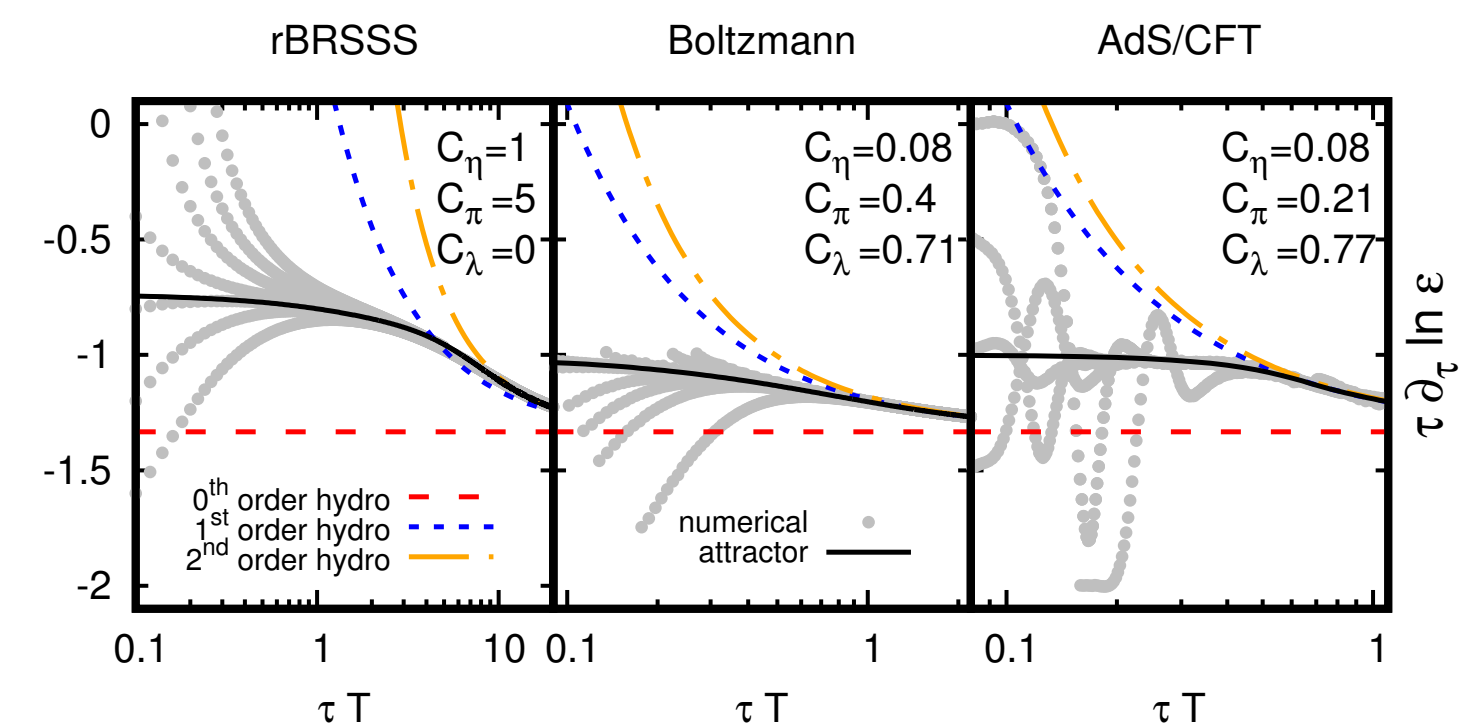


Size of created fire balls ~ 10 fm
 Number of particles $\sim 10^2 - 10^4$

Early Universe is more like a soup



Analysis for flow collectivity manifests
 QGP as a *perfect fluid* [C. Gale et al, 1301.5893](#)

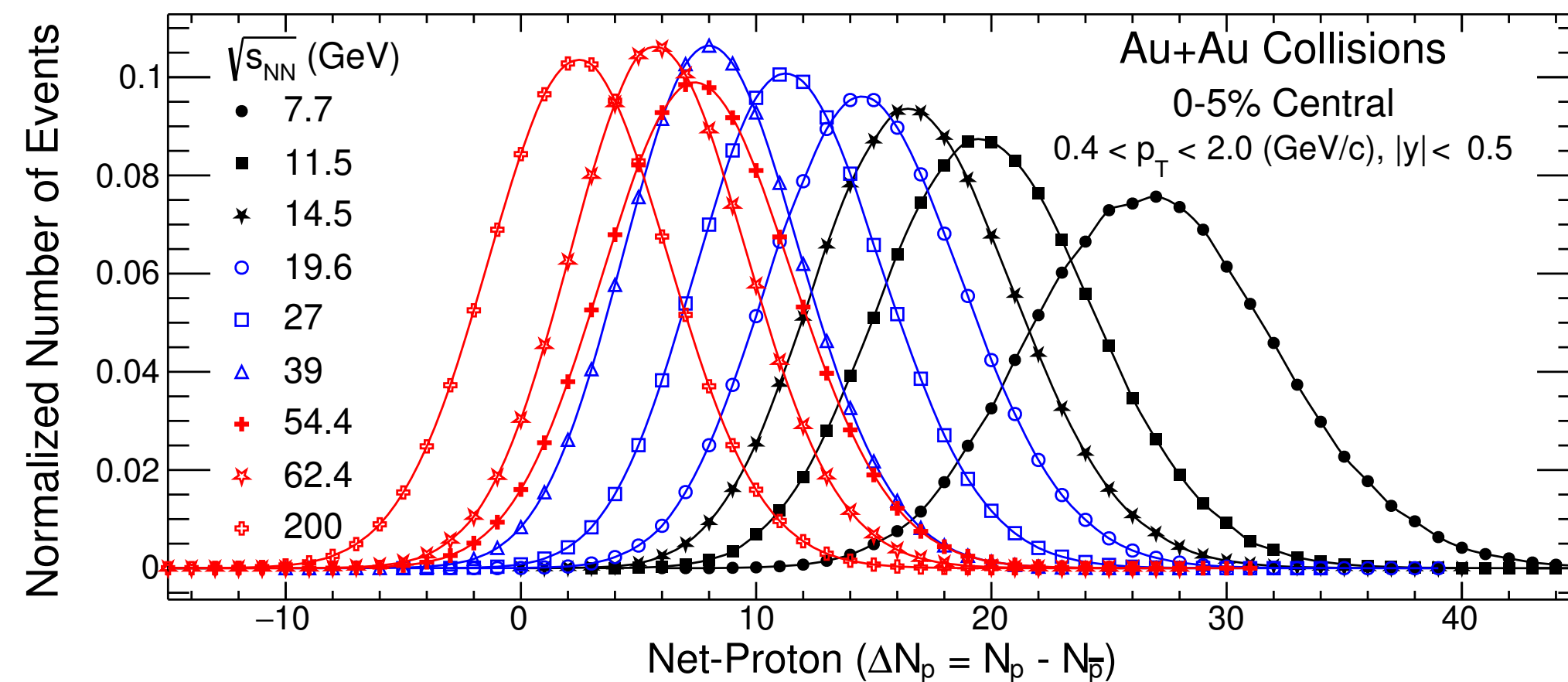


Hydrodynamic *attractor* [W. Florkowski et al, 1707.02282](#),
[P. Romatschke et al, 1712.05815](#)

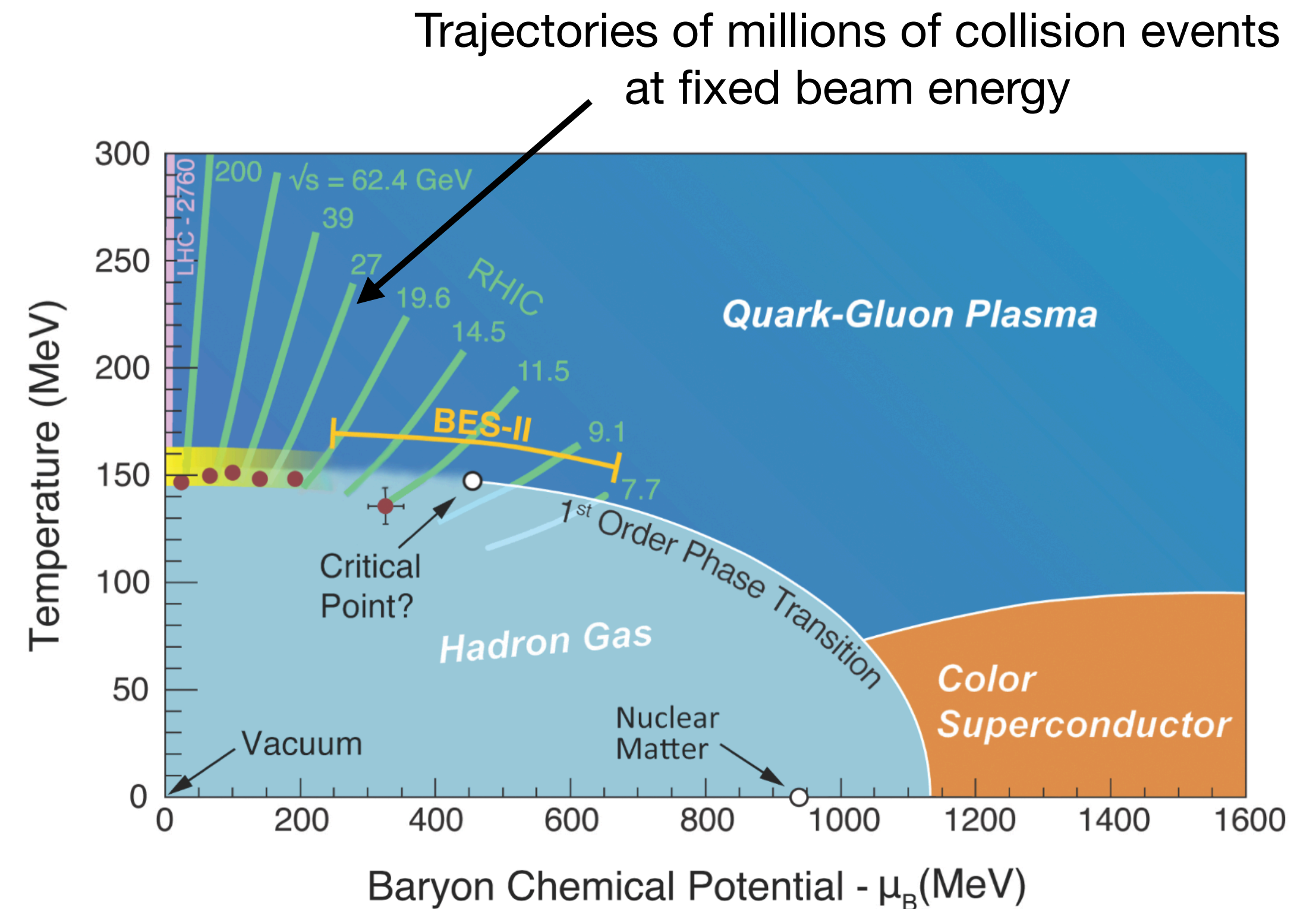
Stochastic process in HIC

- Observables fluctuate event-by-event.
- Beam Energy Scan: 7.7-200 GeV (phase I), and 3.3-19.6 GeV (phase II).

One collision events



Events number vs net-proton yields STAR

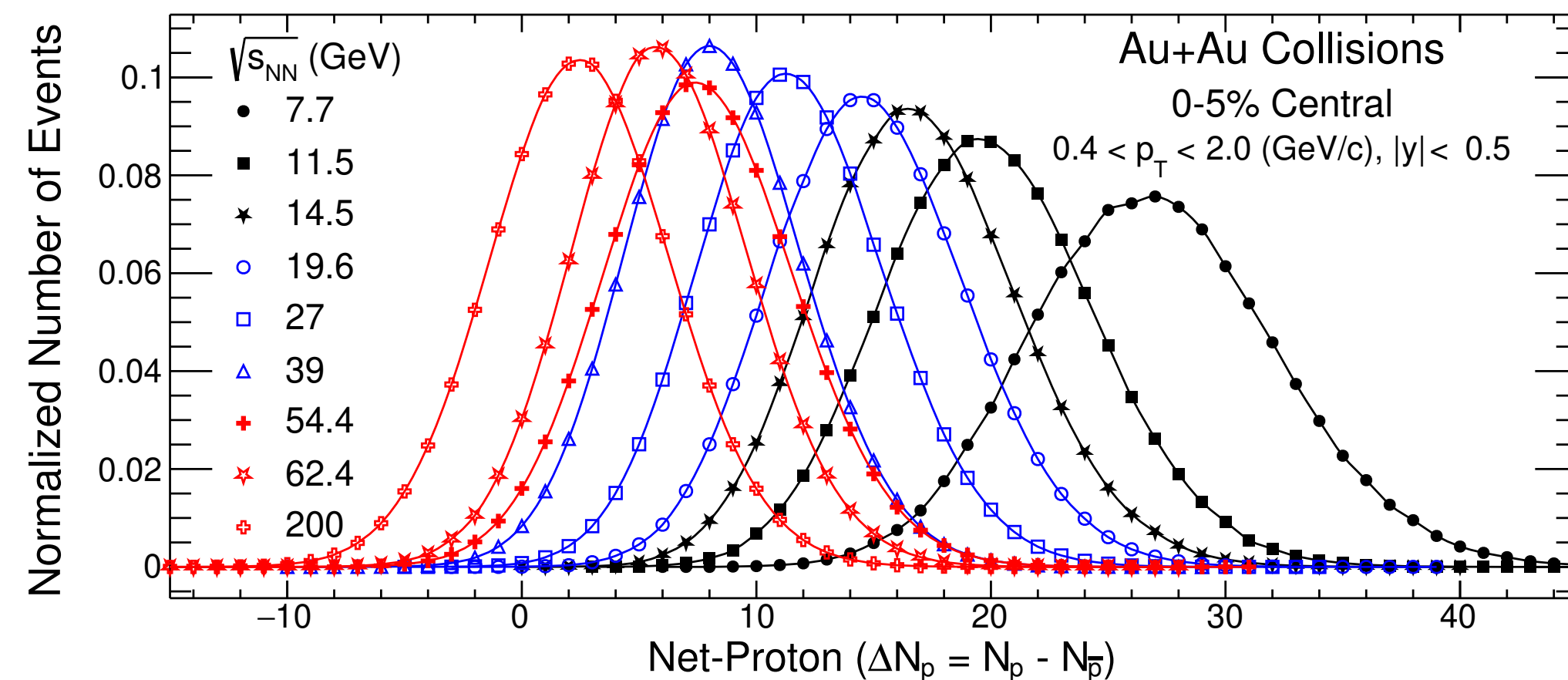
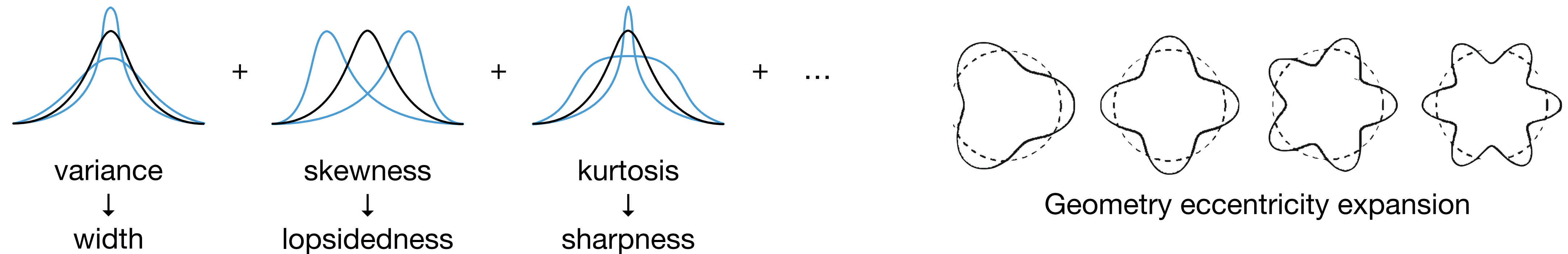


Trajectories of millions of collision events at fixed beam energy

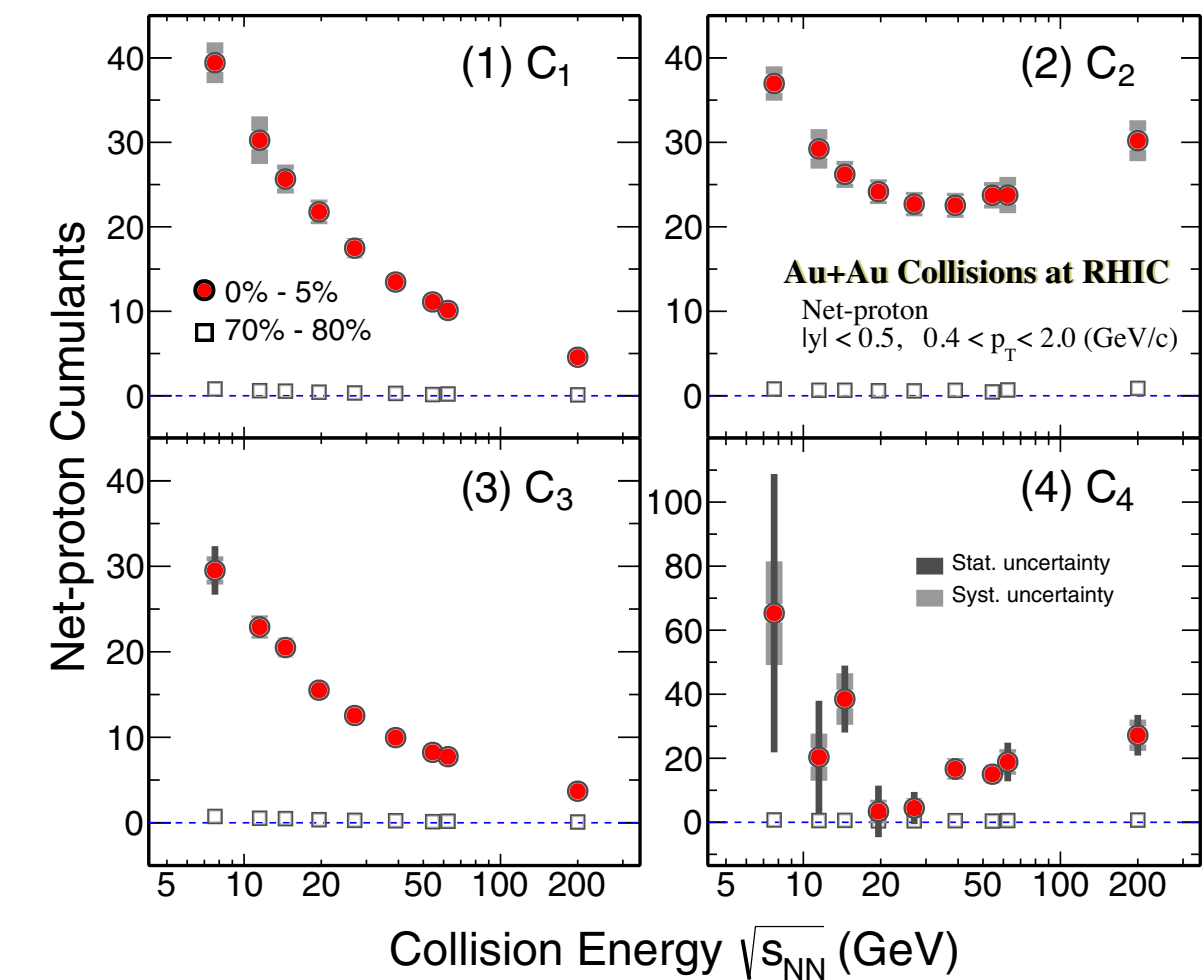
QCD phase diagram

Probability distribution for fluctuating variables

- Fluctuations can be described by probability distribution $P[\psi]$ and its associated *cumulants* (or correlation functions).



Events number vs net-proton yields STAR

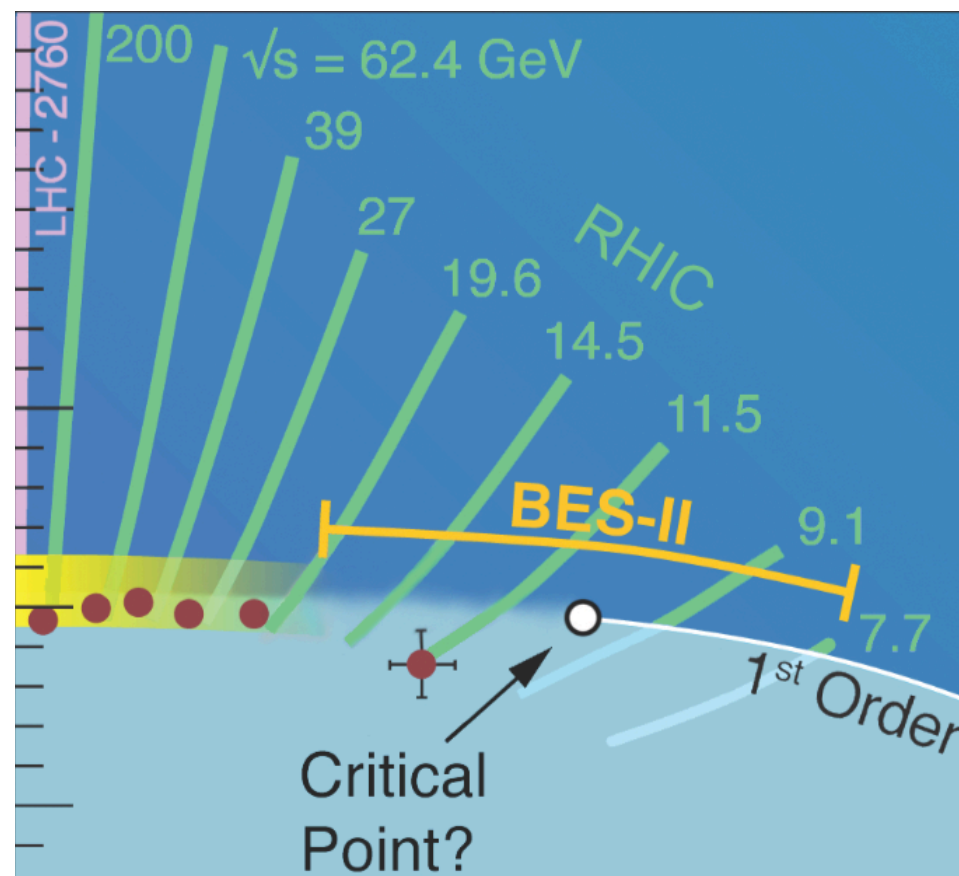


Net-proton cumulants vs energy

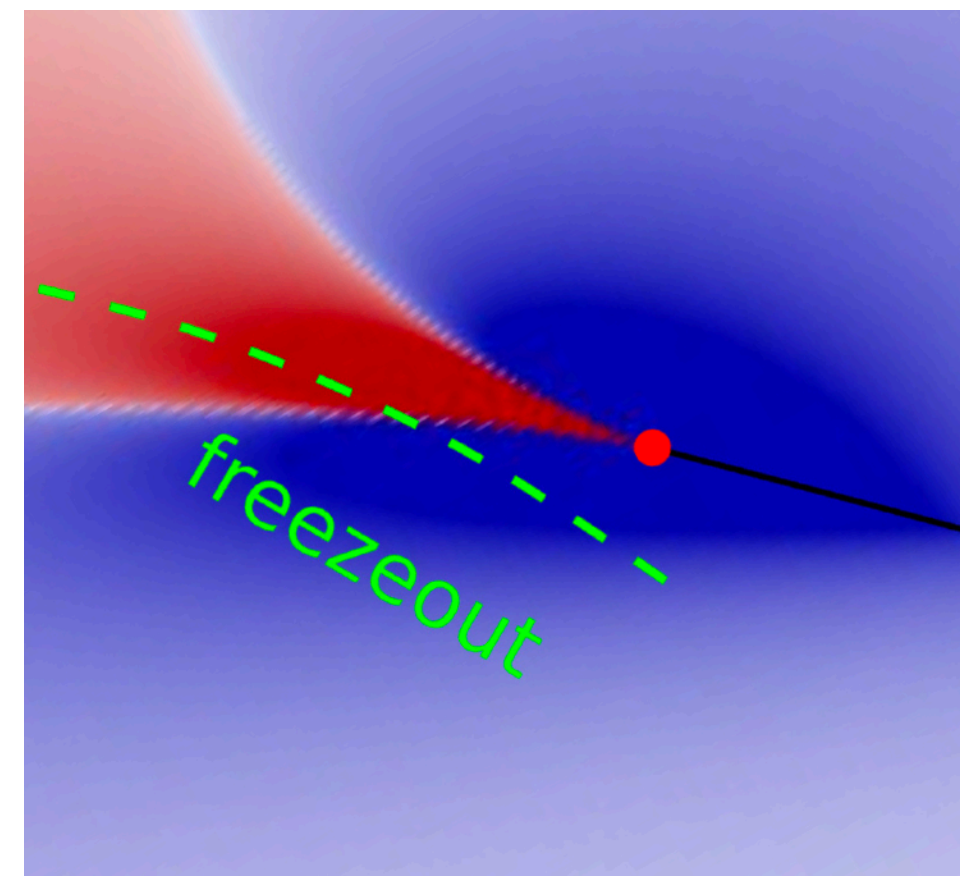
Cumulants and critical point

- In thermodynamics, higher-order cumulants are more sensitive to the correlation length ξ which diverges at critical point. [Stephanov 0809.3450, 1104.1627](#)

$$n\text{-th cumulants} \sim \xi^{\frac{5n-6}{2}}$$

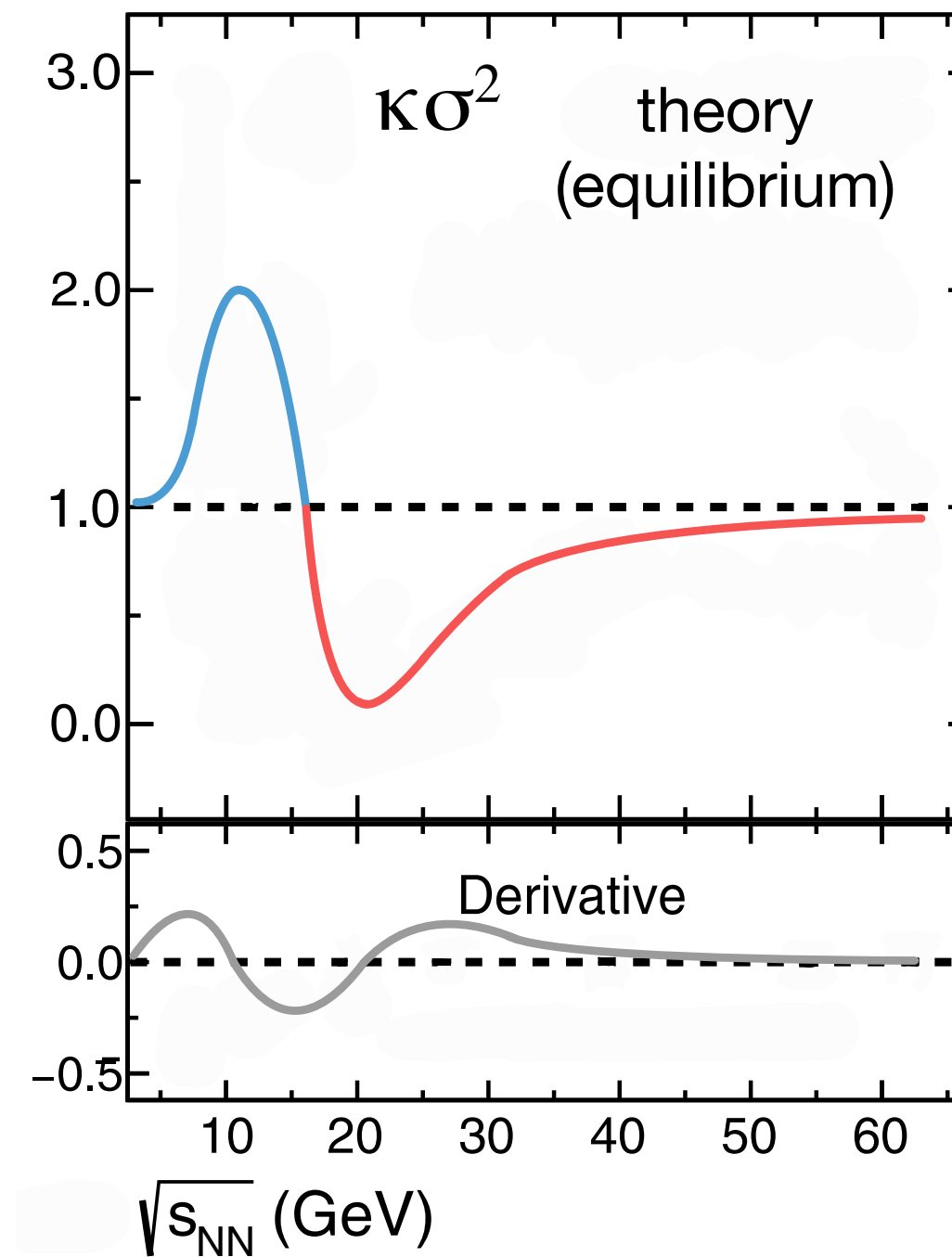


Critical point region



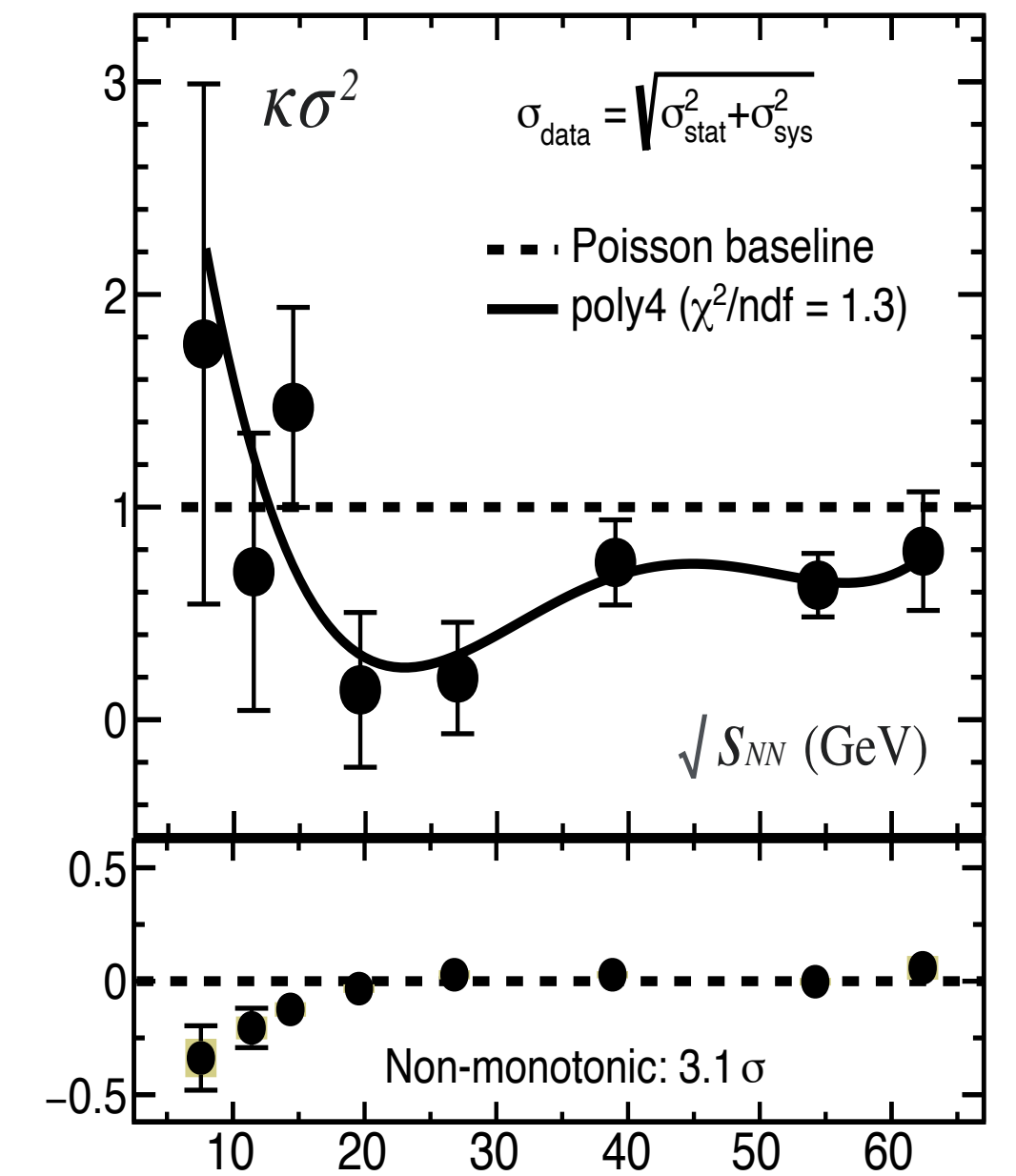
4th cumulant contours

Non-monotonic energy dependence



$\sqrt{s_{NN}}$ (GeV)

Theoretical prediction



Experimental result
(STAR BES-I)

Various theoretical approaches

- **PDEs** (bottom-up like)

Starting from phenomenological equations with required properties

e.g., Langevin equations in *stochastic* description, Fokker-Planck (FP) equations in *deterministic* description.

[Akamatsu et al, 1606.07742](#)

[Nahrgang et al, 1804.05728](#)

[Singh et al, 1807.05451](#)

[Chattopadhyay et al, 2304.07279](#)

...

- **EFTs** (top-down like)

Starting from effective action with first principles

e.g., Martin-Siggia-Rose (MSR), Schwinger-Keldysh (SK), Hohenberg-Halperin (HH), n-particle irreducible (nPI), etc.

[Glorioso et al, 1805.09331](#)

[Jain et al, 2009.01356](#)

[Sogabe et al, 2111.14667](#)

[Chao et al, 2302.00720](#)

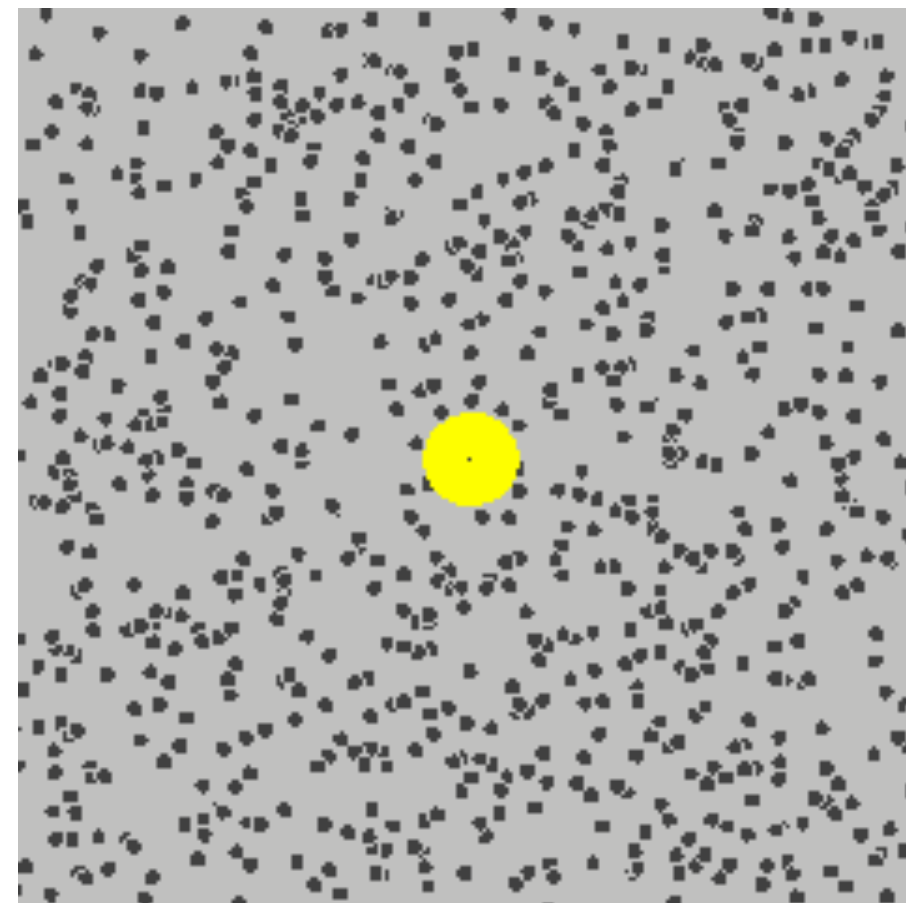
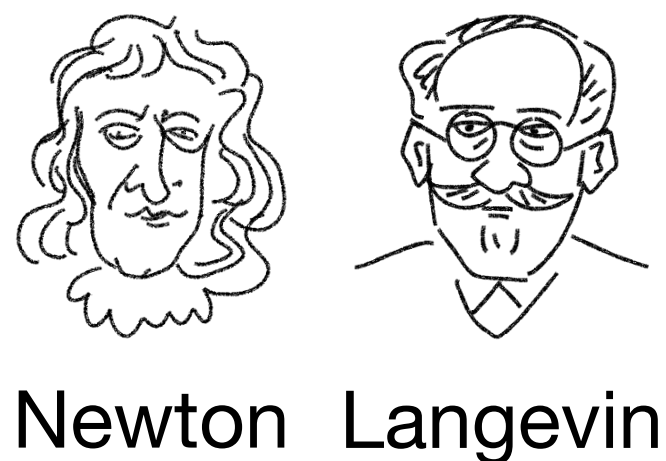
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Stochastic and deterministic description

- Langevin equation
(Newton's equation + noise):

$$\partial_t \check{\psi}_i = F_i[\check{\psi}] + \eta_i$$

$$\langle \eta_i(x_1) \eta_j(x_2) \rangle = 2Q_{ij} \delta^{(3)}(x_1 - x_2)$$

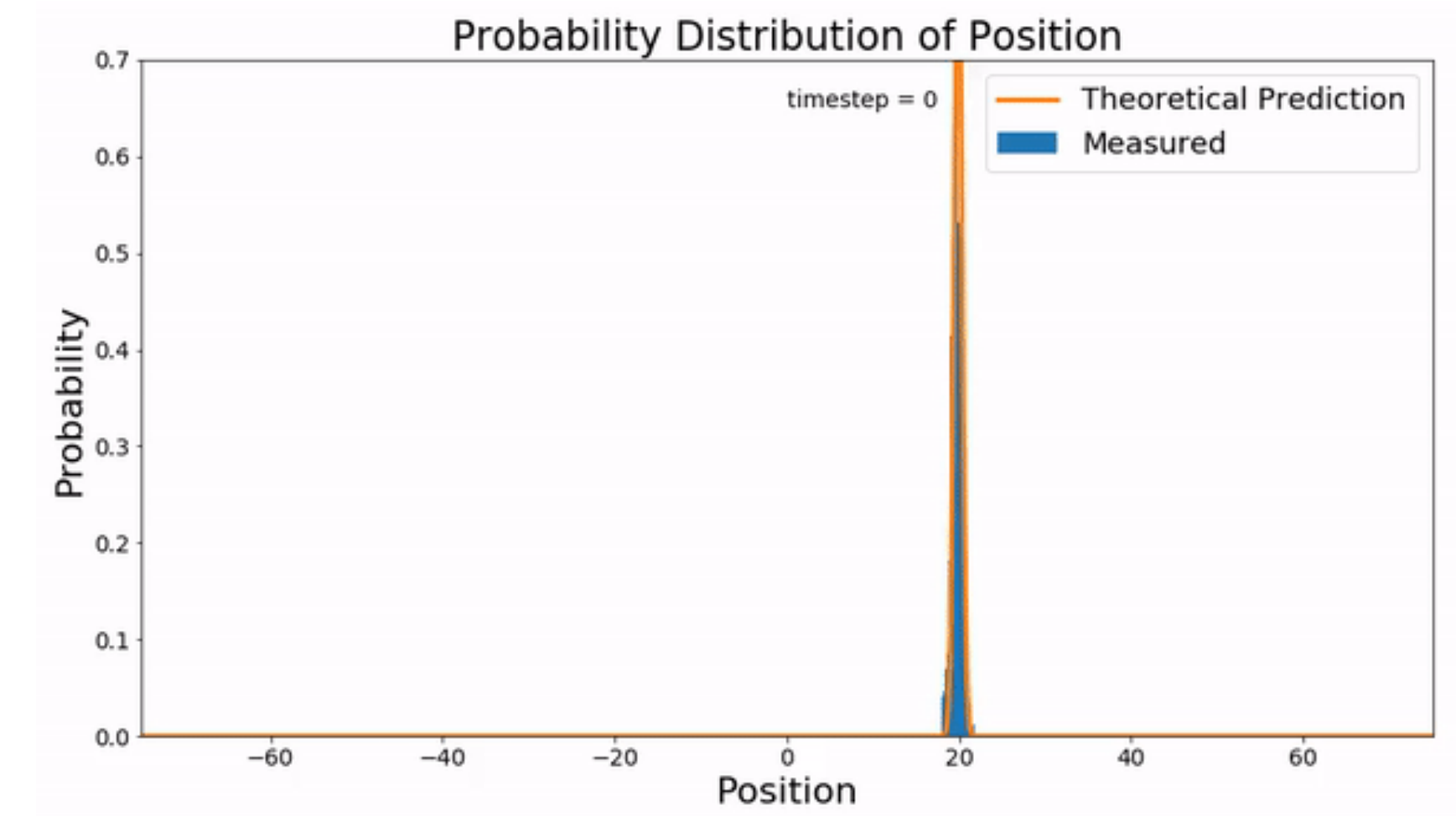
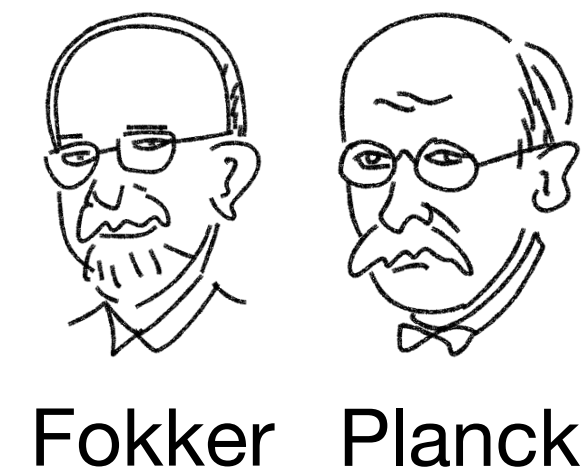


Brownian motion

- Fokker-Planck equation
(probability evolution equation):

$$\partial_t P[\psi] = \frac{\partial}{\partial \psi} (\mathbf{flux}[\psi])$$

$$\mathbf{flux}[\psi] = -FP + \frac{\partial}{\partial \psi} (QP)$$



(Wikipedia)

Correlator evolution equations

- Evolution equations for n-pt correlators $G_n \equiv \langle \underbrace{\phi \dots \phi}_n \rangle \equiv \int d\psi P[\psi] \underbrace{\phi \dots \phi}_n$
 where $\phi \equiv \psi - \langle \psi \rangle$: [XA et al, 2009.10742, 2212.14029](#)

$$\partial_t G_n = \mathcal{F} [\psi, G_2, G_3, \dots, G_n, G_{n+1}, \dots, G_\infty]$$

(with all combinatorial configurations)
 useful for higher order cumulants

$$G_{ij\dots} \equiv \text{vertex with dot} \quad F_i \equiv \text{D-shape} \quad F_{i,j\dots} \equiv \text{D-shape with lines} \quad Q_{ij} \equiv \text{triangle} \quad Q_{ij,k\dots} \equiv \text{triangle with lines}$$

Stochastic:
 One equation
 Millions of samples



Deterministic:
 One sample
 Millions of equations

Truncated evolution equations

- Power counting in loop expansion parameters $\varepsilon \sim (q\xi)^3$ (inverse number of uncorrelated cells in fluctuation scales): [XA et al, 2009.10742](#)

$$G_n \sim \varepsilon^{n-1}, \quad F \sim 1, \quad Q \sim \varepsilon.$$

- Correlator evolution equations can be truncated and iteratively solved:

$$\partial_t G_n = \mathcal{F} [\psi, G_2, G_3, \dots, G_n] + \mathcal{O}(\varepsilon^n)$$

$$\left(\text{---} \bullet \right)' = \text{---} \text{D} + \text{---} \text{D} \text{---} \bullet$$

conventional hydro equations *one loop (renormalization & long-time tails)*

$$\left(\text{---} \bullet \text{---} \right)' = \text{---} \text{D} \bullet \text{---} + \text{---} \triangle \text{---}$$

$$\left(\begin{array}{c} | \\ \bullet \\ | \end{array} \right)' = \text{---} \text{D} \bullet \begin{array}{c} / \\ | \\ \backslash \end{array} + \text{---} \text{D} \begin{array}{c} \bullet \\ / \\ \backslash \end{array} + \begin{array}{c} \triangle \\ | \\ \bullet \end{array}$$

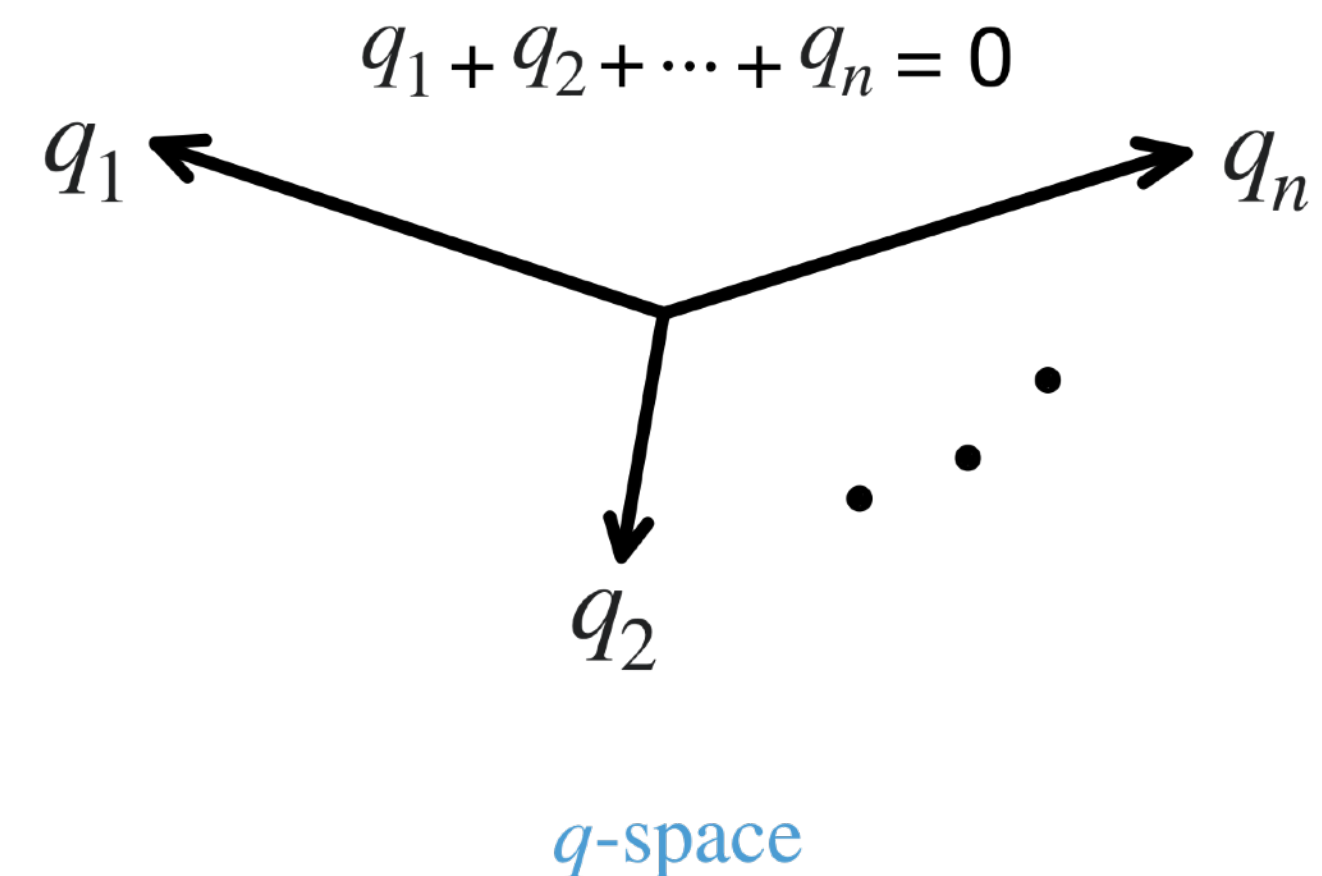
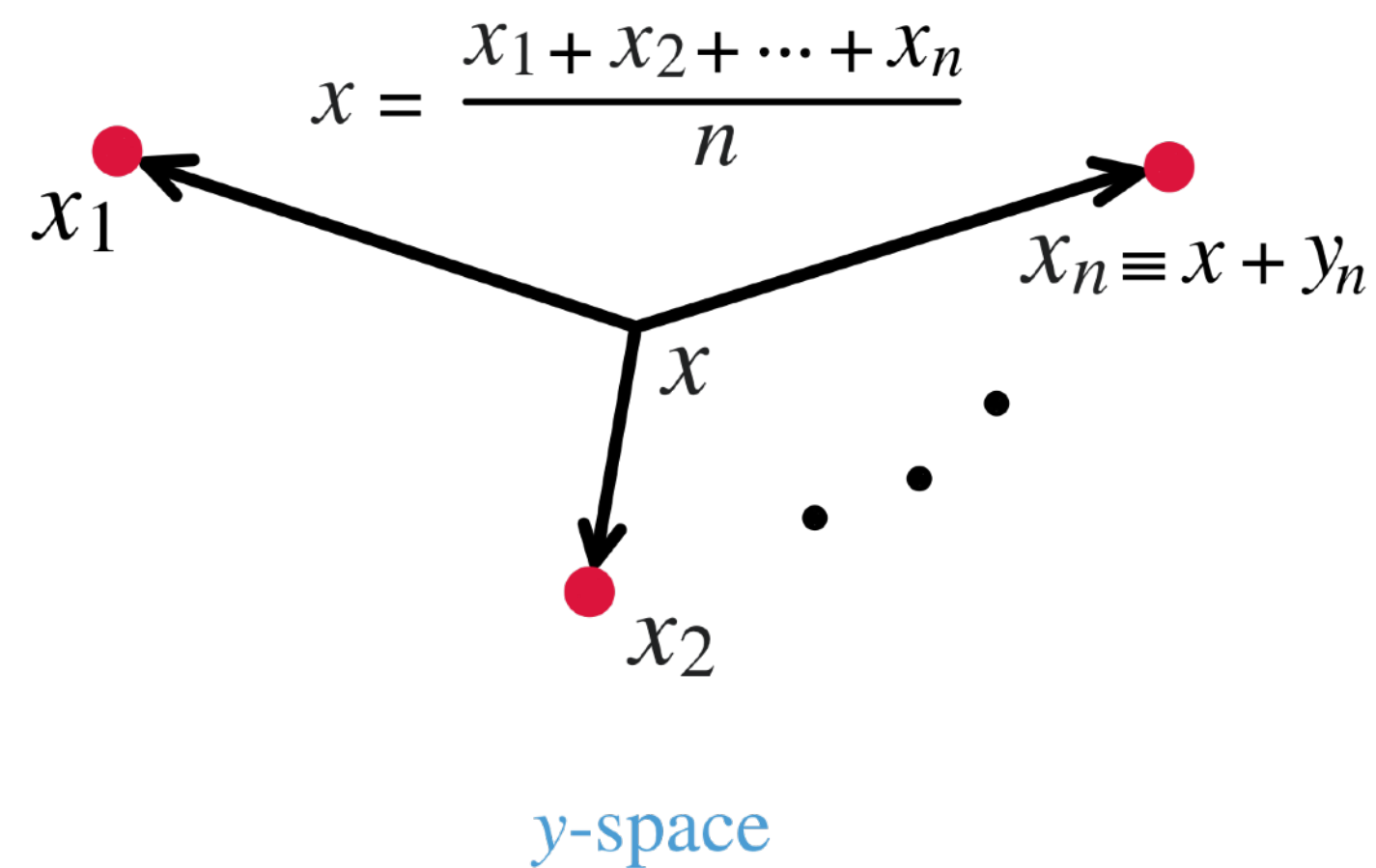
correlator evolution equations

$$\left(\begin{array}{c} | \\ \bullet \\ | \end{array} \right)' = \text{---} \text{D} \bullet \text{---} + \text{---} \text{D} \begin{array}{c} \bullet \\ / \\ \backslash \end{array} + \begin{array}{c} \bullet \\ | \\ \text{D} \\ | \\ \bullet \end{array} + \begin{array}{c} \triangle \\ | \\ \bullet \end{array} + \begin{array}{c} \triangle \\ / \backslash \\ \bullet \bullet \end{array}$$

Multi-point Wigner function

- For fluctuation fields, we introduced n -pt Wigner function [XA et al, 2009.10742](#)

$$W_n(x; q_1, \dots, q_n) = \int d^3y_1 \dots d^3y_n e^{-i(q_1 y_1 + \dots + q_n y_n)} \delta^{(3)}\left(\frac{y_1 + \dots + y_n}{n}\right) G_n(x; y_1, \dots, y_n)$$



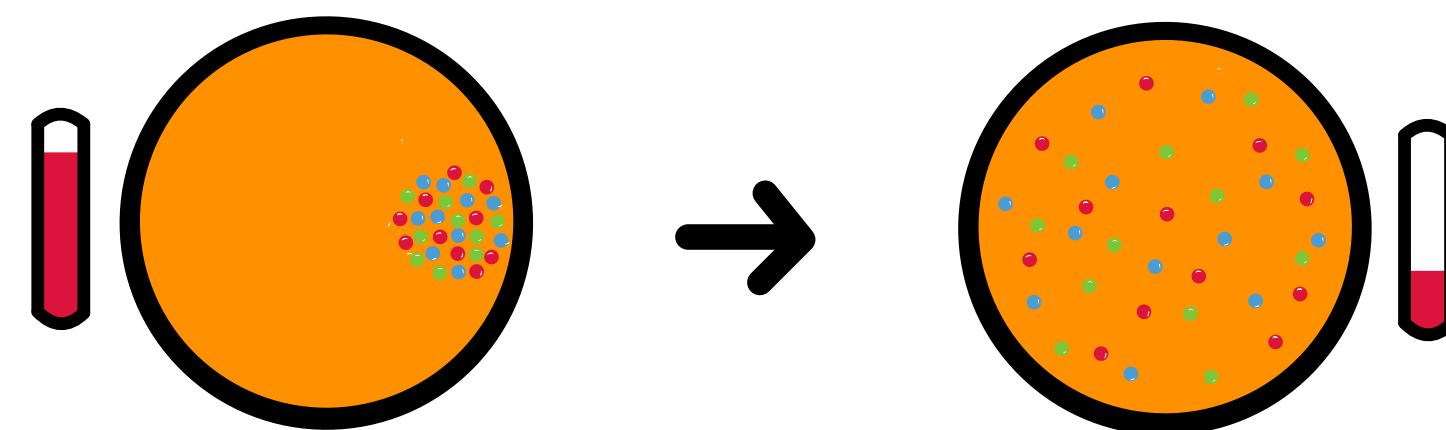
Application: charge diffusion near critical point

- Simple charge diffusion problem: [XA et al, 2009.10742](#)

$$\partial_t n = \nabla \lambda \nabla \alpha + \eta, \quad \langle \eta(x)\eta(y) \rangle = 2 \nabla^{(x)} \lambda \nabla^{(y)} \delta^{(3)}(x - y)$$

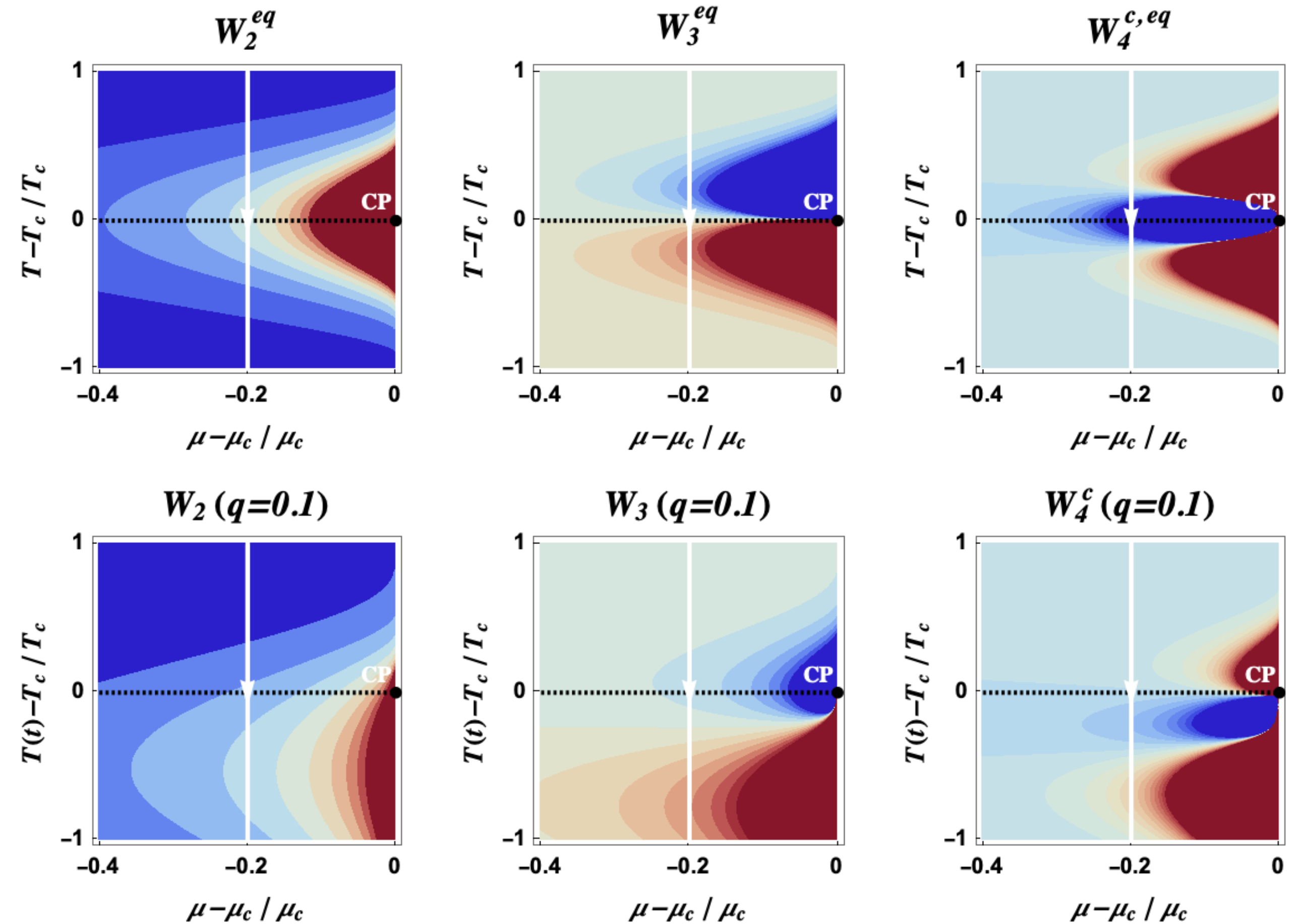
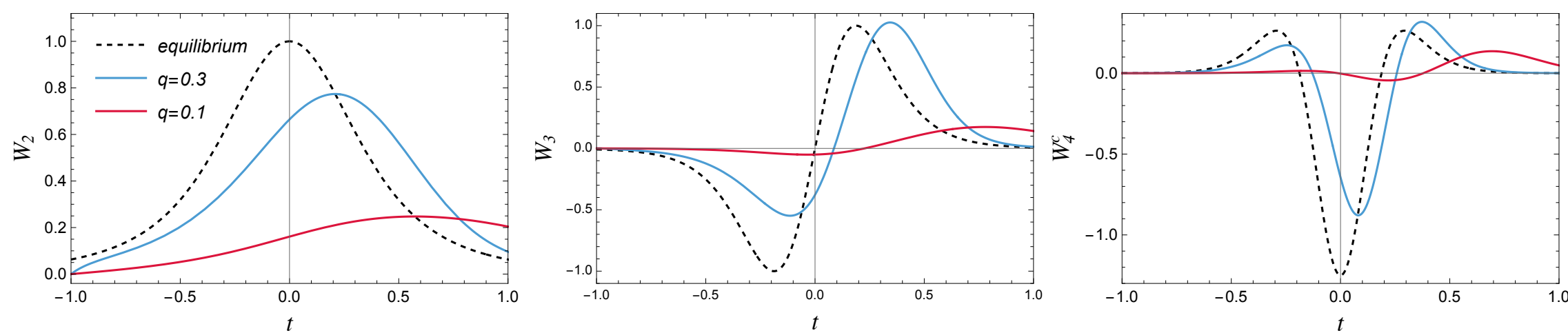
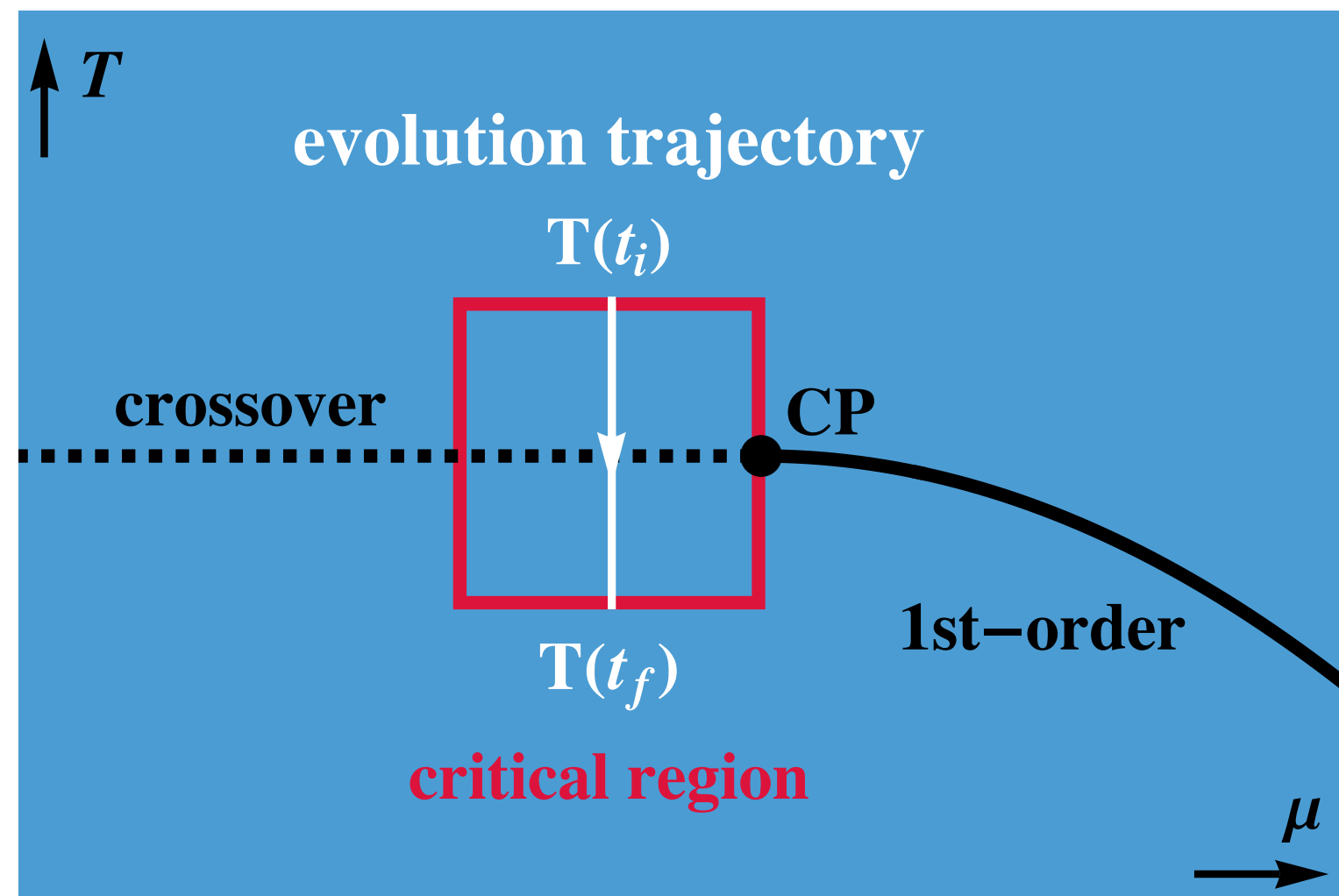
quantities	general	diffusive charge
variable	ψ_i	$n(\mathbf{x})$
variable index	$i, j, k, \text{ etc.}$	$\mathbf{x}, \mathbf{y}, \mathbf{z}, \text{ etc.}$
Onsager matrix	Q_{ij}	$\nabla_{\mathbf{x}} \lambda \nabla_{\mathbf{y}} \delta_{\mathbf{x}\mathbf{y}}^{(3)}$
drift force	F_i	$\nabla_{\mathbf{x}} \lambda \nabla_{\mathbf{x}} \alpha$

$n \equiv$ charge density; $\lambda \equiv$ conductivity; $\alpha \equiv$ chemical potential



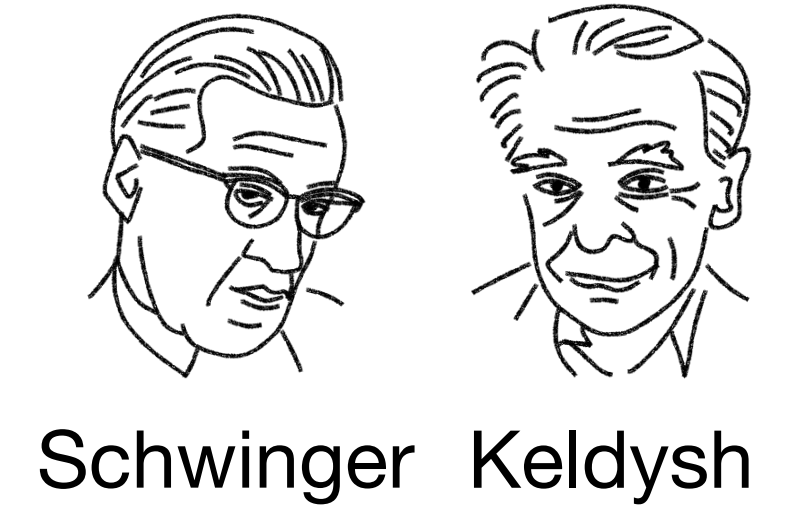
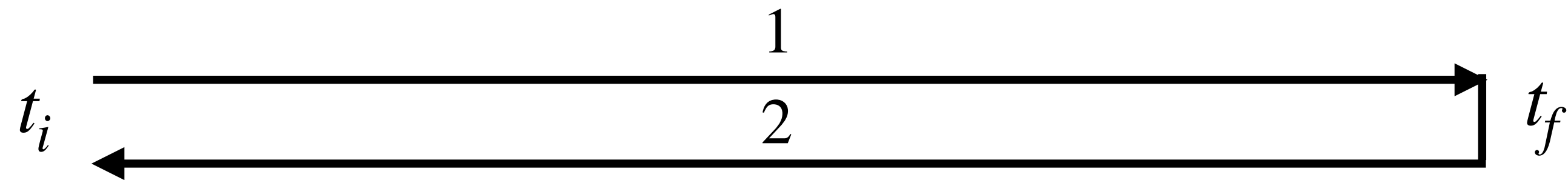
Application: charge diffusion near critical point

- Charge diffusion near QCD critical point: memory effect [XA, 2209.15005](#)



Schwinger-Keldysh approach

- SK formalism



$$Z = \int \mathcal{D}\psi_1 \mathcal{D}\psi_2 \mathcal{D}\chi_1 \mathcal{D}\chi_2 e^{iI_0(\psi_1, \chi_1) - iI_0(\psi_2, \chi_2)} = \int \mathcal{D}\psi_1 \mathcal{D}\psi_2 e^{i\int_\tau \mathcal{L}_{\text{EFT}}}$$

- The effective Lagrangian is constructed following *fundamental symmetries*:

[Glorioso et al, 1805.09331](#); [Jain et al, 2009.01356](#)

$$\mathcal{L}_{\text{EFT}}(\psi_r, \psi_a) = \psi_a Q^{-1} (F - \dot{\psi}_r) + i\psi_a Q^{-1} \psi_a \quad \text{where} \quad \psi_r = \frac{1}{2} (\psi_1 + \psi_2), \quad \psi_a = \psi_1 - \psi_2$$

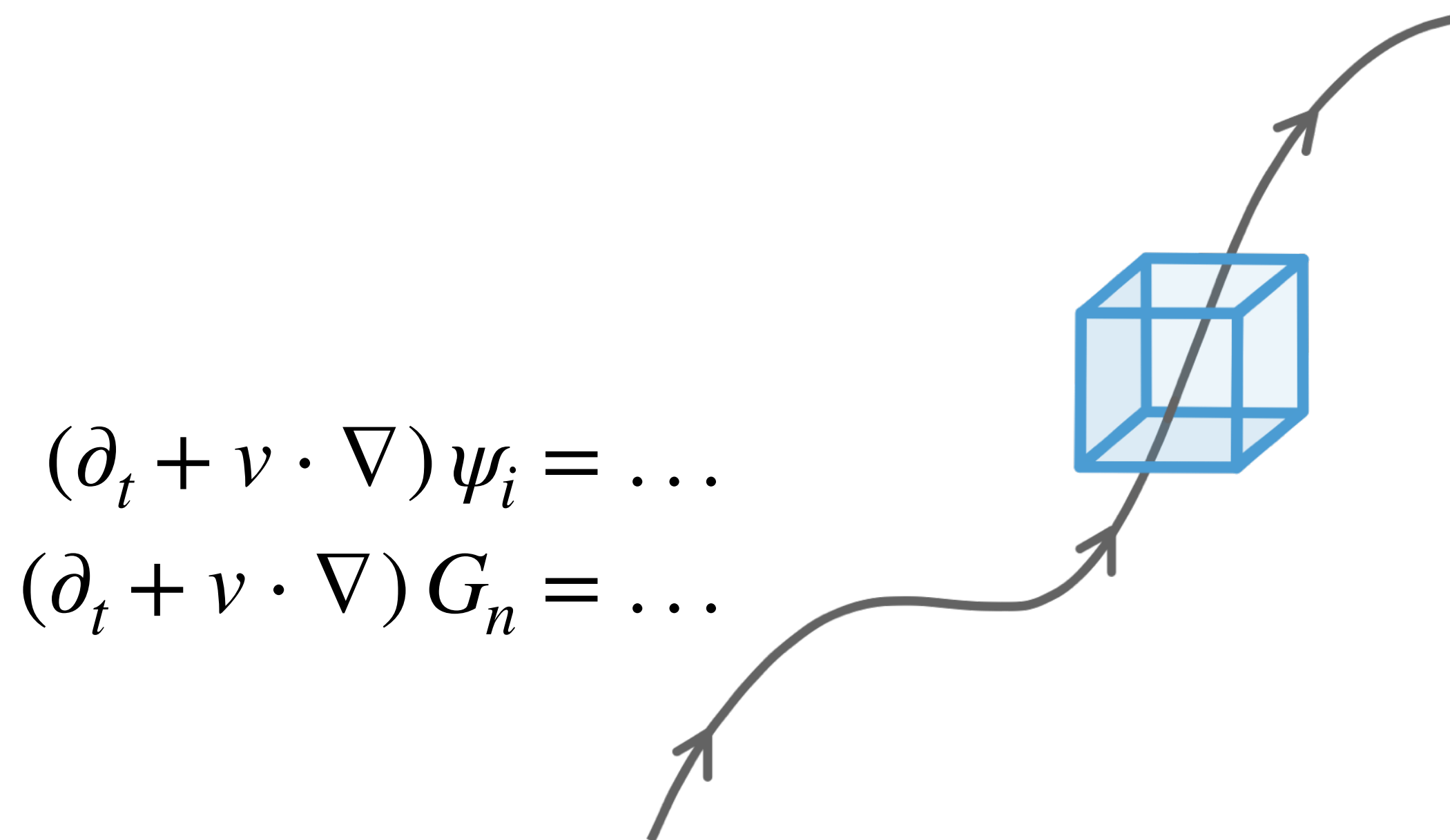
$$P[\psi] = \int_{\psi_r = \psi(t)} \mathcal{D}\psi_r \mathcal{D}\psi_a J(\psi_r) e^{i\int_{-\infty}^t d\tau \mathcal{L}_{\text{EFT}}} \longrightarrow \partial_t P[\psi] = \frac{\partial}{\partial \psi} (\mathbf{flux}[\psi])$$

[XA et al, in progress](#)

Relativistic dynamics

Eulerian specification

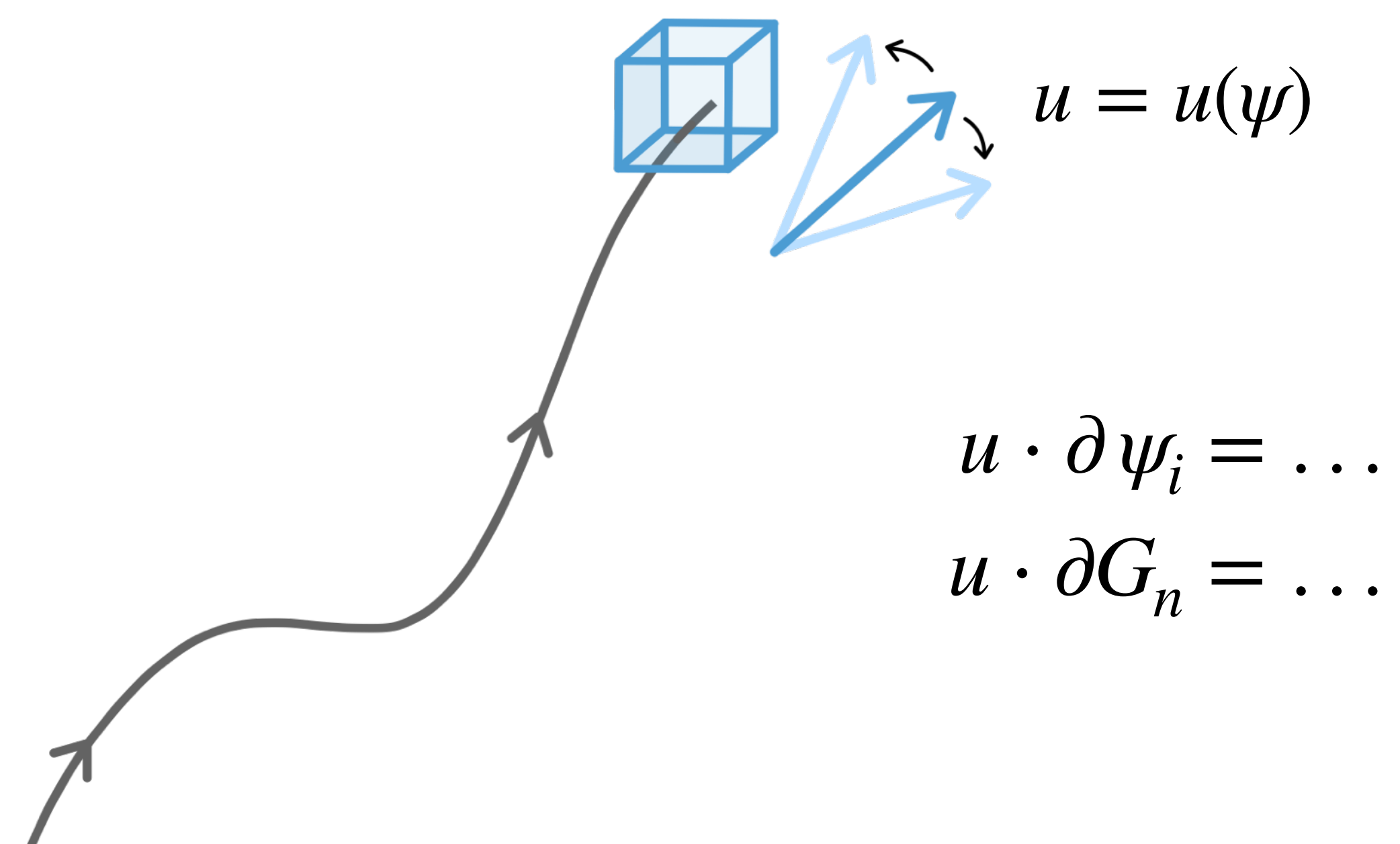
more often used in non-relativistic theory



There is a global time for every observer.
All n-pt correlators G_n can be measured
at the same time of the same lab frame.

Lagrangian specification

more convenient for relativistic theory

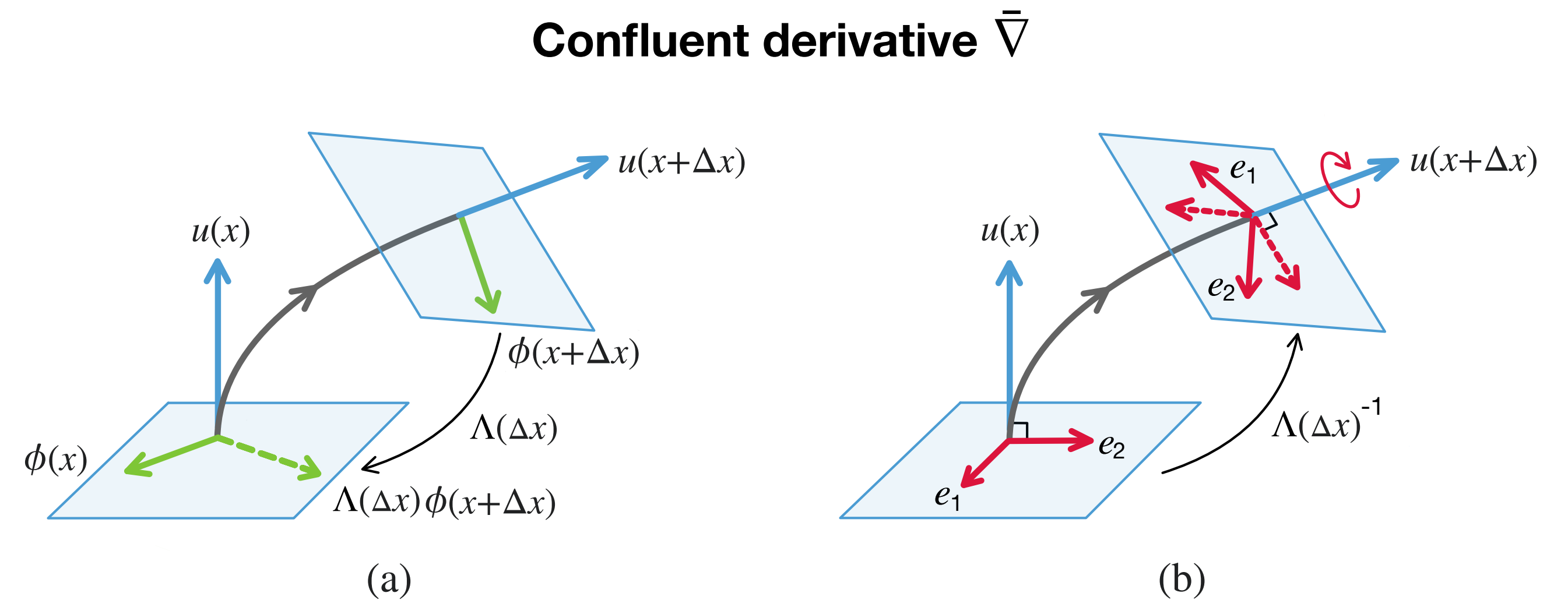
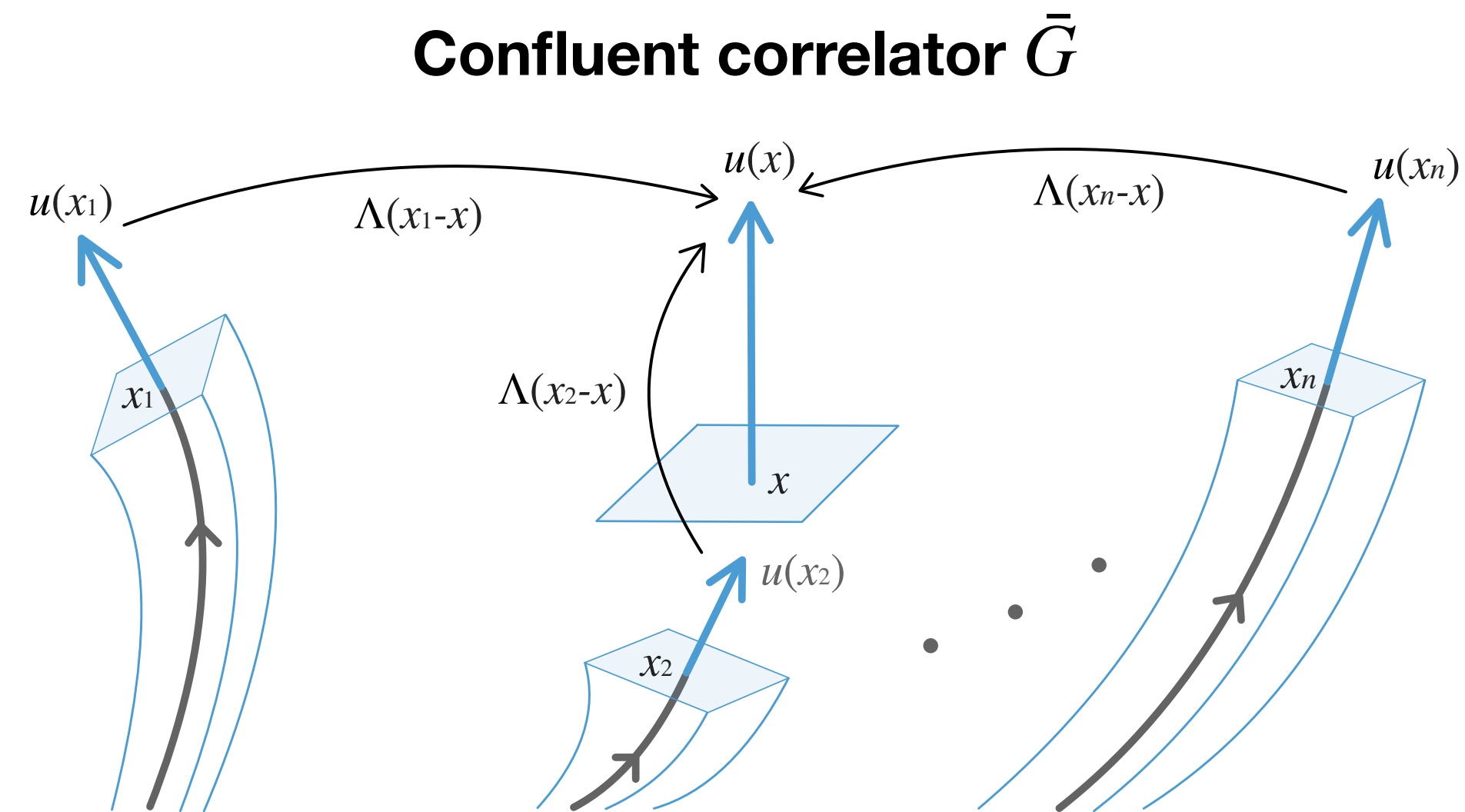


Each fluid cell has its own clock (proper time).
How to define the analogous *equal-time*
correlator G_n in relativistic theory?

Confluent formulation: correlator and derivative

- Confluent formulation: covariant description for the comoving fluctuations.

XA et al, 2212.14029



boost all fields (measured at their own local rest frame) to one common frame (chosen at their midpoint)

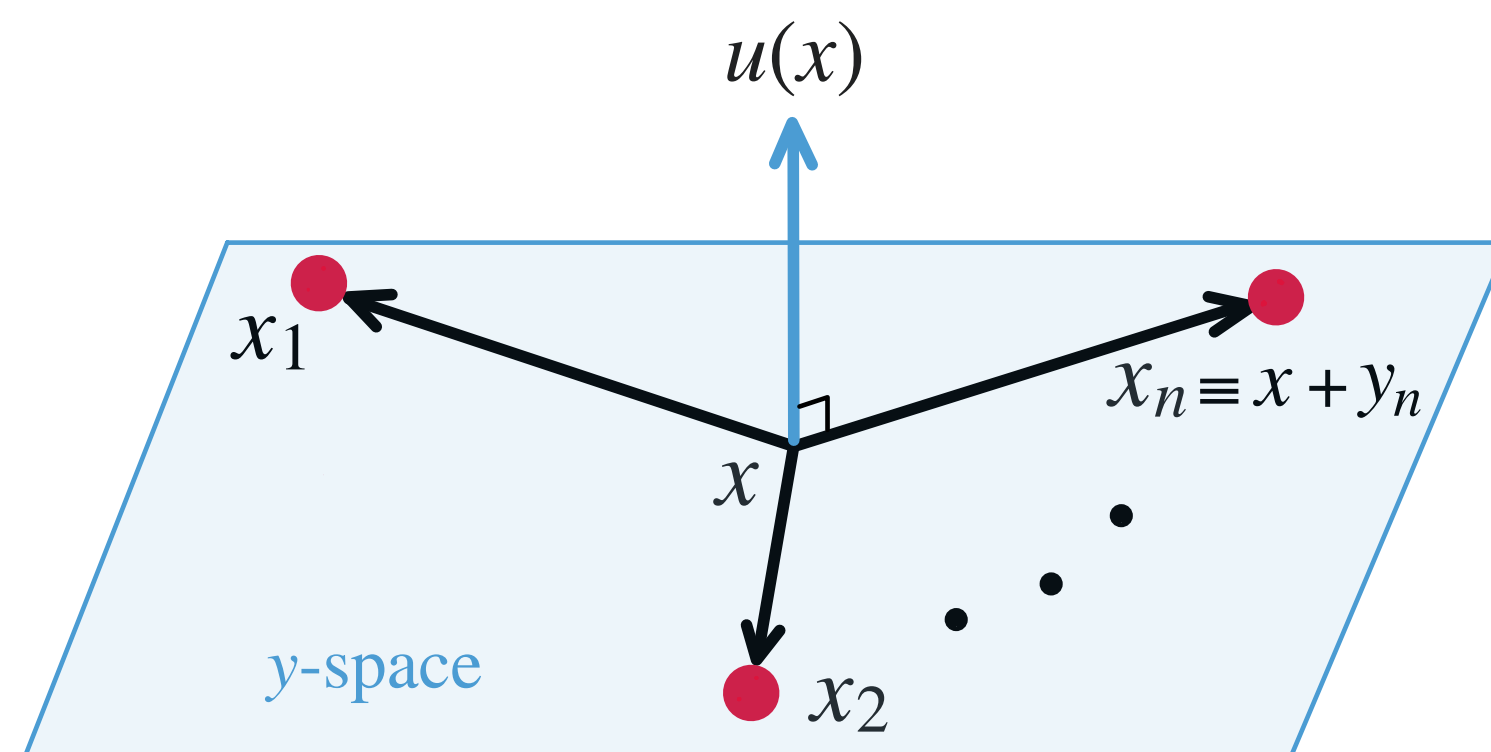
as the n points move, the frame at midpoint moves accordingly, the difference of a given field before and after the movement is calculated in one same frame, with the equal-time constraint preserved by introducing the local triad e_a^μ with $a = 1, 2, 3$

Confluent formulation: Wigner function

- The confluent n -pt Wigner transform is performed from $y^a = e^a_\mu y^\mu$ to q^a

XA et al, 2212.14029

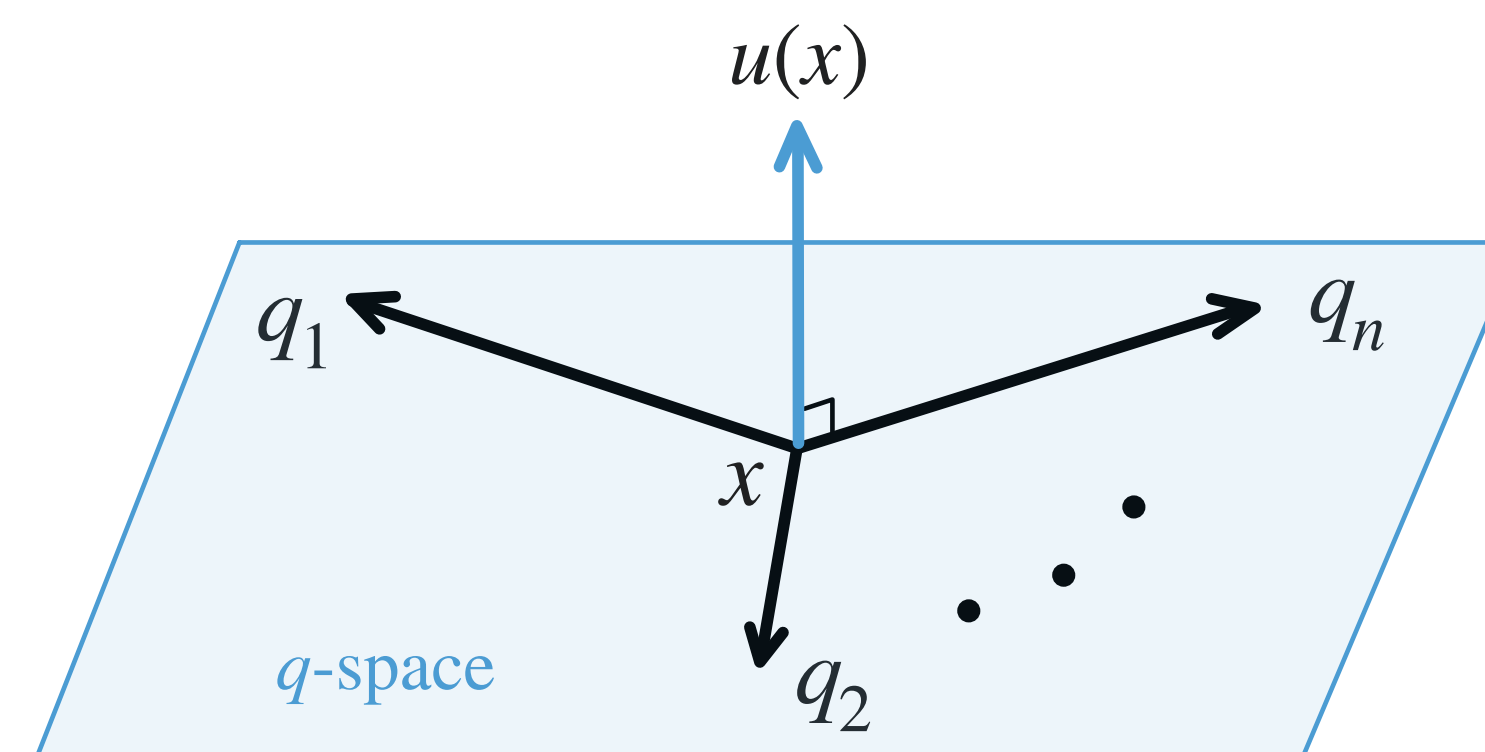
$$W_n(x; q_1^a, \dots, q_n^a) = \int \prod_{i=1}^n (d^3 y_i^a e^{-i q_{ia} y_i^a}) \delta^{(3)} \left(\frac{1}{n} \sum_{i=1}^n y_i^a \right) \bar{G}_n(x + e_a y_1^a, \dots, x + e_a y_n^a)$$



$$u(x) \cdot y_i = 0 \quad \& \quad y_1 + y_2 + \dots + y_n = 0$$

(a)

y-space



$$u(x) \cdot q_i = 0 \quad \& \quad q_1 + q_2 + \dots + q_n = 0$$

(b)

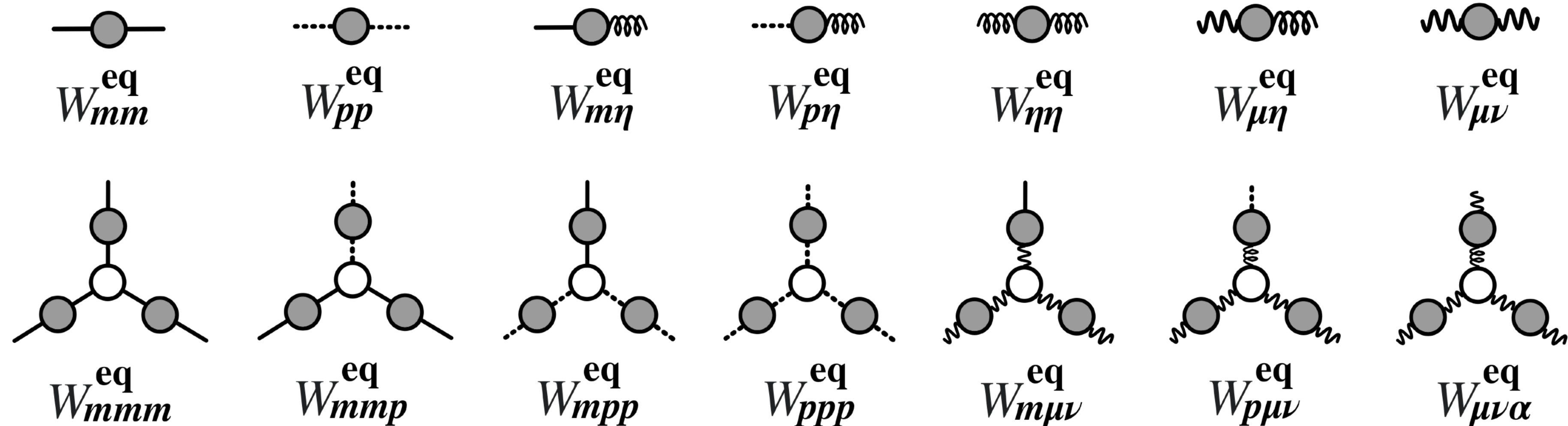
q-space

Fluctuation evolution equations

- Fluctuation evolution equations in the *impressionistic* form: [XA et al, in progress](#)

$$\mathcal{L}W_n = ic_s q(W_n - \dots) + \gamma q^2(W_n - \dots) + kW_n + \dots \quad \text{where} \quad \mathcal{L} = u \cdot \bar{\nabla}_x + f \cdot \nabla_q$$

for which the solutions match thermodynamics with entropy $S(m, p, u_\mu, \eta)$, where $m = s/n$ and η is a Lagrangian multiplier for $u^2 = -1$.

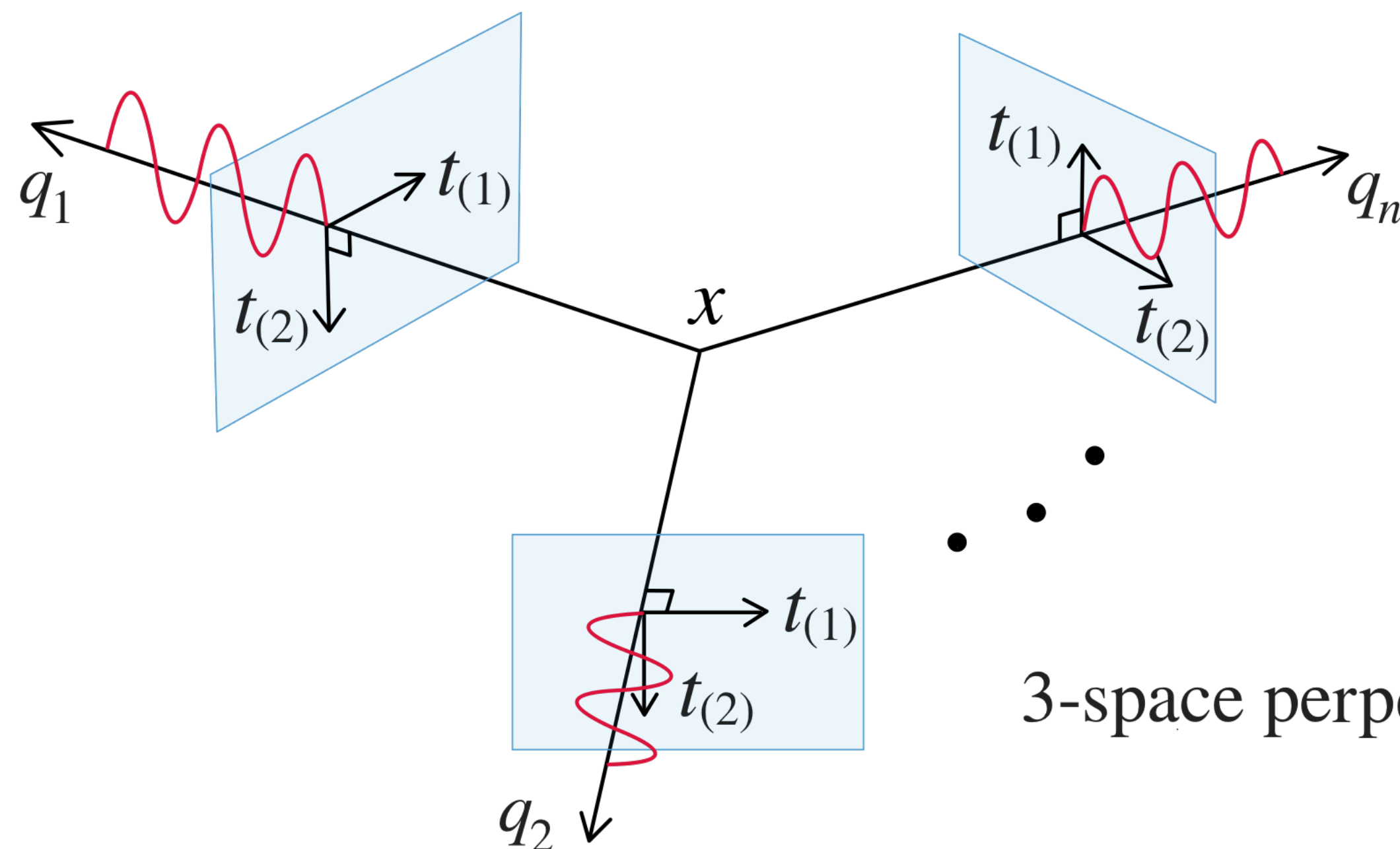


NB: given $\phi = (\delta m, \delta p, \delta u_\mu)$, one can derive $21+56+126=203$ equations for the 2-pt, 3-pt and 4-pt functions — — bite off more than one can chew!

Rotating phase approximation

- n -pt functions are analogous to n -particle states lying in the Fock space. One can choose a set of new bases s.t. the ideal hydro equations are diagonalized:

$$\phi = \begin{pmatrix} \phi_m \\ \phi_p \\ \phi_\mu \end{pmatrix} = \begin{pmatrix} \delta m \\ \delta p \\ \delta u_\mu \end{pmatrix} \longrightarrow \Phi = \begin{pmatrix} \Phi_m \\ \Phi_\pm \\ \Phi_{(i)} \end{pmatrix} \sim \begin{pmatrix} \delta m \\ \delta p \pm c_s \hat{q} \cdot (w \delta u) \\ t_{(i)} \cdot \delta u \end{pmatrix} \quad i = 1, 2$$



RPA: fast modes are averaged out over time scales $1/c_s q \ll \Delta t < 1/\gamma q^2$.

Slow modes

- Under RPA, slow modes are n -pt functions $W_n(q_1, \dots, q_n)$ subject to

$$\sum_{i=1}^n \lambda(q_i) = 0 \quad \text{where} \quad \lambda_{\pm}(q) = \pm c_s |q|, \quad \lambda_m(q) = \lambda_{(i)}(q) = 0.$$

E.g., $W_{+-}(q_1, q_2)$ is a slow mode since $\lambda_+(q_1) + \lambda_-(q_2) = c_s(|q_1| - |q_2|) = 0$.

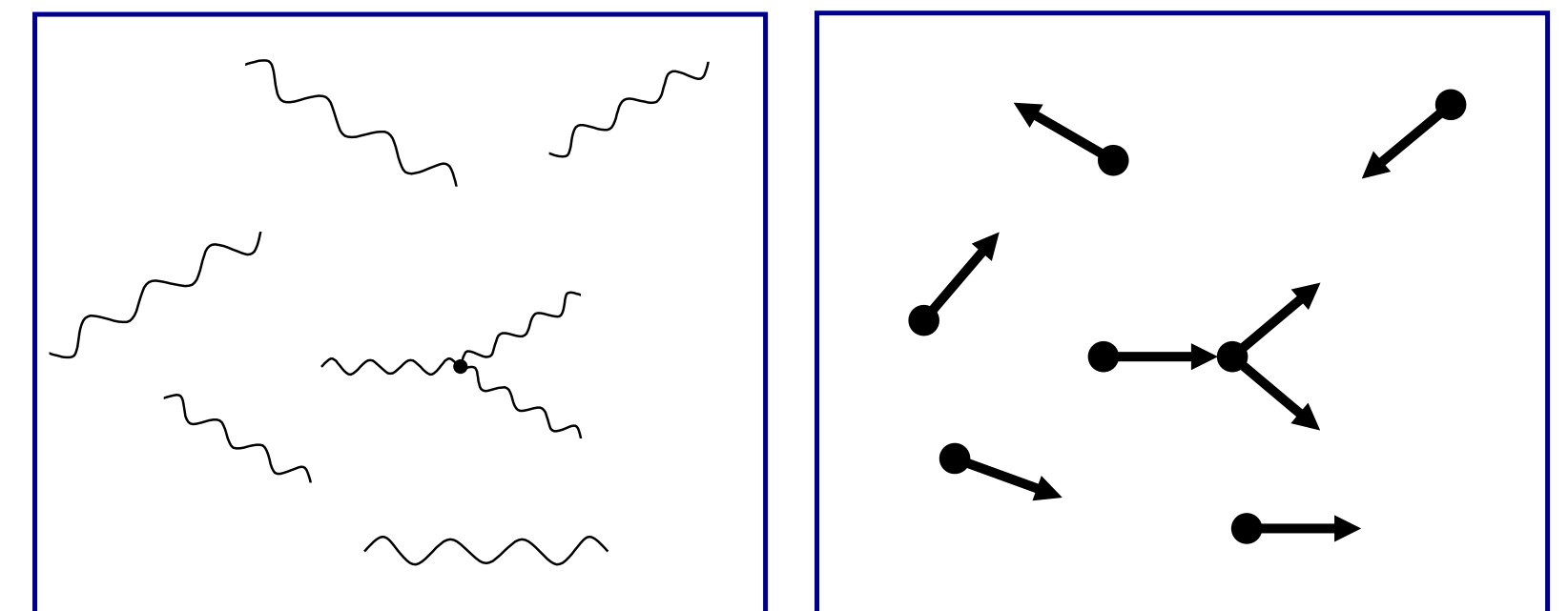
As a result, there are only $7+10+15=32$ equations.

E.g., for 2-pt functions the slow modes are $W_{mm}, W_{m(i)}, W_{(i)(j)}, W_{+-}$.

- Phonon interpretation: [XA et al, 1902.09517](#)

$$\mathcal{L}W_{+-} = -\gamma_L q^2 \left(W_{+-} - \frac{T}{E} \right)$$

Bose-Einstein distribution at high T : $f = \frac{1}{e^{E/T} - 1} \approx \frac{T}{E}$



Kinetic picture (from Schaefer)

NB: there is no analogous phonon interpretation for non-Gaussian fluctuations.

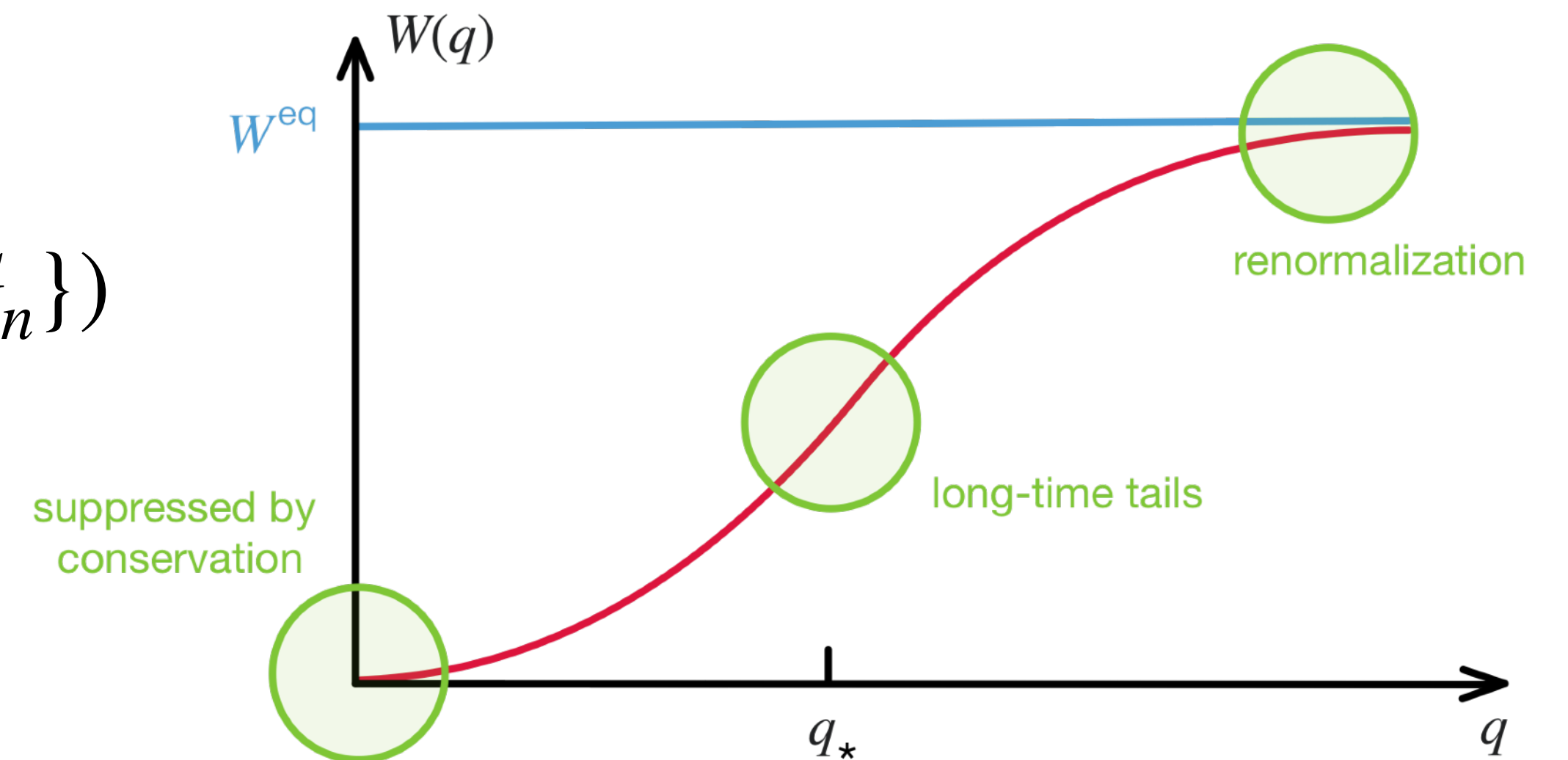
Fluctuation feedback

- Hydrodynamic fluctuations renormalize bare quantities order by order in gradient expansion.

$$\begin{aligned}
 T_{\mu\nu}^{\text{physical}} &= T_{\mu\nu}^{(0)} + T_{\mu\nu}^{(1)} + T_{\mu\nu}^{(2)} + \dots + \delta T_{\mu\nu}(\{G_n\}) \\
 &= T_{\mu\nu}^{R(0)} + T_{\mu\nu}^{R(1)} + T_{\mu\nu}^{R(2)} \\
 &\quad + \tilde{T}_{\mu\nu}^{(3/2)} + \tilde{T}_{\mu\nu}^{(3)} + \tilde{T}_{\mu\nu}^{(9/2)} + \dots
 \end{aligned}$$

where $G_n(x) = \int d^3q_1 \dots d^3q_n \delta^{(3)}(q_1 + \dots + q_n) W_n(x, q_1, \dots, q_n)$

Long-time tail due to n -pt functions is of order $\varepsilon^{n-1} \sim q^{3(n-1)} \sim k^{3(n-1)/2}$, the leading $k^{3/2}$ behavior results from 2-pt functions (via ).



Recap

- Various approaches for fluctuating hydrodynamics are developed, with their own advantages and disadvantages. [See also Schaefer's talk](#)
- Our framework for fluctuations dynamics now incorporates non-Gaussian fluctuations of fluid velocity.

Outlook

- Need efforts to simulate the fluctuation equations with background.
- Need freeze-out prescription for the connection to observables. [See Pradeep's talk](#)
- More...